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SOUTHERN
OBSERVATORY

ESO-MIDAS

MUNICH
IMAGE
DATA
ANALYSIS
SYSTEM

VOLUME A:
SYSTEM

ESO

ESO-MIDAS User Guide

Volume A: System



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Chapter 1

Introduction

ESO-MIDAS¹ is the acronym for the European Southern Observatory - Munich Image Data Analysis System which is developed and maintained by the European Southern Observatory. The official name, ESO-MIDAS, is a registered trademark. In this manual the name MIDAS is used as an abbreviation of ESO-MIDAS. The MIDAS system provides general tools for image processing and data reduction with emphasis on astronomical applications including special reduction packages for ESO instruments at La Silla. The system is available for both VAX/VMS and UNIX systems.

A large number of contributions have been made to MIDAS by people inside and outside ESO. We greatly appreciate and acknowledge these efforts. A full list of acknowledgements can be found in Appendix B.

This manual gives the necessary information to do useful data reduction with the system, whereas a detailed technical description of the design and software interfaces can be found in other documents (see Section 1.6). These documents also describe how users can write and add their own application programs to the system.

1.1 How to use the MIDAS Manual

This document is intended to be a description of how to use the various facilities available in the MIDAS system. The manual consists of three volumes:

Volume A: describes the basic MIDAS system with all general purpose facilities such as MIDAS Control Language, data input/output (including graphics and image display), table system (MIDAS Data Base). A summary of all available commands as well as site specific features are given in appendices.

Volume B: describes how to use the MIDAS system for astronomical data reduction. Application packages for special types of data or reductions (*e.g.* long slit and echelle spectra, object search, or crowded field photometry) are discussed assuming intensity calibrated data. A set of appendices gives a detailed description of the reduction of raw data from ESO instruments.

¹Trade-mark of the European Southern Observatory

Volume C: gives the detailed description for all commands available.

It is intended that users will mainly need Volume A for general reference. For specific reduction of raw data and usage of special astronomical packages, Volume B will be more informative. A printed version of the MIDAS help files is available in Volume C. Users are recommended to use the on-line help facility which always gives a full up to date description of the commands available.

1.1.1 New Users

To be able to use MIDAS, it is a great advantage to have some basic knowledge of computer systems such as how to login, use of the file editor and simple system commands. Such instructions can normally be found in local system documentation or in Appendix C of this volume. After having acquired this knowledge, new users should read Chapter 2 carefully. This will give a basic introduction to the MIDAS system with some examples.

1.1.2 Site Specific Features

MIDAS is used at many different sites on a large variety of configurations. The main part of this manual does not refer to special configurations or hardware devices. Site specific implementations and details of the local installation can be found in Appendix C.

1.2 General Concept of MIDAS

The MIDAS system is built along lines which should allow easy integration of complex analysis algorithms as well as allowing greater flexibility in interactive use and in the creation of user specific procedures from the basic building blocks. The first design proposal for MIDAS, made late 1980, used some ideas from the UK STARLINK project for the software interface definitions. The present version which became available in 1984 follows a similar philosophy in its application program interfaces, but has been expanded to the new Standard Interfaces which have a broader base than previously.

MIDAS has benefitted greatly from the experience gained at ESO using the Hewlett-Packard based image processing system IHAP (see F. Middelburg, IHAP Manual, ESO 1985). Not only have many of the internal design features such as "world coordinates" been incorporated, but also the command language has been designed in such a way that it is similar to the basic philosophy of IHAP.

The MIDAS system can be run in both an interactive and a batch mode. In addition, the interactive user will be able to create batch jobs which will run in parallel with the interactive work.

MIDAS is based on three sets of general interfaces for application programs to data structures, namely: a) the "Standard Interfaces" for general I/O and image access, b) the "Table Interfaces" for access to table structures, and the "Graphics Interfaces" for data representation in graphical format. These interfaces allow easy integration of application programs into MIDAS. To provide a portable system a layer of OS-routines have been

used to shield MIDAS from the local operating system. These routines may only be used at lowest levels and are not available for normal applications.

To facilitate easy implementation of different graphic and display devices, MIDAS has adopted a set of device independent interfaces for plotting and image display. All plotting routines in MIDAS are based on the **ASTRONET Graphic Library**, developed and maintained by the **italian ASTRONET**. Further, the image display applications are using the **Image Display Interface** routines defined in collaboration with **ST ScI, UK** **STARLINK, KPNO** and **Trieste Observatory**.

1.3 Distribution Policy

The MIDAS system is available, free of charge, to all non-profit research institutes, whereas other organisations or companies may be charged a nominal fee to cover distribution. Institutes interested in using MIDAS must sign a User Agreement before they receive MIDAS. The necessary forms can be obtained by contacting the Data Management Division at ESO/Munich. MIDAS is distributed in source code copyrighted by ESO with all rights reserved. Institutes receiving the MIDAS system are not allowed to redistribute it to other sites without explicit written permission from ESO. The use of the MIDAS system for data reduction should be properly acknowledged in papers and publications. It is recommended to refer to the specific MIDAS version used, *e.g.* 95NOV, in the acknowledgement.

The availability of new releases is announced through electronic mail. Requests can be submitted either on special MIDAS Request Forms or through e-mail quoting the user agreement number. Currently, MIDAS has one yearly release in November (*e.g.* the release in 1995 is will be named 95NOV). The current system at ESO/Munich is frozen several months prior to the official release. The official release is based on this version, which is extensively tested both at ESO and at a number of β -test sites. Problems detected during these tests are corrected in the official release version which is then given free for distribution. For institutes with internet access, the release can only be obtained from the MIDAS ftp account with a password (on an exceptional basis it is also available on tape). Institutes without internet access can request MIDAS on a range of tape media.

Binary copies of ESO-MIDAS (already installed and without sources) for specific systems are publicly available in the anonymous ftp account and do not need the signing of an ESO-MIDAS User Agreement. Official user support and documentation service from ESO for these binary copies are however not granted without a prior signed User Agreement.

Updates of the manual and other documentation take somewhat longer to prepare and print and can be obtained from the anonymous ftp account. Only in extreme cases, *e.g.* in cases where an Internet connection is absent, copies and updates of documentation can be obtained by writing to the Data Management Division at ESO/Munich.

1.4 Support

The MIDAS system is supported in a variety of ways. If people encounter problems which cannot be solved locally (*e.g.* through the manual) they can use the MIDAS **Hot-Line**

service. This service will provide answers to MIDAS related questions received through the following list of electronic mail and telex addresses:

- uucp: `midas@eso.uucp`
- internet: `midas@eso.org`
- span: `eso::midas`
- Telefax: +49 89 32006480 (attn.: MIDAS HOT-LINE)
- Telex: 528 282 22 eo d (attn.: MIDAS HOT-LINE)

Requests and questions are acknowledged when received and processed as soon as possible, normally within a few days. Users are also strongly encouraged to send suggestions and comments via the **MIDAS Hot-Line**.

In urgent cases, users can use a special MIDAS Support telephone service at ESO on the number +49-89-32006-456. This line is connected to the MIDAS Users Support which is able directly to answer questions concerning MIDAS or investigate the problem in more complicated cases. Although this telephone service is available we prefer that questions or requests are submitted in writing via the **MIDAS Hot-Line**. This makes it easier for us to process the requests properly. A database with problem reports and answers is available for interrogation using the STARCAT utility at ESO. General information concerning the MIDAS system should be addressed to User Support Group, European Southern Observatory, Karl-Schwarzschild-Straße 2, D-85748 Garching bei München (attn: Midas Support).

Besides these support services, a newsletter, the ESO-MIDAS Courier, is issued twice a year. The ESO WWW Xmosaic server may be accessed for up-to-date information about ESO-MIDAS via URL "<http://http.hq.eso.org/midas-info/midas.html>".

1.5 Requirements for Running MIDAS

MIDAS can run on computers with either VAX/VMS² or UNIX³ operating systems. Details of hardware and software requirements for MIDAS are listed below.

1.5.1 Hardware

- Computer system: any system which can run either VAX/VMS or UNIX operating systems. MIDAS implementations have been made on a large number of systems from vendors e.g. VAX stations, DEC stations, SUN SPARCstations, HP 700 series, IBM RS/6000 systems, PCs. The availability for a specific system can be checked by asking the MIDAS Hot-line.
- Memory: depending on the number of users of the systems but normally at least 8 Mbyte. A physical memory that is too small may significantly reduce the performance due to swapping of data to disk.

²Trademarks of Digital Equipment Corporation

³Trademark of AT&T

- Disk: the full MIDAS system requires of the order of 100 Mbyte of disk storage depending on the type of CPU. During installation an extra 10 Mbyte should be available for temporary files such as object code. The size of the system can be reduced in three ways: a) source files can be deleted after implementation, b) help-files can be removed if on-line help is not required, and c) parts of the system which are not used (*e.g.* crowded field photometry or echelle packages) need not be loaded. Implementations using shared libraries are available on some systems (*e.g.* SPARCstations and PCs). This reduces the needed disk space significantly.

A typical disk size for a single user system is approximately 200 Mbyte assuming 60 Mbyte for the operation system, 80 Mbyte for MIDAS, and 50-100 Mbyte for a user with 2-dimensional data.

- Terminals: any alpha-numeric terminal can be used. MIDAS can either work in a simple line-by-line mode or provide special features such as line editing by using a terminal definition file for special terminal features.
- Graphic display: the graphics software of MIDAS uses the device independent plotting library AGL⁴ made by the Italian ASTRONET. Drivers for a significant number of different devices are available *e.g.* Tektronix 4010/4015, X Window System, version 11 and PostScript (see a complete list in Chapter 6). It is possible to write drivers for devices for which none exist (see the AGL driver manual).
- Image displays: MIDAS uses the Image Display Interfaces which provides a general interface to image display devices. IDI-routines are available for DeAnza IP 8500 and X Window. X Window System, version 11, will be supported as the general interface to workstations. Other devices can be used if an appropriate set of IDI-routines are written. A list of currently available IDI-routines can be obtained from the User Support Group.

1.5.2 Software

- System: VAX/VMS or UNIX operating systems.
- Compilers: C- and FORTRAN-77 compilers (a FORTRAN-to-C conversion package may be used instead of a FORTRAN-77 compiler).
- Libraries: AGL is used for all plotting in MIDAS. This library is normally available on the release media but can also be obtained from the Italian ASTRONET (free of charge for non-profit research institutes).

NAG⁵ is used in a few packages such as fitting. It is under license and can be purchased from the Numerical Algorithms Group. MIDAS can be installed without this library in which case some commands will be unavailable.

⁴Astronet Graphical Library made by the Italian ASTRONET

⁵NAG made by Numerical Algorithms Group

1.6 Other Relevant Documents

There are several other documents relevant to the MIDAS system. General descriptions of the system can be found in the following references:

- Banse, K., Crane, P. Ounnas, C., Ponz, D., 1983 : 'MIDAS' in *Proc. of DECUS*, Zurich, p.87
- Grosbøl, P., Ponz, D. , 1985 : 'The MIDAS Table File System', *Mem.S.A.It.* **56**, p.429
- Banse, K., Grosbøl, P., Ponz, D., Ounnas, C., Warmels, R., 'The MIDAS Image Processing System in Instrumentation for Ground Based Astronomy: Present and Future, L.B. Robinson, ed., New York: Springer Verlag, p.431.

For general bibliographic reference to the MIDAS system (VAX/VMS version), the first reference in the above list should be used.

Detailed technical information of software interfaces and designs used in MIDAS is given in the following documentation:

- MIDAS Environment
- MIDAS IDI-routines
- AGL Reference Manual

Users who want to write their own application programs for MIDAS should read the MIDAS Environment document which gives the relevant information and examples.

For users who have to work with both the IHAP and MIDAS systems a cross-reference document has been made for the most commonly used commands:

- MIDAS-IHAP/IHAP-MIDAS Cross-Reference

All documents, either in postscript or in ascii format, are available via the MIDAS anonymous ftp account.

Chapter 2

Cook-Book

This chapter outlines the necessary information to get started with the MIDAS system. Further details can be found in the appropriate sections of the following chapters.

The essential steps are:

- To login to the computer you want to use.
- Start up the MIDAS monitor.
- Load some data from tape into your disk space.
- Execute the MIDAS commands you want.
- Save processed data on tape.
- Exit from the MIDAS monitor and logout.

These steps are outlined in the following sections.

2.1 Terminology

The following explain various terms used in this manual.

Keywords — global variables in the MIDAS monitor. They can be single numbers or characters or one dimensional arrays used to store input, output, or control information for MIDAS commands.

Frames — arrays of numbers representing data values with uniform sampling. They are used for storage of images or spectra.

Images — used interchangeably with frames.

Descriptors — variables associated to frames, tables or masks describing the contents in them. They are basically the same as keywords just associated to data files instead of the monitor. These descriptors have names like NAXIS (the dimension of the image array), CUNIT (the units of the axes), etc.

Tables — two dimensional arrays organised with rows and columns. They are used for storage of heterogeneous data contrary to frames and masks which store homogeneous data. They are typically used for saving lists of e.g. stellar positions and magnitudes. See Chapter 5 for a full description of the table facilities. Many commands output their results or take their input from tables. They constitute a simple data base system for MIDAS.

Catalogues — a list of frames, tables, or masks which can be used for input to various commands or merely for reference.

Procedures — These are lists of MIDAS commands stored in a file of type *filename.prg* and which can be executed by typing `@@ filename`. See Chapter 3 for further details.

Fit file — a file that contains the function and parameter values for use in conjunction with the fitting commands described in Chapter 8.

2.2 Commands

MIDAS is a command driven system in which the user enters commands followed by parameters. This implies that the user must know a few commands and their structure in order to make effective use of the system. Since most users cannot keep all the commands and their parameters at their finger tips, an extensive on-line help facility has been created as well as a printed version of the help text (see the Appendices).

A MIDAS command has the following structure:

```
COMMAND/QUALIFIER par1 par2 ... par8
```

where *par1* is the first parameter and so on. The important points are:

- Command and qualifier are separated by a / (slash).
- The command/qualifier and the parameters are separated by a *space*.
- Most commands have qualifiers.
- A parameter may contain several sub-parameters which are separated by commas.
- In most cases if the parameters are not specified, the system makes sensible defaults, but the user should **not** always trust these default values to be those he might have chosen.
- Keep these rules in mind, otherwise you will confuse the command.

MIDAS commands divide themselves into three categories:

- MIDAS primitive commands
- MIDAS application commands

- procedure control commands

The MIDAS commands are listed in alphabetical order and explained in detail in Volume C of the MIDAS User Guide. The application commands are developed for special purposes such as CCD or CASPEC reductions. They are described and listed in the various sections related to particular applications. For reductions of data you should refer to Volume B of this manual. The procedure control commands are described in Chapter 3.

2.3 Getting Started

The first thing to do is to login to the computer which you would like to use for your data reductions. The detailed procedure for getting permission to use the computer and getting allocated disk space for your data reductions can be found in Appendix C. Assuming that you have succeeded in logging into the computer, the following subsections describe typical use of MIDAS.

When you have logged in you should check that you have sufficient disk space in the directory in which you are working. For reductions of images an area of the order of 50 Mbytes would be adequate while for spectra reductions and analysis of final results less space is needed. Now you are ready to start the MIDAS system:

- Type **INMIDAS** on a VMS system or **inmidas** on a UNIX system. This will initiate the MIDAS environment and the terminal should respond with:

Midas 001>

- The available commands can be listed out by typing **HELP**. This only gives you the names of the commands presently available in the system. A summary and a subject grouped listing of the commands you will find in Appendix A. You can find the full explanation of the individual commands in Volume C of the MIDAS User Guide. A command can have several qualifiers which will change the mode of execution of the command e.g **STATISTICS** has qualifiers **IMAGE**, **TABLE** and **POISSON**. It is possible, typing **HELP command**, to get a display of all the *command/qualifier* combinations available for the given *command*. Detailed information about a specific *command/qualifier* combination can be obtained by typing **HELP command/qualifier**.
- In some installations a number of tutorial commands are available. (see Table 2.1) They provide an illustration of different parts of the system.

2.3.1 Simple MIDAS Session

This section gives two examples of simple MIDAS sessions. The first one reads some frames from a magnetic tape, displays them on a monitor and performs some simple operations. The second one creates MIDAS tables and performs some simple operations. In the following examples, user input is written in bold face type while comments, (after an exclamation mark) are written in normal roman font.

```

$ inmidas
Midas 001> HELP INTAPE                                !get help for tape input
Midas 002> INTAPE * x TAPE1 FNN                      !lists headers of all files
                                                       !from tape mounted on TAPE1
Midas 003> INTAPE 2,16-17,31-33,52 CCD TAPE1          !read files from TAPE1

Image ccd0002      : FF D V  20S      , naxis: 2, pixels: 337, 520
Image ccd0016      : DK BIAS 1S      , naxis: 2, pixels: 337, 520
Image ccd0017      : DK BIAS 1S      , naxis: 2, pixels: 337, 520
Image ccd0031      : DK      60S      , naxis: 2, pixels: 337, 520
Image ccd0032      : DK      60S      , naxis: 2, pixels: 337, 520
Image ccd0033      : DK      60S      , naxis: 2, pixels: 337, 520
Image ccd0052      : A0532-527 V 300 , naxis: 2, pixels: 337, 520

Midas 004> INTAPE 78 ccd image.fits                  !read fits file from disk

Image ccd0078      : A1029-459 V 40S , naxis: 2, pixels: 337, 520

Midas 005> CREATE/ICAT                                !create a catalogue of images
Image catalog icatalog.cat with 8 entries created...
Midas 006> SET/ICAT                                   !enable catalogue
Midas 007> READ/ICAT                                  !list the catalogue out
Image Catalog: icatalog.cat

No   Name        Ident                         Naxis  Npix(1,2)
#0001 ccd0002.bdf  FF D V  20S                2      337  520
#0002 ccd0016.bdf  DK BIAS 1S                2      337  520
#0003 ccd0017.bdf  DK BIAS 1S                2      337  520
#0004 ccd0031.bdf  DK      60S                2      337  520
#0005 ccd0032.bdf  DK      60S                2      337  520
#0006 ccd0033.bdf  DK      60S                2      337  520
#0007 ccd0052.bdf  A0532-527 V 300            2      337  520
#0008 ccd0078.bdf  A1029-459 V 40S            2      337  520

Midas 008> STAT/IMA ccd0016                         !compute statistics for this frame

frame: ccd0016  (data = R4)
complete area of frame
minimum, maximum:          184.0000      16383.00
at ( 215,    2), ( 337,    520)
mean, standard_deviation:  251.2245      880.0140
3rd + 4th moment:          0.1305419E+11  0.2137182E+15
total intensity:            0.4402457E+08
median, 1. mode:            16287.41      16287.71
total no. of bins, binsize: 256          63.52549
# of pixels used = 175240 or 100.00 % of all possible pixels (= 175240)
from    1      1 to    337    520 (in pixels)

Midas 009> STAT/IMA ccd0017

frame: ccd0017  (data = R4)
complete area of frame

```

```

minimum, maximum:           187.0000      16383.00
at ( 113,    1),( 337,   520)
mean, standard_deviation:  251.2514      880.0129
3rd + 4th moment:          0.1305419E+11  0.2137182E+15
total intensity:            0.4402929E+08
median, 1. mode:            16287.30      16287.73
total no. of bins, binsize: 256           63.51373
# of pixels used = 175240 or 100.00 % of all possible pixels (= 175240)
from      1      1 to    337    520 (in pixels)

Midas 010> CREATE/DISPLAY           !create a display window
Midas 011> LOAD/IMA ccd0017         !display image
Midas 012> GET/CURS                !read some pixels values from the image

cursor #0
frame pixels           world coords      intensity
frame: ccd0017
      334.0      166.0      334.000      166.000      228.000

Midas 013> EXTRACT/IMA ff = ccd0002[<,<:@330,>]           !remove
                                                               !irrelevant columns
Midas 014> EXTRACT/IMA biai1 = ccd0016[<,<:@330,>]
Midas 015> EXTRACT/IMA biai2 = ccd0017[<,<:@330,>]
Midas 016> EXTRACT/IMA dk1 = ccd0031[<,<:@330,>]
Midas 017> EXTRACT/IMA dk2 = ccd0032[<,<:@330,>]
Midas 018> EXTRACT/IMA dk3 = ccd0033[<,<:@330,>]
Midas 019> EXTRACT/IMA ima1 = ccd0052[<,<:@330,>]
Midas 020> EXTRACT/IMA ima2 = ccd0078[<,<:@330,>]
Midas 021> STAT/IMA biai1

frame: biai1  (data = R4)
complete area of frame
minimum, maximum:           184.0000      497.0000
at ( 215,    2),( 147,   408)
mean, standard_deviation:  203.2299      3.752572
3rd + 4th moment:          8402620.      0.1709546E+10
total intensity:            0.3487425E+08
median, 1. mode:            202.0980      204.2529
total no. of bins, binsize: 256           1.227451
# of pixels used = 171600 or 100.00 % of all possible pixels (= 171600)
from      1      1 to    330    520 (in pixels)

Midas 022> STAT/IMA biai2

frame: biai2  (data = R4)
complete area of frame
minimum, maximum:           187.0000      613.0000
at ( 113,    1),( 286,   473)
mean, standard\deviation:  203.2554      3.836732
3rd + 4th moment:          8406459.      0.1710918E+10
total intensity:            0.3487864E+08

```

```

median, 1. mode:          201.6387      202.8706
total no. of bins, binsize: 256      1.670588
# of pixels used = 171600 or 100.00 % of all possible pixels (= 171600)
from 1 1 to 330 520 (in pixels)

Midas 023> READ/DESCR biai1 STATISTIC          !read descriptor statistic

frame: BIAI1  (data = R4)
STATISTIC  :
 184.0000    497.0000    203.2299    3.752572    8402620.
 0.1709546E+10 202.0980    204.2529    256.0000    1.227451
 0.3487425E+08

Midas 024> COMPUTE/IMA ffb = ff-203.          !biais correction
Midas 025> COMPUTE/IMA dk1b = dk1-203.
Midas 026> COMPUTE/IMA dk2b = dk2-203.
Midas 027> COMPUTE/IMA dk3b = dk3-203.
Midas 028> COMPUTE/IMA ima1b = ima1-203.
Midas 029> COMPUTE/IMA ima2b = ima2-203.
Midas 030> AVERAGE/IMA dk = dk1b,dk2b,dk3b      !compute an
                                                     !average of DARK frames
dk1b processed ...
dk2b processed ...
dk3b processed ...
Midas 031> COMPUTE/IMA ffd = ff-dk          !dark subtraction
Midas 032> COMPUTE/IMA ima1bd = ima1b-dk
Midas 033> COMPUTE/IMA ima2bd = ima2b-dk
Midas 034> COMPUTE/IMA ima1bdf = ima1bd/ffd      !flat-field the frame
Midas 035> COMPUTE/IMA ima2bdf = ima2bd/ffd
Midas 036> LOAD/IMA ima1bdf
Midas 037> READ/DESCR ima1bdf LHCUTS

frame: ima1bdf  (data = R4)
LHCUTS  :
 0.0000000E+00  0.0000000E+00 -0.7500000E-01  1.992324  0.0000000E+00
 0.0000000E+00

Midas 038> CUTS ima1bdf 0.,1.992          !modify display cuts
Midas 039> LOAD ima1bdf
Midas 040> READ/ICAT icatalog
Image Catalog: icatalog.cat

No  Name          Ident          Naxis  Npix(1,2)
#0001 ccd0002.bdf  FF D V 20S          2      337  520
#0002 ccd0016.bdf  DK BIAS 1S          2      337  520
#0003 ccd0017.bdf  DK BIAS 1S          2      337  520
#0004 ccd0031.bdf  DK      60S          2      337  520
#0005 ccd0032.bdf  DK      60S          2      337  520
#0006 ccd0033.bdf  DK      60S          2      337  520
#0007 ccd0052.bdf  A0532-527 V 300      2      337  520
#0008 ccd0078.bdf  A1029-459 V 40S      2      337  520

```

#0009 ff	FF D V 20S	2	330	520
#0010 biai1	DK BIAS 1S	2	330	520
#0011 biai2	DK BIAS 1S	2	330	520
#0012 dk1	DK 60S	2	330	520
#0013 dk2	DK 60S	2	330	520
#0014 dk3	DK 60S	2	330	520
#0015 ima1	A0532-527 V 300	2	330	520
#0016 ima2	A1029-459 V 40S	2	330	520
#0017 ffb	FF D V 20S	2	330	520
#0018 dk1b	DK 60S	2	330	520
#0019 dk2b	DK 60S	2	330	520
#0020 dk3b	DK 60S	2	330	520
#0021 ima1b	A0532-527 V 300	2	330	520
#0022 ima2b	A1029-459 V 40S	2	330	520
#0023 dk	average frame	2	330	520
#0024 ffd	FF D V 20S	2	330	520
#0025 ima1bd	A0532-527 V 300	2	330	520
#0026 ima2bd	A1029-459 V 40S	2	330	520
#0027 ima1bdf	A0532-527 V 300	2	330	520
#0028 ima2bdf	A1029-459 V 40S	2	330	520

Midas 041> OUTTAPE icatalog,9-28 TAPE1 !save data on tape

File ff.bdf	written to tape with	241 blocks
File biai1.bdf	written to tape with	243 blocks
File biai2.bdf	written to tape with	243 blocks
File dk1.bdf	written to tape with	241 blocks
File dk2.bdf	written to tape with	241 blocks
File dk3.bdf	written to tape with	241 blocks
File ima1.bdf	written to tape with	241 blocks
File ima2.bdf	written to tape with	241 blocks
File ffb.bdf	written to tape with	241 blocks
File dk1.bdf	written to tape with	241 blocks
File dk2.bdf	written to tape with	241 blocks
File dk3.bdf	written to tape with	241 blocks
File ima1b.bdf	written to tape with	241 blocks
File ima2b.bdf	written to tape with	241 blocks
File dk.bdf	written to tape with	241 blocks
File ffd.bdf	written to tape with	241 blocks
File ima1bd.bdf	written to tape with	241 blocks
File ima2bd.bdf	written to tape with	241 blocks
File ima1bdf.bdf	written to tape with	241 blocks
File ima2bdf.bdf	written to tape with	241 blocks

Midas 042> OUTTAPE icatalog,7-8 ima !save data on disk in FITS format

File ima1.bdf	written to disk> ima0001.mt
File ima2.bdf	written to disk> ima0002.mt

Midas 043> PRINT/LOG !print logfile on the default printer

Midas 042> BYE

```

$ inmidas
Midas 001> CREATE/TABLE flux1 2 30 flux1      ! create table from ASCII file
Midas 002> SHOW/TABLE flux1                  ! list table parameters

Table : flux1                                [Transposed format]
No.Columns :      2   No.Rows      :      30
All.Columns:      3   All.Rows     :      32
Sorted by Sequence   Reference : Sequence
Col.# 1:LAB001          Unit:Unitless      Format:E15.6 R*4
Col.# 2:LAB002          Unit:Unitless      Format:E15.6 R*4
Selection : ALL

Midas 003> READ/TABLE flux1 :LAB001 :LAB002 @1 @9          !read data from table

Table : flux1

Sequence LAB001          LAB002
-----
1   4.71000e+03   1.12850e+04
2   4.77000e+03   1.08300e+04
3   5.05000e+03   8.77000e+03
4   5.09000e+03   8.55000e+03
5   5.12000e+03   8.34000e+03
6   5.19000e+03   7.93000e+03
7   5.23500e+03   7.73500e+03
8   5.26500e+03   7.55000e+03
9   5.39000e+03   7.00500e+03
-----

Midas 004> NAME/COLUMN flux1 :LAB001 "(sec)" F6.0          !change
                                         ! format and define unit
Midas 005> NAME/COLUMN flux1 :LAB002 "(erg/sec)" F6.0
Midas 006> READ/DESCR flux1.tbl HISTORY

HISTORY      :
CREA/TABL flux1 2 30 flux1 NULL TRAN
NAME/COLU flux1 :LAB001 "(sec)" F6.0
NAME/COLU flux1 :LAB002 "(erg/sec)" F6.0

Midas 007> STAT/TABLE flux1 :LAB001          !compute statistics of
                                         !one of the column

Table : flux1
Column # 1 Label :LAB001          Type :R*4
Total no. of entries :      30, selected no. of entries :      30
Miminum value :  0.47100E+04, Maximum value:  0.96800E+04
Mean value   :  0.66437E+04, Standard dev.:  0.15539E+04

Midas 008> $more flux2.fmt

```

```
!
! format file flux2.fmt
!
DEFINE/FIELD 1 6 R F6.0 :LAB001
DEFINE/FIELD 9 13 R F6.0 :LAB002
END

Midas 009> CREATE/TABLE flux2 2 8 flux2 flux2
Midas 010> READ/TABLE flux2
```

Table : flux2

Sequence LAB001 LAB002

1	9705	2750
2	9930	2700
3	10130	2655
4	10150	2650
5	10520	2580
6	10740	2540
7	11040	2485
8	11570	2410

```
Midas 011> MERGE/TABLE flux1 flux2 flux          !merge of two tables
Midas 012> SHOW/TABLE flux
```

```
Table : flux                                [Transposed format]
No.Columns :      2      No.Rows   :      38
All.Columns:      3      All.Rows   :      40
Sorted by # Sequence  Reference : Sequence
Col.# 1:LAB001          Unit:(sec)          Format:F6.0  R*4
Col.# 2:LAB002          Unit:(erg/sec)       Format:F6.0  R*4
Selection : ALL
```

```
Midas 013> SORT/TABLE flux :LAB002          !sort table according to
                                                !increasing values of a column
Midas 014> REGRESSION/POLYNOMIAL flux :LAB001 :LAB002 5  !polynomial fit
```

```
flux
POLYNOMIALS      Input Table : UNION      Type : MUL L-S
N.Cases   :      38 ; N.Ind.Vars   :      1
Dependent variable : column # 1
Independent variable: column # 2      degree : 5

degree
0 4.4818E+04
1 -2.6778E+01
2 7.4671E-03
3 -1.0405E-06
4 7.1309E-11
```

```

5 -1.9126E-15

R.M.S error      : 72.04152

Midas 015> SAVE/REGRESSION flux REGRE          !save the result of
Midas 016> CREATE/COLUMN flux :FIT             !regression in a descriptor
                                         !create a new column
                                         !in the table
Warning: Column overflow mechanism activated
Midas 017> COMPUTE/REGRESSION flux :FIT = REGRE !compute
                                         !the results of the regression
Midas 018> READ/TABLE flux :FIT @1..4,@6,@9

Table : flux

Sequence      FIT
-----
1  1.1337231e+04
2  1.0955642e+04
3  1.0690593e+04
4  1.0505399e+04
6  1.0174703e+04
9  9.7462227e+03
-----

Midas 019> CREATE/GPAPH
!create graphic window
Midas 020> PROJECT/TABLE flux newflux :LAB002      ! project one
                                         !column of a table in a new one
Midas 021> INTERPOLATE/TT newflux :LAB002,:SPLINE flux :LAB002,:LAB001 0.001
                                         !spline interpolation

Midas 022> CREATE/GPAPH
!create graphic window
Midas 023> PLOT/TABLE newflux :LAB002 :LAB001      !plot table columns
Midas 024> SET/PLOT LTYPE=1 STYPE=0               !plot table columns
Midas 025> OVERPLOT/TABLE flux :LAB002 :SPLINE
Midas 026> SELECT/TAB newflux sequence.gt.5       !select part of the table
Midas 027> COPY/TAB newflux result
Midas 028> OUTTAPE result.tbl result.fits        !copy selected table
                                         !save file in FITS
                                         !format on disk

Midas 029> BYE

```

2.3.2 Exit and Logout

To exit from the MIDAS monitor type BYE. You can reenter the MIDAS monitor at any time by typing GOMIDAS (for VMS systems) or \$gomidas (for UNIX systems).

Command	Description
TUTORIAL/ALIGN	Explains the use of the ALIGN command
TUTORIAL/EXTRACT	Demonstrates the extraction of a subimage from a father image
TUTORIAL/FILTER	Displays some of the filtering options
TUTORIAL/FIT	Shows the fitting capabilities
TUTORIAL/HELP	Explains the usage of the HELP command
TUTORIAL/ITT	Shows the effect of various Image Transformation Tables on an image
TUTORIAL/PLOT	Demonstrates the plot package facilities
TUTORIAL/LUT	Shows the effect of various Look-Up Tables on an image
TUTORIAL/PLOT	Demonstrates the plot package facilities
TUTORIAL/SPLIT	Shows the split-screen capabilities of the display
TUTORIAL/TABLE	Demonstrates the table system

Table 2.1: List of Tutorials

2.3.3 Executing System Commands

It is possible to execute commands of the operating system inside MIDAS. This is done by typing a \$ followed by the operating system command you want to have executed. After this command has been finished you can continue your work inside MIDAS.

2.3.4 Some Useful Commands

In table 2.2 you will find a list of some of the most frequently used MIDAS commands. Refer to the detailed command description or the on-line HELP for more details

Command	Description
HELP <i>command</i>	display help for <i>command</i>
HELP/QUAL <i>qualifier</i>	help for all commands with the given <i>qualifier</i> such as IMAGE, TABLE, CATALOGUE, etc.
SET/CATAL	enable cataloging
CREATE/ICAT	create a catalogue of image files
READ/ICAT	list the image frames available
LOAD/IMAGE <i>frame</i>	load and display <i>frame</i> on the image display
MODIFY/LUT	interactively change the look-up table
COMPUTE/IMAGE	perform image arithmetic
EXTRACT/TRACE	interactively extract a line from an image
PLOT/TRACE	plot the extracted line
ZOOM	zoom an image on the cursor
TUTORIAL/ <i>tutorial</i>	execute one of the existing <i>tutorials</i>

Table 2.2: List of Often Used Commands

Chapter 3

Monitor and Command Language

This chapter is organised as follows:

- In the first two sections we describe how to start MIDAS and how the host operating system and MIDAS coexist.
- Section 3 explains the different data structures used in MIDAS and how to access them in a MIDAS session.
- Section 4 describes the syntax of the MIDAS commands, as well as the editing and recalling of commands and also the on-line HELP facility in MIDAS.
- In section 5 you will find some details about how the MIDAS commands are executed.
- Then follows the largest and most detailed section (section 6), which gives in-depth information about the MIDAS command language (MCL). With MCL you can write high level MIDAS “programs” which are called MIDAS procedures.

The topics include:

- the MCL commands
- passing parameters in MIDAS procedures
- symbol substitution
- loops and conditional branching
- special functions

- Section 7 introduces the MIDAS contexts.
- Section 8 explains how to run application programs written in FORTRAN or C inside MIDAS. It also shows how to debug these programs as well as MIDAS procedures.
- All the commands related to MIDAS catalogs are listed in section 9, together with an example of how to use catalogs in MIDAS procedures.

- The MIDAS login procedure and MIDAS user levels are the topics of the last two sections.

The MIDAS directory tree structure is not covered in this chapter. For those interested, please refer to the *MIDAS Environment document*.

3.1 Starting the MIDAS Monitor

In order to get properly initialised, MIDAS needs the following information

- the MIDAS user mode to work in: *Parallel* or *Single User* mode, the default is *Single User* mode
- the MIDAS working directory for internal files and (optionally) private procedures: the default is a subdirectory `midwork` in your login directory, i.e. `SYS$LOGIN:[MIDWORK]` in VMS or `$HOME/midwork` in Unix. This directory does not have to be (or even should not be) the same directory where your data files are stored.

On a VMS system type `SETMIDAS` and follow the dialog if you want to change the defaults for the mode and the MIDAS work directory.

On a Unix system use the option `-m <mid_work>` to change the default working directory, and `-p` or `-P` to run MIDAS in parallel mode when you start MIDAS via the command `inmidas` (see below).

