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Data Collection

VMC\_CEPHEID

**Release Number** 

2

Data Provider Date Maria-Rosa Cioni 02.09.2014

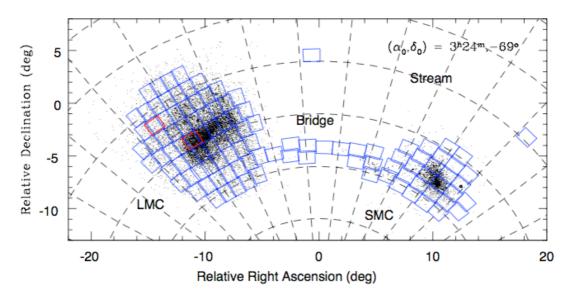
### **Abstract**

Observations were obtained with the VISTA telescope as part of the VISTA survey of the Magellanic Cloud system (VMC; ESO program 179.B-2003) in three filters: Y, J and  $K_s$ . The main goals of the VMC survey are the determination of the spatially resolved star formation history and the three-dimensional geometry of the Magellanic system. The sensitivity of the data is designed to reach sources below the oldest main-sequence turn off point of the stellar population and the multi-epochs to measure accurate  $K_s$  mean magnitudes for pulsating variable stars, e.g. RR Lyrae stars and Cepheids.

This catalogue data release refers to known classical Cepheid stars in two VMC tiles completed by end of September 2011 (Ripepi et al. 2012). These are tiles LMC  $6\_6$  (including the 30 Doradus star forming region) and LMC  $8\_8$  (including the South Ecliptic Pole region). VMC  $K_s$  magnitudes are listed together with V and I magnitudes from OGLE-III (tile LMC  $6\_6$ ) and V magnitudes from EROS-2 (tile LMC  $8\_8$ ) for each Cepheid star. Periods and modes of pulsation are also from the optical-band data. Reddening values were derived from Haschke, Grebel & Duffau 2011 (AJ, 141, 158). The total sky coverage of this release is <3.5 deg² in the LMC, corresponding to the sum of the areas of the two tiles, because tile LMC  $8\_8$  is only partly covered by the EROS-2 survey.

## **Overview of Observations**

The figure below shows the Magellanic system as tiled by the VMC survey (blue) and tiles for which catalogues are released (red). Underlying small dots indicate the distribution of carbon stars, stellar clusters and associations.



Tile numbering begins from the bottom right corner, increasing from right to left and from bottom to right. The first LMC tile is 2\_3, the first SMC tile is 2\_2, the first Bridge tile is 1\_2 and Stream tile 1\_1 is right above the Bridge while 2\_1 is to the right of the SMC.

## **Release Content**

This catalogue release covers two tiles in the Large Magellanic Cloud: LMC 6\_6 and LMC 8\_8.

LMC tiles were oriented with the Y axis more or less along the declination direction and cover about  $1.771 \text{ deg}^2$  each where the central  $(1.475 \times 1.017)=1.501 \text{ deg}^2$  corresponds to the nominal depth of the survey and the remaining area to half the exposure time in each band.

The classical Cepheid stars included in this release are initially extracted from the OGLE-III (tile LMC 6\_6) and EROS-2 (tile LMC 8\_8) surveys. Only 70% of tile LMC 8\_8 is covered by the EROS-2 and the variable stars in the remaining area of the tile have yet to be discovered.

Tile centres, number of records and the size in Mby are listed below.

| Tile    | RA           | Dec           | Records | Mby  |
|---------|--------------|---------------|---------|------|
| LMC 6_6 | 05:37:40.008 | -69:22:18.120 | 321     | 0.06 |
| LMC 8_8 | 05:59:23.136 | -66:20:28.680 | 11      | 0.02 |

#### **Release Notes**

The data for this release were prepared by the Wide Field Astronomy Unit (WFAU) and the VMC team from images processed by the Cambridge Astronomy Survey Unit (CASU).

The main processing steps are described in Cross et al. 2012 (A&A 548, A119) and Cross et al. 2009 (MNRAS 399, 1730). Variability catalogues were extracted from the VISTA Science Archive by WFAU using data in the VMCv20120126 internal release. Full details are given in Ripepi et al. 2012 (MNRAS, 424, 1807).

#### **Data Reduction and Calibration**

The procedures to reduce and calibrate the data are described in detail at: <a href="http://casu.ast.cam.ac.uk/survevs-projects/vista/technical/data-processing">http://casu.ast.cam.ac.uk/survevs-projects/vista/technical/data-processing</a>.

In particular, catalogues were created from images that were filtered for nebulosity with size of the order of 30 arcsec (Irwin 2010, UKIRT Newsletter 26, 14).

The magnitudes were not corrected for reddening.

#### **Data Quality**

The astrometric and photometric quality of the data is described in detail at <a href="http://casu.ast.cam.ac.uk/survevs-projects/vista/technical">http://casu.ast.cam.ac.uk/survevs-projects/vista/technical</a>.

#### **Known issues**

Note that in Ripepi et al. (2012) IAUNAMEs may differ slightly from those included in this release. The right ascension components may differ by 0.01 s and the declination components are rounded to one decimal point.

#### **Data Format**

#### **Files Types**

Two known Cepheid catalogues, one per tile, are released:

```
vmc_er3_05h37-069d22_yjks_cepheidCat_558345748484.fits vmc_er3_05h59-066d20_yjks_cepheidCat_558345748482.fits
```

where the name is constructed as project\_release\_ra/dec\_bands\_typeofCat\_framesetID.fits and framesetID uniquely identifies the tile as follows:

```
558345748484 LMC 6_6
558345748482 LMC 8_8.
```

A MetaData file, vmc\_er3\_yjks\_cepheidCatMetaData.fits, accompanies the release. Its name refers to project\_release\_bands\_typeofCat.fits.

### **Catalogue Columns**

Each known Cepheid catalogue contains the following columns.

- # Number; name; format; description
- 1; IAUNAME; 29A; Unique identifier in IAU naming convention
- 2; SOURCEID; K; UID of this merged detection as assigned by merged algorithm
- 3; VARID; K; UID of VMC variables
- 4; FIELDID; 8A; ID of field
- 5; CUEVENTID; J; UID of curation event giving rise to this record
- 6; RA2000; D; Celestial Right Ascension
- 7; DEC2000; D; Celestial Declination
- 8; CEPHMODE; 5A; Type of Cepheid
- 9; IMEANMAG; E; Mean I band magnititude
- 10; VMEANMAG; E; Mean V band magnititude
- 11; PERIOD; E; Period of first mode of oscillation
- 12; KSMEANMAG; E; Mean Ks band magnititude
- 13; KSAMPL; E; Amplitude of Ks band magnititude variation
- 14; KSMAGERR; E; Error in mean Ks band magnitude
- 15; EVI; E; The dust extinction value E(V-I)
- 16; NOTES; 8A; Additional information

The format refers to the fits notation as follows:

A - string 32 characters; D - double floating point (8 bytes); E - real floating point (4 bytes); I - short integer (2 bytes); J - integer (4 bytes); K - long integer (8 bytes).

# Acknowledgements

Please reference Ripepi et al. 2012, MNRAS, 424, 1807 and use the following statement in your articles when using these data: Based on data products from observations made with ESO Telescopes at the La Silla Paranal Observatory under programme ID 179.B-2003.