Two other variables are very important to MIDAS - `MIDVERS`, which holds the MIDAS version you use at your site, and `MIDASHOME`, the root directory for the MIDAS system code. These variables should have been set up correctly by your system manager when MIDAS was installed or use again `SETMIDAS` in VMS; for Unix these variables can again be chosen within the `inmidas` command. There are many more options for the `inmidas` command in Unix, which can be accessed interactively via the man page of `inmidas` (a complete Midas installation should also include the setting up of the man pages for `inmidas`, `gomidas` and `helpmidas`).

The detailed command description is as follows:

SYNOPSIS

```
inmidas [ unit ] [ -h midashome ] [ -r midvers ] [ -d display ]
[ -m mid_work ] [ -p/-P/-nop ] [ -noh ] [ -j midas-command-line ] [ -help ]
```

Without arguments, `inmidas` initiates a MIDAS session with default definitions. Some of these definitions can be modified with arguments in the command line of `inmidas` or by environment variables. The arguments in the command line override the corresponding environment variables.

OPTIONS

`inmidas` has been configured by the Midas system manager at installation time to start a specific release of MIDAS. However, alternative releases

can be specified using the command line arguments:

-h midashome
Home directory for MIDAS. Absolute pathname containing, at least, one release of MIDAS. It may also contain subdirectories for demo and calibration data.

-r midvers
Release of MIDAS to be executed. It must be a subdirectory under midashome.

-d display
Specifies another X server for the display and graphical MIDAS windows
NOTE: be aware of allowed access to a remote X server using the "xhost" command.

-p/-P/-nop
Options -p and -P set the MIDAS environment variable MIDOPTION to PARALLEL while option -nop sets it to NOPARALLEL (default: NOPARALLEL). In NOPARALLEL mode all intermediate MIDAS files in the MIDAS startup directory are deleted when starting MIDAS via inmidas. In PARALLEL mode no intermediate files are deleted, and this is necessary to run several MIDAS sessions with the same startup directory. With -P option and if unit is not given the system will select automatically one free unit for you. With -p option and no unit, the user will be requested to enter one.

unit
Unit to be associated to the MIDAS session (default: 00 only if MIDAS is working in NOPARALLEL mode). Valid values for this option are in the range (00, 01, ..., 99, xa, ..., zz) where numerical values indicate that the user is working in an X11 environment (DISPLAY environment variable or argument -d should be given), and the others indicate an ASCII terminal.

-m mid_work
Specifies the MIDAS startup directory (default: \$HOME/midwork).

-noh
Starts MIDAS without clearing the terminal and without welcome message.

-j midas-command-line
midas-command-line will be executed in MIDAS as if it were the first command line typed in the MIDAS monitor.
This option sets also the -noh option.
NOTE: midas-command-line should be typed between single quotes to be interpreted by inmidas as a single argument and to be passed to the MIDAS monitor as it is.

-help
Display this help page.

So, to start MIDAS, type **INMIDAS** on a VMS system or **inmidas [arg1] ...** on a Unix system. This will initialize the MIDAS monitor as follows:

- In VMS the logical name **MID_WORK** is assigned to the MIDAS working directory; in Unix the environment variable **MID_WORK** is set accordingly. If the working directory does not yet exist, it is created.

All internal files created by the MIDAS monitor will be stored in the MIDAS working directory. This is also the place to store your own **login.prg** as well as all your other MIDAS procedures which you want to execute from any other directory.

- In *Single User* mode, all MIDAS log- and keyfiles (FORGRxy.LOG, FORGRxy.KEY - where **xy** is the MIDAS unit described below) which exist in the MIDAS working directory as well as all MIDAS internal files are deleted.

In *Parallel* mode no files are deleted.

- In VMS the user process is renamed to **MIDASxy**

- In VMS and in parallel mode in Unix you will be asked to enter the identification of a MIDAS unit as a two-character (case insensitive) string.

Units are in the range (00, 01, ..., 99, xa, ..., zz) where numerical values indicate that the user is working in an X11 environment (DISPLAY environment variable should be set), and the others define a MIDAS session with no image display capabilities.

So 23, xa, yf or Z3 will all be valid units. If you work in *Parallel* mode you have to use different MIDAS units for each session because the MIDAS unit is appended to the names of all MIDAS internal files.

On startup the current MIDAS version and the computer and operating system you are using are displayed together with a copyright notice. Then the prompt string

Midas 001>

appears on the terminal screen and you are ready to execute any of the available MIDAS commands.

The internal MIDAS files all reside in the MIDAS working directory (**MID_WORK**), the data files are taken from the current working directory unless the complete file specification is given in the data file name.

Since MIDAS executes its applications in a child process (subprocess for VMS) which leaves no traces after termination, you cannot simply use the host command **SET DEF** (VMS) or **cd** (Unix) to change the working directory once you are in a MIDAS session. Instead, use the MIDAS command **change/direc** for that purpose.

Another possibility is to set the search path for your data files via the command **set/midas_system DPATH=directory**. Use the MIDAS Help Utility for obtaining detailed

information about these commands, e.g. `HELP change/direc`.

MIDAS is a case insensitive system. That means, you can type your input with upper or lower case characters. There are, however, some pitfalls with respect to the data files that reside in the local file system. In VMS, the system automatically translates all file names to upper case, so `LOLA.BDF` and `lola.bdf` specify exactly the same file. In Unix, file names may be specified using lower and upper case, so `LOLA.BDF` and `lola.bdf` are two different files. The convention in MIDAS is to always use lower case file names (e.g. in tutorial procedures) to guarantee portability between VMS and Unix. Also, all default file types are specified in lower case, e.g. `.bdf` and `.tbl` for images and tables.

Note

All MIDAS commands in the following sections are printed with capital letters. This is just for reasons of readability, i.e., to highlight them. The commands could all be typed in lower case as well.

3.2 MIDAS And the Host Operating System

Care has been taken that MIDAS and the Host Operating System (DCL for VMS and Bourne or C-shell for Unix) co-exist smoothly and complement each other. Migration from one environment to the other is simple:

If you are in the MIDAS environment, type `BYE` to switch back to the Host System.

If you have returned to the host environment from a MIDAS session, (indicated by the \$-prompt in VMS, and by \$ or % in Unix), type `GOMIDAS` (in VMS) or `gomidas` (in Unix) to revive MIDAS. The status of the keywords and the command buffer of the stopped MIDAS session are preserved - if you want to start afresh, use `INMIDAS` (VMS) or `inmidas` (Unix) again.

You may also use host commands directly inside MIDAS by preceding them with '\$'. For instance,

`Midas 027> $DIR` (in VMS) or

`Midas 027> $ls` (in Unix)

will display the contents of the current directory.

Please, note, that currently this mode of operation will only invoke Bourne shell commands in Unix, not C-shell or Kornshell commands. To execute C-shell (or any other Shell) commands you have to insert them in a Bourne shell script which has as the first line: `#!/bin/csh`, or: `#!/bin/ksh`, etc.

Note

If you work on a VMS system, beware of DCL command procedures:

DCL modifies command I/O streams when executing a procedure. This causes problems for the interprocess communication inside MIDAS. When executing a DCL procedure via \$ @ 'procedure' the correct settings will be maintained inside MIDAS.

*However assigning a symbol **MIMI** to the command above and then executing the DCL procedure by just typing **\$ MIMI** will lead to disaster from which only a **BYE** and subsequent **GOMIDAS** will get you going again.*

Since images, tables, etc. are standard disk files, all host commands related to file operations can be employed. However, if a MIDAS catalog is used, care has to be taken that the information in the catalog is not invalidated, when e.g. renaming or deleting data files outside MIDAS (i.e. using commands of the host file system directly).

The output from MIDAS commands can be redirected to ASCII files enabling easy combination of MIDAS and host commands. E.g.

Midas 123> READ/DESCR myimage * >dsc.dat

will send all the output from the **READ/DESCR** command to the ASCII file *dsc.dat* (created in your current work directory) which can then be used by any host command. For example,

Midas 124> \$EDIT dsc.dat (in VMS) or

Midas 124> \$vi dsc.dat (in Unix)

Note

*This mechanism is pretty much like the one used in Unix with the exception that there should be no space between the **>** and the output file name. Furthermore, this output redirection also works on VMS.*

Midas 125> STATISTICS/IMAGE myimage >dsc.dat

always creates a new file *dsc.dat*, if you want to append data to an existing ASCII file use

Midas 126> STATISTICS/IMAGE myimage >>dsc.dat

instead.

As you may have guessed already, there is also input redirection. E.g.

Midas 127> \$ls a*.bdf >dscin.dat

Midas 128> READ/DESCR <dscin.dat

will display the standard descriptors of all Midas images with names beginning with the letter 'a' in the current directory.

Again there should be no space between **<** and the file name.

See also the subsection 3.4.5 for more info about I/O redirection.

On a Unix system you can connect MIDAS and Unix command via the *pipe* symbol , e.g.

Midas 129> READ/DESCR myimage * | \$grep NGC425

Midas 130> \$ls a*.bdf | read/descr

3.3 MIDAS Data Structures

Here we describe and discuss the various data entities (structures) that MIDAS recognizes. They are stored in an internal binary format, accessible only through MIDAS and fall into the following categories:

Images are a set of data of same physical significance in one to three dimensions. The data must be sampled with constant step size along all 1, 2 or 3 axes and are stored in different formats, e.g., as bytes, 16 bit integers, or 32 bit reals on disk. However, most MIDAS applications work on real data, so the image pixels are converted on the fly to real format if necessary. The default file type is **.bdf**.

Tables are a structure for handling data which can be arranged in rows and columns. The data may be of numerical or character type. Numerical data may be sampled in any arbitrary fashion. The default file type is **.tbl**.

Fit-files are “degenerate” image files with just descriptors and no pixels and used to store the parameters needed for the fitting functions. The default file type is **.fit**.

Descriptors are variables attached to the structures mentioned above (i.e. stored in the same file) and describe the structure of the tables, images and fit files. They can also store any other auxiliary information connected to the data such as histograms, coordinates, comments and so on. For fit files they contain the fitting parameters.

Catalogs contain lists of either images or tables or fit files for the purpose of grouping data together within MIDAS. They are exceptional in the sense that they are implemented as ASCII files so you can list and edit them (with care!) outside MIDAS. The default file type is **.cat**.

Keywords are variables which can be used to pass information from one MIDAS program to the next or to temporarily store intermediate results (there are also *reserved* or *system* keywords that keep MIDAS system parameters). They are referred to by a name and can be easily manipulated from the terminal or MIDAS procedures.

The individual data points in an image are referred to as “pixels” and in a table they are called “elements”. The paragraphs below describe the structure of descriptors, and keywords, and the methods for specifying the individual data values in images and tables.

Note

*There is no special syntax for file names in MIDAS. You can use any legal name of your host file system for images, tables and fit files. However, a name beginning with a digit or using any of the characters +, -, *, /, ‘, !, |, (and), should be avoided, because these symbols will cause problems in e.g. the COMPUTE/IMAGE command. If you do want to use a file name with these special characters in a COMPUTE/IMAGE command you have to*

enclose the full name in quotes, i.e.

Midas 456> COMPUTE/IMAGE res = 12 + "quasar01+12.bdf"

The length of these names is, in principle, limited to 60 characters for MIDAS applications (which is the size of keywords *IN_A*, *OUT_A*, used in most procedures to store the image, table names...).

Also file names like *abc.bdf.mine* will not be appreciated by all MIDAS applications.

As mentioned before, file names are case sensitive in MIDAS on Unix systems; names for descriptors and keywords are not. Thus, referring to a keyword with name *KEYA* may be done e.g. via *keyA* or *Keya*.

3.3.1 Specifying a Descriptor

Descriptors have been derived from the concepts used in a FITS file header and have many similarities with the FITS keywords. In particular the names of the MIDAS standard descriptors, e.g. *NAXIS*, *NPIX*, etc., (for details see Appendix “Standard Descriptors” of the MIDAS Environment doc.) correspond to those in the FITS header.

Descriptors come in four flavours: *integer*, *real*, *double precision* and *character*. Mixed types are not possible, i.e. you cannot have a real descriptor TEST and an integer descriptor TEST at the same time. Each descriptor also has a name (max. 15 chars.) and a length (no. of elements). Writing values into positions beyond the current length leads to an automatic extension of the descriptor (and update of its length) just as a text file is extended by the “editor” when you are editing it.

Beware, that you do not create ‘holes’ by writing to descriptor elements which are not immediately following the current last element! MIDAS will not initialize the descriptor elements in between, so their values are unpredictable.

The command to write values into a descriptor requires the name of the data file (which could be an image, table or fit file), the descriptor name, the descriptor type, the first element to be accessed, and the total number of elements to be transferred (all separated by a ‘/’ (slash)). Finally, the data values are given (separated by commas for numeric data, but no spaces). For example,

WRITE/DESCR imgfile Descname/C/1/7 Anyname

would write the ASCII string *Anyname* into the character descriptor *Descname* associated with the data file *imgfile.bdf*. Since spaces serve as parameter delimiters in MIDAS they have to be enclosed by double quotes (“) if used as data. So

WRITE/DESCR imgfile Descname/C/1/7 " "

would fill *Descname* with 7 blanks.

WRITE/DESCR imgfile Descname/R/4/3 17.3,8.8E2,-.3

would write the numbers 17.3, 880.0, -0.3 into elements 4,5 and 6 of real descriptor `Descname`. If the descriptor were created with fewer than 6 elements it would be expanded automatically.

```
WRITE/DESCR tblname.tbl Descname/R/4/3 17.3,8.8E2,-.3
```

would write the numbers 17.3, 880.0, -0.3 into elements 4,5 and 6 of real descriptor `Descname` of the table file `tblname.tbl`.

Note, that we had to add the file extension `.tbl` to the name `tblname`, since the command `WRITE/DESCR` defaults the first parameter to an image and appends the file type `.bdf` if none is given by the user.

Single descriptor elements can also be written in a more direct way, via `frame,descr = value`, e.g. to set `STEP(2)` of image `lola.bdf` to 1.234, use `lola,step(2) = 1.234`

The *value* can also be an *expression* made up of constants and elements of any MIDAS data structure, see the subsection 3.6.2.

This is how descriptors work at the most basic level. However, in many cases, higher level commands have been implemented to update specific descriptors. The MIDAS command `CUTS`, which sets the high and low cuts of an image (in descriptor `LHCUTS`) for displaying or plotting it, is an example of this.

Some of the commands dealing with descriptors are:

```
READ/DESCR, WRITE/DESCR, SHOW/DESCR, DELETE/DESCR, INFO/DESCR, COPY/DD.
```

An optional help text can be attached to each descriptor and is then displayed via the `READ/DESCR` and `SHOW/DESCR` commands. This text is copied from the original FITS file (if existing) when reading in the data file or can be explicitly set via `WRITE/DHELP`.

3.3.2 Specifying Keywords

As is the case for descriptors, keywords also have a name (max. 8 chars.), a type and a length (i.e. no. of elements). However, this length is fixed, and once the keywords are created with a certain size, they cannot be extended. The possible types for keywords are: *real*, *integer*, *character* and *double precision*. Like for descriptors there are no mixed types possible.

In order to write a value to a keyword, the same format as for descriptors is used.

```
WRITE/KEYWORD INPUTC/C/1/8 AKeyword
```

This command would write the ASCII string `AKeyword` into the character keyword `INPUTC` and

```
WRITE/KEYWORD AZTEC/I/1/2 17,-22
```

would write the values 17 and -22 into the integer keyword AZTEC (elements 1 and 2).

Single keyword elements can also be written in a more direct way, via
`key = value`, e.g. to set the 12th element of keyword INPUTR to 1.234, use
`inputr(12) = 1.234`

The *value* can also be an *expression* made up of constants and elements of any MIDAS data structure, see the subsection 3.6.2.

Some of the commands dealing with keywords are:

`READ/KEYWORD`, `WRITE/KEYWORD`, `SHOW/KEYWORD`, `DELETE/KEYWORD`, `COPY/KEYWORD`, `COMPUTE/KEYWORD`. Keywords and descriptors can be copied to each other via `COPY/KD` and `COPY/DK`.

3.3.3 Specifying Elements in a Table

The MIDAS table file system is described in detail in chapter 5 of this volume. Here we just explain briefly how to access the various elements in a table file. To do so, it is necessary to specify the *table name*, the *column* and the *row*. This is done in the following format (order):

`table column row`

where

table is the table name;

column is the desired column which can be referenced by label as `:col` or by sequence number as `#n`;

row is the desired row referenced by number as `@n` or by a value in a predefined reference column.

Like descriptor and keyword names, column labels are case insensitive. The command:

```
READ/TABLE tname #3 @10
```

would display the element in column 3 of row 10 in table file `tname.tbl`.

Similarly, the command:

```
READ/TABLE tname :MAGNITUDE 20.0
```

would access the element in the column labeled 'MAGNITUDE' and value 20.0 in the reference column (this reference column must have been defined before via the `SET/REFCOLNUM` command).

Note, that we need not specify the file extension `.tbl` as in the descriptor related commands. All table commands default the data files to tables with extension `.tbl`.

Single table elements can also be written in a more direct way, via

`table, column, row = value`,

e.g. to set the element in the 3rd row of the column labeled `:XREF` in the table `lola.tbl` to 1.234, use

```
lola,:xref,@3 = 1.234
```

The *value* can also be an *expression* made up of constants and elements of any MIDAS data structure, see the subsection 3.6.2.

Some of the commands dealing with tables are:

READ/TABLE, WRITE/TABLE, COMPUTE/TABLE, SHOW/TABLE, EDIT/TABLE, COPY/TABLE.

3.3.4 Specifying Pixels in an Image

In some commands it is necessary to specify the columns and rows of an image to which that command should refer.

This is done in the following way for e.g. a 2-dim frame: `frame[x1,y1:x2,y2]` where the column specification, `x` or the row specification, `y` can be any of

- world coordinates, indicated via real or integer numbers: 20.0,300
- pixel numbers, indicated via integers preceded by `@`: `@35,@200`
- or a special symbol to indicate *start* (`<`), or *end* (`>`) of a row or column; thus `[@20,<:@20,>]` specifies the complete 20th column of a 2-dimensional image

World coordinates are the physically meaningful coordinates with units such as wavelengths or arc seconds (which are defined in the descriptor CUNIT). Pixel numbers (starting with 1 for each dimension) are the indices of an image seen as an array.

For example, extracting the complete 12th plane from the 3-dim image stored in `cube.bdf` is done via

`EXTRACT/IMAGE plane12 = cube[<,<,@12:>,>,@12]`

Single pixels can also be written in a more direct way, via

`frame[x,y,z] = value`, e.g. to set the pixel in row 27 and column 1023 of the 2-dim image `lola.bdf` to 1.234, use `lola[@1023,@27] = 1.234`

The *value* can also be an *expression* made up of constants and elements of any MIDAS data structure, see the subsection 3.6.2.

Some of the commands dealing with images are:

READ/IMAGE, WRITE/IMAGE, COMPUTE/IMAGE, STATIST/IMAGE, DELETE/IMAGE, COPY/II.

3.3.5 Specifying Sub-Image

In all MIDAS commands which accept images as parameters you can also provide subimages in the syntax specified above.

For example, the command `LOAD/IMAGE` loads an image into the MIDAS display (window). Then,

`Midas 567> LOAD/IMAGE mygalaxy`

would try to display the whole 2-dim image `mygalaxy.bdf`, whereas

`Midas 568> LOAD/IMAGE mygalaxy[@100,@200:@199,@299]`

would only load a 100x100 subimage of `mygalaxy.bdf` beginning at the 100th x-pixel and 200th y-pixel.

3.4 Command Syntax

After start-up, the MIDAS monitor prompts you to interact with MIDAS by entering commands on the terminal which are terminated by **Return**(Enter key).

To enter a command-line on more than one terminal/window line, use the continuation character, a minus sign (-), as the last character of a line. The command line is limited to 256 characters. If you want to enter more than one command on a single line, separate the commands with a semicolon and space (;).

A MIDAS command line is structured as

```
command/qualifier par1 par2 ... par8 !comments
```

The command describes the general action you want to perform (a verb) and the qualifier usually specifies the object of that action, e.g. **WRITE/DESCRIPTOR**. The command parameters (max. 8) hold all other information needed to perform the required action. All parameters are separated by spaces.

Currently the following “objects” exist in MIDAS:

- keywords
- descriptors
- bulk data frames
 - images
 - tables
 - fit files
 - ASCII files
- catalogs
 - for images
 - for tables
 - for fit files
- auxiliary image–display data structures (where applicable)
 - LUTs (Colour Look–Up Tables)
 - ITTs (Intensity Transfer Tables)
 - (internally these structures are stored in MIDAS table files)

All user input and output from MIDAS commands is recorded in an ASCII file, named **FORGRxy.LOG** (with **xy** the MIDAS unit) and stored in the **MID_WORK** directory. This MIDAS logfile serves as a hardcopy of a full MIDAS session. Terminating MIDAS with

BYE and continuing later on via `gomidas` will not restart a new logfile but append to the existing one. The logging in MIDAS can be controlled via commands like `LOG/OFF`, `LOG/ON`, `LOG/TOF`, `DELETE/LOG`.

Comments may be appended to the command string and are separated by at least one white space and ‘!’ (exclamation mark) from it. To give a complete line of comments, enter ‘!’ as the first character of the input line (useful for structuring the contents of the MIDAS logfile).

Commands and qualifiers may be abbreviated to the number of significant characters needed to distinguish them from the rest. At most 6 characters are necessary for the command and 4 characters for the qualifier. Command and qualifier are separated by a ‘/’ (slash). Nearly all commands need a qualifier, but there is only one qualifier per command (e.g. `comm/qual1/qual2` is unsupported in MIDAS). In case you omit the qualifier, the default qualifier of that command is used by MIDAS. The default qualifier of a MIDAS command may be displayed via `SHOW/COMMAND` command. For example, the default qualifier for the `LOAD` command is `IMAGE`, so typing `LOAD/IMAGE` or `LOAD` will have the same effect.

The parameters depend on the actual command. A space (blank) is the delimiter for parameters in the command-line. Commas are used to subdivide parameters. If you need a space inside a parameter, this parameter has to be enclosed in double quotes (“”). Normally, parameters are position dependent, i.e. `par1` is the first, `par2` the second, and so on. This may be overridden by using the following syntax:

```
command/qualifier P4=par4 P1=par1 P7=par7 ... !comments
```

If the command procedure which is activated by a MIDAS command uses the `CROSSREF` command, it is also possible to execute that command via:

```
command/qualifier label4=par4 label1=par1 label7=par7 ... !comments
```

The help text of each command specifies whether such a cross referencing of parameters is possible and if so, which labels to use. For details about the command `CROSSREF` see the description of it in the section below on MIDAS procedures.

Whenever possible parameters have defaults. If you do not want to override them use the symbol ‘?’ (question mark) for a parameter if you use the position dependent format. Therefore,

`command/qualifier P4=22.345` is equivalent to

`command/qualifier ? ? ? 22.345`

The preset default values of MIDAS commands can be overridden with `CREATE/DEFAULTS`. For example, the default for the descriptor name(s) in the `READ/DESCR` command is `NAXIS,NPIX,START,STEP,IDENT,CUNIT,LHCUTS` for an image frame. Thus, typing

`Midas 234> READ/DESCR myimage`

will display the contents of the descriptors `NAXIS`, `NPIX`, `START`, `STEP`, `IDENT`, `CUNIT` and `LHCUTS` of image `myimage.bdf`. After changing this default via

`Midas 235> CREATE/DEFAULTS READ/DESCR ? HISTORY`

the same command

Midas 236> READ/DESCR myimage

will display the contents of the descriptor HISTORY of myimage.bdf.

To abort a MIDAS command, use **Ctrl/C** (sometimes you also have to hit **Return**), which will return control to you.

Note

Be careful when aborting commands which interact with a display/graphics window in the X11-Environment. For, you run the risk of losing the synchronisation with the MIDAS display server, which must then be re-initialized via RESET/DISPLAY (see chapter 6 for details).

3.4.1 Command Recalling

By default, the last 15 commands entered on the terminal are kept in an internal buffer (the no. of commands saved can be changed via SET/BUFFER). To recall (and execute) any of these commands, simply type the associated command number. This is the number "xyz" appearing in the prompt Midas xyz> when that command was entered. To display the command buffer, simply hit **Return**.

If you want to recall more than one command at once, enter all the relevant command numbers (separated by a semicolon and space), e.g. enter 14; 17; 22 if you want to repeat the commands numbered 14, 17 and 22. Also 14; read/keyw in_a; 17 is possible.

To recall commands not by number but by pattern, use :pattern to repeat the last command matching the specified pattern. For example, if the last two commands in your command buffer are:

```
22 READ/IMAGE supernova
23 show/commands
```

Then, typing 22 as well as :READ or :nova will execute the command READ/IMAGE supernova again. Note that for the pattern matching MIDAS *does* make a distinction between upper and lower case.

You can also use the vertical arrow keys to navigate up and down through the command buffer.

Besides repeating complete input lines it is also possible to just use parts of the last command line. Each "token" of the last command line is saved internally until the next input. A "token" is the information separated by spaces in the command line. To repeat the tokens on a subsequent command line merely type a '.' For example, if you have in the command buffer:

```
READ/KEYWORD in_a
LOAD/IMAGE myframe 0 2,2
```

Then typing ‘. yourframe . .’ as the next command is equivalent to typing ‘LOAD/IMAGE yourframe 0 2,2’.

All features described so far apply to genuine MIDAS commands as well as to host system commands (where the first character of the command line is the \$ sign).

Some words of caution:

In VMS the version number of files may be specified using a semicolon, e.g.

```
$ RENAME file.typ;7 lola.bdf.
```

Typing such a command inside MIDAS will not work, since the monitor will interpret this input as two Midas commands. Instead, use a dot to separate the version number, e.g.

```
Midas 234> $ RENAME file.typ.7 lola.bdf.
```

In Unix the repetition of tokens may cause trouble. Consider the following:

```
Midas 123> load/image vaca
```

```
Midas 124> $cp /elsewhere/toro.bdf .
```

The intention was to simply copy the file toro.bdf from somewhere else to the current directory. But instead of toro.bdf you will find a strange file named ? in your directory... In the line ‘123’ only two tokens are entered, so all other 8 tokens are set to the default value ‘?’ . In line ‘124’ the third token will be set to the third token in the line above, so it changes to:

```
Midas 124> $cp /elsewhere/toro.bdf ?
```

Instead, specify also the result frame completely, e.g.

```
$cp /elsewhere/toro.bdf toro.bdf.
```

Preceding a host command by \$\$ disables the interpretation of specific symbols by MIDAS, thus

```
Midas 124> $$cp /elsewhere/toro.bdf .
```

will actually do the expected copy.

3.4.2 Command Line Editing

The commands in the internal command buffer may also be edited.

On Unix systems MIDAS comes (since the 94NOV release) with two different line-editors: *Term Windows* (the one developed at ESO) and *readline* (which is from the GNU project). The default line-editor is *readline* which provides a history stack of commands, *emacs* or *vi* editing functions, command and filename completion functions, and a communication channel to the MIDAS GUI Xhelp for the On-Line Help utility. See the man page of *readline* for a complete list of options.

To use the old *Term Windows* line-editor just set the environment variable TERMWIN to yes. Please, note, that the *Term Windows* editor will be phased out in a future MIDAS release.

On VMS systems MIDAS comes only with the *Term Windows* line-editor.

To edit a MIDAS command, type the command number preceded by a dot (period) or followed by a dot. So ‘.xyz’ or ‘xyz.’ will both display the command ‘xyz’ and put

you into the edit mode where you can modify that command.

If you employ the pattern matching style, use ‘.:pattern’ or ‘::pattern’ to edit a previous command containing that pattern.

You edit that line using the *arrow* keys and *delete* key of the keyboard and retyping the characters. On VMS systems to toggle between *Replace* and *Insert* mode use **CtrlA**.

Using ‘.xyz’ (‘.:pattern’) will lead to the creation of a new MIDAS command with new command number, whereas ‘xyz.’ (‘::pattern’) modifies the specified command directly (and keeps this command number).

Also the commands recalled via the arrow keys can be edited.

As mentioned above only the 15 most recently used commands are kept in the command buffer on a first-in, first-out basis. So if you repeat or edit a certain command via its MIDAS command-number at least once in 15 command inputs, this command will always be kept in the buffer.

However, you may wish to make sure that a command remains always in the buffer. Entering ‘xyz/LOCK’ will lock the command with number ‘xyz’ in the buffer; to unlock the command, use ‘xyz/UNLOCK’.

The command **CLEAR/BUFFER** empties the command buffer and resets the command counter to 1. Since 3 digits are used for the command count, the counter is also reset to 1 after MIDAS command no. 999.

3.4.3 Command Line Suspension

If, while entering a command, you realize that you forgot the full command syntax or want to check something else, a mechanism has been introduced to let you interrupt the command line, execute another command or commands, and then resume with the interrupted line. To interrupt a command line enter ‘\’ (back-slash) as the last character and hit **Return**. The command string is then saved internally. To resume entering the interrupted command line, type ‘\’ (back-slash) again followed by **Return**. The saved command line will be displayed on the terminal and you may add more input.

Note

You cannot edit or change the saved portion of the command after reentering the interrupted string, since your new input is handled as if it were a continuation of the original command line.

3.4.4 On-Line Help

The help facility of MIDAS (command **HELP**) provides detailed descriptions of all supported commands and qualifiers. This applies also to the **HELP** command itself.¹

If you work in an X-Window environment we suggest to use **XHelp**, the graphical user interface to the MIDAS Help facility, by executing the MIDAS command **CREATE/GUI HELP**. Besides providing a separate and convenient interface to the MIDAS Help utility this GUI also supports a *feedback* facility for reporting errors, problems and suggestions

¹All MIDAS commands are described in detail in Volume C of this User Manual.

to the MIDAS group at ESO (this problem-report mechanism is based on the GNATS system from the GNU project).

And if you use the *readline* line-editor you can also update the MIDAS Help facility with the current command-line by typing **Ctrl/X** or **F1**.

Also, the TUTORIAL commands will help you in exploring the MIDAS system. Use the MIDAS command **HELP TUTORIAL** to find out which tutorials exist and try them out.

There is also a tutorial about the **HELP** command itself. Use **TUTORIAL/HELP** to exercise many of the features described in this section.

Some of the other **HELP** features are given in the following table.

Command	Description
HELP	To display all currently existing MIDAS commands and topics
HELP <i>comnd</i>	To display all the <i>comnd/qualif</i> combinations available for the given <i>comnd</i>
HELP <i>comnd/qualif</i>	To get detailed information about the specified <i>comnd/qualif</i> combination
<i>comnd/qualif ??</i>	To display full command syntax (one line) of specified <i>comnd/qualif</i>
<i>pattern?</i>	To list all commands which begin with given <i>pattern</i>
HELP/QUALIF <i>qualif</i>	To list all commands which may use the given <i>qualif</i>
HELP/SUBJECT <i>subject</i>	For any info related to <i>subject</i>
HELP/CL <i>comnd</i>	To get detailed information on the MIDAS command language <i>comnd</i>
HELP/KEYWORD <i>keyword</i>	To get detailed information about the specified <i>keyword</i>
HELP <i>[Topic]</i>	To get information about a <i>topic</i> , e.g. the standard descriptors or available contexts
HELP/APPLIC	To get information about available application procedures
HELP/CONTRIB	To get information about available contributed procedures

Table 3.1: Help Features

Normally, the **HELP** text is not written to the MIDAS logfile (to save space in the logfile) but if you wish to include this, you may do so by setting keyword **LOG(3)** to 1 via e.g. the MIDAS command **LOG(3) = 1**. To print out the help text, use **PRINT/HELP**.

3.4.5 Input/Output Redirection in Midas

Similar to Unix MIDAS commands also accept input from an ASCII file instead of an input line typed on the terminal. For example, you can use host commands to create a list of files and then execute a MIDAS command on all files in that list:

Midas 045> \$ ls bb*.bdf >input.dat

Midas 046> STATISTICS/IMAGE <input.dat

This command sequence will execute the STATISTICS/IMAGE command for all frames in the file *input.dat*, i.e. all images beginning with 'bb'.

For more elaborate processing of groups of files you should use MIDAS catalogs and the STORE/FRAME command (see section 3.9 for detailed info about catalogs).

Different ways exist to save the output of a MIDAS command in an ASCII file besides just using the logfile. There is a set of PRINT/... commands to print out the contents of the different MIDAS data objects, e.g. PRINT/TABLE. By default the output from these commands is sent to a printer but this can be changed to a file via the ASSIGN/PRINT command. So, if you want an ASCII copy of a MIDAS table, you do:

Midas 088> ASSIGN/PRINT file mytable.dat

Midas 089> PRINT/TABLE mytable

Another possibility which is applicable to all MIDAS commands, not just the PRINT commands, is to specify the output file directly in the command. Thus,

Midas 089> WRITE/TABLE mytable >mytable.dat

is equivalent to the two commands above. The file *mytable.dat* is created in the current directory. If you want to append the data to an existing file, use

Midas 090> WRITE/TABLE mytable >>mytable.dat

Midas 090> WRITE/TABLE mytable >terminal

only produces output on the terminal, i.e. it's the same as

Midas 090> WRITE/TABLE mytable

but is useful if you want to provide optional file output in a procedure (see section 3.6.2 for more details about that).

In case, the output should go to a file and also be displayed in the MIDAS command window, use

Midas 091> WRITE/TABLE mytable >mytable.dat+terminal

If you want to suppress the output completely use the special name Null for the output file, e.g.

Midas 092> WRITE/TABLE mytable >Null

will omit all output. No file Null is created.

Currently it is not possible to redirect the input as well as the output in the same command line, e.g.

Midas 046> STATISTICS/IMAGE <input.dat >output.dat

is not possible.

Note

No space should be between the <, > or >> and the in/output ASCII file

names!

This output redirection scheme has been modeled after the way it is done in Unix but it also works on VMS systems.

3.5 Execution of Commands

MIDAS commands fall into two categories: the *basic* commands and all other *application* commands. The *basic* commands are executed inside the MIDAS monitor, which is the program you are interacting with. All other commands are implemented by executing a MIDAS procedure which runs one or more programs in a subprocess (child process). During the time a command is being processed in the subprocess, the MIDAS monitor is suspended until the corresponding program terminates in the subprocess. Only then control is returned to the user. To stop a command prematurely, type **Ctrl/C**.

Since process creation is much more expensive in VMS than in Unix these subprocesses are handled differently in VMS and Unix:

In VMS, the subprocess, named FORGRxy (with *xy* the MIDAS unit specified at start-up), is created at MIDAS initialisation time and kept alive until you exit from MIDAS via the command **BYE**.

In Unix, the child process is created each time the MIDAS command executes an application program. Upon termination of that program the child process dies. This also applies to commands of the host system – they are executed in a subshell.

Therefore, issuing ‘\$ cd /elsewhere’ inside MIDAS does not change your current directory permanently...

See also section 3.1 about the *change-directory* problem.

Some internal files are created when starting a MIDAS session in the directory specified via **MID_WORK:**. The most important ones are the *keyword file* and the *logfile*.

The *keyword file* is named **FORGRxy.KEY** (*xy* the MIDAS unit) and holds the keyword data base accessible by all programs running in the MIDAS environment.

The *logfile* is named **FORGRxy.LOG** and receives a log of all user input and all MIDAS output on the terminal (except HELP text, as explained before, and output from the host system). The *logfile* serves also as a “fall back” utility in case of system crash or other breakdown. In such a case the command **PLAYBACK/LOG ‘logfile’** may be used to regenerate the complete MIDAS session.

Note

*In order to use the playback facility, you have to rename the original logfile before restarting MIDAS via **INMIDAS** or **inmidas**. Remember that **INMIDAS** deletes old MIDAS logfiles unless you run in parallel mode.*

3.6 MIDAS Command Language

The MIDAS command language (MCL) consists of all the commands which you enter interactively, and an additional set of commands to provide the necessary tools to write MIDAS “programs”, called MIDAS *procedures*.

The MCL is a flexible and powerful tool to integrate application modules into MIDAS and to do rapid prototyping. But it is not intended to be a full blown programming language - for programming tasks MIDAS supports the standard interfaces in FORTRAN 77 and C (cf. the MIDAS Environment document). It is an interpreted language, so you do not need to compile MIDAS procedures. It is also a “Macro” language in the sense that you can build complex procedures, attach these procedures to a MIDAS command and qualifier combination and then put a single line with that command name into yet another procedure (up to 20 levels deep).

MIDAS procedures are handled in the following way:

The ASCII procedure file is read in by the MIDAS monitor and translated into an internal more compact format. This translated code is then executed inside the Monitor.

The individual lines of code are parsed and decoded in two passes: In the first pass, all symbol substitutions are done using the specified formats to convert from binary to ASCII. In the second pass, all control and conditional statements are processed directly by the Monitor (e.g. positioning the internal program pointer to the command line referred to by a GOTO statement) until an “executable” command line is found which is passed on to the usual command input pipeline of MIDAS as if it were typed in by the user.

For a detailed explanation of all the MIDAS Command Language commands see the appendix of this volume or use the MIDAS command **HELP/CL**.

The following *Command Language* commands provide the necessary programming constructs like looping and conditional branching for MIDAS procedures, they cannot be used interactively:

BRANCH *variable* *comparisons* *labels*

Compare *variable* with *comparison* values and branch to related *labels*

CROSSREF *label1* ... *label8*

Define cross reference *labels* for parameters *par1* ... *par8*

DEFINE/LOCAL *key* *data* *all_flag* *level_flag*

Define local keyword *key* and initialize it using *data*

DEFINE/PARAMETER *par* *def* *type* *prompt* *limits*

Declare *default* value, *type*, *promptstring* and *limits* for parameter *par*

DO *loopvar* = *begin* *end* *step_size*

...command body...

ENDDO
Execute a do-loop (as in FORTRAN)

ENTRY procedure
Define the beginning of a MIDAS *procedure* in procedure file with a different name

GOTO *label*
Jump to a *label* defined as *label*: see below

IF *par1 op par2 command*
Execute conditional statement (as in FORTRAN)

IF *par1 op par2 THEN*
...if-sequence...
ELSEIF *par1 op par2 THEN*
...else if-sequence...
ELSE
...else-sequence...
ENDIF
Execute a conditional statement (as in FORTRAN)

INQUIRE/KEYWORD *key* *prompt-string*
Demand value for *key* from the user

label:
Declare a *label*, e.g. **HOME**:

RETURN *par1 ... par3*
Return to calling procedure or terminal and pass up to 3 parameters

PAUSE
interrupt the current procedure and return to interactive level

DEFINE/MAXPAR *nopar*
Indicate that max. *nopar* parameters are expected

The following commands may also be used interactively, but are especially useful inside MIDAS procedures:

@ (or: **@@**, or: **@a**, or: **@s**, or: **@c** *proc par1 ... par8*)
Execute the MIDAS procedure *proc* which is stored
in **MID_PROC:**, (or in the current directory or **MID_WORK:**),

or in APP_PROC:, or in STD_PROC:, or in CON_PROC:, respectively

ECHO/qualif levela,levelb

Control the display of MIDAS commands (*qualif* = ON, OFF, FULL) for procedures executing at a level in the interval [*levela,levelb*]

COMPUTE/KEYWORD reskey = expression

Evaluate an algebraic *expression* involving keywordss and constants, store result in *reskey*

SET/FORMAT I-format E-format

Define *formats* used for replacements of keyword and descriptor names in procedures with their actual values

WRITE/OUT text

Display *text* on terminal

! *comment*

Indicate beginning of a *comment* line

Note

It is good practice not to abbreviate the commands and qualifiers of a MIDAS command inside a procedure. Since new MIDAS commands can be created any time an abbreviated command may work at one time but become ambiguous at other times and cause the procedure to fail.

The command TRANSLATE/SHOW myproc X will check that all commands and qualifiers are fully specified in procedure myproc.prg.

3.6.1 Passing Parameters in MIDAS Procedures

A MIDAS command procedure may be created with an editor or via the command WRITE/COMMANDS which constructs a MIDAS procedure from the current command buffer. Default type for such a procedure file is .prg . This MIDAS procedure can then be executed with the commands:

@ file par1 par2 ... par8	! if the procedure is in MID_PROC
@@ file par1 par2 ... par8	! if in current directory or MID_WORK
@a file par1 par2 ... par8	! if in APP_PROC
@s file par1 par2 ... par8	! if in STD_PROC
@c file par1 par2 ... par8	! if in CON_PROC

where *par1 ... par8* are the actual parameters which may be accessed within the command procedure through the character keywords P1 ... P8.

As with data files you can specify a search path for procedures via the command SET/MIDAS_SYSTEM

PATH=directory.

The maximum size of a single parameter is 80 characters, but all parameters together may not exceed 256 characters (which is the maximum size of a command line). The size of the code of a procedure is not limited.

In the following, let us assume that all procedures are stored in the directory specified by MID_WORK so that we always use the MIDAS command @@ to execute them. A command procedure in turn can execute another command procedure (or itself) – up to 20 procedure levels deep. The end of a procedure file or the commands RETURN or ENTRY will bring you back up to the next higher level.

To pass parameters back to a higher level command, use the command

RETURN rtpar1 ... rtpar3 . These return values can then be accessed via the character keywords Q1, Q2, Q3. This technique is an alternative to using global keywords for that purpose.

To use the actual values of a parameter in the procedure, the formal parameters P1,...,P8 have to be enclosed in curly brackets ({, }):

```
!+
! Example 1, MIDAS procedure exa1.prg
!+
READ/KEYWORD {P1}      ! read keyword the name of which is given as par1
@@ test {P2}            ! execute test.prg and pass par2 as first parameter
WRITE/KEYWORD INPUTC {P2} ! write contents of par2 into keyword INPUTC
```

Entering the MIDAS command @@ exa1 OUTPUTC ESO-Garching will lead to the execution of:

```
READ/KEYWORD OUTPUTC
@@ test ESO-Garching
WRITE/KEYWORD INPUTC ESO-Garching
```

The command @@ always passes 8 parameters to a command procedure. If fewer than 8 parameters are specified in the command line, dummy parameters (indicated by the special character '?' (question mark)) are internally appended.

Therefore, @@ exa1 OUTPUTC will put the character '?' into the first element of the character keyword INPUTC.

If we enter the command ECHO/ON before executing the procedure we would actually see the above commands displayed on the terminal (cf. subsection 3.8.1).

Note

Up to MIDAS release 88NOV apostrophes were used for symbol substitutions (e.g. 'P1'). Because of the backward compatibility of MIDAS you could still use apostrophes to indicate symbol substitutions, which is, however, discouraged. The main reason being that using { and } instead, makes nesting of substitutions possible.

The command **DEFINE/PARAMETER** should be used for each parameter that is referenced in the procedure. This command will set the defaults, the type, and the prompt string for each parameter. For numeric values passed as parameter also lower and upper limits can be specified in the **DEFINE/PARAMETER** command.

The default values defined inside the procedure will be used in case the parameters are not explicitly provided (i.e. entered as '?'):

```
!+
! Example 2, MIDAS procedure exa2.prg
!+
DEFINE/PARAM P1 999 NUMBER "Enter first input number: " 22,1024
DEFINE/MAXPAR 1                                ! only 1 parameter expected
WRITE/KEYWORD INPUTI/I/7/1 {P1}      ! store contents of P1 in INPUTI(7)
```

The MIDAS command: **@@ exa2 77** will set INPUTI(7) to 77, whereas **@@ exa2** will set INPUTI(7) to 999.

Entering **@@ exa2 17** will result in an error since the valid interval for the number passed as the first parameter is [22,1024].

If you do not want to give default values for a parameter (in other words, if specific input is required for this parameter), use the symbol '?' as default. In that case, and if the relevant parameter is not given, the user will be prompted for this parameter (using the prompt string specified in the **DEFINE/PARAMETER** command) .

The **DEFINE/PARAMETER** line above also demonstrates how to put a character string with embedded blanks into a single parameter (remember that blanks are parameter delimiters in MIDAS) by enclosing the prompt string with double quotes.

The **DEFINE/PARAMETER** command also checks the type of the parameter. The types which may be tested are: I(mage), T(able), F(itfile) , N(umber), C(haracter).

If for any reason you do not want type checking, use the character '?' instead of any of the types listed above.

For file-type parameters the following translations are executed:

catalog entry numbers, e.g. #27 are replaced by the corresponding file name in the catalog (if that catalog is active!) and the asterisk ('*') is substituted by the currently displayed image, if any.

For numerical parameters it is tested if the input is a number; for character strings it is only checked that the first character is a non-numeric character.

Using the plus sign ('+') as default value is another way to disable parameter type checking. This is the correct way to test inside a procedure whether a certain parameter has been entered or not, because it is impossible to distinguish between a parameter defaulted to '?' and an explicitly entered '?' parameter. For an example see example 14a, 14b in subsection 3.6.5.

The system keyword **PARSTAT** holds 8 flags (for P1,...,P8) which are set to 1 or 0, if the

type of the *i*th parameter conforms to the specified type or not. If PARSTAT(i) is 0 for any *i* the MIDAS procedure is aborted.

However, if /C(ONTINUE) is appended to any of the types listed above, the keyword PARSTAT will only be set to 0 or 1 and the execution of the procedure continues, leaving it to the user to test PARSTAT(i) and decide how to go on.

So in our example above the command @@ exa2 KB will result in an error message and the procedure is aborted.

If we change the procedure to:

```
!+
! Example 3, MIDAS procedure exa3.prg
!+
DEFINE/PARAM P1 999 N/CONT "Enter first input number:"
DEFINE/MAXPAR 1                               ! only 1 parameter expected
IF PARSTAT(1) .EQ. 1 -
    WRITE/KEYWORD INPUTI/I/7/1 {P1} !store contents of P1 in INPUTI(7)
```

then @@ exa3 KB will not yield any error.

If we enter @@ exa3 RW PG CG KB MP PB the message

Warning: 5 parameter(s) more entered than required...

will be displayed but the execution of the procedure continues (the additional parameters are ignored). Note also the use of the continuation character (-) in the IF statement above.

The MIDAS command CROSSREF defines labels (of maximum 10 characters) for the parameters P1,...,P8 to enable cross-referencing of parameters if they are passed in arbitrary order.

Note

The command CROSSREF has to be the first executable command (i.e. any command but a comment line) in a MIDAS procedure!

The command DEFINE/MAXPAR provides an additional consistency check and helps to detect erroneous usage of MIDAS procedures. Therefore, it's highly recommended to include it in all procedures.

If we modify exa3.prg to:

```
!+
! Example 4, MIDAS procedure exa4.prg
!+
CROSSREF INFILE OUTFILE METHOD ALPHA
DEFINE/PARAM P1 ? IMA "Enter name of input file: "
DEFINE/PARAM P2 ? IMA "Enter name of result file: "
DEFINE/PARAM P3 ? C "Enter method: "
DEFINE/PARAM P4 999 NUM "Enter alpha value: " 22,1024
DEFINE/MAXPAR 4                               ! max. 4 parameters expected
WRITE/KEYWORD INPUTI/I/7/1 {P4}
```

then the following command lines will all be equivalent:

```
@@ exa4 in out FILTER 33
@@ exa4 P2=out P1=in P4=33 P3=FILTER
@@ exa4 OUT_FILE=out IN_FILE=in alpha=33 METHOD=FILTER
```

The labels may be truncated, so also

```
@@ exa4 OUT=out IN_F=in al=33 METH=FILTER
```

is o.k.

If you do not know a parameter value at the time you execute a MIDAS procedure, e.g. the value depends on the execution inside the procedure itself, use the command INQUIRE/KEYWORD in the procedure. The execution of the procedure is then interrupted and the user is prompted for a value before continuing. For example,

```
!+
! Example 5, MIDAS procedure exa5.prg
!+
CROSSREF INFILE OUTFILE
DEFINE/PARAM P1 ? IMA "Enter name of input file: "
DEFINE/PARAM P2 ? IMA "Enter name of result file: "
DEFINE/MAXPAR 2 ! max. 2 parameters expected
WRITE/KEYWORD IN_B " " all ! fill keyword IN_B with blanks
INQUIRE/KEYWORD IN_B "Which filter, enter LOW or HIGH: "
IF AUX_MODE(7) .EQ. 0 IN_B = "LOW" ! LOW is the default
```

The command `@@ exa5 old new` will stop with the message

Which filter, enter LOW or HIGH:

and wait for user input. The 7th element of keyword `AUX_MODE` will contain the number of characters typed in response to the `INQUIRE/KEYWORD` command. `AUX_MODE(7)` is set to 0 if the user just types `Return`.

3.6.2 Symbol Substitution in Command Procedures

As mentioned before, the Monitor performs symbol substitutions on MIDAS command lines in the first pass by replacing symbol names in the command line with their current value. For character symbols just the string is put in; for symbols of other types the binary data are converted to ASCII using the formats specified in the `SET/FORMAT` command. This substitution is iterated until no more symbol substitutions are possible. Keywords, descriptors, pixel values of an image or elements of a table are valid symbols in the MIDAS command language.

The following syntax is used to distinguish among keywords, descriptors, pixel values and table elements:

{star}	refers to the value stored in the keyword star
{galaxy,disk}	refers to the contents of descriptor disk of frame galaxy.bdf
{galaxy[x,y]}	refers to the value of the image pixel at coordinate x,y of the 2-dimensional frame galaxy.bdf
{dust,:particles,7}	refers to the element of the table dust.tbl in column labeled :particles and row 7
{dust,#2,77}	refers to the element of the table dust.tbl in the second column and row 77

Elements of numerical keywords with more than one element are specified like elements in a FORTRAN vector, e.g. INPUTR(7). Also substrings of character keywords are indicated as in FORTRAN, e.g. INPUTC(2:5). These features are also implemented for descriptors but not for table entries (yet).

Any algebraic expression using the operators +, -, *, / and parentheses (,) and constants as well as any symbol above which defines MIDAS data is supported by the command COMPUTE/KEYWORD and its short form key = expression. This also applies to all the other direct assignments of single values to MIDAS data structures we had described above in section 3.3, e.g. image[x,y] = expression.

Let us look at an example of this:

```
!+
! Example 6, MIDAS procedure exa6.prg
!+
DEFINE/PARAM P1 ? N "Enter alpha value: " -88.5,912.4
DEFINE/PARAM P2 ? N "Enter loop_count: " 1,999
DEFINE/MAXPAR 2 ! max. 2 parameters expected
WRITE/KEYWORD VAR/R/1/1 0. ! init key VAR
VAR = {P1} * 3.3 ! set VAR to 3.3 * (contents of P1)
WRITE/DESCR myframe rval/r/1/2 0.0,0.0 !
LOOP: ! declare label LOOP
VAR = 1.+VAR ! set VAR = 1.0 + VAR
myframe,rval(2) = var+12.99
WRITE/OUT {myframe,rval(2)}
myframe[@10,@20] = 20.0-{myframe,rval(2)}
WRITE/OUT {myframe[@10,@20]}
mytable,:DEC,@7 = {myframe[@10,@20]>*2.0
WRITE/OUT {mytable,:DEC,@7}
WRITE/OUT ""
IF VAR .LE. {P2} GOTO LOOP ! go to label LOOP, if VAR ≤ contents of P2
```

Then the command `@@ exa6 1.0 5.2` will yield:

```
1.72900E+01
2.71000E+00
5.42000E+00
```

```
1.82900E+01
1.71000E+00
3.42000E+00
```

Note

For character keywords `COMPUTE/KEYWORD` only supports character concatenation ('//'). If you want to write a character string into a character keyword, use `WRITE/KEYWORD` instead. Therefore, if we had written `VAR = P1 * 3.3` instead of `VAR = {P1} * 3.3` in the procedure `exa5.prg`, MIDAS would have protested because no multiplication is permitted for character keywords.

You may want to use symbol substitutions for sending the output of a MIDAS command to an ASCII file or to the terminal depending upon the contents of e.g., the character keyword `mykey`.

Setting keyword `mykey` once to "`>outfile`" or "`>terminal`" if you want output to a file or not together with the command line `WRITE/TABLE mytab {mykey}` in your procedure will *not* work!

For, the check for the output redirection is done at the very first parsing of the command line before any symbols in that line are replaced...

Instead, setting `mykey` to "`outfile`" or "`terminal`" and changing the command line to: `WRITE/TABLE mytab >{mykey}`

will do the intended switching of output to a file or terminal.

Since symbols may be tested in conditional statements and thus change the control flow of a MIDAS procedure, they provide the link between application programs and the MIDAS command language.

The number of characters used in the ASCII representation of a numerical symbol may be controlled via the command `SET/FORMAT I-format` for integer symbols and `SET/FORMAT x-format,y-format` (where x or y can be E, G or F) for real (x-format) and double (y-format) precision symbols. Integer symbols are then encoded via I-format (with leading zeroes not suppressed) and real or double precision symbols as E-format, G-format or F-format (used as in FORTRAN 77):

```
!+
! Example 7, MIDAS procedure exa7.prg
!+
WRITE/KEYWORD INPUTI 12 ! set INPUTI(1) to 12
```

```

WRITE/KEYWORD INPUTR 12.345           ! set INPUTR(1) to 12.345
                                         ! and set INPUTD(1) to 123456.98765432
WRITE/KEYWORD INPUTD 123456.98765432
WRITE/OUT {inputi(1)} {inputr(1)} {inputd(1)}
SET/FORMAT I2                         ! use format I2.2 for integer symbols
                                         ! and use format E12.8 and G22.8 for real and double symbols
SET/FORMAT E12.8,G22.8
WRITE/OUT {inputi(1)} {inputr(1)} {inputd(1)}
SET/FORMAT i5                         ! use format I5.5 for integer symbols
                                         ! and use format F12.4 and E22.13 for real and double symbols
SET/FORMAT f12.4,e22.13
WRITE/OUT {inputi(1)} {inputr(1)} {inputd(1)}

```

The command @C exa7 will yield:

0012 1.23450E+01 1.23457E+05	default is I4 and E15.5,E15.5
12 1.23450003E+01 1.23456988E+05	uses I2 and E12.8,G22.8
00012 12.345 1.2345698765432E+05	uses I5 and F12.4,E22.13

If you want to omit any leading zeroes for integer symbols use SET/FORMAT I1, then only the necessary digits will be displayed.

Note

Until MIDAS version 94NOV the same format was used for real and double symbols. This led to problems when real and double symbols in the same command line had to be substituted.

Use SET/FORMAT f-format to only change the format for real symbols and SET/FORMAT ,f-format to only change the format for double symbols.

Substitution begins inside the curly brackets, starting at the deepest nested level:

```
WRITE/OUT {IN_A}{INPUTC(1:3)}
```

will display SPIRALABC on the terminal, if key IN_A contains the string SPIRAL and key INPUTC(1:3) the string ABC.

It is sometimes necessary to substitute symbols in a nested order:

```

!+
! Example 8, MIDAS procedure exa8.prg
!+
DEFINE/PARAM P1 myframe IMA "Enter name for input frame: "
SET/FORMAT F5.1
WRITE/OUT {{P1},STEP(1)}

```

the command @C exa8 will force the Monitor to substitute the last command line in exa8.prg first to: WRITE/OUT {myframe,STEP(1)} and then yield: 20.5 assuming that descriptor STEP of myframe.bdf contains 20.5 as first element. This example also illustrates the concept of recursive substitution.

3.6.3 DO Loops

Loops are supported in MIDAS procedures like the DO loops in FORTRAN (but note that loops are always executed at least once):

```
!+
! Example 9, MIDAS procedure exa9.prg
!+
WRITE/KEYWORD N/I/1/1 0           ! keywords serve as loop variables
DO N = 1 6 2                     ! loop from N=1 until N≤6 in steps of 2
  WRITE/OUT N = {N}
ENDDO
```

A keyword of integer type (called N in our example) must be used to store the loop variable. The parameters follow the standard FORTRAN conventions with start (=1 in `exa9.prg`), end (=6) and in/decrement (=2) values given as shown above. DO loops may be nested up to 8 levels deep in a procedure.

The command `@@ exa9` will yield

```
N = 0001
N = 0003
N = 0005
```

Assume we have images `imag0001.bdf` to `imag0100.bdf` and want to add successive pairs and store the results into images `res0001.bdf` to `res0050.bdf`:

```
!+
! Example 10, MIDAS procedure exa10.prg
!+
DEFINE/PARAM P1 ? IMA "Enter root_name for input frames: "
DEFINE/PARAM P2 ? IMA "Enter root_name for output frames: "
DEFINE/MAXPAR 2                  ! max. 2 parameters expected
SET/FORMAT I4                    ! we need 4 digits
WRITE/KEYWORD N/I/1/1 0
WRITE/KEYWORD NN/I/1/2 0,0
!
DO N = 1 50                     ! default increment is 1
  NN(1) = 2*N
  NN(2) = NN(1)-1
  COMPUTE/IMAGE {P2}{N} = {P1}{NN(2)}+{P1}{NN(1)}      ! sum up
  LOAD/IMAGE {P2}{N}                ! display the result frame
ENDDO
```

Then, the MIDAS command `@@ exa10 imag res` will do the required task.

3.6.4 Local Keywords

Because keywords are implemented as a global data structure, different MIDAS procedures can access the same keyword. This useful feature can cause problems, however, if these keywords just serve as local, temporary variables like the DO variables. Consider the procedures below:

```
!+
! Example 11, MIDAS procedure exa11.prg
!+
WRITE/KEYWORD N/I/1/1 0
DO N = 1 10
  @@ test
ENDDO

!+
! MIDAS procedure test.prg
!+
WRITE/KEYWORD N/I/1/1 0
DO N = 1 12
  WRITE/KEYWORD INPUTI/I/12/1 {N}
ENDDO
```

Executing @@ exa11 will give some unexpected results, since both procedures access the same integer keyword N as a common variable.

Therefore, procedures should use *local* keywords for DO loops and internal working storage. Local keywords are defined inside a MIDAS procedure via the command **DEFINE/LOCAL**. They are only known inside the procedure where they are defined (if the *lower_levels* flag is set, they are also defined in all procedures called from this procedure). Local keywords may have the same name as an existing global keyword (except the system keyword names as stored in **MID_MONIT:syskeys.dat**) or local keyword of any other procedure, since local keywords are searched before the global ones. The above example will work, if modified as follows:

```
!+
! Example 12, MIDAS procedure exa12.prg
!+
DEFINE/LOCAL N/I/1/1 0
DO N = 1 10
  @@ test
ENDDO

!+
! MIDAS procedure test.prg
```

```
!+
DEFINE/LOCAL N/I/1/1 0
DO N = 1 12
  WRITE/KEYWORD INPUTI/I/12/1 {N}
ENDDO
```

Local keywords are deleted when returning to the next higher level at the end of a procedure.

3.6.5 Conditional Statements, Branching

As in FORTRAN 77 any of the following forms of the IF statement may be used:

IF log_exp command	! command = any MIDAS command ! with at most 4 params.
--------------------	---

IF log_exp THEN	! xyz = any logical expression
...	
ELSEIF log_exp THEN	! uvw = any logical expression
...	
ELSE	
...	
ENDIF	

IF log_exp THEN	
...	
ENDIF	

IF blocks may be nested up to 8 levels deep in a procedure.

The logical expression 'log_exp' is of the form:

arg1 op arg2

where *arg1*, *arg2* are either names of keywords (this includes also the names P1, ..., P8) or constants, and *op* may be any of .EQ., .NE., .GT., .GE., .LT. or .LE. (with the same meaning as in FORTRAN 77). As with symbol substitution, specify single elements of an array and substrings via, e.g., OUTPUTI(7) and IN_B(2:15).

If we do

```
WRITE/KEYWORD INPUTC beer
WRITE/KEYWORD OUTPUTC wine
WRITE/KEYWORD INPUTR/R/1/3 1.,2.,3.
```

Then,

INPUTC .EQ. "beer"	is TRUE
INPUTC .EQ. "BEER"	is also TRUE
INPUTC .EQ. OUTPUTC	is FALSE
INPUTC(2:2) .EQ. OUTPUTC(4:4)	is TRUE
INPUTR(2) .GT. 5.4	is FALSE

In string comparisons upper and lower case characters are not distinguished in order to guarantee case insensitivity.

Character keywords can only be compared to character keywords or character constants (which are enclosed by double quotes). This can become tricky in conjunction with symbol substitution:

```
!+
! Example 13a, MIDAS procedure exa13a.prg
!+
DEFINE/PARAM P1 ? N "Enter number: "
DEFINE/MAXPAR 1                                ! only one parameter expected
IF {P1} .EQ. 1 THEN
  WRITE/OUT P1 = 1
ELSE
  WRITE/OUT P1 is not = 1
ENDIF
```

Entering @@ exa13a 1 as well as @@ exa13a 001 will give the expected output message P1 = 1 since the line IF {P1} .EQ. 1 THEN has been converted in the first pass by the Monitor to

IF 1 .EQ. 1 THEN or IF 001 .EQ. 1 THEN

and the two integer constants are equal. Now, consider the almost identical procedure exa13b.prg:

```
!+
! Example 13b, MIDAS procedure exa13b.prg
!+
DEFINE/PARAM P1 ? N "Enter number: "
DEFINE/MAXPAR 1                                ! only one parameter expected
IF P1 .EQ. 1 THEN
  WRITE/OUT P1 = 1
ELSE
  WRITE/OUT P1 is not = 1
ENDIF
```

Entering @@ exa13b 1 will return the error message invalid IF statement... and abort. Why?

Well, in the IF statement above the contents of the character keyword P1, which is the character '1', is compared to the integer constant 1, an invalid comparison.

We modify the procedure once more:

```
!+
! Example 13c, MIDAS procedure exa13c.prg
!+
DEFINE/PARAM P1 ? N "Enter number: "
DEFINE/MAXPAR 1                                ! only one parameter expected
IF P1 .EQ. "1" THEN
  WRITE/OUT P1 = 1
ELSE
  WRITE/OUT P1 is not = 1
ENDIF
```

Now, entering @@ exa13c 1 will work and yield P1 = 1 but @@ exa13c 001 will output P1 is not = 1 since the string "001" is not equal to "1".

As another example let us see, how we can check if a parameter has been entered at all:

```
!+
! Example 14a, MIDAS procedure exa14a.prg
!+
DEFINE/PARAM P6 + NUMBER "Enter first input number: "
IF P6(1:1) .EQ. "+" THEN
  WRITE/KEYWORD INPUTC NONE      ! no P6 entered, set INPUTC accordingly
ELSE
  WRITE/KEYWORD INPUTI/I/7/1 {P6}  ! store contents of P6 in INPUTI(7)
  WRITE/KEYWORD INPUTC YES      ! indicate, that INPUTI holds a valid number
ENDIF
```

However, if we also want to check the limits of the given number we have to use the DEFINE/PARAMETER command again, because testing "+" against a numerical interval would lead to an error:

```
!+
! Example 14b, MIDAS procedure exa14b.prg
!+
DEFINE/PARAM P6 + NUMBER "Enter first input number: "
IF P6(1:1) .EQ. "+" THEN
  WRITE/KEYWORD INPUTC NONE      ! no P6 entered, set INPUTC accordingly
ELSE
  DEFINE/PARAM P6 + NUMBER "Enter first input number: " 22,1024
  WRITE/KEYWORD INPUTI/I/7/1 {P6}  ! store contents of P6 in INPUTI(7)
  WRITE/KEYWORD INPUTC YES      ! indicate, that INPUTI holds a valid number
ENDIF
```

Since, in the ELSE branch we know that parameter P6 is given, the default value "+" itself is never tested against the interval [22,1024].

For testing multiple alternatives use the BRANCH command. It has the syntax:
 BRANCH variable casea,caseb,...,casez labela,labelb,...,labelz.

```
!+
! Example 15, MIDAS procedure exa15.prg
!+
DEFINE/PARAMETER P1 ? C "Enter method: "
DEFINE/MAXPAR 1           ! only one parameter expected
!
! Use the first 2 characters of parameter P1 to distinguish the methods
BRANCH P1(1:2) AN,DI,HY ANALOG,DIGIT,HYBRID
!
! fall through if no match ...
WRITE/OUT Invalid option - please try again
RETURN
!
ANALOG:
RUN ANALO
RETURN
!
DIGIT:
RUN DIGI
RETURN
!
HYBRID:
RUN HYBRI
```

Then, @@ exa15 ANALOG will execute the command RUN ANALO and @@ exa15 digital or @@ exa15 di will run the program digi.exe.

3.6.6 Special Functions

Special functions may be used with the command COMPUTE/KEYWORD. The currently supported functions are listed in the following table (on the next page). Note, that *arg1*, *arg2* may either be the name of a keyword, the contents of which are used, or a constant. Character constants have to be enclosed in double quotes to distinguish them from a keyword name. On-line help for these functions is available via HELP COMPUTE/KEYWORD.

As an example we want to to display the header of a FITS file stored on disk (without converting the data), the FITS file name may be entered with or without the file extension .mt; if not given we append the type inside the procedure:

```
!+
! Example 16, MIDAS procedure exa16.prg
```

Command	Description
M\$ABS(<i>arg1</i>)	returns the absolute value of integer/real/double <i>arg1</i> as integer/real/double
M\$EXIST(<i>arg1</i>)	returns 1 or 0, if file <i>arg1</i> exists or not
M\$EXISTD(<i>arg1, arg2</i>)	returns 1 or 0, if descriptor <i>arg2</i> of frame <i>arg1</i> exists or not
M\$EXISTK(<i>arg1</i>)	returns 1 or 0, if keyword <i>arg1</i> exists or not
M\$EXISTC(<i>arg1, arg2</i>)	returns the number of the column specified in <i>arg2</i> of the table <i>arg1</i> ; returns -2 or -1, if table <i>arg1</i> or column <i>arg2</i> doesn't exist
M\$FILTYP(<i>arg1, arg2</i>)	returns a type_no for file in <i>arg1</i> , if the file name does not include a file type, the type definition in <i>arg2</i> is appended to the file name
M\$INDEX(<i>arg1, arg2</i>)	returns index of string <i>arg2</i> in string <i>arg1</i> as integer value (same as function INDEX of FORTRAN 77)
M\$INDEXB(<i>arg1, arg2</i>)	same as M\$INDEX but search is done backwards, starting at the end of the string
M\$LEN(<i>arg1</i>)	returns length of string <i>arg1</i> (without trailing blanks)
M\$NINT(<i>arg1</i>)	returns nearest integer of real/double <i>arg1</i>
M\$SYMBOL(<i>arg1</i>)	returns the translation of DCL symbol (VMS) or environment variable (Unix) <i>arg1</i> as a character string
M\$LOWER(<i>arg1</i>)	returns character string <i>arg1</i> in lower case
M\$UPPER(<i>arg1</i>)	returns character string <i>arg1</i> in upper case
M\$TSTNO(<i>arg1</i>)	returns 1 or 0, if string <i>arg1</i> is a number or not
M\$TIME()	returns current date and time as string of 24 characters
M\$SECS()	returns the current time as no. of seconds elapsed since 1st Jan. 1970 (as an integer)
M\$AGL(<i>arg1</i>)	returns contents of AGL definitions file agldevices.dat related to <i>arg1</i> as an ASCII string
M\$LN(<i>arg1</i>)	returns natural logarithm of real/double <i>arg1</i>
M\$LOG(<i>arg1</i>)	returns base-10 logarithm of real/double <i>arg1</i>
M\$EXP(<i>arg1</i>)	returns exponential of real/double <i>arg1</i> (base e)
M\$SIN(<i>arg1</i>)	returns sine of real/double angle <i>arg1</i> (angle in degrees)
M\$COS(<i>arg1</i>)	returns cosine of real/double angle <i>arg1</i> (angle in degrees)
M\$TAN(<i>arg1</i>)	returns tangent of real/double angle <i>arg1</i> (angle in degrees)
M\$ASIN(<i>arg1</i>)	returns arcsine of real/double <i>arg1</i> in degrees
M\$ACOS(<i>arg1</i>)	returns arccosine of real/double <i>arg1</i> in degrees
M\$ATAN(<i>arg1</i>)	returns arctangent of real/double <i>arg1</i> in degrees
M\$SQRT(<i>arg1</i>)	returns square root of real/double <i>arg1</i>

Table 3.2: Special Functions available for operations on keywords

```

!+
DEFINE/PARAM P1 ? ? "Enter FITS file name:"
!
DEFINE/LOCAL INA/C/1/80 " " all      ! that fills all elements of INA with
blanks
DEFINE/LOCAL K/I/1/2 0,0
!
K = M$INDEX(P1,".mt")                  ! test, if type of FITS file entered
IF K .LT. 2 THEN
    WRITE/KEYW INA {P1}.mt             ! if not, append type
ELSE
    WRITE/KEYW INA {P1}                 ! if yes, no need to append type
ENDIF
INTAPE/FITS 1 midd {ina} fnn | $more

```

The MIDAS commands `@@ exa16 test` as well as `@@ exa16 test.mt` will both display the header of the FITS file `test.mt`. Note, that this procedure will display the header in a user friendly way, i.e. one screen at a time (and only work for Unix). One of the MIDAS verification procedures, `verify3.prg` shows the usage of all currently available functions. Enter `@ vericopy` to copy this procedure into your current directory (also the usage of `verify3` will be shown then).

3.6.7 Interrupting Procedures

Sometimes, it may be necessary to interrupt the execution of procedures. One way to do this is via the command `INQUIRE/KEYWORD` which was already discussed before; depending upon the user input the procedure could continue or stop. But while the procedure is waiting for input, MIDAS is blocked, no other command can be executed.

With the command `PAUSE` a procedure is stopped and saved; MIDAS returns to the interactive level and you can execute any other command. To resume the stopped procedure at a later time, enter `CONTINUE`. Then, the procedure continues with the next command after the `PAUSE` line. Only one procedure can be in the 'PAUSED' state at a time, in other words it is not possible to stop and save several procedures together.

As an example, consider the case where after some tricky operations on an image you want to get a grayscale copy of the result on a Postscript Laser printer. Since the grayscale plot is quite a time consuming operation you want to make sure that the frame is really o.k. before sending that job to the printer queue.

```

!+
! Example 17, MIDAS procedure exa17.prg
!+
DEFINE/PARAM P1 ? IMA "Enter input frame: "
DEFINE/PARAM P2 ? IMA "Enter output frame: "
DEFINE/MAXPAR 2                                ! max. 2 parameters expected
WRITE/KEYWORD IN_A {P1}

```

```

DEFINE/LOCAL MYRESULT/C/1/80 {P2}
RUN tricky.exe
PAUSE
!
INQUIRE/KEYWORD INPUTC "Result frame o.k.? Enter YES or NO: "
IF INPUTC(1:1) .EQ. "Y" THEN
  ASSIGN/DISPLAY LASER
  LOAD/IMAGE {MYRESULT}
ENDIF

```

With `@@ exa17 venus jupiter` the procedure will start the program `tricky` to operate on `venus.bdf` and produce the frame `jupiter.bdf`, and then it will stop. Now, you can check the result by e.g. calculating the statistics of `jupiter.bdf` or simply displaying it. Then, resume the procedure via `CONTINUE` and type `YES` if you are satisfied with the result and want the hardcopy or `NO` if not.

Note also, that we used a local keyword to hold the name of the result frame and not the usual keyword `OUT_A`. Thus, we are sure that the result name is not accidentally overwritten by another command which also uses `OUT_A`.

3.6.8 Entry points

It is sometimes desirable to group several related procedures into a single file. In MIDAS, the `ENTRY` command defines entry points for different procedures in the same file. These individual procedures are executed by specifying also their entry point besides the file name in the '`@@`' command.

```

!+
! Example 18, MIDAS procedure exa18.prg
!+
DEFINE/PARAM P1 11 NUMBER "Enter input number: "
DEFINE/MAXPAR 1                                ! only one parameter expected
WRITE/OUT "Parameter 1 = {P1}"
!
ENTRY 2
DEFINE/PARAM P1 new C "Enter input: "
DEFINE/MAXPAR 1                                ! only one parameter expected
WRITE/OUT "Parameter 1 = {P1}"
!
ENTRY third
DEFINE/PARAM P1 spiral IMA "Enter input image: "
DEFINE/MAXPAR 1                                ! only one parameter expected
WRITE/OUT "Parameter 1 = {P1}"

```

The string following the ENTRY command (max. 8 characters) is used in the '@@' command to select the code segment in the file exa18.prg. Thus, @@ exa18,2 old will result in the display of the line: 'Parameter 1 = old'; the following ENTRY statements indicates the end of this code segment and acts like a RETURN statement. Entering @@ exa18,third produces the output: 'Parameter 1 = spiral'; and @@ exa18 -12 will execute the lines with no preceding ENTRY statement, i.e. write: 'Parameter 1 = -12'. This example also shows that parameter P1 is not global, that means it has to be defined in each ENTRY segment of the procedure file.

Entries may also be used to structure the contents of a MIDAS procedure. In the following example, the procedure exa19.prg executes different code segments according to its first parameter.

```
!+
! Example 19, MIDAS procedure exa19.prg
!+
DEFINE/PARAM P1 000 C "Enter control flags for entries: "
DEFINE/PARAM P2 sombrero IMA "Enter image to work with: "
DEFINE/MAXPAR 2                                ! max 2 parameters expected
!
DEFINE/LOCAL LOOP/I/1/1 0
DEFINE/LOCAL CCC/C/1/3 {P1(1:3)}
SET/FORMAT I1
DO LOOP = 1 3
  IF CCC({LOOP}:{LOOP}) .EQ. "1" @@ exa19,000{LOOP} {P2}
ENDDO
!
! here the different sub-procedures
!
ENTRY 0001
!
CREA/IMAGE {P1} 2,256,256 ? gauss 128.5,128,128.5,128
!
ENTRY 0002
!
READ/DESCR {P1}
!
ENTRY 0003
!
STATIST/IMAGE {P1}
```

Then, to read the standard descriptors of image frame luna.bdf we would enter the command @@ exa19 010 luna; to create the frame sol.bdf we enter @@ exa19 100 sol.

Finally, in order to create a gaussian image `estrella.spc`, and read its standard descriptors and do the statistics on the newly created image, we type the command
`@@ exa19 111 estrella.spc`.

3.7 Context Levels

Besides the fixed (general) MIDAS commands, the user may dynamically create new commands any time during a MIDAS session. Context files provide a way to group commands which relate to a specific reduction sequence or application package.

For example, the command `SET/CONTEXT applic1` will execute the MIDAS procedure `applic1.ctx`, which would contain all the new command definitions for the application package `applic1` as well as any new keyword definitions and default settings.

Each enabled context has a corresponding context no. which links new commands to the context in which they were created. The command `SHOW/COMMANDS` displays all additional MIDAS commands together with their context no. The context no. 0 is used for all commands which are created outside a given context.

Once the user has finished his/her data reduction with `applic1`, he/she may want to work with package `applic2` on some of his/her data as well. One could either add all commands of `applic2` on top of the ones from `applic1` or first remove all the commands from the currently enabled context, i.e. `applic1`, in one go via the command `CLEAR/CONTEXT`. `SET/CONTEXT applic2` will then create all the new commands of the package `applic2`.

Use `SHOW/CONTEXTS` to display all the currently enabled contexts.

Up to 8 different contexts may be enabled at any time (assuming that all enabled commands fit in the MIDAS command table).

Some of the currently available contexts (application packages) are:

<code>astromet</code>	Astrometric Package
<code>ccd</code>	Reduction of CCD data
<code>cloud</code>	Model for absorption lines
<code>daophot</code>	Object detection and classification, the DAOPHOT-2 package
<code>do</code>	Data organizer for astronomical observations
<code>echelle</code>	Reduction of echelle spectra
<code>exsas (*)</code>	Analysis of X-ray data from the ROSAT satellite
<code>geotest</code>	Utilities to create geometric test frames and other artificial images
<code>imres</code>	Programs related to image restoration
<code>invent</code>	Object detection and classification, the INVENTORY package
<code>irac2</code>	Reduction programs for the IRAC2 camera
<code>irspec</code>	Reduction of IRSPEC spectra
<code>iue (*)</code>	IUE-tape reader
<code>long</code>	Reduction of long slit spectra
<code>lyman</code>	Package for multiple-components fitting of interstellar absorption lines

mva	Package for multivariate data analysis
optopus	Package to prepare observations with the Optopus facility at La Silla
pepsys	Photometric planning and reductions (extinction correction + transformation to std. system)
pisco	Reduction package for data obtained with the PISCO instrument at La Silla
romafot	Photometric extraction package, the ROMAFOT software
spec	Package for 1-dim spectra
statist	Statistical tests on tables
surfphot	Deconvolution and rebinning
tsa	Package for analysis of astronomical time series
wavelet	Image processing tools using the wavelet transform

For example, the command `SET/CONTEXT invent` will activate the commands related to the `INVENTORY` photometric package.

`HELP [CONTEXT]` will display all currently available contexts at your site.

Note

The contexts `exsas` and `iue` have been developed at the Max Planck Institute for Extraterrestrial Physics in Garching, Germany and ESA Vilspa, Villafranca Satellite Tracking Station, Spain, respectively. They may be obtained on request from these institutions.

3.8 Running a Program within MIDAS

To execute a user-written MIDAS application program (coded in FORTRAN or C), employ the command `RUN`. The command `RUN MYPROG` or `RUN myprog` will execute `myprog.exe` in a subprocess like any other MIDAS command.

It is better practice to embed the command `RUN MYPROG` in a MIDAS command procedure. Typical tasks of this procedure would be to provide default values for all parameters, to check the validity of parameter values, and to store the parameters into the keywords your program will use.

Let us assume you have written your special filter program and stored the executable module as `bestfilt.exe` on disk. Program `bestfilt` just needs the names of the input and output image which are obtained inside the program from the keywords `IN_A` and `OUT_A`.

The following MIDAS procedure:

```
!+
!  MIDAS procedure bestfilt.prg
!+
CROSSREF INPUT RESULT
```

```

DEFINE/PARAMETER P1 ? IMA "Enter input frame: "
DEFINE/PARAMETER P2 ? IMA "Enter output frame: "
DEFINE/MAXPAR 2                                ! max. 2 parameters expected
!
WRITE/KEYWORD IN_A {P1}
WRITE/KEYWORD OUT_A {P2}
RUN BESTFILT                                    ! .exe is the default type

```

will check, that the two parameters are valid MIDAS file names and prompt for input if any parameter is not given. Together with the MIDAS command

```
CREATE/COMMAND BESTFILT/IMAGE @@ bestfilt
```

your application will then be integrated smoothly into MIDAS.

Now, you can apply your own filtering algorithm to the image `lobo.bdf` by typing e.g. `BESTFILT/IMAGE res=perrito in=lobo`.

As with data files you can specify a search path for executables via the command `SET/MIDAS_SYSTEM EPATH=directory` which is then used by the `RUN` command. Furthermore, if your module is written in C you can pass parameters to the executable in the usual way (`argc, argv` stuff) via, e.g.

```
RUN BESTFILT par1 par2 ...
```

3.8.1 Debugging of Procedures and Modules

Normally, the command lines of a MIDAS procedure are not displayed on the terminal. To control the display of the lines of a MIDAS procedure, use the command `ECHO`. With `ECHO/ON` the lines of a MIDAS procedure are displayed on the terminal as they are read from the file and executed. This way, it is possible to get an impression of how much time various parts of a procedure need.

With `ECHO/FULL` the lines are displayed as they are read and if symbols have to be substituted, the lines are again displayed after substitution. To avoid echoing and return to a *silent* mode, enter `ECHO/OFF`.

The `ECHO` command has as parameter the procedure-level-interval where it should be applicable. Thus you can, e.g., display only the lines of a MIDAS procedure executing at level 2, etc. Echoing each command line of a MIDAS procedure will identify most of the syntax and other obvious errors. However, this may not be sufficient for long and complicated procedures.

For these cases use the Midas Command Language Debugger:

DEBUG/PROCEDURE levla,levlb ON/OFF	!en/disable procedure debugging
DEBUG/MODULE levla,levlb ON/OFF	!en/disable module (F 77, C) debugging
SHOW/CODE comnd/qualif	!display the code of related procedure

Once procedure debugging is switched on, e.g., via DEBUG/PROC 1,3 ON, all MIDAS procedures executing at level 1, 2 or 3 start up in stepwise debugging mode. The prompt changes to **Mdb** and each command line is displayed on the terminal, and only executed when you hit [Return]. Furthermore, a set of basic debugging commands may be executed, e.g. listing the preprocessed procedure code, setting and clearing break points, and switching from stepwise to continuous mode. Also the keywords may be inspected at any moment. This is an important tool because local keywords cannot be checked otherwise; once the procedure terminates, all local keywords disappear.

If you want to execute any other command, enter PAUSE to interrupt the procedure you're debugging, execute any other command you want, and enter CONTINUE to continue debugging the procedure.

When you are in the debugger (indicated via the **Mdb** prompt), use the command 'h' (for HELP) to display all the available debug commands.

To switch the debugging mode for procedures off, use DEBUG/PROC 1,3 off.

If you must debug your application program, first compile and link that program with the debugger of your host system. Make sure, that this is the same debugger as the one stored in a system keyword of MIDAS (via the command SET/MIDAS.SYSTEM debug=...). Enter the command DEBUG/MODULE to switch on the debugging mode for applications. Subsequently, your application (as well as all other programs activated via the MIDAS RUN command) will be started with the debugger of your system and you can debug it in the usual way.

Note

Typing \$dbx myprog.exe (e.g. on a SUN) would also start up program myprog.exe in debug mode. But that would not tie the application into the MIDAS environment, i.e. the keywords would not be set correctly.

If you just want to list the preprocessed code of a MIDAS procedure use the command TRANSLATE/SHOW proc. TRANSLATE/SHOW proc X will also check all commands in the procedure for completeness, so it's a good idea to execute that command for all your MIDAS procedures.

The command SHOW/CODE comnd/qualif will display the code of the procedure which is actually executed when you enter comnd/qualif as a MIDAS command.

Note

For a detailed description of the integration of user applications into MIDAS as well as complete examples (in FORTRAN and C) see the MIDAS Environment Document.

3.9 Catalogs in MIDAS

MIDAS catalogs are best described as a list/collection of one of the supported data structures, e.g. Images, Tables or Fit files. Catalogs are implemented as ASCII files with

the file type `.cat`. A MIDAS catalog has entries either for images or tables or fit files currently existing in `MID_WORK` and is then referred to as an Image, Table or FitFile catalog, accordingly. If a catalog is enabled (via `SET/ICAT`, `SET/TCAT`, `SET/FCAT`), new entries are added automatically whenever new frames are created, otherwise entries in a catalog have to be explicitly created via `ADD/ICAT`, `ADD/TCAT`, etc.

Frames that are listed in a catalog may then be referenced by their name, as well as via `#n`, if `n` is the entry_no. of the frame in the currently enabled catalog, or `#n,cat_name` if the entry is in catalog `cat_name.cat` (only one catalog of each type can be enabled at any one time).

The following commands are related to the use of catalogs:

`CREATE/xCAT cat_name dir_specs x = I, T, F` for images, tables, fit files
 create catalog `cat_name` for images/tables/fit files, using `dir_specs`,
 which are the options of the host commands `DIRECTORY` (VMS) or `ls` (Unix).

`SET/xCAT cat_name; CLEAR/xCAT`
 enable/disable automatic addition of entries for images/ tables/fit files in `MID_WORK`
 to catalog `cat_name`

`READ/xCAT cat_name low,hi`
 display all entries of image/table/fit file catalog `cat_name` within given range `[low,hi]`

`ADD/xCAT cat_name frame_list`
 add image/table/fit file entry(ies) specified in `frame_list` to catalog `cat_name`

`SUBTRACT/xCAT cat_name frame_list`
 remove entry(ies) from image/table/fit file catalog `cat_name`

`EXECUTE/CATALOG proc P1 P2 ... P7`
 execute the MIDAS procedure `proc` which was written for a single frame
 for all frames in a catalog (this command currently only implemented for image catalogs).

`WRITE/SETUP CATALOG`
 setup the necessary keywords for an `EXECUTE/CATALOG` command

3.9.1 Using Catalogs in MIDAS Procedures

Assume we have written a specific application program, `pearl`, within the MIDAS environment, that processes an input image and produces some numbers as a result. We would like this program also to work on a sequence of images, not just on one input image:

```
!+
! Example 20, MIDAS procedure exa20.prg
```

```

!+
DEFINE/LOCAL CATAL/I/1/1 0           ! define local key CATAL
!
LOOP:
STORE/FRAME IN_A {P1}               ! fill key IN_A with parameter 1
RUN pearl                           ! run our application
GOTO LOOP

```

If P1 contains the name of an image, the command STORE/FRAME works exactly like WRITE/KEYWORD. The keyword CATAL is not modified.

However, if P1 contains the name of a catalog of the form *file.cat*, this catalog (which has to contain images) is opened and the first entry in the catalog is stored into the keyword IN_A.

The number of the next entry is saved in the keyword CATAL (so this name is fixed!). In the loop, the entry number is taken from CATAL and the corresponding frame name put into keyword IN_A (in our example). If there are no more entries in the catalog, control is either transferred to a label which may be specified in the command line of STORE/FRAME or if not given, the procedure is terminated.

So @@ exa20 myframe will work on the single frame myframe, whereas @@ exa20 mycatal.cat works on all frames with entries in the image catalog mycatal.cat.

If the program pearl produces also output frames, you should not have the catalog enabled (cf. SET/ICAT command). Because each new frame gets added to the enabled catalog and you would end up with an infinite loop!

Furthermore you may write your application and procedure just to work on single frames and then execute this procedure on all frames in a catalog via the command EXECUTE/CATALOG. Note, that you have to set up some special keywords in advance for that via WRITE/SETUP CATALOG; for details see the HELP of EXECUTE/CATALOG.

3.10 Adapting MIDAS to your personal needs

There are commands in MIDAS which help you in tailoring MIDAS to your personal taste and needs.

Maybe the most important one is CREATE/COMMAND with which you can add abbreviated and alias command names. As a next step, try out CREATE/DEFAULTS in order to set up your own defaults for frequently used commands, e.g. size and location of display and graphics windows in an X11-environment. The command SET/MIDAS_SYSTEM has an extended set of options to let you change internal MIDAS features, ranging from selecting your preferred text editor (to be used e.g. in REPORT/PROBLEM) to choosing your own MIDAS prompt. With the command SET/BUFFER you modify the size of the internal command buffer.

If you have a MIDAS command procedure named 'login.prg' in the directory specified

by MID_WORK, this procedure will be automatically executed whenever you get into the MIDAS environment, i.e. when you type INMIDAS (inmidas) or GOMIDAS (gomidas). Therefore, the procedure login.prg is the place where you should put all the commands needed to adapt MIDAS.

```
!+
! MIDAS procedure login.prg
! personal set up file for A. S. Tronomer      940815
!+
CREATE/COMMAND RK READ/KEYWORD           !define abbreviations
CREATE/COMMAND WK WRITE/KEYWORD
CREATE/COMMAND RD READ/DESCR
CREATE/COMMAND WD WRITE/DESCR
CREATE/COMMAND XH CREATE/GUI HELP
CREATE/COMMAND SMOOTH/SPECIAL @@ mysmooth !define a new command
!
CREATE/DEFAULT CREATE/GRAPH ? 400,800      !size for graphic window
CREATE/DEFAULT CREATE/DISP ? 600,600,400,400 !size+loc for display
window
!
SET/MIDAS_SYS edit=vi user=user
SET/MIDAS_SYS prompt=Mid{mid$sess(11:12)}
```

Assuming you are working with MIDAS unit 22, this procedure will change the MIDAS prompt to Mid22, use 'vi' as editor when you run the REPORT/PROBLEM command, define the commands RK, WK, RD, WD, XH, SMOOTH/SPECIAL, and override the preset defaults for CREATE/GRAPHICS, CREATE/DISPLAY. Also, the user level is set to USER; cf. the following section.

3.11 MIDAS User Levels

Three different user-levels are maintained within the MIDAS system.

- NOVICE (beginning MIDAS user)
- USER (normal MIDAS user)
- EXPERT (expert MIDAS user)

These user-levels are set via the command SET/MIDAS_SYS user=level, e.g. SET/MIDAS_SYS us=EXPERT.

The level NOVICE is the default level assigned to you when getting into MIDAS.

The following functions of the MIDAS system are affected by the user-level:

HELP facility:

Help text for NOVICE- and USER- level MIDAS users is limited to the screen size of the terminal, i.e. you have to hit **Return** to scroll through the help text, if it is longer than single screen size.

For EXPERT users the help text is typed out completely.

ERROR reporting:

For NOVICE- and USER- level MIDAS users the full error text is displayed.

For EXPERT users only one-line error messages are displayed.

CREATE/COMMAND command:

EXPERT users may create commands which override already existing MIDAS commands (be careful ...).

All other users may only create *new* commands.

Note

Currently, there is no difference between the level NOVICE and USER, but this may change in a future release of MIDAS.

Chapter 4

Data Structures

This chapter will contain information on the data structures used in MIDAS such as **keys** (**keywords**), **descriptors**, **frames** (**bulk data frames**), **catalogs**, and **tables**.

At the present moment we refer the reader to the previous chapter 3 which gives a basic description of data structures used.

Chapter 5

Table File System

This Chapter describes the structure and use of tables in the MIDAS system. Section 5.1 is a general introduction. The table structure is outlined in section 5.2. Different functional aspects of the tables are described in sections 5.3 (Input/Output), 5.4 (Management) and 5.5 (Operations). A review of the commands is given in section 5.6, a more detailed explanation is included in appendix A. Section 5.7 describes format files to control input/output operations. Finally, section 5.8 contains an example that can be run at a terminal with graphic capabilities as **TUTORIAL/TABLE**.

5.1 Tables in Image Processing

The purpose of image processing systems is to extract information from image data. Systems which are designed to treat large numbers of images must also be able to analyse the data extracted from the images. The standard Database Management Systems (DBMS) provide many of the facilities needed; however, some of the desired interactive graphics, statistics, and mathematics are not always available. Therefore, a dedicated table system has been made to serve these purposes in MIDAS (Grosbøl and Ponz, 1985, *Mem.S.A.It.*, 56, 429).

The use of tables can be divided into three main categories: internal, external, and user applications. One of the advantages of the MIDAS Table System (MTS) is that tables for these different purposes have the same structure and can be treated with the same set of routines. The three categories are discussed separately, although there is some overlap in the applications.

Internal tables are mainly used by MIDAS to compute transformations which later will be applied to images. Typical examples are dispersion relations, characteristic curves and coordinate transformations.

External tables: During a reduction procedure, data from external catalogs or data base may be needed. This information can be made available by transferring them to the MTS format. Examples are given: catalogs of photometric data which can be used to

establish transformations from internal magnitudes to a standard photometric system, or astrometric catalogs for computations of accurate reference frames for images.

User tables are used for storage of values computed during the reduction (e.g. stellar magnitudes, line intensities, or isophotal diameters of galaxies). This provides an easy way to save such heterogeneous data in a computer readable format. Further, the user can investigate the properties of the data (e.g. distributions and correlations of different values, and so on).

5.2 Structure of Tables

Table data are arranged in columns and rows, and stored in MIDAS files with the extension **.tbl**. The entry at a given row and column may be either a single value or an array. The items in one row may describe different properties of the same object or feature. All elements in a given column must be of the *same* type and thus be associated with the *same* property. For instance, a table with stellar data could contain the following items in each row: identification, right ascension, declination, magnitude, and spectral type. The first column would then contain all the stellar identifications, the second the right ascensions, etc.

The supported column data types are numerical data (8/16/32-bit integers or 32/64-bit reals) and character strings.

Each column is tagged with a user-defined label, a display format and optional physical units and can be referred to either by its absolute number or its label.

An item in a table is accessed by giving its column and row in addition to the table name. The row number can either be given as an absolute value (i.e. the sequence number) or indicated by the value in a previously defined reference column.

In addition to the normal columns all tables contain a **SELECT** and a **SEQUENCE** column:

- The **SELECT** column enables the user to define and work with a subset of his table by flagging the rows that satisfy a selection criteria. The subtable will be used by commands that do not modify the table information whereas the selection flag will be reset by commands that modify table information and this before taking any action. The values of this column can be accessed in the **COMPUTE** command by using the name **SELECT** (short form **SEL**).
- The **SEQUENCE** column contains the sequence number of each row. The values of this column can be accessed in the **COMPUTE** command by using the name **SEQUENCE** (short form **SEQ**).

If an element is not defined it will be a **NULL** entry and will be listed as a "*" for all data types except for character strings. In that case it will be listed as an empty field.

The tables may be physically stored on disk in two formats: by records corresponding to the natural way of storing sequentially the rows and transposed, where all the values of a given column are stored together (default mode). A table can be always expanded in the sense that its number of columns and rows is automatically increased when the allocated space is exceeded.

5.3 Input/Output of Tables

The exchange of table data to and from the MTS is mainly done through standard ASCII files. This makes it easy for any program to get data from the MTS and to transfer data into it. Thus, output files from text editors and Database systems containing table data in a fixed format can directly be transferred into the MTS format.

Conversion between the ASCII data file and the table is defined by a format file (see section 5.7). If the format file does not exist, the conversion is done automatically via list-direct input in free format. In this case only REAL*4 and integer data, without **NULL** values are allowed.

Command	Description
CREATE/TABLE	Convert from ASCII files to MTS format
PRINT/TABLE	Transfer MTS information to the ASCII file assigned as output. The printer is used by default.

Table 5.1: Conversion between ASCII Files and MIDAS Tables

Normally, table files should be copied to magnetic tape in the FITS format for tables (Harten et al., 1985, *Mem.S.A.It.*, 56, 437) to make it easy to read them again on other computers. The conversion to FITS is done by the MIDAS command **OUTTAPE**; FITS tables are loaded onto disk by the command **INTAPE**.

5.4 Management of Tables

The management of tables is divided into four tasks: *defining, displaying, modifying* and *interactive editing* of tables. The commands that define or modify a table will update its descriptor **HISTORY**. As the length of this descriptor is limited, if you are doing a lot of operations on the same table, you may get a descriptor overflow. In that case you can turn off the automatic addition of history_lines by adding an integer descriptor named **HISTORY_UPDA** to the table and setting it to 0.

5.4.1 Definition of Tables

External tables are created as described in the previous section and the definition of their content is taken from the format file specified. To *create* a user table one can also use the **CREATE/TABLE** command by giving **NULL** as input. This will create a table of the specified size where all elements are **NULL**. Columns in a table can be created or deleted by using the commands **CREATE/COLUMN** and **DELETE/COLUMN**. The available commands are collected in Table 5.2. Some commands use internal tables to store results. In such cases the tables will be created and defined by the system according to defaults. Labels, display formats and units in an existing column are modified by the **NAME/COLUMN** command.

Command	Description
CREATE/TABLE	Create a table with specified size.
CREATE/COLUMN	Create a column.
DELETE/COLUMN	Delete column(s).

Table 5.2: Commands to Define Tables

5.4.2 Displaying Tables

Both the table parameters and the elements values can be *displayed*. The former are shown using the **SHOW** command. Table values are listed out by the **PRINT** and **READ** commands, the output formatting being done using the display format associated with each column. Supported formats are Fortran-77 standard formats and special display formats to accommodate sexagesimal and time values. Finally table values can be plotted on a graphic device or display unit using the **PLOT** or **OVERPLOT** and **LOAD** command respectively. A list of these commands is given in Table 5.3.

Command	Description
SHOW/TABLE	Show table characteristics
PRINT/TABLE	Print elements in table
READ/TABLE	Read elements in table and display them on the terminal
PLOT/TABLE	Plot table elements on graphic device
OVER/TABLE	Plot table elements on top of a previous plot on the graphic device
LOAD/TABLE	Load table elements on the overlay plane of the display

Table 5.3: Commands to Display a Table

5.4.3 Modification of Tables

Elements in a table can be *inserted*, *changed*, and *deleted*. These functions are all performed by the **WRITE/TABLE** or **COPY** commands (See Table 5.4). The element to be modified must be defined by giving its column and row location. An element is deleted if the value is set to **NULL**. A whole row is considered deleted if the element in the reference column is **NULL**. The data type of a column cannot be changed once the column has been created. However, the command **COPY/TT** can be used to copy and convert the values of a column of a certain type into a column of an another type.

It is possible to *define a "subset"* of a table by the **SELECT** command. All commands that do not change a table element will only use the subset selected. By selecting **ALL** the whole table is selected.

Command	Description
WRITE/TABLE	Write value into a table element.
COPY/KT	Copy a keyword into a table element.
COPY/TK	Copy a table element into a keyword.
COPY/TT	Copy columns values into another column.
COPY/TI	Transform the format of the file from table into image.
COPY/IT	Transform the format of the file from image into table.

Table 5.4: Commands to Modify a Table

It is also possible to *transfer* data from one table to another. The four commands described in Table 5.5 can be used. *Interactive identification* of table entries is done with

Command	Description
COPY/TT	Copy all selected elements with identical reference values.
COPY/TABLE	Copy all selected elements from one table into another.
MERGE/TABLE	Merge common columns in several tables.
PROJECT/TABLE	Copy a set of columns from one table into another.

Table 5.5: Commands to Transfer Table Data

the command **IDENTIFY/xxx**, where **xxx** is CURSOR for the image display and GCURSOR for the graphic screen.

5.4.4 Interactive Editing of Tables

An interactive editing facility **EDIT/TABLE** exists in MIDAS to *create* and *modify* tables. The editor works in a “page-oriented” form, a “page” consisting of 20 rows and several columns to fill the screen format, using a Keypad mode or a command mode to perform the editing functions. The command mode is accessible by hitting CNTL-Z. Most of the editing functions are implemented on the right keypad of the keyboard (Table 5.2) as well as on the left keypad if it exists (e.g on Sun workstations). Some keys of the central keyboard are also recognized (Table 5.7). The functions only available in command mode are listed in Table 5.6

Command	Description
EXIT	to finish the editing session and produce the edited table
QUIT	to finish the editing session without producing the output table

Table 5.6: Table Editor COMMAND Functions

Command	Description
Tab	put the cursor in the next column field.
Return	next row.
Delete	delete previous character.
Backspace	move the cursor to beginning of line.
Cursor Arrows	move the cursor \uparrow , \downarrow , \leftarrow , or \rightarrow .

Table 5.7: Layout of the Table Editor Central Keypad

Page (L1)	Next (L2)	Command	Find	Gold (F1)	H Functions (F2)
Advance (L3)	Backup (L4)		Top		H Keypad
Bottom	Right P (L5)	Crea Col (L6)			
	Left P	Del Col			
Word (L7)	Change (L8)				
	Show	Sort			
Line (L9)	Row (L10)				
Screen	Status				

Figure 5.1: Layout of the Table Editor Left Keypad

Page (KP 7)	Section (KP 8)	Right P (KP 9)	Command	Gold (F1)	H Left Key (F2)
		Left P			H right Key
Advance (KP 4)	Backup (KP 5)	Crea Col (KP 9)	Bottom		
Bottom	Top	Del Col			
Word (KP 1)	Eol (KP 2)	Change (KP 3)	Show		
		sort			
Line (KP 0)		Row (KP .)			
Screen		Status			

Figure 5.2: Layout of the Table Editor Right Keypad

5.5 Operations on Tables

In this section we describe several of the mathematical operations that can be performed on table data. More specialised topics are described in Chapter 8, "Fitting of Data", and in Chapter 11 Vol B, "Multivariate Analysis Methods".

Arithmetic operations between columns in a table are done with the command **COMPUTE/TABLE**. The selection flag is reset by this command. Special functions, named according to the FORTRAN mathematical library, are also available.

Simple statistical descriptors are displayed with the command **STATISTICS/ TABLE**. These descriptors are stored in output keywords for further usage in a procedure.

A set of histogram-related commands, with qualifier **HISTOGRAM**, allow the graphic display of the histogram of a column (**PLOT** and **OVERPLOT** commands), the printout of histogram values (**READ** and **PRINT** commands), or the generation of a 1D image with the histogram of a column (command **COMPUTE/HISTOGRAM**).

Linear or polynomial fits in one or two dimensions can be performed on table columns with the command **REGRESSION**, qualifiers **LINEAR** and **POLY** respectively. The coefficients and error estimations are kept in output keywords that can be stored as table descriptors with the command **SAVE/REGRESSION**. Fitted values are calculated with **COMPUTE/REGRESSION**.

A topic of special interest is the generation of table data from an image and vice versa. A transformation was already described in section 5.4.3, and consists in copying the data from one format to the other, using the commands **COPY/IT** and **COPY/TI**. Tabular data can be converted into 1D or 2D image data with the command **CONVERT/TABLE**. This command works in several modes controlled by a parameter. In all cases the sampling domain of the result is defined by a reference image. The modes currently available are : **POLY** (polynomial fit to table data), **SPLINE** (spline approximation), **PLOT** (scattergram of the data in the table) and **FREQ** (2D histogram of the data in the table).

For more specific resampling and interpolating algorithms, the commands **REBIN** and **INTERPOLATE** will provide full conversion between image and table formats (qualifiers **TT**, **TI**, **IT** and **II**).

5.6 Command Overview

In this section we include a short description of the table commands in alphabetical order, (see the Detailed Command Description in Volume C of the MIDAS User Guide for a more detailed explanation).

Reference to tables is done by the filename. The extension **.tbl** must be appended to the filename in commands which can work both on images and tables (e.g.: **READ/_DESCRIPTOR table.tbl**).

Reference to columns can be done either by "name" or by "number". Columns are referred to by name as **:label**, where **label** is a character string (Note the starting colon '**:**' in front of '**label**'.) The string (max. 16 characters, case insensitive) should start with a letter and may contain alpha-numeric characters and the underscore symbol.

Command	Description
COMPUTE/TABLE	Compute numeric expression of columns.
COMPUTE/HISTOGRAM	Compute column histogram, result in table or image format.
REBIN	Resampling data in table/image formats.
INTERPOLATE	Spline interpolation of data.
REGRESSION/LINEAR	Compute linear regression.
REGRESSION/POLY	Compute polynomial fit.
SAVE/REGRESSION	Store regression coefficients as table descriptors.
COMPUTE/REGRESSION	Compute fitted values using the regression coefficients.
STATISTICS/TABLE	Simple statistics on a table column.

Table 5.8: Operations on Table Data

Columns are referred by number as `#n`, where `n` is the integer defining the column position.

Access to rows can be done in two modes, “sequential” or “direct”.

- Sequential access is defined by the row number as `@n`, where `n` is an integer constant.
- Direct access is done through the values in the reference column.

5.6.1 List of Commands

Table 5.9 contains a list of table commands.

Other table related commands are described in Chapter 8, “Fitting of Data”, and in Chapter 11, Vol B, “Multivariate Analysis Methods”.

5.7 Table Format Files

The conversion of ASCII data into table data can be done automatically (default option) for tables with `REAL*4` columns. In the case of more complex tables, a format file has to be provided to control this conversion.

Format files are ASCII files with an extension `.fmt`, used optionally by the commands `CREATE/TABLE`, `READ/TABLE` and `PRINT/TABLE` to control the input/output conversion. They may contain first a `FS` statement, they must contain then one `DEFINE/FIELD` statement for each column of the table and optional comment statements. `DEFINE/FIELD` statements follow the syntax:

`DEFINE/FIELD pos1 pos2 type [format] label [unit]`

where:

```

COMPUTE/TABLE table column = expression
COMPUTE/REGRESSION table column = name[(ind-vars)]
COMPUTE/HISTOGRAM image = table column
COMPUTE/HISTOGRAM table/TABLE = table column
CONVERT/TABLE image = table indv[,indv] depv refima method [par]
COPY/ET keyword table column row
COPY/TK table column row keyword
COPY/TI in-table out-image
COPY/IT in-image out-table
COPY/TT in-table column [out-table] column
COPY/TABLE in-table out-table
CREATE/COLUMN table column [unit] [format] [type]
CREATE/TABLE table ncol nrow filename [formatfile]
DELETE/COLUMN table column [...]
EDIT/TABLE table [ncol nrow]
IDENTIFY/CURSOR table identifier x [y] [tolerance]
IDENTIFY/GCURSOR table identifier x [y] [tolerance]
INTERPOLATE/IT out-table i,d in+image 5 [degree]
INTERPOLATE/TI out-image in-table i,d refima 5 [degree]
INTERPOLATE/TT out-table i,d in-table i,d s [degree]
JOIN/TABLE tab1 col1,col2 tab2 col1,col2 outtab tol1,tol2
LOAD/TABLE table1 column1 column2 [column3] [p1 [p1] [p3]]
MERGE/TABLE table1 [table2 ...] out-table
NAME/COLUMN table column [column] [unit] [format]
OVERPLOT/HISTOGRAM table column [bin [min [max [LOG10]]]]
OVERPLOT/TABLE table column1 column2 [s-type]
PLOT/HISTOGRAM table column [sc-x,sc-y] [bin [min [max]]] [LOG10]
PLOT/TABLE table column1 column2 [sc-x,sc-y]
PRINT/HISTOGRAM table column [bin [min [max]]]
PRINT/TABLE table [column1 ...] [row1 [row2]] [file [format]]
PROJECT/TABLE in-table out-table column [column ...]
READ/HISTOGRAM table column [bin [min [max]]]
READ/TABLE table [column1 ...] [row1 [row2]] [format]
REBIN/IT out-table i,d[,b] in-image func parm intop
REBIN/TI out-image in-table i,d[,b] refima func parm intop
REBIN/TT out-tb i,d[,b] in-table i,d[,b] func parm intop
REGRES/LINEAR table dep-var ind-var1,ind-var2, ...
REGRES/POLYN table dep-var ind-var1[,ind-var2] degree1[,degree2]
RETRO/TAB table
SAVE/REGRESSION table name
SELECT/TABLE table logical-expression
SET/REFCOLUMN table column
SHOW/TABLE table
SORT/TABLE table column
STATISTICS/TABLE table column
WRITE/TABLE table column row value

```

Table 5.9: Table Commands

pos1 — INTEGER, is the optional *starting position* of the field.

pos2 — INTEGER, is the optional *last position* of the field.

type — defines the *type* of information as:

- R — REAL number, single precision,
- D — real number, DOUBLE PRECISION,
- I — INTEGER number,
- C — CHARACTER string.

format — defines the *format* associated with that field and used for listing out its values.

Supported format are FORTRAN 77 standard format or special formats to accommodate sexagesimal values (*Sww.dd*) and time values (*Tww.dd*). These formats may be defaulted, the defaults being defined as:

- A*w* — for CHARACTER string, where *w* = *pos2-pos1+1*
- I*11* — for INTEGER
- E*12.6* — for REAL in single precision
- D*24.17* — for REAL in double precision

label — defines the associated *label*, according to the rules in section 5.6.

unit — defines, optionally, the associated *units*.

The statement FS defines the list of field separators used in the ASCII data file. It is only used when *pos1* and *pos2* are not specified in the DEFINE/FIELD statement. This statement should be written as follows: FS = "f1f2f3". The number of field separators is not limited. If the blank is used as field separator and if the ascii data file contains character strings, the strings have to be enclosed by double quotes. Per default, FS = "\t", i.e TABS and blanks are used as field separators.

The following format file

```
!+
! Example format file test1.fmt
!+
DEFINE/FIELD 1 9 C :NAME "NGC"
DEFINE/FIELD 10 14 R F5.2 :RA "HOUR"
DEFINE/FIELD 16 20 R F5.2 :DEC "DEGREE"
DEFINE/FIELD 22 22 C :TYPE ""
DEFINE/FIELD 24 26 I :RV "KM.SEC-1"
END
```

corresponds to an ASCII file, test1.dat say, with the following record structure:

.....1.....2.....3
123456789012345678901234567890

NGC 3379 10.75 12.85 E 893

(The ruler, of course, is not part of the data file.)

The following format file, using FS statement,

```
!+
! Example format file test2(fmt
!+
FS = "" DEFINE/FIELD C :NAME "NGC"
DEFINE/FIELD R F5.2 :RA "HOUR"
DEFINE/FIELD R F5.2 :DEC "DEGREE"
DEFINE/FIELD C :TYPE " "
DEFINE/FIELD I :RV "KM.SEC-1"
END
```

can be used to create a table from the ASCII file test2.dat

NGC 3379<TAB>10.75<TAB>12.85<TAB>E<TAB>893

5.8 Example

As an example of use of the table file system, we here describe the tutorial procedure executed via the TUTORIAL/TABLE command. This procedure uses a subset of the Uppsala General Catalogue. The format of the catalogue is defined in the file ugc fmt as follows:

```
DEFINE/FIELD 9 20 R G11.6 :RA "HOUR"
DEFINE/FIELD 21 32 R G11.6 :DEC "DEGREE"
DEFINE/FIELD 33 44 R G11.6 :DB "ARC.MIN."
DEFINE/FIELD 45 56 R G11.6 :DR "ARC.MIN."
DEFINE/FIELD 57 68 R G11.6 :BT "MAGNITUDE"
DEFINE/FIELD 69 80 R G11.6 :RV "KM.SEC-1"
END
```

This file and the actual ASCII data in ugc.dat will be copied into your workspace. The first four lines of the data file have the following layout:

1234567890123456789012345678901234567890123456789012345678901234567890
0.0117 15.87 6.500 6.300 12.00 1047.
0.0233 20.47 4.000 3.800 12.70 *
0.2933 59.03 8.000 10.00 * *
0.3583 16.20 5.800 5.100 14.60 *

(The ruler, of course, is not part of the data file.)

This tutorial shows the usage of some of the basic table file commands to analyse and display the data set.

```

CREATE/TABLE ugc 10 700 ugc      ! create the table file (UGC.tbl)
NAME/COL ugc :RA F11.3          ! change format
NAME/COL ugc :DEC G12.6          ! change format
SHOW/TABLE ugc                  ! display structure
READ/TABLE ugc @1 @30           ! display a few entries
!
PLOT/TABLE ugc :DR :DB          ! plot diameters in red and blue bands
!
REGR/LINEAR ugc :DB :DR          ! linear regression on these variables
READ/KEY    OUTPUTD             ! and display stored coefficients
!
READ/HIST  ugc :BT              ! display results on terminal
PLOT/HIST  ugc :BT              ! and plot device
!
SELECT/TAB ugc :BT.LT.13.5      ! select brightest objects
!
STAT/TAB   ugc :DR              ! do statistics on the subset,
READ/HIST  ugc :DR              ! display the result
!
PLOT/TAB   ugc :BT :RV          ! and plot the selected set
!
SELECT/TAB ugc :RV.GT.4000.0    ! select new subset with largest rad.vel.
!
PRINT/TAB  ugc                  ! print them
!
COMPUTE/TAB ugc :MBT = :BT-25.-5.*LOG10(:RV/50) ! compute abs.magnitude
NAME/COL   ugc :MBT "ABS.B.MAG."          ! include units
!
COMPUTE/TAB ugc :SIZE = :RV*SIN(0.000291*:DB)*20 ! diameter
NAME/COL   ugc :SIZE "KPC"                ! include units
!
PLOT/TAB   ugc :MBT :SIZE          ! display result

```


Chapter 6

Graphic and Image Display

6.1 Graphic Facilities

This section describes the facilities of the graphics package in MIDAS. The package makes use of the Astronet Graphic Library (AGL), which has been accepted as the standard for maintenance and development of the MIDAS graphics software. An overview of the graphic commands currently available in MIDAS is presented.

6.1.1 Introduction

In order to provide a device-independent graphics package, the Italian Astronet Graphic Library has been adopted as the standard low level library for the plot package in MIDAS. One of the main reasons for using the AGL package rather than one of the more evolved and sophisticated packages (*e.g.* GKS) was the fact that the AGL package is simple; meanwhile AGL is still capable of doing the things one needs for reasonably advanced graphics, in particular with respect to interactive facilities. The AGL graphics library is fully integrated within the MIDAS directory structure and is generated like every other MIDAS subroutine library during installation. For an extensive description of the package we refer to the AGL User's and Installation Guides for Version 3.

The graphic facilities available in MIDAS can be divided into three main categories of commands, based on their functionality:

- general plot commands that deal with the setup of graphic packages: the assignment of the graphics device, routing a plot file to a graphic device, and general (over)plot commands for text, line, symbols, etc.;
- main plot commands which do the actual plotting and overplotting of data (*i.e.* images, tables, descriptors or keywords);
- cursor commands which use the graphics cursor.

Below, subsection 6.1.2 first describes how graphic display units (*e.g.* terminals, workstations) can be activated inside the MIDAS environment. Thereafter, in the subsec-

tions 6.1.3 to 6.1.5 the plot commands available in MIDAS will be discussed. In subsection 6.1.6 information can be found about how plot files can be manipulated. In section 6.1.8 a number of examples are presented to illustrate the available commands and their syntax. Finally, in subsection 6.1.9 an overview is given of all the available graphic commands in MIDAS.

6.1.2 Graphic Devices

The characteristics of graphic devices differ from device to device. Therefore, in order to obtain a useful plot, the characteristics of the graphics device in use have to be known in advance. MIDAS can run on various combinations of alpha-numerical terminals, image display systems and graphic devices. How to specify the graphic output device for these possibilities is described below.

If you are using the standard MIDAS configuration (alpha-numerical and graphic terminal together with an image display), the MIDAS start-up procedure MIDAS takes care of the proper assignment. If you want to obtain graphic output on a workstation running under X-Windows you have to issue the command **CREATE/GRAFIC**. This command creates a window on the workstation where subsequent plot and overplot commands will write. Up to 4 graphic windows can be created this way. Removal of a window can be done with **DELETE/GRAFIC**.

To get plot output on a graphics device (the graphic terminal included), a proper assignment for that device has to be done in advance. Obviously, the assignment depends on the type of device in use and hence may differ from system to system. If your institute mainly uses graphic terminals of brand "abc", life would be much simpler if this device were the default one and therefore the assignment to be included in the MIDAS startup procedure. You can ask your local MIDAS support to do so. The assignment to be made is:

- for VAX/VMS systems: **ASSIGN AGL_type AGL3DEV**;
- for UNIX systems (C-Shell): **setenv AGL3DEV AGL_type**;
- for UNIX systems (Bourne-Shell): **AGL3DEV=AGL_type**.

Changing the assignment from the default device to another one can always be done in the **login.prg** file (see Chapter 3). This of course would be useful if a particular device is not assigned by the MIDAS startup procedure but is used regularly. Finally, if you run MIDAS from a stand-alone (non-graphic) terminal, an assignment to the **NUL** device or to the **postscript** device can be made. In the first case only a graphic meta file will be created; in the later case a postscript file. In case of problems consult your local MIDAS support or your system manager.

Table 6.1 contains the most commonly used graphics devices (the *AGL_type*'s) currently supported by MIDAS.

Device Identification	AGL_type
DEC VT125 terminal	vt125
DEC VT240 terminal	vt125
DEC VT125 terminal with Retrogr.	tkg.vt640
DEC VT100 terminal with Retrogr.	tkg.vt640
CIT101 with CIG 201 card	tkg.cit101
GraphOn GO-250	tkg.vt640
HDS 2200	tkg.hds22
HPGL plotters	hpgl
LN03 plus laser printer	tkg.ln03
Tektronix 4010	tkg.t4010
Tektronix 4014	tkg.t4014
Tektronix 4100 series	tkg.t4100
QMS laser printer	tkg.qms
X-Windows	idi
Apple Laser Writer	pscript
Postscript printers	pscript
Postscript file	pscript
Versatec V-80	raster
Raster devices	raster
Null device	null

Table 6.1: Supported Devices

6.1.3 General Commands

The general commands in MIDAS mostly concern setting the plot characteristic, displaying the setup, assigning the graphic output device, and sending an existing plot to a device. These commands are:

- CREATE/GP** create a graphic window on the workstation
- DELETE/GP** delete a graphic window from the workstation
- CLEAR/GP** clear the graphic screen or window
- SET/GP** set the graphic characteristics
- SHOW/GP** show the graphic characteristics
- ASSIGN/GP** assign the graphic device
- COPY/GP** route the plot file to a graphic device

The first two commands in this table are meant for users with workstations running

the X-Window display software. With these commands one can create and delete graphics window(s). The create command allows control over the size of the graphics window as well as the position where it is to be put. CLEAR/GRAFIC erases the graphics window or graphics terminal screen.

SET/GRAFIC and SHOW/GRAFIC

SET/GRAFIC plays an important role in the plot package. This command gives the user control over the plot size, line types (LTYPE), line thickness (LWIDTH), symbol type (STYPE) and size (SSIZE), etc. In Table 6.2 the various options that can be set with the SET/GRAFIC command are listed, together with the default settings. Below we briefly discuss some of the options. More extended documentation can be found in the help file of the SET/GRAFIC command.

By default, all data points in frames, keywords and descriptors are connected with a line; data points in tables are plotted individually. This setting can be changed with the LTYPE and STYPE options in SET/GRAFIC. Normally, when plotting data points in a frame, descriptor, or keyword, the plot package first looks for the line type. If the line type is set to 0 (LTYPE=0) it looks for the symbol (STYPE). If both LTYPE and STYPE are found to be 0 a fatal error occurs.

In case of table plotting the package first looks for the symbol type. When STYPE=0 a line will be drawn corresponding to LTYPE. An error occurs if both LTYPE and STYPE are 0. For histogram plotting or when the bin mode @BIN is on (BIN=ON), the package needs a line type greater than zero; an error occurs when LTYPE=0. Table data can not be plotted with BIN=ON.

By default, before a PLOT command is executed the graphics window is erased. To switch off the erase, you can use the option CLEARGRAFIC=OFF. By doing so, subsequently issued plot commands will run in overplot mode: the screen content is kept. Hence, by issuing a number of PLOT commands you can easily produce several plots on one screen (page). To help you more in designing the layout of your plot scales (and sizes) as well as the position on the screen (paper) can be pre-defined by XSCALE and YSCALE, and XOFFSET and YOFFSET.

The default font used by MIDAS is a simple one but is plotted fast. More fonts are available to enable you to obtain publication quality graphics output. With the command SET/GRAFIC FONT=n, where n is larger than 0, you can use a different (nicer) font than the default one. Currently, the following font sets are available:

- 0 Default built-in font
- 1 High quality roman font
- 2 Greek font
- 3 Script font
- 4 Old English font with astronomical symbols
- 5 Tiny roman font, simpler than 1

To display these fonts and the associated character and symbol sets you can run the tutorial **TUTORIAL/GRAF GENERAL**.

If a colour postscript printer is at your disposal you may use the colour setting options **COLOUR=n** and **BCOLOUR=n**.

In the **SET/GRAF** command defaults for single parameters can be (re)set by: **SET/GRAF par_name=value**. The reinitialization of all plot parameters can be done by: **SET/GRAFIC** or **SET/GRAFIC DEF**.

ASSIGN/GRAFIC and COPY/GRAFIC

Table 6.1 contains the graphic output devices on which MIDAS plots can be produced. The user can specify one of these output devices in advance by the **ASSIGN/GRAFIC** command. In addition to the hardcopy devices available, the command can also be used to reassign the graphics window. For example, the user can indicate that the plot has to be produced on a second graphics window, or on a display window of his/her workstation. After the plot command has finished and a plot file is produced, this plot file can be sent to a device by the **COPY/GRAF** command. This command accepts the same graphics devices as the **ASSIGN/GRAF** command. For workstations this offers the possibility to copy a graph from one window to another. In section 6.1.6 more information can be found about how MIDAS takes care of your plot files.

Example:

```
assign/gra laser nospool
plot/tab example ? #1 -50,-70,10,90
overplot/tab example ? #2 -50,-70,70,90
copy/graphic laser
copy/graphic g,0
assign/gra g,0
```

In this example we first assign the graphics output to become the output device. However, the plot file is kept on disk and not spooled to the printer. After the plot is finished, it is sent to the printer. Hereafter, we also send a copy to the graphics window (provided one exists). Finally, we assign the graphics window as the output device.

6.1.4 Plot Commands

As described in Chapter 3, the MIDAS data structures include frames, masks, tables, catalogues, descriptors, and keywords. With the exception of the masks and catalogues, the plot package is able to plot the data stored in these structures. Data can be plotted with **PLOT/** as well as with **OVERPLOT/** commands. In the first case, MIDAS will start a complete new plot (*e.g.* a graphic terminal screen will be erased and old plotfiles will be deleted); in the latter one MIDAS will extend the existing plot information without erasing the results of previous plot commands.

In addition of the plot commands that enable the user to plot data command are available to produce coordinate box(es) within subsequently data can be overplotted, and com-

Option	Value and meaning and defaults
DEFAULT	no value; sets plot package in default mode
XAXIS=	AUTO or <code>xstart, xend, xbig.tick, xsmall.tick</code> in world coordinates; when <code>xsmall.tick < 0</code> a logarithmic axis is plotted; the default is AUTO
YAXIS=	AUTO or <code>ystart, yend, ybig.tick, ysmall.tick</code> in world coordinates; when <code>ysmall.tick < 0</code> a logarithmic axis is plotted; default is AUTO
FRAME=	RECT or SQUA; default is RECT
XSCALE=	AUTO, scale in world units/per mm or size of plot
YSCALE=	AUTO, scale in world units/per mm or size of plot
XOFFSET=	NONE or the offset of the left y axis to left device boundary
YOFFSET=	NONE or the offset of the lower x axis to lower device boundary
XFORMAT=	NONE, AUTO or format description (see below)
YFORMAT=	NONE, AUTO or format description (see below)
PMODE=	0 (plot without frame and legend), or 1 (plot with frame and some information), or 2 (plot with frame and full legend) which is default
FONT=	font to be used to write text; default is 1; (see below)
LTYPE=	1 (solid line) to 6 (long dash - short dash); (see below) default is 1; 0 corresponds with no line at all
LWIDTH=	set line width; 0 or 1 for single width; 2, 3 and 4 for increasing thickness
STYPE=	1 (dot) to 21 (left arrow); default 4 (cross); 0 corresponds with no symbol at all
SSIZE=	value; set the scaling factor of symbols; default is 1
TSIZE=	value; set the scaling factor for text strings; default is 1
TWIDTH=	value; set the line width for text strings; default is 1
BINMODE=	OFF or ON; default is OFF
COLOUR=	number ranging from 0 to 7; the default is black (1) The setting has only effect on graphic display devices supporting colour
BCOLOUR=	number ranging from 0 to 7; set the background colour; the default is black (1) The setting has only effect on graphic display devices supporting colour
CLEARGRA=	ON or OFF. OFF will not clear graphic screen for a PLOT command; default ON

Table 6.2: SET/GRAPIHC Options

mands to add text, symbols, and (grid) lines. An overview of the commands is given below.

PLOT/AXES	plot a box with tick marks, tick labels and axes labels
OVERPLOT/AXES	overplot a box with tick marks, tick labels and axes labels
PLOT/CONTOUR	plot contours of a two-dimensional image
OVERPLOT/CONTOUR	overplot contours of a two-dimensional image
PLOT/COLUMN	plot a column of an image
OVERPLOT/COLUMN	overplot a column of an image
PLOT/DESCRIPTOR	plot an entry in a descriptor
OVERPLOT/DESCRIPTOR	overplot an entry in a descriptor
PLOT/GRAY	plot gray scale map of a two-dimensional image
OVERPLOT/GRAY	overplot gray scale map of a two-dimensional image
PLOT/HISTOGRAM	plot a histogram of a table column or image
OVERPLOT/HISTOGRAM	overplot a histogram of a table column or image
PLOT/KEYWORD	plot the contents of a keyword
OVERPLOT/KEYWORD	overplot the contents of a keyword
PLOT/PERSPECTIVE	perspective plotting (3-dim.) of an image
PLOT/ROW	plot a row (line) of an image
OVERPLOT/ROW	overplot a row (line) of an image
PLOT/TABLE	plot table data
OVERPLOT/TABLE	overplot table data
OVERPLOT/ERROR	overplot table column containing errors
PLOT/VECTOR	plot vector map from two 2-dim. images with smoothing option
OVERPLOT/VECTOR	overplot vector map from two 2-dim. images with smoothing option
LABEL/GRAFIC	plot text in an existing plot
OVERPLOT/LINE	overplot a line in an existing plot
OVERPLOT/SYMBOL	overplot a symbol in an existing plot
OVERPLOT/GRID	overplot a grid, by connecting tickmarks

Plotting and overplotting axes

The command **PLOT/AXES** offers the users the possibility to draw a frame with certain ranges in x and y. The command is very flexible, since it accepts both the ranges in x and y and the scaling factors as input parameters. Also, the user has the freedom to select the location where the frame is to be drawn. The actual data points can be plotted with subsequent overplot commands (see example below). More coordinate boxes can be plotted using the command **OVERPLOT/AXES** with the same parameter list as in the **PLOT/AXES** command. In the example below a series of plots is produced with plot and

overplot commands. First, we start with a PLOT/AXES and an OVERPLOT/TABLE command, and then continue with three OVERPLOT/AXES and OVERPLOT/TABLE commands. The result is four graphs in the graphics window.

Example:

```
assign/graph g,0                                ! assign graphic window 0
plot/axes 0,10 -1,1 -50,-70,10,80             ! plot the first axes
overplot/tab example ? #1                      ! plot first coord. box
overplot/axes 0,10 -1,1 -50,-70,70,90          ! overplot second box
overplot/tab example ? #2                      ! overplot third box
overplot/axes 0,10 -1,1 -50,-70,130,90
overplot/tab example ? #3                      ! overplot fourth box
overplot/axes 0,10 -1,1 -50,-70,190,90
overplot/tab example ? #4
```

Alternatively, the system also offers a more simpler way of doing the same thing: instead of the PLOT/AXES and OVERPLOT/AXES commands we switch off the clearing of the graphics window first and continue with simple PLOT/TABLE commands.

Example:

```
clear/graph set/graph clear=off                ! erase switched off
plot/tab example ? #1 -50,-70,10,10           ! plot the fifth box
plot/tab example ? #2 -50,-70,70,10
plot/tab example ? #3 -50,-70,130,10
plot/tab example ? #4 -50,-70,190,10          ! plot the last box
```

For both the plot and overplot commands one can use the command SET/GPGRAPH XFORMAT=none YFORMAT=none to switch off the tick mark labels along the axes. Of interest, especially for overlays, is another syntax of the PLOT/AXES and OVERPLOT/AXES command: it offers the possibility of drawing axes around (part of) an image displayed in the display window.

PLOT and OVERPLOT commands for plotting data

In general all commands for plotting data, have a well defined syntax:

PLOT/QUALIFIER P1 P2 P3 P4 P5 P6 P7 P8,

where:

P1 = table, image, descriptor or keyword name

P2 = columns, area, or indices of P1

P3 = scales in world coordinates/mm or size of the plot; only for PLOT commands

The meanings of the remaining parameters on the command lines vary from command to command; in most cases they are used for options. Obviously, in case of overplotting, the parameter for the scales is absent.

Plotting text, symbols, and lines

For overplotting of text one can use the command `LABEL/GRAFIC`. The text will be plotted in the font style set with the `FONT` keyword in the `SET/GRAFIC` command. *E.g.* with `LABEL/GRAFIC` and running in `FONT=1` text will be generated in the roman font type. `LABEL/GRAFIC` make use of the built-in features of AGL. These features allow to change font, draw subscripts and superscripts, scale the text size, or draw various symbols, all within the text string. All these possibility become available by including metacharacters in the text string. Currently, AGL knows the metacharacter set listed in Table 6.3.

Metacharacter	Meaning
<code>\{</code>	begin grouping
<code>\}</code>	end grouping
<code>\^</code>	move the following part of the string up by half a character
<code>\!u</code>	move the following part of the string up by half a character
<code>_</code>	move the following part of the string down by half a character
<code>\!d</code>	move the following part of the string down by half a character
<code>\<</code>	backspace by a single character
<code>\+</code>	increase character size by 20%
<code>\-</code>	decrease character size by 20%
<code>\!</code>	force interpretation of the following part as a metasequence (this is needed to allow metasequences starting with 'n' not to be interpreted as newlines)
<code>\0</code>	select font 0 (Default font)
<code>\1</code>	select font 1 (Quality roman font)
<code>\2</code>	select font 2 (Greek font)
<code>\3</code>	select font 3 (Script font)
<code>\4</code>	select font 4 (Old English)
<code>\5</code>	select font 5 (Tiny roman font)
<code>\[</code>	increase line width (bolding; optional)
<code>\]</code>	decrease line width (bolding; optional)
<code>\#<n></code>	draw marker number $< n >$ into the line
<code>\n</code>	perform a 'newline'
<code>--</code>	draw a single '--' character (must be following by a blank)
<code>\~</code>	draw a single '\~' character (must be following by a blank)
<code>\`</code>	draw a single '\` character (must be following by a blank)
<code>\\"</code>	draw a single '\\" character (must be following by a blank)

Table 6.3: Meta Character in AGL and MIDAS

The character '~~' can also be used instead of '\' as metacharacter flag. The '~~' is more suited to C programs where '\' has a special meaning. All selections made by metacharacters are valid from the point in the string where they are defined up either

to the end of current group (the part of the string enclosed in $\{\dots\}$) or to the end of the string. If the metacharacter sequence is more than one character long (escape not included, of course), it must be followed by a blank space.

Example:

```
LABEL/GRAP "e\{\!\u(x\{\!\u2\}+y\{\!\u2\})\}= -  
(\alpha +\beta ) \sin\{\!\u2\}\theta "
```

This command will produce the label $e^{(x^2+y^2)} = (\alpha + \beta) \sin^2 \theta$ at a position which the user should give via cursor input.

MIDAS/AGL also interprets a set of 'TeX-like' keywords as listed in table 6.4. Due to the fact that most of them represent special characters and symbols to be printed, only the names are listed; the symbols can only be seen by LABEL/GRAPHIC or the command TUTORIAL/GRAPH.

Besides overplotting of text strings, the user can also overplot lines (up to six different line types, depending on the device), and symbols (more than twenty). Depending on the device, up to four different line widths can be used. The selection of line properties and of symbol type can be done with SET/GRAPH, or, at least for line and symbol type, on the command line.

The command OVERPLOT/LINE offers the possibility to overplot a line. Similarly, OVERPLOT/SYMBOL overplots symbols. Both commands can be used interactively using the graphics cursor, or by giving the coordinates on the command line. Grid lines, connecting the major and/or minor tick marks can be drawn with the command OVERPLOT/GRID.

6.1.5 Graphic Cursor Commands

In some of the analysis programs in MIDAS the graphic cursor is a powerful tool. For example, using the cursor one can retrieve wavelengths and line intensities in a plotted spectrum, integration of emission or absorption lines over a wavelength range selected by the cursor, compute the line width and center, etc. With the general GET/GCURSOR command the user can retrieve information from plotted data and store this in a table. Listed below are some of the core and application commands which use cursor interaction. Many additional graphics commands, including those that use cursor interaction, are available in the various contexts, *e.g.* SPEC and ECHELLE.

GET/GCURSOR	read coordinates from graphics screen
CENTER/GAUSS	computes center of a 1-dim. or 2-dim. feature
MODIFY/GCURSOR	change data line of an image interactively
INTEGRATE/APERTURE	compute flux inside an aperture
INTEGRATE/LINE	integrate row of a frame using the cursor
INTEGRATE/STAR	compute flux, radius and background of stars

\AA	\Alpha	\Aquarius	\Aries
\Beta	\Cancer	\Capricorn	\Chi
\Delta	\Earth	\Epsilon	\Eta
\Gamma	\Gemini	\Iota	\Jupiter
\Kappa	\Lambda	\Leo	\Libra
\Mars	\Mercury	\Moon	\Mu
\Neptune	\Nu	\Omega	\Omicron
\Pi	\Phi	\Pisces	\Pluto
\Psi	\Rho	\Sagittarius	\Saturn
\Scorpio	\Sigma	\Sqrt	\Tau
\Taurus	\Theta	\Upsilon	\Uranus
\Venus	\Virgo	\Xi	\Zeta
\aleph	\alpha	\asteroid	\beta
\bigcirc	\black	\blue	\cent
\chi	\circ	\cyan	\clover
\clubsuit	\comet	\dag	\ddag
\default	\delta	\diamond	\div
\downarrow	\epsilon	\equinox	\equiv
\eta	\firtree	\gamma	\ge
\greek	\green	\hbar	\heart
\infty	\int	\iota	\italic
\kappa	\lambda	\larrow	\le
\magenta	\mp	\mu	\nabla
\ne	\nu	\odot	\oint
\old	\omega	\omicron	\oplus
\otimes	\palm	\paragraph	\parallel
\partial	\perp	\phi	\pi
\pm	\propto	\psi	\red
\rho	\rightarrow	\roman	\script
\shield	\sigma	\snow	\spade
\sqrt	\sum	\tau	\theta
\times	\tiny	\uparrow	\upsilon
\varepsilon	\varphi	\vartheta	\white
\xi	\yellow	\zeta	

Table 6.4: TeX-like Characters for text strings in MIDAS Graphics

6.1.6 Handling of Plotfiles

Plotting in MIDAS will create a MIDAS plotfile which contains all essential plot information. By default this plotfile (metafile) is always created by the execution of the main plot commands. The plotfile will carry the name of the data structure that has been plotted: the name of a frame, table, descriptor or keyword. The plotfile has the extension ".plt". MIDAS keeps track of what the user has plotted. The SHOW/GRAPH command shows the user which is the last created plotfile. Subsequent overplot commands will append to this plotfile. Names of plotfiles are a composition of the device names that is currently assigned and the MIDAS session number. As an example, take the situation were the user runs MIDAS under the unit number 04 and the graphics window 1 (graph_wnd1) is assigned. In this case all plots created will be stored in the file graph_wnd104.plt. If the device "lp" was assigned, plot files will go into the file 1p04.plt. Names of plot files are unique and do not have version numbers. Hence, MIDAS will delete an old plotfile if a new one with the same name is created. With the command SHOW/GRAPH you can display the name of the plot file.

There are several ways to obtain a hardcopy of a plot. Below you will find a few examples.

1. The user works with a graphic terminal or a workstation and has made this plot on the graphic terminal first. He/she can now send the plot to a hardcopy device using the COPY/GRA command.

Example:

```
MIDAS 001> PLOT/TABLE mytable :velocity :distance
MIDAS 002> COPY/GRAPH LASER
```

2. The user does not have a graphics terminal (or does not want to use it), and wants to dump his plot directly onto a hardcopy device. In this case, the hardcopy device has to be assigned first as the output device by the ASSIGN/GRAPH command. Now all the plot(s) (including the overplot !!!) will be sent directly to the hardcopy device.

Example:

```
MIDAS 003> ASSIGN/GRAPH LASER      ! assign LASER as output device
MIDAS 004> PLOT/TABLE mytable :velocity :distance  ! make plot
```

3. In a MIDAS plot command sequence (with many *e.g.* OVERPLOT and LABEL commands) intermediate output is not always wanted, in some cases even undesirable. In order to switch off the direct routing of plots to a device users can specify the extra switch NOSPOOL in the ASSIGN/GRAPH command. Using this switch the plotfile(s) will be stored on disk first. Once the user has finished his sequence of plot commands, he/she can create the complete plot on the hardcopy device using the command COPY/GRAPH. Intermediate results can be obtained using the same command.

Example:

```
MIDAS 005> ASSIGN/GRAPH LASER NOSPOOL      ! plot file, don't send
MIDAS 006> PLOT/ROW frame [@100,@150:@150,@250] 20.0,20.0
```

```

MIDAS 007> OVERPLOT/TABLE table #1           ! overplot
MIDAS 008> LABEL/GRAFIC "THIS IS AN EXAMPLE" 90 4 400,300
MIDAS 009> COPY/GRAF LASER                  ! send the plot file

```

4. The user wants to send a previously created MIDAS plotfile (*e.g.* "midas.plt", and different from the last created plotfile) to a device.

Example:

```
MIDAS 010> COPY/GRAF LASER frame.plt ! send plotfile to LASER
```

As can be seen in section 6.1.9, in most cases the user can produce a plot with certain scales of the x- and y-axis. In the current version routing the plot file (with COPY/GRAF) to a device different from the one pre-specified (with the ASSIGN/GRAF command) may lead to incorrect scales. In case the pre-specified device is the same as the device to which the plot is sent the scales will be correct.

Example:

```

MIDAS 005> ASSIGN/GRAF VERSA NOSPOOL
MIDAS 006> PLOT/ROW image [0100,0150:0150,0250] 20.0,20.0
MIDAS 007> OVERPLOT/TABLE table #1
MIDAS 008> LABEL/GRAFIC "THIS IS A EXAMPLE" 90 4 400,300
MIDAS 009> COPY/GRAF LASER                 ! plot will have incorrect scales
MIDAS 010> COPY/GRAF VERSA                  ! with correct scales

```

6.1.7 Encapsulated PostScript Files

For any PostScript hardcopy printer that has been assigned by ASSIGN/GRAF, MIDAS produces a so-called encapsulated PostScript file. Using a public domain macro package (*e.g.* Psfig/TeX) this PostScript plot file can, with a minimum of effort, be included in TeX or LATEX documents. To do so, in the TeX or LATEX document one should refer to the (possibly renamed) MIDAS PostScript file (normally pscrplot.ps). Below, follows a simple LATEX example that shows how it works.

Example:

```

MIDAS 005> ASSIGN/GRAF postscript
MIDAS 006> PLOT/ROW image [0100,0150:0150,0250] 20.0,20.0
MIDAS 007> OVERPLOT/TABLE table #1

```

In your directory we now have a PostScript file postscript.ps, containing the complete plot information written by the commands 006 and 007. Now, you can include this MIDAS PostScript file in our LATEX document, in this case using psfig, developed by Trevor Darrell (trevor@media.mit.edu). Here is how the LATEX text file with the included MIDAS plot then looks like.

Example:

```

\documentstyle[11pt,psfig]{article}
\begin{document}
\section*{Abstract}
We show a simple example of how one can include a PostScript figure,
generated by MIDAS, into a existing \LaTeX document.
\nopagebreak

\begin{figure}[h]
\centering{
\hspace*{-1.cm}
\vbox{\psfig{figure=postscript.ps,width=10cm,height=5cm}}\par
}
\end{figure}
\end{document}

```

6.1.8 Examples

To show the usage and the possibilities of the plot package, a number of examples are built are available below. These examples are available via the TUTORIAL/PLOT P1 command, where P1 can be:

- **GENERAL** to show an example for the graphics utilities, like drawing lines, symbols, text, changing fonts, etc ...
- **AXES** to show an example of plotting several axes in one graphics window and using different axes;
- **TABLE** to show an example of table plotting using different symbols and axes;
- **1DIM** to show an example of (spectral) line plotting;
- **2DIM** to show two-dimensional gray scale and contour plotting.

6.1.9 Command Summary

Table 6.5 on the next page shows an overview of the graphic commands in MIDAS, listed in alphabetical order.

```

ASSIGN/GRAF [device_name] [spool_option]
CENTER/GAUSS in_specs out_specs [out_opt]
CLEAR/GRAFIC
COPY/GRAFIC [device_name] [plot_file]
CREATE/GRAFIC [graph_id] [graph_spec]
DELETE/GRAFIC [graph_id]
GET/GCURSOR [out_specs] [app_flag] [max]
INTEGRATE/LINE frame [y_coord] [x_start,x_end] [no_curs,degree] [batch_specs]
LABEL/GRAFIC label [x_pos,y_pos[,mm]] [angle] [size] [centering]
MODIFY/GCURSOR frame [y_coord] [x_start,x_end] [no_curs,degree]
OVERPL/AXES [x_axis_spec] [y_axis_spec] [x_scale,y_scale] [x_label] [y_label] [x_off,y_off]
OVERPLOT/AXES [coord_str]
OVERPLOT/COLUMN frame [x_coord] [x_sta,x_end] [offset] [l_type]
OVERPLOT/CONTOUR frame coord_str [contours] [sm_par]
OVERPLOT/DESCRIPTOR frame [descriptor] [start,end] [offset]
OVERPLOT/ERROR table column1 [column2] column_err [orient]
OVERPLOT/GRAY frame [coord_str] [gray_lev] [sm_par] [gray_ness] [options]
OVERPLOT/GRID [tick_opt]
OVERPLOT/HISTOGRAM table column [bin [min [max]]] [offset] [log_flag]
OVERPLOT/HISTOGRAM frame [offset] [log_flag]
OVERPLOT/KEY [key_word] [start,end] [offset]
OVERPLOT/LINE [line_type] [x_sta,y_sta [x_end,y_end]]
OVERPLOT/ROW frame [y_coord] [x_sta,x_end] [offset] [l_type]
OVERPLOT/SYMBOL [x_coord,y_coord] [s_type] [s_size]
OVERPLOT/TABLE table [column1] [column2] [s_type]
OVERPLOT/VECTOR frame_a frame_b [coord_str] [sc_x,sc_y] [scale_r] [pos_range] [sm_par] [head]
PLOT/AXES [x_axis_spec] [y_axis_spec] [sc_x,sc_y] [x_label] [y_label] [x_off,y_off]
PLOT/AXES [coord_str]
PLOT/COLUMN frame [x_coord] [x_sta,x_end] [sc_x,sc_y]
PLOT/CONTOUR frame [coord_str] [x_scale,y_scale] [contours] [sm_par]
PLOT/DESCRIPTOR frame [descriptor] [start,end] [x_scale,y_scale]
PLOT/GRAY frame [coord_str] [x_scale,y_scale] [gray_lev] [sm_par] [gray_ness] [options]
PLOT/HISTOGRAM frame [x_scale,y_scale] [log_flag]
PLOT/HISTOGRAM table column [x_scale,y_scale] [bin [min [max]]] [log_flag]
PLOT/VECTOR frame_a frame_b [coord_str] [sc_x,sc_y] [scale_r] [pos_range] [sm_par] [head]
PLOT/KEY [keyword] [start,end] [x_scale,y_scale]
PLOT/PERSPECTIVE frame [coord_str] [azi_angle,alt_angle] [sm_par] [xy_flag]
PLOT/ROW frame [y_coord] [x_sta,x_end] [sc_x,sc_y]
PLOT/TABLE table [column1] [column2] [x_scale,y_scale]
SET/GRAFIC option1[=value1] [option2[=value2] ...]
SHOW/GRAFIC

```

Table 6.5: Graphic Commands

6.2 Image Displays

This section describes the setup of the image displays used by MIDAS and the functionality provided by MIDAS to interact with these displays. For a description of the conceptual model for an image display device see the definition document for the IDI-routines. MIDAS supports peripheral displays like *e.g.* the Gould IP8000 (former DeAnza) series and XWindow displays. We describe here the IP8500 display specifically and a generic XWindow display.

Note

The Gould image display systems are not manufactured anymore and hardly in use nowadays. Therefore, the 94NOV release of MIDAS will be the last release where we support these devices.

6.2.1 IP8500 display

Each image memory or “channel” has independent scroll and zoom capabilities as well as an intensity transformation table which can be used like a colour look-up table and can also be used to change the output values that are fed to the look-up tables in the video output controller. This allows fast displays of log or histogram equalized images without having to reload the entire image.

One image channel is designated as the graphics (or overlay) channel. This also has its own zoom and scroll capabilities. In addition, the colour of the overlay can be selected. In MIDAS the last available channel is always used as the graphics channel. Also an alphanumeric memory is associated with the image display station.

The video output controller (VOC) selects which image memory is to be displayed on which colour channel as well as performing the task of overlaying the graphics plane. It also takes care of integrating the cursor and alphanumeric data into the video output. Finally, the VOC supports *Split Screen* mode where parts of 2 or 4 image channels are displayed together on the screen.

Using Image Memories

Several image memories are associated with each image display station. Thus it is possible to have several images loaded in the image display at the same time and to switch quickly from one channel to the other. Images can be loaded into any of the image memories. They are referenced by numbers 0, 1, 2, or 3. A typical command sequence would be:

LOAD/IMAGE galaxy 0	! load an image in channel 0
DISPLAY/CHANNEL 0	! display memory channel 0

The most important commands associated with handling the image memories are:

LOAD/IMAGE - load an image into image memory
GET/IMAGE - read an image from the image memory
GET/CURSOR - read pixel values of displayed image
DISPLAY/CHANNEL - display the image in the selected channel
CLEAR/CHANNEL - erase contents of an image channel
SET/SPLIT - enable split screen
CLEAR/SPLIT - disable split screen
ZOOM/CHANNEL - zoom in integer steps 1 to 8
SCROLL/CHANNEL - scroll the image
BLINK - blink between two different image memories
SHOW/CHANNEL - show status of image channel

Look-Up Tables

Look-up tables or LUTs are the tables that map the data in the image memory into colours on the display when the system is used in pseudo-colour mode. Commands exist to load LUTs into the image display, to modify LUTs interactively and to read back LUTs from the image display. Interactive modification is done via the joystick or trackball device.

Some of the existing LUTs are:

```

backgr
color
heat
light
pastel
pseudo1, pseudo2
rainbow, rainbow1 ... rainbow4
random, random1 ... random4
smooth
staircase
stairs8

```

Use the command **TUTORIAL/LUT** to see what some of the available LUTs actually look like and how to modify the LUTs interactively.

The main commands available for handling LUTs are:

LOAD/LUT - load a look-up table
GET/LUT - read back a look-up table
MODIFY/LUT - interactively modify a look-up table
CLEAR/LUT - bypass the look-up table
SET/LUT - pass through a look-up table

DISPLAY/LUT - show the look-up table as a colour bar

CREATE/LUT - create a look-up table using the HSI colour model

Intensity Transformation Tables

The intensity transformation tables (ITTs) come between the image memories and the look-up tables. Using ITTs in the display mode allows special modifications to be applied to the displayed data without modifying the look-up tables or the data in the image memories. Interactive modification is done via the joystick or trackball device.

The main commands which control the ITT functions are:

LOAD/ITT - load an ITT

CLEAR/ITT - bypass the ITT

SET/ITT - pass through an ITT

GET/ITT - read back an ITT

MODIFY/ITT - modify the ITT interactively

To display the ITT, use **DISPLAY/LUT** which shows the combined effect of LUT and ITT. Some of the currently available ITT tables are:

```
ramp
neg
expo
log
neglog
jigsaw
staircase
```

With the command **TUTORIAL/ITT** you can see the effect of ITTs and modify the ITTs interactively.

Using the Cursors

Each image display has two independent cursors available. In addition each cursor shape can be defined. The variety of possibilities available for various cursor forms and types defies a simple explanation here that would make much sense. The interested user is referred to the command **TUTORIAL/CURSOR** which gives a demonstration of the possible cursor shapes.

The cursor(s) are controlled interactively by a special device *e.g.* tracker ball, joystick or mouse. Detailed information on how to operate the cursors can be found in Appendix C. In general there is a switch for each cursor to define its status ON/OFF. When a cursor is active (ON) its position can be read by pressing the **Enter** key. To exit an interactive cursor command one normally has to switch the cursors off and press the **Enter** key.

The commands associated with cursor control operations are:

SET/CURSOR - enable selected cursor
LOAD/CURSOR - load a programmable cursor
GET/CURSOR - read cursor positions
CLEAR/CURSOR - disable cursors

Graphics

Each user has a graphics (or overlay) channel associated with the image display. The currently available commands associated with the use of the overlay channel are listed below:

SET/OVERLAY - enable the overlay of graphics on top of the image
CLEAR/OVERLAY - disable the overlay memory
LOAD/OVERLAY - load a LUT for the graphic channel (for experts only!!)
SCROLL/OVERLAY - scroll the graphics channel with the image channel
ZOOM/OVERLAY - zoom the overlay with the image channel
DRAW/... - draw a geometric shape like CIRCLE, RECTANGLE, etc. in the overlay plane
SET/CHANNEL - designate a channel as image or graphics

Turning off the overlay/graphics via CLEAR/OVERLAY only disables the overlaying of the graphics. To really get rid of what is in the overlay channel you must use CLEAR/CHANNEL OVERLAY.

Furthermore, due to the internal hardware of the IP8500, disabling the overlay will also turn off the visibility of the cursors!

Alphanumerics

The alphanumerics memory is divided up into 22 lines of 80 characters. (see Appendix C). The alphanumeric characters that are available are alphabetical upper case and numbers plus several special characters. Options exist to choose the colour and priority of the alphanumeric display.

Commands associated with alphanumeric display are:

LABEL/DISPLAY - load a string into the alphanumeric memory
CLEAR/ALPHA - clear the alphanumeric display

There is also an option in the LABEL/DISPLAY command to use the overlay channel for text (with higher resolution) instead of the alphanumeric memory.

True Colour or RGB Operations

The IP8500 allows pictures to be displayed in *true* colour using three image memories (channel 0, 1 and 2) simultaneously for the red, green and blue images needed to make up a true colour image. Channel 0, 1 and 2 may also be referred to as Red, Green and Blue. To start using the RGB mode of the display, execute the command:

SET/DISPLAY RGB

The CUTS used for mapping the image into the range of image memory need to be set individually for each of the images to be used. Next, each channel must be loaded individually with the appropriate image. This is done as follows:

LOAD/IMA rpict R	! load the red picture in channel 0
LOAD/IMA gpict G	! load the green picture in channel 1
LOAD/IMA bpict B	! load the blue picture in channel 2

It is now possible to use several of the other commands that control the image display, but they may perform in slightly different ways than when the system is used for pseudo-colour displays. The following comments are intended to give some guidelines.

ZOOM — This command will zoom all three channels together.

SCROLL — This scrolls only one channel at a time. The choice is governed by the input parameter which can be R, G, or B.

SET/SPLIT — This will show the three channels in their native colours.

GET/CURSOR — Generally cursor operations will not perform the operation you expect in RGB mode. Nevertheless, cursor values can be extracted. They will come from the last accessed image channel.

LOAD/LUT — This command should be avoided since LUTs are handled in a very special way in RGB mode.

To exit from the RGB mode, execute the command **SET/DISPLAY PSEUDO**.

6.2.2 XWindow display

With the term XWindow display we refer to a bitmapped screen supporting the XWindow environment. These displays have less functionality provided in hardware than the “classical” peripheral image displays. On the other hand they offer much more flexibility via software. For example, display screens of different sizes may be created and different number of image channels may be connected to any one display.

Another important difference results from the way XWindows works: when an X application program terminates, all the connected windows and data structures disappear. Therefore, MIDAS starts up an independent server process, the *display server*, which owns all X11 related data structures. The MIDAS applications do not interact directly with the windows but send messages to the server which then performs the actual task. Like this we can keep the windows alive while the different applications are executed and terminated, one by one.

Also, keep in mind that all interaction with the display will only work while the input focus is in the display window (either enforced by clicking the mouse in that window or just moving the cursor into it – that depends on how your window manager is set up).

Image displays are created on the screen via the CREATE/DISPLAY command. An “image display” is then represented by a window on the bitmapped screen. It may have one or several image channels associated with it. The image channels may have the same size as the display window or could be larger. These channels are not realised in hardware (*e.g.* video memory) like the peripheral image displays, but exist as data structures in main memory. Also an overlay channel and an alphanumerics memory are emulated for each image display. Initially each display is provided with a grayscale LUT.

You may create several image displays at the same time on your bitmapped screen even though only one display can be the current active display at any time. With the command ASSIGN/DISPLAY you switch from one display to the next.

Each image channel has independent scroll (also emulated in software) but no zoom capabilities. There are special commands like GET/CURSOR and VIEW/IMAGE which provide zoom in a special *zoom window*. Also available is an intensity transformation table but only one per image display and not one per image memory since the ITTs are emulated by convolving the ITT values with the current LUT.

Using Image Memories

Several image memories may be associated with each image display. However, at the moment the blinking between different image memories is so slow that is it not very useful to create image displays with many image channels. Images can be loaded into any of the image memories. They are referenced by numbers 0, 1, A typical command sequence would be:

LOAD/IMAGE galaxy 0	! load an image in channel 0
LOAD/LUT heat	! load a colour look-up table
DISPLAY/CHANNEL 0	! display memory channel 0

The most important commands associated with handling the image memories are:

CREATE/DISPLAY - *create an image display with image channels(s)*
VIEW/IMAGE - *explore an image ...*
LOAD/IMAGE - *load an image into image memory*
GET/IMAGE - *read an image from the image memory*
GET/CURSOR P5=W - *read pixel values of displayed image using a zoom window*
DISPLAY/CHANNEL - *display the image in the selected channel*
CLEAR/CHANNEL - *erase data in an image channel*
SCROLL/CHANNEL - *scroll the image*
SHOW/CHANNEL - *show status of image channel*
DELETE/DISPLAY - *delete display window*

Using the Cursors

Each image display has two independent cursors available. The first cursor (cursor 0) is moved via the mouse, the second cursor (cursor 1) is moved via the *Arrow* keys on the keyboard.

Using both cursors a region of interest (ROI) is supported. The ROI can have rectangular or circular shape and is moved via the mouse, its size is adjusted via the *Arrow* keys. The resizing of the ROI may be done in small or large increments. This is controlled via the number keys 0, 1, 2, ..., 9 on the keyboard. Pressing the 0-key corresponds to an increment of a single screen pixel, whereas 1, 2, ..., 9 lead to larger increments.

The cursor position can be read by pressing the **ENTER** button which is the leftmost button on the mouse. To exit from a command which uses interactive cursor input, press the **EXIT** button which is the rightmost button on the mouse. The middle button is currently not used in MIDAS (it behaves like the **EXIT** button), but may be employed in the future. The **RETURN** key on the keyboard serves as an **EXECUTE** button. The **EXECUTE** button works usually like the **ENTER** button, only in some very special cases its functionality is different from the **ENTER** button. If so, it is explained in the relevant help info, an example are the **MAGNITUDE/...** commands.

Pressing the **ENTER** button on the mouse requires a stable hand. If you press the **ENTER** button and only slightly move the mouse by doing so, this will be interpreted by the display server as a *Cursor Move* instead. Therefore, it may be safer to move the cursor via the mouse but get the cursor input with the **EXECUTE** button.

The commands associated with cursor control operations are:

GET/CURSOR - *read cursor positions*
SET/CURSOR - *define specific cursor shape*

Look-Up Tables

Look-up tables or LUTs are the tables that map the data in the image memory into colours on the display when the system is used in pseudo-colour mode. In contrast to hardware display systems, the size of the LUT is not constant but depends upon how many colours are already used by other X applications running already. Commands exist to load LUTs into the image display, to modify the LUTs interactively and to read back LUTs from the image display.

Some of the existing LUTs are:

```
backgr
color
heat
light
pastel
pseudo1, pseudo2
rainbow, rainbow1 ... rainbow4
random, random1 ... random4
smooth
staircase
stairs8
```

Use the command **TUTORIAL/LUT** to see how some of the available LUTs actually look like and how to modify a LUT interactively.

The main commands available for handling LUTs are:

```
LOAD/LUT - load a look-up table
GET/LUT - read back a look-up table
MODIFY/LUT - interactively modify a look-up table
CLEAR/LUT - bypass the look-up table
SET/LUT - pass through a look-up table
DISPLAY/LUT - show the look-up table as a colour bar
CREATE/LUT - create a look-up table using the HSI colour model
```

Commands like **MODIFY/LUT** need a displacement. Use the *Arrow* keys to move left, right and down (to thin) or up (to thicken) a colour bar. Again the speed of the movements is controlled via the keys 0, 1, ..., 9.

Intensity Transformation Tables

The intensity transformation tables (ITTs) come between the image memories and the LUT and are emulated in the XWindow environment by convolving the ITT values with the current LUT. Interactive modification is done via the arrow keys.

The main commands which control the ITT functions are:

LOAD/ITT - *load an ITT*

CLEAR/ITT - *bypass the ITT*

SET/ITT - *pass through an ITT*

GET/ITT - *read back an ITT*

MODIFY/ITT - *modify the ITT interactively*

To display the ITT, use **DISPLAY/LUT** which shows the combined effect of LUT and ITT. Some of the currently available ITT tables are:

```
ramp
neg
expo
log
neglog
jigsaw
staircase
```

With the command **TUTORIAL/ITT** you can see the effect of ITTs and modify the ITTs interactively.

Graphics

A graphics (overlay) channel is provided with the image display. However, it is not implemented like an image channel but emulated in software. Only drawing functions are supported (i.e. you cannot load an image into the overlay channel).

Three different font sizes are supported for writing text. Fonts can be chosen via the **INITIALIZE/DISPLAY** command.

The currently available commands associated with the use of the overlay channel are:

SET/OVERLAY - *enable the overlay of graphics on top of the image*

CLEAR/OVERLAY - *disable the overlay*

CLEAR/CHANNEL overlay - *clear the overlay channel*

DRAW/... - *draw a geometric shape like CIRCLE, RECTANGLE, etc. in the overlay plane*

LABEL/DISPLAY - *write a string into the overlay channel*

Alphanumerics

Each display window may or may not have an associated alphanumerics memory depending on the options of the **CREATE/DISPLAY** command.

The alphanumerics memory is attached to the bottom of the display window and consists of 3 lines, the number of characters per line depends on the width of the image display. The same three different fonts which are used for the overlay channel can be used for alphanumeric characters.

Commands associated with the alphanumeric memory are:

LABEL/DISPLAY - write a string into the alphanumeric memory
CLEAR/ALPHA - clear the alphanumeric display

Command Interruption

MIDAS commands are usually aborted by entering **Ctrl/C**, but be careful when aborting commands which interact with a display/graphics window, e.g. GET/CURSOR. For you run the risk of losing the synchronisation with the MIDAS display server, which must then be re-initialized via the RESET/DISPLAY command.

Graphical User Interface – XDisplay

There exists a graphical user interface (GUI), XDisplay for interaction with a MIDAS display window which is activated via the command **CREATE/GUI display**. This GUI is aimed at the casual MIDAS user who is not very familiar with the commands which provide specific display-related functions. It supports most of the display related commands of MIDAS and also lists all available LUTs and ITTS.

6.2.3 Image Hardcopy

A hardcopy of a frame shown on the image display or stored on disk can be made only if the site has appropriate devices (see Appendix C which gives the detailed description of the available options).

Grayscale and colour hardcopies can be produced only on Laser printers which support PostScript. Typical command sequences for getting an image hardcopy are:

LOAD/IMAGE frame	! load image into image display
COPY/DISPLAY	! make hardcopy of screen

or

ASSIGN/DISPLAY device	! assign hardcopy device as display
LOAD/IMAGE frame	! load image to hardcopy

where available devices and special parameters are described in Appendix C. In addition, a set of device specific commands are normally defined to set it up and to get status information (see Appendix C).

Chapter 7

Data Exchange Format

This chapter describes how to exchange data between MIDAS and other systems. MIDAS supports the Flexible Image Transport System (FITS) as the official format for data interchange and long term storage. The FITS format is recommended by the IAU for exchange of digital information between astronomical institutes and provides a computer independent description of the data.

It is important to recognise that MIDAS officially only supports the FITS format for data exchange. For backwards compatibility, MIDAS can also read the IHAP format which was used by the old ESO image processing system and some data acquisition systems at La Silla. Other formats are not supported officially and no help will be provided to access such data.

Note

The data representation of internal MIDAS files is machine dependent in order to optimise performance. Further, their layout may change with time. The use of the FITS format for storage and exchange will always ensure that proper conversions are made. This is not the case with other formats. It is safer for you to save your data in FITS format, especially if you intend keeping them for a longer time.

For standard text files such as programs, procedures and ASCII data files, operating systems utilities can be used (e.g. tar). Text files can also be saved as FITS headers using ASCII-catalogues in MIDAS, however, such files can only be decoded by the MIDAS FITS reader.

The MIDAS commands for data exchange (INTAPE, OUTTAPE) can convert the FITS files directly to/from either sequence tape devices or disks. It is therefore also possible to convert to/from FITS on disk and afterward use general utilities (e.g. tar or ftp) to do the actual transfer. In that case, proper options must be used to ensure that the files are transferred in binary mode and with correct blocking factors.

7.1 FITS Format

The FITS format provides a general way to encode both a definition of data and the data themselves in a machine independent form. The original definition is published in a number of articles, see Wells et al. [1], Greisen et al. [2], Grosbøl et al. [3] and Harten et al. [4]. A good general introduction to FITS can be found in [5] while a concise description is given in [6] both available through the NASA FITS Support Office and its anonymous ftp archive at Internet node `nssdca.gsfc.nasa.gov`.

7.1.1 Structure of FITS files

A FITS file contains a sequence of logical header/data units (HDU) which all start with a set of header records describing the following data records. The logical record length of a FITS file is always 2880 bytes of 8 bits. Both header and data sections start in a new logical record. FITS headers are encoded in ASCII as 80 character card images each starting with an 8 character keyword defining the type of information contained on the card. The card images follow each other directly without any end-of-line character which means that many standard text processing tools may have problems. Values of parameters are decoded using standard FORTRAN-77 rules. They describe in detail the data following the header records. Since a single FITS file may have many HDU's each corresponding to a data set (*e.g.* an image or a table), the translation of it may produce several result frames. After the last HDU in the file additional records may exist.

The basic FITS paper [1] specified both a logical and physical record length of 2880 bytes. The increasing volume of data and higher recording densities made this physical record size inefficient. To increase storage efficiency and make use of new recording media such as optical disks and helical scan devices, the FITS standard was extended to allow physical blocking factors different from one [3]. The allowed range of blocking factors is explicitly defined for a given media. For most tape media, factors between 1 and 10 are allowed giving a maximum physical block length of 28800 bytes. Each file terminates with a tape-mark, and the last file on tape terminates with a double tape-mark *i.e.* end of information.

7.1.2 FITS data-types and extensions

The FITS header specifies both the format and size of the data records following. The data representation is defined by the value of the `BITPIX` keyword and can have the values given in Table 7.1.2. The MIDAS data format closest matching it is chosen by default. Single precision real values are used when the explicit scaling is given in the FITS file (*i.e.* by the keywords `BSCALE` and `BZERO`).

When writing FITS files, their data type will also be the one closest to the internal representation except if *basic* FITS formats are explicitly requested by an option. In that case, `BITPIX` will be one of the originally allowed values (*i.e.* 8, 16 or 32) or for tables the ASCII format.

Each HDU in a FITS file will normally correspond to a single MIDAS file. HDU's with no associated data are not stored. Besides the prime HDU which either is a simple

BITPIX value	Data representation	Scaling	MIDAS data type
8	8-bit unsigned integer (ASCII)	No	D_I1_FORMAT
		Yes	D_R4_FORMAT
16	16-bit twos-complement integer	No	D_I2_FORMAT
		Yes	D_R4_FORMAT
32	32-bit twos-complement integer	No	D_I4_FORMAT
		Yes	D_R4_FORMAT
-32	32-bit IEEE floating point	-	D_R4_FORMAT
-64	64-bit IEEE floating point	-	D_R8_FORMAT

Table 7.1: Relation between FITS and MIDAS data types

data matrix or a random group structure a number of extension are defined. Currently, the extensions listed in Table 7.1.2 are translated by MIDAS whereas other 'unknown' extensions are skipped. Text and MIDAS-fit files can also be stored in FITS. The content of these file types is stored as FITS headers using a special MIDAS conversion. Thus, other FITS readers may not be able to retrieve the full information.

7.1.3 FITS keywords

Keywords in the FITS headers are stored in MIDAS files as descriptors. The naming conventions are not identical. It is therefore necessary to perform a mapping between them. In general, the MIDAS descriptors will get the same name as the FITS keyword except that illegal characters for descriptors (e.g. '-') will be substituted by underscore ('_'). Keywords which describe the data structures are handled in a special way. Special name mapping is given in the following sections.

It is simple to map FITS keywords into MIDAS descriptors because the latter can have longer names and store arrays of values. The other way the translation is non-unique. Thus, general MIDAS descriptors are encoded in a special section of the header as *HISTORY* cards. The format used preserves all the descriptor information but is only

FITS Extension type → MIDAS frame	
prime	image (.bdf)
random groups	image (.bdf) + table (.tbl)
TABLE	table (.tbl)
BINTABLE	table (.tbl)
IMAGE	image (.bdf)

Table 7.2: Relation between FITS Extensions and MIDAS frame types

FITS keyword	MIDAS descriptor	FITS keyword	MIDAS descriptor
NAXIS	NAXIS	OBJECT	IDENT
NAXISn	NPIX(n)	DATE-OBS	O_TIME(1)
CRPIXn	REFPIX(n)	DATAMIN	LHCUTS(3)
CRVALn	START(n)*	DATAMAX	LHCUTS(4)
CDELTh	STEP(n)	EPOCH	O_POS(3)
CTYPEn	CTYPE(n)	EQUINOX	O_POS(3)
CROTAn	ROTA(n)		
CUNITn	CUNIT(n+1)		
BUNIT	CUNIT(1)		

Table 7.3: Name translation between standard FITS keywords and MIDAS descriptors. Asterisk indicates that data value is transformed (see text).

known to MIDAS and therefore cannot be decoded by other readers.

Standard keywords

The FITS format specifies a number of “standard” keywords with well defined meaning. They can be classified in two groups namely: a) data structure definition, and b) general description. The first group is translated into special MIDAS descriptors for the data structure in question. The conversion for image files is given in Table 7.1.3. The **START** descriptor of a MIDAS image refers to the first pixel and is computed from the values **CRVAL** and **CRPIX** for the reference pixel as given in the FITS header. The actual reference pixel is saved in the descriptor **REFPIX** which is necessary for non-linear coordinate systems.

Tables have more complicated structures which make the translation less direct. With respect to other keywords, images and tables are treated in the same way. The data types of these descriptors are checked and must correspond to the FITS standard.

Non-standard keywords

Besides the standard FITS keywords defined in the FITS documents, it is possible to use other, non-standard keywords. Such non-standard keywords have no global definition and may be interpreted differently by different FITS readers.

A number of non-standard keywords has been used by ESO *e.g.* for definition of frames acquired at the La Silla observatory. They are given in Table 7.1.3 and should not be used with other definitions.

Due to the freedom to use non-standard keywords it is important to know their exact definitions when transferring data between different systems.

FITS keyword	Data type	MIDAS descriptor	Remarks
MJD-OBJ	Real	O_TIME(4)	Modified Julian Date of start of exposure
TM-START	Real	O_TIME(5)	UT of start of exposure, in sec. after midnight
TM-END	Real	-	UT of end of exposure
EXPTIME	Real	O_TIME(7)	Exposure time in sec.
RA	Real	O_POS(1)	Right Ascension in degrees
DEC	Real	O_POS(2)	Declination in degrees
POSTN-RA	Real	O_POS(1)	Right Ascension in degrees
POSTN-DE	Real	O_POS(2)	Declination in degrees
AIRMASS	Real	O_AIRM	Mean airmass of exposure
FILENAME	String	FILENAME	Original file name
ESO-LOG	String	ESO_LOG	Observing log

Table 7.4: Name translation between special non-standard FITS keywords and MIDAS descriptors.

Hierarchical keywords

To avoid possible misinterpretations and naming conflicts for keywords describing acquisition parameters, ESO has adopted a hierarchical keyword convention for this purpose using the keyword **HIERARCH**. Such keywords have the following syntax:

```
HIERARCH domain level-1 ... level-n keyword = value / comment
```

where the *domain* always is **ESO**. Several hierarchical levels may exist with a keyword and associated value at the lowest level. Naming and parameter follow the general FITS rules when applicable (e.g. max. 8 characters). The levels defined for the **ESO** domain are given in Table 7.1.3 where also an abbreviation character is specified. FITS readers which do not know this convention should save the **HIERARCH**-keywords as comments following standard FITS rules. The standard specification of the FITS Acquisition and Archive format used by ESO can be found in [7]. This document also gives a full definition of all hierarchical keywords in the **ESO**-domain.

The hierarchical structure provides a convenient and clear way to separate information concerning different subsystems. The full concatenated name can however be much longer than the 15 character limit for MIDAS descriptors making it necessary to use a name translation scheme.

Descriptor names of hierarchical keywords start with an underscore character **'.'**. Domain and level names are abbreviated to single characters where **E** is used for the **ESO**-domain. Other abbreviations are listed in Table 7.1.3. Some levels may also have an associated index which is maintained in the translation. The domain and level part is

ESO hierarchical keyword levels					
Level-1	Abbr.	Level-2	Abbr.	Remarks	
GEN	G			General	
		PROJ	P	Project	
		TARG	T	Target	
		EXPO	E	Exposure	
		CAT	C	Catalogue	
		WIND	W	Wind	
TEL	T			Telescope	
		FOCU	F	Focus	
		TRAC	T	Tracking	
		DOME	D	Dome	
		PARANG	P	Parallactic Angle	
		AIRM	A	Airmass	
ADA	A			Adaptor	
		LAMP-#	L#	Lamp no.	
		GUID-#	G#	Guide probe no.	
INS	I			Instrument	
		FOCU	F	Focus	
		COMP	C	Computer	
		MIRR-#	M#	Mirror no.	
		GRAT-#	G#	Grating wheel no.	
		SLIT-#	S#	Slit wheel no.	
		OPTI-#	O#	Optical wheel no.	
DET	D			Detector	
		FRAM	F	Frame	
		SHUT	S	Shutter	
		COMP	C	Computer	
		OUT-#	O#	Output no.	
ARC	C			Archive	

Table 7.5: Levels defined for ESO hierarchical FITS keywords and the abbreviations used for the translations to MIDAS descriptor names

separated with an underscore character from the keyword which follows the rules for non-standard keywords. The hierarchical keyword:

HIERARCH ESO TEL FOCU SCALE = 1.489 / Focus length (deg/m) = 5.36"/mm

will yield the MIDAS descriptor name `_ETF_SCALE` while

HIERARCH ESO INS OPTI-3 ID = 'ESO#427' / Optical element identifier

will become the descriptor `_EIO3_ID`. When writing such descriptors out again, the reverse translation is applied.

7.1.4 Restrictions

The MIDAS reader of FITS tapes accepts most of the FITS formats including standard extensions. The MIDAS implementation has the following restrictions:

- the maximum number of axes in images is 16,
- maximally 512 columns are allowed for tables
- matrix and variable size array conventions for BINTABLE are not available,
- information in headers without associated data is not stored to avoid creation of empty data files (*e.g.* for simple table files).

7.2 IHAP Format

The IHAP format is defined in the IHAP manual (see version April 1990, Appendix A). It is the internal format of the IHAP system and may also be used by some old data acquisition systems at the La Silla observatory.

7.2.1 Translation of IHAP header

The IHAP header contains the description of the data in a fixed, binary format. With reference to the IHAP manual, the translation of this information into MIDAS descriptors are given in Table 7.2.1. IHAP defines the end of exposure which is converted to exposure time. Also, start and end are given for image axes instead of start and number of pixels. In addition to the main header, keywords and comments may exist. They are saved in the descriptor `O_COM` in ASCII format.

7.2.2 Restrictions

The main limitations of the MIDAS reader of IHAP formats are the following:

- only standard image formats are decoded,
- only tapes written with code 1 specifications are decoded.

Other formats such as tables and special data acquisition, can only be read by an IHAP system after which they can be converted to FITS by IHAP itself.

IHAP header	MIDAS descriptor	IHAP header	MIDAS descriptor
DRASC	O_POS(1)	JUSR1	IHAPUSER(1)
DECLN	O_POS(2)	JUSR2	IHAPUSER(2)
JYEAR/JDAT	O_TIME(1)	DUSR1	IHAPUSER(3)
DCRAT	O_TIME(5)	DUSR2	IHAPUSER(4)
DSTTM	O_TIME(6)	AHGC	LHCUTS(2)
DENTM	O_TIME(7)*	ALWC	LHCUTS(1)
JDENT	IDENT	DAIRM	O_AIRM
DXSTR/DYSTR	START		
DXSTP/DYSTP	STEP		
DXEND/DYEND	NPIX*		

Table 7.6: Name translation between IHAP header and MIDAS descriptors. Asterisk indicates that the value is transformed (see text).

7.3 Conversion between FITS and internal format

The FITS conversion commands **INTAPE** and **OUTTAPE** can both access FITS files on disk or directly on tape devices. Thus, it is in principle possible to use general system utilities to transfer the FITS files between disk and tape. When using operating system utilities special care must be taken to ensure that the files on the tape conform to FITS blocking specification - it is by no means obvious!

The syntax of FITS conversion commands is as follows:

```
INTAPE file_list id device [flag]
OUTTAPE cat[,list] device [flag] [dens,block] [type]
```

These are described in more detail in the following sections.

In order to save disk space, it may be desirable to compress the files. For this purpose, operating system utilities may be used (*e.g.* **compress** for UNIX). The steps to achieve this are as follows:

- Use **OUTTAPE** to write file to disk in FITS format.
- Compress FITS file on disk *e.g.* **compress**.
- Use system utility to save compressed file on tape *e.g.* **tar**.

To retrieve files execute the above process in reverse order.

7.3.1 Devices

The *device* in the **INTAPE/OUTTAPE** commands may either refer to a tape device, a disk file or a disk file prefix (possibly including the full path name). If the first 4 characters of the

device name are the string 'tape' (case independent), it will be assumed to be a logical tape name. The environment variable of the same name will then be used for the logical to physical name translation. Otherwise, the name will be assumed to be a physical name.

The name of a tape device may contain the host name in which case a remote device will be accessed. This will work only if the MIDAS tape-server has been installed. The full name is then given as 'host:device' e.g. for a device on a UNIX system *ws1* it could be '*ws1:/dev/nrst8*'. For access to local devices the physical device name can be given directly. When writing on a tape device, it must be set to *write enable* either by inserting a write ring or changing a switch on the cartridge.

A single file can be specified by including its extension otherwise a name is assumed to be a file prefix. When a file prefix is used the full file is made up of the prefix, the four digit sequence number and the file extension '.mt'.

7.3.2 File naming

The file name is generated from the prefix and the sequence number. A FITS file may contain several extensions each of which may result in a MIDAS file. If more than one file of the same type (e.g. image or table) is associated to the same FITS file, an additional letter is appended to the name. The fourth table file in the FITS file '*obs0023.mt*' would get the name '*obs0023c.tbl*'.

The original file name is normally available in the descriptor **FILENAME** which can be used for renaming it (see **RENAME/FILE**). This is not done by default because it could overwrite existing files.

7.3.3 Reading FITS

The **INTAPE**-command is used to convert FITS (or IHAP) files to the internal MIDAS format:

```
INTAPE file_list id device [flag]
```

where **file_list** is a list of absolute position numbers of files (if input is from tape) to be read with 1 being the first file (e.g. 1, 3, 5, 50-60 would read from the 1st, 3rd, 5th file plus files 50 through 60 included), and **id** is an identification prefix of the names of the files created. Tape devices will be rewound if the driver does not maintain the absolute position of the tape. This may take a significant amount of time on some devices. It is easier to read many files from disk (with **INTAPE**) if their names are constructed with the given prefix, a four digit number and the extension '.mt'. It is also possible to read a single file from disk by specifying its full name including extension in which case any extension may be given.

The **flag** is a list of three one character flags which specifies the amount of information listed on the senior terminal and in the log file and the storage format of the data (see help file for full description).

7.3.4 Writing FITS

Translation of MIDAS files to FITS format is done by the OUTTAPE command:

```
OUTTAPE cat[,list] device [flag] [dens,block] [type]
```

where the `cat[,list]` is a catalogue of files to be written with an optional list of numbers (see CREATE/xCAT for creation of catalogues). It can be defaulted by giving either '*' or '?' in which case all files in catalogues set by the SET/xCAT command are written out. A single file can be written by specifying it with its full name including extension. The `device` may specify an actual tape device, a prefix for disk files or an explicit file name for a single file.

The `flag` is an optional list of three one-character flags specifying the append mode, the amount of information listed on the senior terminal and in the log file and if the LHCUTS descriptor in the file should be used for scaling (see help file for full description).

Note

The default options of the OUTTAPE command will start at the current tape position. This may over-write previous data on the tape. Be sure to use the append flag if files have to be added to the tape. Or position the tape at end-of-information using operating system utilities before writing new files.

The `dens,block` parameter can specify the tape density (e.g. 1600 or 6250bpi) and a physical blocking factor in the range 1-10. By default, a blocking factor of 10 is used. Note that some old FITS readers may not be able to read blocked FITS files (e.g. IHAP) in which case a blocking on 1 must be given explicitly. The tape density is used only for 9-track 1/2-inch tapes. Some 1/2-inch tape devices require the density also to be set manually on the drive.

The `type` flag is used to specify the type of FITS format to write where 'B' indicates basic FITS i.e. with integer format only. The default is '0' for original including floating point representation.

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Chapter 8

Fitting of Data

This chapter deals with the modelling and the analysis of image and table data by fitting non-linear functions, using least squares approximation. The different non-linear least squares methods implemented in MIDAS are first shortly described and discussed. The MIDAS commands dealing with functions or linear combination of functions and with the modelling process are then presented.

The basic scheme under these commands is to provide the necessary tools to define the functions entering in the fit, to give initial guesses for the parameters and, in iterations controlled by the user, find the optimal parameters of the functions. These parameters can be used to generate fitted data either as images or as columns in tabular form.

Due to the nature of the methods, it is recommended to use these commands in fitting problems involving small amounts of data. For analysis involving large amounts of data, like full CCD images, there are algorithms, in the context of 2D-photometry, optimized for special purpose analyses. A tutorial command (TUTORIAL/FIT) has been introduced in order to show the capabilities of the package.

A brief description of the implemented methods is included in section 8.1. Section 8.2 describes how to specify functions in the fit. Section 8.3 describes how to include external functions. The usage of the commands is illustrated in section 8.4. The output of the programs and their possible interpretation are discussed in section 8.5. An example is presented in section 8.6, it may be convenient for first time users to run the command TUTORIAL/FIT while reading this section. Section 8.7 contains a summary of the commands. Finally, the functions supported in the current version are listed in section 8.8. References can be found in section 8.9.

8.1 Outline of the Available Methods

Let $y(x, a)$ be a function where $x = (x_1, \dots, x_n) \in \mathbb{R}^n$ are the independent variables and $a \in A \subset \mathbb{R}^p$ are the p parameters lying in the domain A . If A is not the whole space \mathbb{R}^p , the problem is said to be constrained.

If a situation can be observed by a set of events $(y^{(i)}, x^{(i)}) i = 1, \dots, m$, i.e. a set of couples representing the measured dependant and variables, it is possible to deduce the

value of the parameters of your model $y(x, a)$ corresponding to that situation. As the measurements are generally given with some error, it is impossible to get the exact value of the parameters but only an estimation of them. Estimating is in some sense finding the most likely value of the parameters. Much more events than parameters are in general necessary.

In a linear problem, if the errors on the observations have a gaussian distribution, the "Maximum Likelihood Principle" gives you the "best estimate" of the parameters as the solution of the so-called "Least Squares Minimization" that follows:

$$\min_{a \in A} \chi^2(a)$$

with

$$\chi^2(a) = \sum_i w^{(i)} [y^{(i)} - y(x^{(i)}, a)]^2$$

The expected variance of the so-computed estimator is minimum among all approximation methods and is therefore called in statistics an "efficient estimator".

The quantities

$$r^{(i)}(a) = \sqrt{w^{(i)}} [y^{(i)} - y(x^{(i)}, a)]$$

are named the residuals and $w^{(i)}$ the weight of the i^{th} observation that can be, for instance, the inverse of the computed variance of the observation.

If $y(x, a)$ depends linearly on each parameter a_j , the problem is also known as a Linear Regression and is solved in MIDAS by the command REGRESSION. This chapter deals with $y(x, a)$ which have a non-linear dependance in a .

Let us now introduce some mathematical notations. Let $g(a)$ and $H(a)$ be respectively the gradient and the Hessian matrix of the function $\chi^2(a)$. They can be expressed by

$$g(a) = 2 J(a)^T r(a) \quad \text{and}$$

$$H(a) = 2 (J(a)^T J(a) + B(a))$$

where $r(a)$ is the residuals vector

$$r(a) = (r^{(1)}(a), \dots, r^{(m)}(a)) ,$$

$J(a)$ the Jacobian matrix of $r(a)$ i.e.

$$J(a)_{ij} = \frac{\partial r^{(i)}}{\partial a_j}$$

and $B(a)$ is

$$B(a) = \sum_i r^{(i)}(a) H_i(a)$$

with $H_i(a)$, the Hessian matrix of $r^{(i)}(a)$.

In the rest of the chapter, all the functions are supposed to be differentiable if they are applied the derivation operator even when this condition is not necessary for the convergence of the algorithm.

A certain number of numerical methods have been developed to solve the non-linear least squares problem, four have so far been implemented in MIDAS. A complete description of these algorithms can be found in [1] and [3], the present document will only give a basic introduction.

8.1.1 The Newton-Raphson Method.

This is the simplest one. The necessary condition for the function $\chi^2(a)$ to have an extremum is that the partial derivatives vanish i.e.

$$\sum_i r^{(i)} \frac{\partial r^{(i)}}{\partial a_j} = 0 \quad (j = 1, \dots, p)$$

or, equivalently,

$$J(a)^T r(a) = 0$$

This is usually a system of non-linear equations that, numerically, can be solved using the Newton-Raphson's method also called in the one-dimensional case the tangents method. The Taylor development of the function limited to the first order is taken around some initial guesses of the parameters. The resulting linear system

$$J(a^{(k)})^T J(a^{(k)}) \Delta a^{(k)} = -J(a^{(k)}) r(a^{(k)})$$

gives thus a correction to the solution and

$$a^{(k+1)} = a^{(k)} + \gamma \Delta a^{(k)}$$

is taken as the new approximation of the optimum. The relaxation factor γ is a parameter of the method. The convergence of the process towards the solution of the non-linear minimization problem has been proven for locally convex $\chi^2(a)$ or under other assumptions impossible to detail here. These conditions are not generally fulfilled in real problems. Moreover, the algorithm ignores the second order conditions and therefore, may end on a saddle point or never converge. Two different relaxation factors may lead to different solutions or one may give convergence and the other one not. No general rule can be given for the choice of a good relaxation factor.

8.1.2 The Modified Gauss-Newton Method.

From a first guess of the parameters $a^{(1)}$, a sequence $a^{(2)}, a^{(3)}, \dots$ is generated and is intended to converge to a local minimum of $\chi^2(a)$. At each iteration, one computes

$$a^{(k+1)} = a^{(k)} + \alpha^{(k)} d^{(k)}$$

where $d^{(k)}$ is a certain descent direction and $\alpha^{(k)}$ is a real coefficient which is chosen such that $\chi^2(a^{(k)} + \alpha^{(k)} d^{(k)})$ is approximately minimum. The direction $d^{(k)}$ is ideally the solution of the Newton equation

$$H(a^{(k)}) d^{(k)} = -g(a^{(k)})$$

which can also be rewritten

$$[J(a^{(k)})^T J(a^{(k)}) + B(a^{(k)})] d^{(k)} = -J(a^{(k)}) r(a^{(k)}) .$$

Neglecting the second derivatives matrix $B(a^{(k)})$, we obtain the “normal equations” and the Gauss–Newton direction

$$J(a^{(k)})^T J(a^{(k)}) d^{(k)} = -J(a^{(k)}) r(a^{(k)})$$

This so-called Gauss–Newton method is intended for problems where $\|B(a)\|$ is small. If the Jacobian $J(a)$ is singular or near singular or if $\|r(a)\|$ is very large (the so-called large residuals problem), the Gauss–Newton equation is not a good approximation of the normal equations and the convergence is not guaranteed.

The algorithm implemented here is a modification of that Gauss–Newton method, that allows convergence even for rank deficient Jacobians or for large residuals. The Gauss–Newton direction is computed in $V_1 = \text{Im} [J(a^{(k)})^T J(a^{(k)})]$, the invariant space corresponding to the non-null eigenvalues. A correction is taken in V_2 , the orthogonal of V_1 , according to the second derivatives if the decrease of the objective function at the last iteration is considered too small. The Hessian matrix is estimated using finite differences of the gradient.

This method requires the availability of the derivatives and as the number of gradient evaluations is almost p at each iteration, it is recommended for problems with a small number of parameters, let us say $p \leq 10$

8.1.3 The Quasi–Newton Method.

This is identical to the modified Gauss–Newton method, except in the way that the Hessian matrix is approximated.

This matrix is first initiated to zero. At each iteration, a new estimation of the Hessian is obtained by adding a rank one or two correction matrix to the last estimate such that $H^{(k+1)}$, the estimate of the Hessian matrix at the $k+1^{\text{th}}$ iteration, satisfies

$$(J(a^{(k+1)})^T J(a^{(k+1)}) + H^{(k+1)}) (x^{(k+1)} - x^{(k)}) = J(a^{(k+1)}) r(a^{(k+1)}) - J(a^{(k)}) r(a^{(k)})$$

The so-called BFGS updating formulas are applied in this algorithm

$$H^{(0)} = 0 \quad H^{(k+1)} = H^{(k)} + C^{(k)}$$

$$C^{(k)} = \frac{1}{\alpha^{(k)} y^{(k)T} d^{(k)}} y^{(k)} y^{(k)T} - \frac{1}{d^{(k)T} W^{(k)} d^{(k)}} W^{(k)} d^{(k)} d^{(k)T} W^{(k)}$$

where

$$W^{(k)} = J(a^{(k+1)})^T J(a^{(k+1)}) + H^{(k)}$$

and

$$y^{(k)} = J(a^{(k+1)}) r(a^{(k+1)}) - J(a^{(k)}) r(a^{(k)}) ,$$

please see Gill, Murray and Pitfield (1972) for more details. After some iterations and around the optimum, $H^{(k)}$ converges to the Hessian.

This method requires the knowledge of the derivatives and, as the gradients are only computed once per iteration and consequently, the Hessian is more roughly approximated than with the modified Gauss-Newton method, this is better designed for a great number of parameters i.e. $p > 10$.

8.1.4 The Corrected Gauss-Newton No Derivatives.

This method is identical to the Gauss-Newton method where the Jacobian is estimated by finite differences and the Hessian by second order differences.

It does not require the programming of the derivatives but makes a lot of function computations. Its use has to be restricted to problems where the derivatives are really too difficult to write. It is slower and less precise than the two last algorithms.

8.2 Function Specification

The functions to be fitted to data are linear combinations of a set of, so called, "basic" functions. Basic functions are either defined in the system or defined by the user as external FORTRAN routines. The actual combination of basic functions is defined via an interactive editor, (MIDAS command **EDIT/FIT**)

Basic functions are specified by the name, the independent variable(s) and parameter(s), with optional guesses for the parameters, following the syntax:

```
name(var1[...];par1[...]) [par1=value] ...
```

The function name **name** defines the basic function used, it can be a system function, as defined in the table 8.1, or a external function with name **USER00**, ..., **USER09**. In this case, the corresponding file(s) **USER00.FOR**, ..., **USER09.FOR** will exist in the working area and will contain the definition of the routines following the syntax described in the next section.

The number of independent variables of the function is determined by the string **var1[...]**. The actual names of the independent variables are considered as dummy names but their number has to coincide with the actual number of parameters of the function. All the functions defined in a given fit must have the same number of independent variables.

Parameters are defined by unique names after the semicolon in the function specification. Parameters are interpreted according to the position and to the number of independent variables in the function.

A parameter is generally given a first guess on the same line, as **par1=value**, it can also be fixed to a given value or kept proportional to another parameter. The parameter is defined as fixed with the symbol **@** immediately following the value as **par1=value@**. Linear constrains between parameters are defined as **pari=parj*value** or **pari=parj/value**.

According to these rules, a one dimensional gaussian function is specified with the **EDIT/FIT** command as

```
1 GAUSS(X;A,B,C) A=10. B=3200.0 C=1.
```

where **X** is the dummy name of the only independent variable, the first parameter, defining the maximum of the function, is called **A**, initialized to 10, the second parameter, defining the position of the gaussian, is called **B**, and its fixed value is 3200 in world coordinates, and the FWHM is the parameter **C**, with initial value 1.

A linear combination of a gaussian and a Cauchy distribution, centered at the same position is specified as

```
1 GAUSS(X;A1,B1,C1) A1=10. B1=3200. C1=1.
2 CAUCHY(X;A2,B2,C2) A2=A1/10. B2=B1 C2=4.
```

in this case, the maximum of the Cauchy distribution is determined by the corresponding parameter of the Gaussian.

We include in table 8.1 a summary of the system basic functions; the actual mathematical expressions, with the meaning of the function parameters are given in section 8.8.

POLY(X;A,B,...)	polynomial (1D, 2D)
LOG(X;A,B,C)	natural logarithm
EXP(X;A,B,C)	exponential
SIN(X;A,B,C)	sinus
TAN(X;A,B,C)	tangent
SINC(X;A,B,C)	sinc
SINCS(X;A,B,C)	sinc square
GAUSS(X;A,B,C)	(FWHM) Gaussian distribution (1D)
GAUSS(X,Y;A,B,C,D,E,F)	(FWHM) Gaussian distribution (2D)
GAUSSA(X;A,B,C)	(Standard) Gaussian distribution (1D)
CAUCHY(X;A,B,C)	Cauchy distribution (1D)
CAUCHY(X,Y;A,B,C,D,E,F)	Cauchy distribution (2D)
LORENTZ(X,Y;A,B,C,D,E,F)	Modified Cauchy (Lorentz) distribution
LAPLACE(X;A,B,C)	Laplace distribution
TRIANG(X;A,B,C)	Triangular distribution
POISSON(X;A,B,C)	Poisson distribution
IGAUSS(X;A,B,C)	Integrated (FWHM) Gaussian distribution (1D)
IGAUSSA(X;A,B,C)	Integrated (Standard) Gaussian distribution (1D)

Table 8.1: Basic Fit Functions

8.3 External Functions

If the set of basic functions provided by the system as listed below is not sufficient for your own purpose, it is possible to define user functions. To do this, the user has to provide the code of the function(s) as a FORTRAN routine, in his own area, in files named **USER00.FOR**, ..., **USER09.FOR**. The command **CREATE/FUNCTION** will compile the routine(s) and link them with the corresponding system programs (primitives). A library with the local definitions of the routines **USER00**, ..., **USER09** and the executable code will

be created in the user area. With this scheme, it is possible to fit the external functions **USER0i** as if they where basic functions.

Here is a template to write a user defined function:

```

C+
C.NAME
C      USER0i
C
C.DESCRIPTION
C      ...
C
C.INPUT ARGUMENTS:
C      NIND      INTEGER      Number of independent variables
C      X (NIND)  REAL        Array of NIND elements with the
C                           independent variable
C      NPAR      INTEGER      Number of parameters
C      PARAM (NPAR) DOUBLE PRECISION  Array of NPAR elements with the
C                           values of the parameters
C
C.OUTPUT ARGUMENTS:
C      Y          DOUBLE PRECISION  Value of the function
C      DERIV (NPAR) DOUBLE PRECISION  Array of NPAR elements with the
C                           partial derivatives of the
C                           function for each parameter
C-
      SUBROUTINE USER0i(NIND,X,NPAR,PARAM,Y,DERIV)
      IMPLICIT NONE
C      ..
C      .. Scalar Arguments ..
      INTEGER NIND,NPAR
      DOUBLE PRECISION Y
C      ..
C      .. Array Arguments ..
      REAL X(NIND)
      DOUBLE PRECISION DERIV(NPAR),PARAM(NPAR)
C      ..
C      .. Local Scalars ..
C      ..
C      .. Local Arrays ..
C      ..
C      .. Executable Statements

      RETURN
      END

```

The variable Y must contain the value of the basic function at the parameter value PARAM and the array DERIV has to receive the value of the partial derivatives, except if the method used is CGNND (the abbreviation of Corrected Gauss–Newton No Derivative). In the user functions, it is recommended to scale the parameters in such a way that their absolute values lies in a small scale range let us say in $[0.1 + 10.]$. It is advised to use

this scheme to test and debug new functions that can later on be included in the system supported set.

8.4 The Fitting Process.

The typical sequence of operations for a FITting process would be first to create the approximating function, to choose relatively to your problem and your needs the FIT options, then to execute the real least squares approximation and finally to store and view the results. This corresponds to the typical sequence of MIDAS instructions :

```
EDIT/FIT fitname
SET/FIT options
FIT/... nfeval,prec image (or table and cols)
COMPUTE/FIT output
```

The commands have been designed so that defaults exists for almost all the parameters, (see description in Volume C of the MIDAS User Guide).

EDIT/FIT has been described in the last section.

The MIDAS command SET/FIT is used to specify the different options of the FIT command, for instance the method to be applied or the type of used functions. The instruction

```
SET/FIT METHOD=CGNND PRINT=1 WEIGHT=S FUNCT=BLACBODY FCTDEF=USER
```

declares that the Corrected Gauss-Newton no derivatives method is to be applied, that at each iteration, a display of the intermediate result will be performed, that the weighting factors are statistical, that the name of the approximating function is BLACBODY, and that this BLACBODY function which contains user defined functions has already been built in the user area. The appendix or the MIDAS interactive HELP facility will give you the complete description of the SET/FIT command.

The command SHOW/FIT displays the actual selected FIT options.

The FIT instruction is performing the least squares approximation itself. It has a slightly different syntax if the fitting concerns a table or an image.

```
FIT/TABLE nfeval[,prec,[metpar]] table :depcol[,:wgt] :indcol, ...
FIT/IMAGE nfeval[,prec,[metpar]] image
```

nfeval is the maximum number of function evaluations that can be performed, **prec** is the precision on the parameters i.e. the program stops if

$$\|a^{(solution)} - a^{(found)}\| \leq prec (1 + \|a^{(found)}\|) ,$$

and **metpar** are the specific method parameters (for instance in NR : the relaxation factor). The latter have not generally to be given as they can be deduced by the program. For stiff problems, they can thus be overwritten by the user. Any non-given parameter is defaulted. Consult the appendix or use the MIDAS interactive HELP to get a complete description. For instance, the instruction

FIT/IMAGE 100,0.001 PROFILE

executed after the preceding SET/FIT, will execute a non-linear least squares approximation using the CGNND method. The program will stop if more than 100 computations of the approximating function have been performed or if the solution has been found with a precision of 10^{-3} .

8.5 Outputs

To check any typing error or missing specification, first are displayed the options, the required precision, the maximum number of function evaluations and the method parameters.

The frequency of the intermediate displays are controlled by the SET/FIT PRINT=iter. It includes the display of the iteration number, the actual number of function evaluations, the sum of the squares of the residuals, the so-called reduced chi, the percentage of decrease of the reduced chi since last iteration, and, except for the NR method, the norm of the gradient and the dimension of the space V_1 spanned by the Jacobian. The reduced chi square is the

$$\frac{\chi^2(a^{(k)})}{\text{degree of freedom}}$$

In any case, this is followed by the value of the parameters. Except for the NR method, the value of the gradient and the singular values of the Jacobian matrix are added.

At the end, a diagnostic message telling you if the convergence was reached or if any numerical failure occurred during the algorithm. The different messages are:

```
--> METHOD : Convergence achieved <--  
*** ERR-1-METHOD : Bad initializations ... Aborting ***  
*** ERR-METHOD : Likely an error in forming the derivatives ***  
*** ERR-NR : Problems in inverting matrix ***  
*** WARN-2-METHOD : No convergence reached ***  
*** WARN-3-METHOD : Final parameters not really satisfactory ***  
*** ERR-4-METHOD : No convergence in singular value decomposition ***  
*** WARN-5-METHOD : parameters only a good estimation ***  
*** ERR-i-METHOD : Final parameters are not satisfactory ***
```

In the last message i varies from 6 to 8 and the greater i is, the less reliable are the final value of the parameters. For warnings and errors numbered more than 3, it is recommended to perform another FITting with different initial guesses. If warning 2 is displayed, do again the FITting starting with the last computed value of the parameters (nfeval < 0 in the FIT/... command). If an error in the derivatives is reported, check your user functions code.

The diagnostic will be followed by the covariance matrix if you set iter to a negative value in SET/FIT PRINT=.

Finally, the found optimal value of the parameters with their estimated standard deviation are listed.

8.6 Tutorial

A tutorial procedure (TUTORIAL/FIT) shows how to use the fitting package in the simple case of a 1D-image consisting of two overlapping gaussians on top of a non-linear background with additional noise. It is recommended to run the tutorial while reading this section and if possible, on a graphic terminal.

Two functions are copied into your area if you run the example: TEST to generate the artificial data, and FUNCTION with the "model" to be fitted.

The artificial image, to be used in the example, is created as follows: First it creates a reference image, called REF, to provide the definition interval of the independent variable. Then the command COMPUTE/FUNCTION creates the 1D frame with the gaussian profiles on top of the background. Finally, some noise is added to the data. The resulting frame, PROFILE is displayed on the graphic screen.

Now the 'model' FUNCTION will be fitted to the frame PROFILE, using the command FIT/IMAGE. The 'model' was copied into your area already, but you could define it using the editor as:

EDIT/FIT FUNCTION

This command allows you to create or modify FIT-files (ref Volume C of the MIDAS User guide and Appendix C for use of the EDIT command on your terminal). In our example, the user will edit the three basic functions e.g. as follows:

```
1 GAUSS(X;A1,A2,A3) A1=50. A2=95. A3=45.
2 GAUSS(X;A4,A5,A6) A4=A1 A5=135. A6=A3
3 POLY(X;A,B,C) A=0. B=0. C=0.
```

where two of the parameters of the second gaussian, height and FWHM, are related to the parameters of the first Gaussian.

The different fitting methods are then successively applied, changing regularly the options through the SET/FIT. The exact sequence of instructions is:

```
SHOW/FIT
FIT/IMAGE 11,1.,0.5 PROFILE FUNCTION
SET/FIT METHOD=CGNNND
SET/FIT PRINT=0
FIT/IMAGE 30,.5 PROFILE FUNCTION
SET/FIT METHOD=QN
SET/FIT PRINT=-4
FIT/IMAGE 30,.5 PROFILE FUNCTION
SHOW/FIT
SET/FIT METHOD=MGN
FIT/IMAGE 30,.5 PROFILE FUNCTION
```

It is possible to compare efficiency, precision and effects.
Finally, the fitted result is computed as:

COMPUTE/FIT FITTED = FUNCTION

and the fitted frame is plotted on top of the original data.

Individual components of the fit can be selected with the command **SELECT/FUNCTION**. In the example, the sequence of commands

```
SELECT/FUNCTION FUNCTION 1,3
COMPUTE/FIT FIT1 = FUNCTION
SELECT/FUNCTION FUNCTION 2,3
COMPUTE/FIT FIT2 = FUNCTION
SELECT/FUNCTION FUNCTION ALL
```

is used to compute the two gaussian components on top of the background. The results are also plotted.

The full compatibility between image and tabular formats for input and output means that, in our example, the fitted parameters can be used to compute fitted values in a table, using the **COMPUTE/FIT** command as follows:

```
COMPUTE/FIT table :outcol = fitname(:incol)
```

where **table** is the name of a table containing the independent variable in the column **:incol**, fitted values are stored in the column **:outcol**.

To begin with, it is advised to consult the appendix or use the MIDAS interactive **HELP** about **EDIT/FIT**, **SET/FIT**, **CREATE/FUNCTION**, **REPLACE/FUNCTION**, **FIT**, **FIT/IMAGE**, **FIT/TABLE**, **COMPUTE/FIT**.

8.7 Command Summary

Table 8.2 summarizes the commands which are implemented in the context of functions and least squares fitting.

8.8 Basic Functions

8.8.1 Polynomials (1D and 2D)

$$\begin{aligned} POLY(x; a, b, c, \dots) &= a + bx + cx^2 + \dots \\ POLY(x, y; a, b, c, \dots) &= a + bx + cy + \dots \end{aligned}$$

8.8.2 Logarithmic and Exponential Function

$$\begin{aligned} LOG(x; a, b, c) &= a \ln(b + cx) \\ EXP(x; a, b, c) &= a \exp(b + cx) \end{aligned}$$

```

COMPUTE/FIT outima [= funct[(refima)]]
COMPUTE/FIT table:out[,:error] [= funct[(:,col1,...)]]
COMPUTE/FUNCTION outima = funct(refima)
COMPUTE/FUNCTION table:out = funct(:,col1,...)
CREATE/FUNCTION userfunc1[...]
EDIT/FIT [funct]
FIT/IMAGE [nfeval[,prec[,metpar]]] [image[,wgt]] [funct]
FIT/TABLE [nfeval[,prec[,metpar]]] table dep[,wgt] ind [funct]
MODIFY/FIT table seqno [funct]
REPLACE/FUNCTION userfunc1[...]
SAVE/FIT table seqno [funct]
SELECT/FUNCTION funct number[...]
SELECT/FUNCTION funct ALL
SET/FIT [METHOD=mname] [PRINT=iter] [WEIGHT=wgttyp] [FUNCT= fname] [FCTDEF=where]
SHOW/FIT

```

Table 8.2: Fitting Commands

8.8.3 Trigonometric Functions

$$SIN(x; a, b, c) = a \sin(b + cx)$$

$$TAN(x; a, b, c) = a \tan(b + cx)$$

8.8.4 Sinc and Sinc Square

$$SINC(x; a, b, c) = a \sin(b + cx)/(b + cx)$$

$$SINCS(x; a, b, c) = a \sin^2(b + cx)$$

8.8.5 Distributions

$$GAUSS(x; a, b, c) = a \exp \left[-\ln 2 \left(\frac{2(x-b)}{c} \right)^2 \right]$$

$$GAUSS(x, y; a, b, c, d, e, f) = a \exp \left[-\ln 2 \left(\frac{(x-b)^2}{d^2} + \frac{(y-c)^2}{e^2} - \frac{2f(x-b)(y-c)}{de} \right) \right]$$

$$GAUSSA(x; a, b, c) = \frac{a}{\sqrt{(2\pi)c}} \exp \left[-\frac{1}{2} \left(\frac{(x-b)}{c} \right)^2 \right]$$

$$IGAUSS(x; a, b, c) = a \int_{-\infty}^x \exp \left[-\ln 2 \left(\frac{2(u-b)}{c} \right)^2 \right] du$$

$$IGAUSSA(x; a, b, c) = \frac{a}{\sqrt{(2\pi)c}} \int_{-\infty}^x \exp \left[-\frac{1}{2} \left(\frac{(u-b)}{c} \right)^2 \right] du$$

$$CAUCHY(x; a, b, c) = a \left[1 + \left(\frac{2(x-b)}{c} \right)^2 \right]^{-1}$$

$$LORENTZ(x; a, b, c, d) = a \left[1 + \left(\frac{2(x-b)}{c} \right)^2 \right]^{-d}$$

$$POISSON(x; a, b, c) = \frac{ab^x \exp(-b)}{\Gamma(x+1)}$$

$$LAPLACE(x; a, b, c) = a \exp \left[-\ln 2 \left(\frac{2|x-b|}{c} \right) \right]$$

$$TRIANG(x; a, b, c) = a \left(1 - \frac{|x-b|}{c} \right)$$

8.9 References

A good introduction to optimization theory and a description of the the modern minimization techniques can be found in Gill, Murray and Wright [1]. Bard [2] deals with the particular problem of parameter estimation; chapters concerning the different estimators and their properties, and the interpretation of the estimates are remarkable. Updating formulas for Quasi-Newton methods are discussed in [4].

The reading of the cited chapters of [2] will allow an error-free interpretation of the results of the optimization algorithms. It is therefore recommended.

- [1] Gill P.E., Murray W and Wright M.H. . Practical Optimization. Academic Press. London. 1981.
- [2] Bard Y. . Non--linear Parameter Estimation. Academic Press.

London. 1974.

[3] Gill P.E., Murray W. . Algorithms for the solution of non-linear least squares problems. SIAM J. of Num. An., vol 15, pp 977-992, 1978

[4] Gill P.E., Murray W. and Pitfield . The implementation of two revised algorithms for unconstrained optimization. Rep. NAC 11. Nat. Phys. 1972. Lab., Teddington. England.

15-January-1988

Appendix A

Command Summary

Below is an alphabetical list of all basic MIDAS commands. The following abbreviations have been used where appropriate for the command parameters:

input image:	<i>infr infr_1, infr_2...</i>
output image:	<i>outfr outfr_1, outfr_2...</i>
input tables:	<i>intab intab_1, intab_2...</i>
output tables:	<i>outtab outtab_1, outtab_2...</i>
display channel:	<i>chan chan_1</i>
device:	<i>device</i>

A.1 Core Commands

@@ proc <i>[par1] ... [par8]</i>	execute a MIDAS procedure
@ proc <i>[par1] ... [par8]</i>	execute a procedure in MID_PROC (MIDAS core procedures)
@a proc <i>[par1] ... [par8]</i>	execute a procedure in APP_PROC (MIDAS application procedures)
@s proc <i>[par1] ... [par8]</i>	execute a procedure in STD_PROC (MIDAS standard reduction procedures)
@c proc <i>[par1] ... [par8]</i>	execute a procedure in CON_PROC (MIDAS contributed procedures)

ADD/ACAT <i>[cat_name] frame_list</i>	add entries to an ASCII file catalog
ADD/FCAT <i>[cat_name] file_list [lowstr,histr]</i>	add entries to a fitfile catalog
ADD/ICAT <i>[cat_name] frame_list [lowstr,histr]</i>	

add entries to an image catalog
ADD/TCAT [*cat_name*] *table_list* [*lowstr,histr*]
 add entries to a table catalog
ALIGN/CENTER *inframe* *refframe* *incent_x,-y* *refcent_x,-y*
 compute start coordinates for inframe to match with refframe center
ALIGN/IMAGE *intab* *reftab* [*option*] [*overlay_flag*] [*residual_flag*]
 compute transformation coefficients for two rotated images
APPLY/CONVERSION *IMTB* *ima* *tab* *threshold*
 convert a "mask" image to a table
APPLY/CONVERSION *TBIM* *tab* *ima* *npx1,npx2* *sta1,sta2,stp1,stp2* *bg,fg*
 convert a table to a "mask" image
APPLY/EDGE *inframe* *outframe* [*thresh*]
 do edge detection on an image
APPLY/MAP *outframe* = *inframe* *mapframe* *control_flags*
 use an image frame like a Lookup Table
APPLY/THIN *inframe* *outframe*
 apply thinning algorithm
ASSIGN/DEFAULT
 assign default devices
ASSIGN/DISPLAY [*dev*] [*file_name*]
 define output device for display
ASSIGN/GRAPHICS [*device*] [*option*]
 define the graphic device for plot output
ASSIGN/INPUT [*dev*] [*file_name*]
 define input device for writing
ASSIGN/PRINT [*dev*] [*file_name*]
 define output device for printing
AVERAGE/AVERAGE [*in_specs*] [*out_specs*] [*out_opt*] [*draw_flag*]
 compute average over subimage
AVERAGE/COLUMN *out* = *in* [*start,end*] [*SUM*]
 average image columns
AVERAGE/IMAGES *out* = *in_specs* [*merge*] [*null*] [*av_option*] [*dat_intval*]
 average images
AVERAGE/KAPPA [*in_specs*] [*out_specs*] [*out_opt*] [*draw_flag*] [*no_iter*]
 compute average (kappa-sigma clipping) over subimage
AVERAGE/MEDIAN [*in_specs*] [*out_specs*] [*out_opt*] [*draw_flag*]
 compute average (median value) over subimage
AVERAGE/ROW *out* = *in* [*start,end*] [*SUM*]
 average image rows
AVERAGE/WEIGHTS *out* = *in_specs* [*merge*] [*null*] [*av_option*] [*dat_intval*]
 average weighted images
AVERAGE/WINDOW *out* = *in_specs* [*meth*] [*bgerr,snoise*]
 compute average of (consistent) pixel values

BLINK/CHANNEL [*chan1,chan2,...*] [*intval*]
 blink between Image Display channels
 BYE [*proc*]
 terminate a MIDAS session + return to the host system

CENTER/GAUSS [*in_specs*] [*out_specs*] [*out_opt*] [*curs_specs*]
 [*wsize*] [*zw_option*] [*invert_flag*]
 find intensity weighted center
 CENTER/IQE [*in_specs*] [*out_specs*] [*out_opt*] [*curs_specs*]
 [*wsize*] [*zw_option*] [*invert_flag*]
 find intensity weighted center + get angle of major axis
 CENTER/MOMENT [*in_specs*] [*out_specs*] [*out_opt*] [*curs_specs*]
 [*wsize*] [*zw_option*] [*invert_flag*]
 find intensity weighted center
 CHANGE/DIRECTORY *direc*
 change the default (current) directory for MIDAS
 CLEAR/ACAT
 deactivate the ASCII file catalog
 CLEAR/ALPHA
 clear the alpha-numerics memory of the image display
 CLEAR/BACKGROUND
 put Midas session into "foreground" mode
 CLEAR/BUFFER
 clear the command buffer + reset command numbers
 CLEAR/CHANNEL [*chan1*]
 clear + initialize memory channel
 CLEAR/CONTEXT [*context*]
 remove command definitions of a context
 CLEAR/CURSOR
 disable cursors
 CLEAR/DISPLAY
 reset image display
 CLEAR/FCAT
 disable automatic catalog functions for fit files
 CLEAR/GRAPHIC
 erase the screen of the graphic window or terminal
 CLEAR/ICAT
 disable automatic catalog functions for image frames
 CLEAR/ITT [*chan1*]
 bypass ITT on display of memory
 CLEAR/LUT [*screen_segm*]
 bypass LUT in screen_segment on image display
 CLEAR/OVERLAY

disable graphics/overlay plane of display
CLEAR/SCROLL [chan1]
 reset scroll values
CLEAR/SPLIT
 disable split screen
CLEAR/TCAT
 disable automatic catalog functions for table files
CLEAR/ZOOM [chan1]
 clear zoom
CLOSE/FILE file_id
 close an ASCII file
COMPUTE/AIRMASS frame [long] [lat]
COMPUTE/AIRMASS alpha delta ST [exptime] [long] [lat] [date] [UT]
 compute airmass (from sec z)
COMPUTE/BARYCORR date UT alpha delta [longitude] [latitude]
COMPUTE/BARYCORR table.tbl [longitude] [latitude]
COMPUTE/BARYCORR image alpha delta [longitude] [latitude]
 correct universal times and radial velocities to center of sun or
 barycenter of solar system
COMPUTE/COLUMN res_frame.column = arithmetic_expression
 do arithmetics on columns of an image
COMPUTE/HISTOGRAM result = table col [bin [min [max]]]
 table-to-image or table-to-table histogram transformation
COMPUTE/IMAGE [outspec =] expression
 compute arithmetic expression
COMPUTE/KEYWORD key = arithmetic_expression
 compute values of a keyword
COMPUTE/PIXEL [outspec =] expression
 compute expression on pixel basis
COMPUTE/PRECESSION alpha delta equinox0 equinox1
COMPUTE/PRECESSION table.tbl equinox0 equinox1
 precess equatorial coordinates from one epoch to another
COMPUTE/REGRESSION table column = name[(ind)] [d_type]
 compute result of a regression
COMPUTE/ROW res_frame.row = arithmetic_expression
 do arithmetics on rows (lines) of images
COMPUTE/ST date UT [longitude]
COMPUTE/ST table.tbl [longitude]
COMPUTE/ST image [longitude]
 calculate geocentric Julian date (JD) and local mean sidereal time (ST)
 from civil date and universal time (UT)
COMPUTE/TABLE table column = expression
 compute arithmetic or string operations on table columns
COMPUTE/UT date ST [longitude]

```

COMPUTE/UT table.tbl [longitude]
COMPUTE/UT image [longitude]
    calculate geocentric Julian date (JD) and universal time (UT) from
    civil date and local mean sidereal time (ST)
COMPUTE/WEIGHTS input_specs [window_specs]
    determine weights for command AVERAGE/WEIGHTS
COMPUTE/XYPLANE result_cube = expression
    compute arithmetic expression on xy_planes of cubes
COMPUTE/XZPLANE result_cube = expression
    compute arithmetic expression on xz_planes of cubes
COMPUTE/ZYPLANE result_cube = expression
    compute arithmetic expression on zy_planes of cubes
CONNECT/BACK_MIDAS unit wait_specs b_char method
    connect "command syntax" to another MIDAS
CONVERT/TABLE image = table x[y] z refima [method] [par]
CONVERT/TABLE image = table x[y] refima FREQ
    table to image conversion
CONVOLVE/IMAGE frame psf result
    convolve image with point spread function
COPY/DD source_frame source_desc dest_frame dest_desc
    copy descriptors of source frame to destination frame
COPY/DIMA source_frame source_desc dest_frame
    copy descriptor of source frame to new image
COPY/DISPLAY [out_dev] [stop_flg] [ITTdef] [LUTnam] [prfflag] [prmode]
    make a hardcopy of the display on output_device
COPY/DK source_frame source_desc dest_key
    copy descriptor of source frame to keyword
COPY/DK source_frame source_desc dest_key
    copy descriptor of source frame to keyword
COPY/GRAPHICS [device] [plotfile]
    copy the existing plot file to the graphic device
COPY/ID source_frame dest_frame dest_desc
    copy image data to descriptor of destination frame
COPY/II source_frame dest_frame dest_format delete_flag
    copy source frame to destination frame
COPY/IT inframe outable [column]
    copy image into table
COPY/KD source_key dest_frame dest_desc
    copy keyword to descriptor of destination frame
COPY/KEYWORD source_key dest_key [M_unit]
    copy keywords of same type
COPY/KI source_key dest_frame
    copy keyword to new frame
COPY/KT keyword table [column ...] element

```

copy keyword into table element
COPY/LSDD *list source_frame dest_frame*
 copy list of descriptors of source frame to descriptors of dest_frame
COPY/LSDK *list source_frame*
 copy list of descriptors of source frame to keywords
COPY/LSKD *list dest_frame*
 copy list of keywords to descriptors of destination frame
COPY/TABLE *intable outable [organization]*
 copy source table to destination table
COPY/TI *intable outimage*
 copy table into image
COPY/TK *table [column ...] element keyword*
 copy table element into keyword
COPY/TT *intable incolumn [outable] outcolumn*
 copy a table column to an other existing table
COPY/ZOOM *[out_dev] [stop_flg] [ITTnam] [LUTnam] [prflag] [prmode]*
 make a copy of the zoom window on output device
CREATE/ACAT *[cat_name] [dir_spec]*
 create a catalog of files in the current directory
CREATE/COLUMN *table column [unit] [format] [type]*
 create a table column
CREATE/COMMAND *comnd text*
 create a "user" command
CREATE/CURSOR *[dspid] [wind_specs] [Xstation]*
 create a cursor window
CREATE/D_COMMAND *comnd text*
 create a directory "user" command
CREATE/DEFAULT *comnd def1 def2 ... def8*
 create special defaults for MIDAS command
CREATE/DISPLAY *[dspid] [dspinfo] [meminfo] [alph_flag] [gsiz] [Xstation]*
 create a display window
CREATE/FCAT *[catname] [dir_spec]*
 create a catalog of fit files in the current directory
CREATE/FILTER *frame [dim_specs] [frame_specs] [filt_type] [coefs]*
 create filter frame
CREATE/GRAFICS *[graph_id] [graph_spec] [gsiz] [Xstation]*
 create a graphics window
CREATE/ICAT *[catname] [dir_spec]*
 create a catalog of images in the current directory
CREATE/IMAGE *frame [dim_specs] [frame_specs] [func_type] [coefs]*
 create an image
CREATE/LUT *LUT_table H_specs S_specs I_specs cyclic_flag*
CREATE/LUT *LUT_table CURSOR [start_LUT] [cursor_LUT]*
 create a colour lookup table

```

CREATE/RANDOM_IMAGE name [dims] [starts,steps] [func_type] [coefs]
  [seed]
CREATE/RANDOM_IMAGE name = ref_frame [func_type] [coefs] [seed]
  create a random image
CREATE/ROW table row_position number_of_rows
  add one or several rows at a given position of a table
CREATE/TABLE table ncol nrow file [format_file] [organization]
  create a table
CREATE/TCAT [catname] [dir_spec]
  create a catalog of tables in the current directory
CREATE/VIRTUAL virtual table
  create a virtual table from a physical table
CREATE/ZOOM [dspid] [wind_specs] [Xstation]
  create a zoom window
CUTS/IMAGE frame [cut_specs]
  display or set low + high cut values of an image frame

DEBUG/MODULE [low_lev,hi_lev] [switch]
  run MIDAS modules in debug mode
DEBUG/PROCEDURE [low_lev,hi_lev] [switch]
  run MIDAS procedures in debug mode
DECONVOLVE/IMAGE frame psf result [no_iter] [cont_flag]
  deconvolve image with point spread function
DELETE/ACAT [catalog] [conf_flag] [range]
  delete files with entry in ASCII file catalog
DELETE/COLUMN table column_sel
  delete table column(s)
DELETE/COMMAND [comnd]
  delete user defined command
DELETE/CURSOR [disp]
  delete cursor window(s) on XWindow displays
DELETE/DEFAULTS [comnd]
  delete special defaults for command
DELETE/DESCRIPTOR frame descr
  delete descriptor of frame
DELETE/DISPLAY [disp]
  delete display window(s) on XWindow displays
DELETE/FCAT [catalog] [conf_flag] [range]
  delete fit files with entry in catalog
DELETE/FIT name [conf_flag]
  delete a fit file
DELETE/GRAPHICS [grap]
  delete graphic window(s) on XWindow displays

```

DELETE/ICAT [*catalog*] [*conf_flag*] [*range*]
 delete image frames with entry in catalog

DELETE/IMAGE *name* [*conf_flag*]
 delete an image frame

DELETE/KEYWORD *key*
 delete user defined keyword

DELETE/LOGFILE
 delete current logfile

DELETE/ROW *table* *row_position* *number_of_rows*
 delete one or several rows of a table

DELETE/TABLE *name* [*conf_flag*]
 delete a table file

DELETE/TCAT [*catalog*] [*conf_flag*] [*range*]
 delete table files with entry in catalog

DELETE/TEMP
 delete temporary MIDAS files

DELETE/ZOOM [*disp*]
 delete zoom window(s) on XWindow displays

DISCONNECT/BACK_MIDAS *unit*
 disconnect from a background MIDAS

DISPLAY/CHANNEL [*chanl*] [*LUT_sect*]
 display contents loaded in an Image Display channel

DISPLAY/LUT [*switch*] [*intens*]
 en/disable display of current LUT

DRAW/ANY [*intens*]
 draw manually in the overlay channel

DRAW/ARROW [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw arrows in the overlay channel

DRAW/CIRCLE [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw circles in the overlay channel

DRAW/CROSS [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw crosses in the overlay channel

DRAW/ELLIPSE [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw ellipses in the overlay channel

DRAW/IMAGE *frame* [*chanl*] [*scale*] [*center*] [*cuts*] [*over*] [*iaux*] [*fix*]
 draw intensities of a line of an image into display channel

DRAW/LINE [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw straight line in the overlay channel

DRAW/RECTANGLE [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw rectangles in the overlay channel

DRAW/SLIT [*in_spec*] [*coord_ref*] [*draw_opt*] [*intens*] [*nocurs*] [*key_flag*]
 draw IUE slits in the overlay channel

```

ECHO/FULL [levla,levlb]
    show substitutions in MIDAS procedure files
ECHO/OFF [levla,levlb]
    suppress display of input from MIDAS procedure files
ECHO/ON [levla,levlb]
    display input from MIDAS procedure files
EDIT/TABLE table [edit_option] [col] [row]
    interactive table editor
EQUALIZE/HISTOGRAM frame descr itt_name
    perform histogram equalization
EXECUTE/CATALOG com_string parm1 ... parm7
    execute a MIDAS procedure or command for all entries in a catalog
EXECUTE/TABLE table command-string
    execute command on all rows of table
EXTRACT/CTRACE [step] [frame] [plot_flag] [zw_option]
    extract a column from displayed image
EXTRACT/CURSOR [subfr] [xpx,ypx] [loop_flag]
    extract a subframe via cursor
EXTRACT/IMAGE subframe = frame[....:....]
EXTRACT/IMAGE subframe = frame[...] loffsets roffsets
    extract a subimage from an image frame
EXTRACT/LINE out = in[...] step
    extract a 1-dim line from a 2-dim frame
EXTRACT/REFERENCE_IMAGE in ref out thresh
    extract subimage according to reference image
EXTRACT/ROTATED_IMAGE steps frame
    extract a rotated subimage from displayed image
EXTRACT/RTRACE [step] [frame] [plot_flag] [zw_option]
    extract a row from displayed image
EXTRACT/SLIT [in_option] [resframe] [slit_specs]
    extract a subimage defined by a fixed slit from image
EXTRACT/TRACE [step] [frame] [plot_flag] [cut_option] [zw_option]
    extract a line from displayed image

FFT/FINVERSE inr ini outr outi
    make inverse discrete Fourier transform
FFT/FPOWER inr ini outr outi pow_spec
    make discrete Fourier transform and power spectrum
FFT/FREQUENCY inr ini outr outi
    make discrete Fourier transform with frequency scaling
FFT/IMAGE inr ini outr outi
    make discrete Fourier transform
FFT/INVERSE inr ini outr outi

```

make inverse discrete Fourier transform
FFT/POWER *inr ini outr outi pow_spec*
 make discrete Fourier transform and power spectrum
FILTER/COSMIC *inframe outframe sky,gain,ron,[ns],[rc] [mask]*
 remove cosmic ray events.
FILTER/DIGITAL *frame outframe [filter_specs] [subimage] [options]*
 use digital filter on an image
FILTER/GAUSS *in out [radx,rady] [gauss_specs] [subima] [filtnam] [options]*
 use Gaussian filter on an image
FILTER/MAX *frame outfram [xyradius] [subima] [options]*
 apply maximum filter to an image
FILTER/MEDIAN *frame outfram [filt_specs] [flag] [subima] [options]*
 smooth an image with median filter
FILTER/MIN *frame outfram [xyradius] [subima] [options]*
 apply minimum filter to an image
FILTER/SMOOTH *frame outfram [filter_specs] [flag] [subima] [options]*
 smooth an image by averaging
FIND/MINMAX *frame*
 find min, max of frame and corresponding pixel numbers
FIND/PIXEL *frame low,high [inout_flag] [first_flag] [table] [rowmax]*
 find first/all pixel(s) with a value in/outside interval [low,high]
FIT/FLAT_SKY *outframe = inframe [in_specs] [order] [back_surface]*
FIT/FLAT_SKY *inframe [in_specs] [order] [back_surface]*
 Approximate background of image by a surface
FLIP/IMAGE *frame [flag]*
 flip an image around an axis

GET/CURSOR *[output] [option] [marker] [curs_specs] [zw_option]*
 read cursor coords from display
GET/GCURSOR *[output_spec] [app_flag] [max]*
 read and store cursor coordinates from the graphics display
GET/IMAGE *frame [input_source] [ITT_flag]*
 read image from displayed image channel
GET/ITT *out_specs [chan1] [sect]*
 read currently active ITT from image display
GET/LUT *out_specs [get_specs] [ITT] [format] [range]*
 read currently active LUT from image display
GROW/CUBE *frame no_planes frame_list*
 expand 2-dim/3-dim frame
GROW/IMAGE *out = in [start,step,no] [lincol_specs] [lincol_flag]*
 expand single line into 2-dim image

```

HELP [help_topic]
    display info about help_topic
HELP/APPLIC [proc]
    display header information of application procedure
HELP/CL [command]
    display help for commands only used in MIDAS procedures
HELP/CONTRIB [proc]
    display header information of procedures in the Midas 'contrib' area
HELP/KEYWORD key
    explain contents of given key
HELP/QUALIF [qualif]
    display all commands with given qualifier
HELP/SUBJECT [topic]
    display information related to given topic

```

```

INDISK/ASCII in_file [out_file] [npix_string]
    read ASCII file from disk + convert to Midas image
INDISK/FITS in_files [out_files] [option]
    read FITS files from disk
INFO/DESCR frame descr
    get type and size of descriptor
INFO/IMAGE frame
    get internal info of frame
INFO/SETUP [setup]
    display all the information about a Setup
INITIALIZE/DISPLAY [noLUT,LUTsz] [ownLUT] [M_unit] [fonts]
    initialize the image display
INITIALIZE/SETUP [setup]
    initialize the variables of a Setup
INSERT/IMAGE subframe modframe [startx,y,z]
    insert a subframe into another frame
INTAPE/FITS file_specs file-id device [flags]
    read frames from magtape in FITS/IHAP format
INTERPOLATE/II outima inima refima [s] [degree]
    interpolate Image to Image
INTERPOLATE/IT outtab i,d inima [s] [degree]
    interpolate Image to Table
INTERPOLATE/TI outima intab i,d refima [s] [degree]
    interpolate Image to Image
INTERPOLATE/TT outtab i,d intab i,d [s] [degree]
    interpolate Table to Table
ITF/IMAGE inframe table col1,col0 scal outframe
    ITF correction

```

JOIN/TABLE *intab1* [:*X1*,[:*Y1*] *intab2* [:*X2*,[:*Y2*]] *outtable* [*tolX*,*tolY*]
 join table files

LABEL/DISPLAY *label* [*position*] [*mode*] [*option*] [*size*] [*key_flag*]
 write a label on the image display

LABEL/GRAPHIC *label* [*x_pos*,*y_pos*[,*mm*]] [*angle*] [*size*] [*pos_ind*]
 write a label_string on the graphics device

LOAD/CURSOR *curs_table* *curs_no*
 load programmable cursor into the Image Display

LOAD/IMAGE *frame_spec* [*chan1*] [*scale*] [*center*] [*cuts*] [*dirs*] [*fix*]
 load image into display device

LOAD/ITT *in_specs* [*chan1*] [*load_specs*]
 load intensity transfer table to Image Display

LOAD/LUT *in_specs* [*load_specs*] [*disp_flag*] [*format*]
 load colour lookup table into Image Display

LOAD/OVERLAY *overlay_table* *load_specs*
 load lookup table for overlay + graphics

LOAD/TABLE *table* *x* *y* [*ident*] [*symbol* [*size* [*intens*]]]
 load table into overlay channel of Image Display

LOCK/KEYWORD *key_list* *lockno*
 lock keyword(s)

LOG/OFF
 suppress logging

LOG/ON
 enable logging

LOG/TOF
 write top_of_form into logfile

MAGNITUDE/CENTER [*in_specs*] [*out_specs*] [*Fsiz*,*Nsiz*,*Bsiz*] [*out_opt*]
 [*center_params*] [*curs_specs*] [*zw_option*]
 compute magnitude in center

MAGNITUDE/CIRCLE [*in_specs*] [*out_specs*] [*Fsiz*,*Nsiz*,*Bsiz*] [*out_opt*]
 [*center_params*] [*curs_specs*] [*zw_option*]
 compute magnitude within circular aperture

MAGNITUDE/RECTANGLE [*in_specs*] [*out_specs*] [*Fsiz*,*Nsiz*,*Bsiz*] [*out_opt*]
 [*center_params*] [*curs_specs*] [*zw_option*]
 compute magnitude within square aperture

MERGE/TABLE *intable* [*intable* ...] *outable*
 merge table files

MODIFY/AREA [*source*] [*resfram*] [*degree*] [*constant*] [*drawflg*]
 remove bad data from a circular pixel-area in an image

MODIFY/COLUMN *source_def res_frame [col_type] column_coords*
 approximate values in a column
 MODIFY/CUTS *[image] [cursor_spec]*
 modify cut values of full frame or in cursor selected windows
 MODIFY/GCURSOR *frm_in frm_out y-coord xstart,xend no_curs,degree*
 interactive modification of pixel values in a frame
 MODIFY/ITT *[method] [value] [prflag]*
 modify the currently active ITT
 MODIFY/LUT *[method] [colour] [prflag]*
 modify the currently active LUT
 MODIFY/PIXELS *[source] [resfram] [arfacts] [xdeg,ydeg,niter] [drawflg]*
 [noise]
 approximate pixel-area in an image
 MODIFY/ROW *source_def res_frame row_type row_coords*
 approximate values in a row

NAME/COLUMN *table column [new-column] [unit] [format]*
 redefines label/unit/format of a column
 NORMALIZE/SPECTRUM *inframe outframe [mode] [table] [batch_flag]*
 approximate continuum of 1-D spectra for later normalization

OPEN/FILE *filename flag file_control_key*
 open an ASCII file for reading or writing
 OUTTAPE/FITS *[catalog[,list]] device [flags] [density,block] [type]*
 write to device in FITS format
 OVERPLOT/AXES *[x_axis_spec] [y_axis_spec] [x_sc,y_sci[,x_off,y_off]]*
 [x_lab] [y_lab]
 overplot a coordinate box with tickmarks and labels
 OVERPLOT/AXES *[coord_str] [x_lab] [y_lab]*
 overplot a coordinate box around a displayed frame
 OVERPLOT/COLUMN *frame [x-coord] [y_start,y_end] [offset] [l_type]*
 overplot a column of a frame on a graphic device
 OVERPLOT/COLUMN *frame [x-coord] [y_start,y_end] [offset] [l_type]*
 overplot a column of a frame on a graphic device
 OVERPLOT/CONTOUR *frame [coord_str] [contours] [sm_par]*
 overplot contour map of 2-dim. frame with smoothing option
 OVERPLOT/DESCRIPTOR *frame [descr] [start,end] [offset]*
 overplot the contents of a descriptor
 OVERPLOT/ERROR *table [col1] [col2] col3 [direct] [bar]*
 overplot table error column
 OVERPLOT/GRAY *frame [coord_str] [gray_lev] [sm_par] [gray_ness] [options]*
 overplot gray scale map of 2-dim. frame with smoothing option

OVERPLOT/GRID *grid*
 overplot a grid on an existing coordinate box

OVERPLOT/HISTOGRAM *tab col [offset] [bin[,min[,max]]] [exc] [log] [opt]*
 overplot histogram of a column in the table

OVERPLOT/HISTOGRAM *frame [offset] [log] [opt]*
 overplot the histogram of an image

OVERPLOT/KEYWORD *[key_name] [start,end] [offset]*
 overplot the contents of a keyword

OVERPLOT/LINE *[l_type] [x_sta,y_sta [x_end,y_end]]*
 overplot a line on a graphic device

OVERPLOT/ROW *frame [y-coord] [x_start,x_end] [offset] [l_type]*
 overplot a row (line) of a frame on a graphic device

OVERPLOT/SYMBOL *[s_type] [x_coord,y_coord] [s_size]*
 overplot a symbol

OVERPLOT/TABLE *table [plane1] [plane2] [x_sc,y_sc[,x_off,y_off]] [symbols] [lines] [flag_dir]*
 plot table data on selected plotting device

OVERPLOT/VECTOR *fram_a fram_b [coord_str] [scale_r] [pos_range] [sm_par] [head]*
 overplot vector map from two 2-dim. images with smoothing option

PLAYBACK/FILE *name*
 playback MIDAS commands from an ASCII file

PLAYBACK/LOGFILE *file*
 playback MIDAS commands from a previous logfile

PLOT/AXES *[x_axis_spec] [y_axis_spec] [x_sc,y_sc[,x_off,y_off]] [x_lab] [y_lab]*
 plot a coordinate box with large and small tickmarks and labels

PLOT/AXES *[coord_str] [x_lab] [y_lab]*
 plot a coordinate box around a displayed frame

PLOT/COLUMN *frame [x_coord] [y_sta,y_end] [x_sc,y_sc,x_off,y_off]*
 plot a column of an image on a plotting device

PLOT/CONTOUR *frame [coord_str] [x_sc,y_sc[,x_off,y_off]] [contours] [c_type] [sm_par]*
 plot contour map of 2-dim. image with smoothing option

PLOT/DESCRIPTOR *frame [descr] [start,end] [x_sc,y_sc[,x_off,y_off]]*
 plot a descriptor on plotting device

PLOT/GRAY *frame [coord_str] [x_sc,y_sc[,x_off,y_off]] [gray_lev] [sm_par] [gray_ness] [gray_opt]*
 plot gray scale map of 2-dim. image with smoothing option

PLOT/HISTOGRAM *tab col [x_sc,y_sc[,x_off,y_off]] [bin[,min[,max]]]*

```

[exc] [log] [opt]
    plot histogram of a column in the table
PLOT/HISTOGRAM frame [x_sc,y_sc[,x_off,y_off]] [exs] [log] [opt]
    plot the histogram of an image
PLOT/KEYWORD [key_name] [start,end] [x_sc,y_sc[,x_off,y_off]]
    plot the contents of a keyword
PLOT/PERSPECTIVE frame [coord_str] [alt,azi] [scal,offs] [sm_par] [xy_flag]
    tree dim. representation of a 2-dim. frame, with smoothing option
PLOT/ROW frame [y_coord] [x_sto,x_end] [x_sc,y_sc[,x_off,y_off]]
    plot a row (line) of an image on a plotting device
PLOT/TABLE table [plane1] [plane2] [x_sc,y_sc[,x_off,y_off]]
[symbols] [lines] [flag_dir]
    plot table data on selected plotting device
PLOT/VECTOR frame_a frame_b [coord_str] [x_sc,y_sc[,x_off,y_off]] [scale_r]
[range] [sm_par] [head]
    plot vector map from two 2-dim. images with smoothing option
PRINT/ACAT [cat_name] [lowno,hino]
    print ASCII file catalog entries
PRINT/DESCR frame [descr_list] [disp_flag]
    print descriptor values
PRINT/FCAT [cat_name] [lowno,hino]
    print fit file catalog entries
PRINT/HELP [help_topic]
    print info about help_topic
PRINT/HISTOGRAM table column [bin [min [max]]]
    print statistics of a column
PRINT/ICAT [cat_name] [lowno,hino]
    print image catalog entries
PRINT/IMAGE frame_specs [pixel_specs] [hide_header_flag]
    print image data values
PRINT/KEYWORD [key_list] [since]
    print contents of keywords
PRINT/LOGFILE [page_specs]
    print contents of logfile
PRINT/TABLE table [column ...] [elem1 [elem2]] [N] [width]
PRINT/TABLE table [elem1 [elem2]] [form] [N]
    print table values on the device/file specified via ASSIGN/PRINT
PRINT/TCAT [cat_name] [lowno,hino]
    print table catalog entries
PROJECTION/TABLE intable outable column_selection
    projection of one or more columns from a table

```

READ/ACAT [*cat_name*] [*lowno,hino*]
 read ASCII Catalog entries
 READ/COMMANDS [*proc*]
 read commands from a procedure + store into command buffer
 READ/DESCR *frame* [*descr_list*] [*disp_flag*]
 display descriptor values
 READ/FCAT [*cat_name*] [*lowno,hino*]
 read fit file catalog entries
 READ/FILE *file_id* *cbuf_key* [*maxrd*]
 read an ASCII file
 READ/HISTOGRAM *table column* [*bin [min [max]]*]
 display statistics of table column
 READ/ICAT [*cat_name*] [*lowno,hino*]
 read Image Catalog entries
 READ/IMAGE *frame_specs* [*pixel_specs*] [*hide_header_flag*]
 display image data values
 READ/KEYWORD [*key_list*] [*disp_flag*] [*since*] [*Midunit*]
 display contents of keywords
 READ/SETUP *setup*
 read the contents of the variables related to a Setup
 READ/TABLE *table* [*column_sel*] [*row_sel*] [*form*]
 display table elements
 READ/TCAT [*cat_name*] [*lowno,hino*]
 read Table Catalog entries
 REBIN/II *outima inima refima* [*func*] [*param*] [*intop*]
 nonlinear rebin Image to Image
 REBIN/IT *outtab i,d[,b] inima* [*func*] [*param*] [*intop*]
 nonlinear rebin Image to Table
 REBIN/LINEAR *in out* [*stepx,stepy*] [*offx,offy*] [*startx,starty*] [*fluxcons*]
 REBIN/LINEAR *in out* [*refframe*] [*fluxcons*]
 rebin an image linearly
 REBIN/ROTATE *in out* [*rot_specs*] [*ref_frame*] [*ref_flag*]
 rotate + rebin an image
 REBIN/SPLINE *in out* [*stepx,stepy*] [*offx,offy*] [*startx,starty*]
 REBIN/SPLINE *in out* [*refframe*]
 rebin an image using cubic splines
 REBIN/TI *outima intab i,d[,b] refima* [*func*] [*param*] [*intop*]
 nonlinear rebin Table to Image
 REBIN/TT *outtab i,d[,b] intab i,d[,b]* [*func*] [*param*] [*intop*]
 nonlinear rebin Table to Table
 REGRESSION/LINEAR *table y x1,x2,...*
 linear regression on table columns
 REGRESSION/POLY *table y x1[,x2] d1[,d2]*
 polynomial fit on table columns

```

REGRESSION/TABLE table1 x1[,x2] table2 y1[,y2] degree tol [guess]
    polynomial fit of variables in two tables (not yet implemented)
RENAME/FIT old new [history] [overwrite]
    rename a fit file
RENAME/IMAGE old new [history] [overwrite]
    rename an image frame
RENAME/TABLE old new [history] [overwrite]
    rename a table frame
REPLACE/IMAGE in out [test/]low,hi=express1[,express2]
    replace pixels according to intensity
REPLACE/POLYGON in,intab out test/low,hi=value
    replace pixels inside polygon
REPORT/PROBLEM [errfile]
    send error reports and comments to the person(s) in charge of MIDAS
RESET/DISPLAY
    reset Xwindow display after Control C
RESTORE/NAME [file_spec] [verbose] [history] [overwrite] [descr]
    change file name according to descr. FILENAME
RETRO/TAB table
    retrofit 3-dim table to old 90NOV format
ROTATE/1DIM in out nop_flag
    rotate a 1-dim profile around its startpoint
ROTATE/CLOCK in out [factor]
    rotate an image by multiples of 90 degrees clockwise
ROTATE/COUNTER_CLOCK in out [factor]
    rotate an image by multiples of 90 degrees counter_clockwise
RUN prog
    execute a program inside the MIDAS environment

```

```

SAVE/REGRESSION table name
    save results of a regression
SCROLL/CHANNEL [chan1] [scrolx,scroly]
    scroll given ImageDisplay channel
SEARCH/FCAT [cat_name] search_string [options]
    search in fit file catalog for frame with matching descriptor IDENT
SEARCH/ICAT [cat_name] search_string [options]
    search in image catalog for frame with matching descriptor IDENT
SEARCH/LINE frame w,t[,nscan] [table] [meth] [type]
    search for spectral lines
SEARCH/TCAT [cat_name] search_string [options]
    search in table catalog for table with matching descriptor IDENT
SELECT/TABLE table logical-expression
    select table entries

```

SET/ACAT [*cat_name*]
 make given catalog the "active" ASCII file catalog

SET/BACKGROUND [*method*] [*echo*] [*sleep_time*]
 put Midas session into "background" mode

SET/BUFFER [*no_lines*]
 set up command buffer for MIDAS

SET/CONTEXT *ctxt*
 enable new context

SET/CURSOR [*curs_no*] [*curs_form*] [*curs_coords*] [*flag*]
 set cursor form and position

SET/DISPLAY [*colour_mode*]
 set up Image Display for RGB or pseudo colours

SET/FCAT [*cat_name*]
 make given catalog the "active" fit file catalog

SET/FORMAT [*format_specs*]
 set formats for replacement of Midas data

SET/GCURSOR [*curs_no*] [*curs_form*]
 set cursor form in graphics window

SET/GRAFICS *option1[=value1]* *option2[=value2]* ...
 set plot characteristics

SET/ICAT [*cat_name*]
 make given catalog the "active" image frame catalog

SET/ITT [*chan1*] [*sect*]
 enable ITT for *Image display channel*

SET/LUT [*sect*]
 enable usage of colour lookup tables

SET/MIDAS_SYSTEM *option=value*
 set different modes and options for Midas

SET/OVERLAY
 enable graphics overlay

SET/REFCOLUMN *table column*
 define column as reference in table access

SET/SPLIT [*chanls*]
 enable split screen

SET/TCAT [*cat_name*]
 make given catalog the "active" table file catalog

SHIFT/IMAGE *inframe outframe [x,yshift]*
 shift the pixels in an image

SHOW/ACAT [*cat_name*] [*display_flag*]
 show no. of entries in an ASCII file catalog

SHOW/BACK_MIDAS [*option*]
 show info related to background MIDAS sessions

SHOW/CHANNEL [*chan1*]
 show info related to ImageDisplay channel

```

SHOW/CODE command_string [flag]
    display the procedure which implements the command_string
SHOW/COMMAND [comnd/qualif]
    display MIDAS commands
SHOW/DEFAULTS
    display all special defaults
SHOW/DESCR frame [dsclist] [flag]
    show existing descriptors with name, type and size
SHOW/DISPLAY
    show current status of ImageDisplay + Graphics
SHOW/FCAT [cat_name] [display_flag]
    show no. of entries in a fit file catalog
SHOW/GRAPHICS device_name
    show the setup parameters for plotting
SHOW/ICAT [cat_name] [display_flag]
    show no. of entries in an image catalog
SHOW/KEYWORDS [keyword]
    display contents of keyword data base
SHOW/TABLE table
    display table parameters
SHOW/TCAT [cat_name] [display_flag]
    show no. of entries in a table catalog
SORT/FCAT [cat_name]
    sort a fit file catalog
SORT/ICAT [cat_name]
    sort an image catalog
SORT/TABLE table keys
    sort table according to (ascending) values
SORT/TCAT [cat_name]
    sort a table catalog
STATISTICS/IMAGE [frame] [area] [bins] [lo,hi-exc] [options] [outtab]
 [plotflg] [format]
    calculate statistics of a frame
STATISTICS/TABLE table column
    simple statistics on a table column
STORE/FRAME key frame [indx] [exit_label]
    store frame or entries of catalog into key
SUBTRACT/ACAT [cat_name] frame_list
    remove entries from an ASCII file catalog
SUBTRACT/FCAT [cat_name] frame_list
    remove entries from a fit file catalog
SUBTRACT/ICAT [cat_name] frame_list
    remove entries from an image catalog
SUBTRACT/TCAT [cat_name] frame_list

```

remove entries from an table catalog

SYNCHRONIZE/MIDAS
write keyfile and logfile to disk

TRANSLATE/SHOW *proc option*
translate MIDAS procedure and display resulting code

TRANSPOSE/CUBE *inframe* [*outframe*] [*plane_spec*]
rearrange the planes of a cube

TRANSPOSE/IMAGE *inima* *outima* [*diagonal*]
transpose image

TUTORIAL/EXTRACT
demonstrate some of the different EXTRACT commands

TUTORIAL/FILTER
explain the usage of filters

TUTORIAL/GRAFICS *option*
explain the use of the graphics packages

TUTORIAL/HELP
explain usage of the HELP command

TUTORIAL/ITT [*plotflag*]
explain the usage of ITT's

TUTORIAL/LUT [*plotflag*]
show some standard LUT's and related MIDAS commands

TUTORIAL/SPLIT
explain the usage of split screen

TUTORIAL/TABLE
explain usage of tables

\$ comnd
execute a host system command

VIEW/IMAGE [*frame*] [*out_tab*] [*plot_option*] [*g,zhardcopy*]
view an image with a "looking glass"

WAIT/BACK_MIDAS [*unit*]
wait until command in background MIDAS terminates

WAIT/SECS [*no_of_secs*]
suspend MIDAS monitor for *no_of_secs* seconds

WRITE/COMMANDS [*procnam*] [*par1*] [*par2*] ... [*par8*]
save commands from command buffer + write into a procedure

WRITE/DESCR *frame descr data [flg]*
store values into a descriptor

WRITE/DHELP *frame descr text*
store help-text/comments for an existing descriptor

```

WRITE/FILE file_id charbuf
    write into an ASCII file
WRITE/IMAGE frame_specs [pixel_specs] data [flg]
    store values into image pixels
WRITE/KEYWORD key data [flg]
    write values into a keyword
WRITE/OUT text_spec [section] [label]
    display text on terminal
WRITE/SETUP [setup]
    modify the variables of a Setup
WRITE/TABLE table column row_sel value
    Store a value into a table

XCORRELATE/IMAGE temp spec result shift
    correlate 2 similar 1-dim frames over 2*(shift) bandwith

```

```

ZOOM/CHANNEL [zoom_fact] [center]
    zoom image on image display
ZOOM/OVERLAY [zoom_fact] [center]
    zoom image + overlay together

```

A.2 Application Commands

```

ASSOCIATE/RANK table col1 col2 [action]
    rank-order correlation coefficient

```

```

BIN/TABLE table col1 col2 [bin] [min] [max] [sigma]
    creates a table, bin.tbl with averages of col2 in bins of col1

```

```

COMPARE/2SAM table col1 col2
    kolmogorov two-sample test
COMPUTE/FIT image [, error] [= function[(refima)]]
    compute fitted image values
COMPUTE/FIT table y [, error] [= function[(ind)]]
    compute fitted table values
COMPUTE/FUNCTION image = function[(refima)]
    compute function values, result as image
COMPUTE/FUNCTION table y = function[(ind)]
    compute function values, result in table column

```

```

CREATE/FUNCTION fun1[,fun2...] [library_specs]
    define user functions for fitting
CREATE/GUI [name]
    Creates graphical user interfaces
CREATE/STAR in_frm in_tab out_frm [n_size] [frm_specs] [dmin,dmax] [radius]
    create the profile of a reference star by adding and averaging

EDIT/FIT function
    interactive function editor

FILTER/ADAPTIV frame outframe [maskframe] [type] [shape] size k noise
    adaptive filtering of an image
FIT/IMAGE [nfeval[,prec[,metpar]]] [image[,wgt]] [funct]
    fit image values
FIT/TABLE [nfeval[,prec[,metpar]]] table :dep,[:wgt] :ind [funct]
    fit table
FTEST/VAR table col1 col2
    f-test for different variances

GET/FIT table [image]
    create a table for fitting subimages

IDENTIFY/CURSOR table ident x [y] [error]
    identify table entries from display
IDENTIFY/GCURSOR table ident x [y] [error]
    identify table entries from graphic terminal
INTEGRATE/APERTURE [in_specs] [out_tab] [radius]
    integrate the flux within an aperature
INTEGRATE/LINE frame [y_coo] [x_sta,x_end] [n_cur,deg] [batch] [x-pos,range]
    integrate area in a (spectral) line
INTEGRATE/STAR [in_specs] [out_tab] [parameters] [mode]
    computes flux, radius and background of stars previously centered

KSTEST/1SAM table col distri coeffs
    kolmogorov one-sample test

MODIFY/FIT table seq_no [name]
    modify fit parameters

```

```

PRINT/FIT func_name
    print function parameters

READ/FIT func_name
    display fitted function parameters
REGISTER/SESSION session directory file table
    Register a session in the session manager
REPLACE/FUNCTION fun1[,fun2...]
    replace user functions for fitting

SAVE/FIT table seq_no [name]
    save results of a regression
SELECT/FUNCTION name number[,...]
    select function components
SET/FIT par=value [par=value ...]
    set parameters for the FITTING package
SHOW/FIT
    display parameters used in FITTING package
SORT/COLUMN input output
    column oriented sorting of the pixels of a frame.
SORT/ROW input output
    row oriented sorting of the pixels of a frame.
STEST/MEAN table col1 col2
    student t-test for different means

TUTORIAL/ALIGN
    explain the alignment of two images
TUTORIAL/FIT
    explain the modelling of table and image data by fitting non-linear
    functions.

```

A.3 Standard Reduction Commands

A.3.1 ccdred

```

ALIGN/MOSAIC in_frm in_tab out_frm method,data [nxrsub,nyrsub] [xref,yref]
[x_size,y_size]
    Align the elements of the mosaiced frame
BIAS/CCD [in_fram] [out_fram] [bs_fram]
    Correct the input frame for the bias offset using a bias frame
COMBINE/CCD exp [in_spec] [out_fram]

```

Combined a number of CCD frames of the same exposure type

CREATE/MOSAIC *in_cat* *out_frm* *out_tab* *nx_sub,ny_sub* [*not1,not2,...*]
[*nocol,norow*]

Mosaic a set of (infrared) ccd frames

DARK/CCD [*in_fram*] [*out_fram*] [*dk_fram*]

Correct input frame for dark current offset using a dark current frame

FIT/MOSAIC *in_frm* *in_msk* *in_tab* *out_frm* [*match*] [*nxrsub,nyrsub*] [*xref,yref*]
[*x_size,y_size*]

Align and match the elements of the mosaiced frame

FIXPIX/CCD [*in_fram*] [*out_fram*] [*fix_table*] [*fix_meth*]

Do a correction of bad pixels in the input frame

FLAT/CCD [*in_fram*] [*out_fram*] [*ff_fram*]

Do a flat field correction of the input frame

FRCOR/CCD [*in_spec*] [*out_frm*] [*xboxmn,xboxmx*] [*yboxmn,yboxmx*] [*clip*]
[*lowsig,higsig*]

Make fringe correction frame(s) from sky frames

FRINGE/CCD [*in_fram*] [*out_fram*] [*fr_fram*] [*fr_scale*]

Do a fringe correction of the input frame

HELP/CCD [*keyword*]

show the parameter setting of the current CCD session

ILLCOR/CCD [*in_spec*] [*out_frm*] [*xboxmn,xboxmx*] [*yboxmn,yboxmx*] [*clip*]
[*lowsig,higsig*]

Make flat field illumination correction frame(s)

ILLFLAT/CCD [*in_spec*] [*out_frm*] [*xboxmn,xboxmx*] [*yboxmn,yboxmx*] [*clip*]
[*lowsig,higsig*]

Apply correction to a flat field to remove illumination pattern

ILLUMINATION/CCD [*in_fram*] [*out_fram*] [*il_fram*]

Do an illumination correction of the input frame

INIT/CCD [*name*]

Initialize the CCD package, optionally using the setting of a saved session

LOAD/CCD [*intr*]

Load instrument/detector specifications into the CCD context

MATCH/MOSAIC *in_frm* *in_tab* *out_frm* *method,data* [*match*] [*nxrsub,nyrsub*]
[*xref,yref*] [*x_size,y_size*]

Align and match the elements of the mosaiced frame

MKREDT/CCD *out_tab*

Create CCD empty table with columns for science and calibration frames

OVERSCAN/CCD [*in_fram*] [*out_fram*] [*sc_area*] [*mode*]

Correct the input frame for the bias offset in the overscan region

REDUCE/CCD [*in_spec*] [*out_frm*]

Do the (partial) calibration of one or more frames

SAVE/CCD *name*

```

        save current CCD session
SET/CCD keyw=value [...]
        Define the values of parameters in the current CCD session.
SHIFT/MOSAIC out_tab [curs_opt] [csx,csy] [clear_opt]
        Get x and y shifts of the subraster in the mosaic frame
SHOW/CCD [subject]
        Show (part of) the setup of the CCD package
SKYCOR/CCD [in_spec] [out_frm] [xboxmn,xboxmx] [yboxmn,yboxmx] [clip]
 [lowsig,higsig]
        Make sky illumination correction frame(s)
SKYCOR/CCD [in_spec] [out_frm] [xboxmn,xboxmx] [yboxmn,yboxmx] [clip]
 [lowsig,higsig]
        Apply sky observation to flat field to remove illumination pattern
TRIM/CCD [in_fram] [out_fram] [im_sec] [del_flg]
        Extract the useful data from the ccd frame.

```

A.3.2 ccdtest

```

TEST1/CCD in_cat [out_id] [meth] [option]
        Combine bias frames stored in a catalogue and display it
TESTB2/CCD in_frm [out_id] [row_ran] [col_ran]
        Compute row and column average of a (averaged) bias frame
TESTB3/CCD in_frm [out_id] [area] [size] [option]
        Find the hot pixels in a (combined) bias frame
TESTB4/CCD in_frm [out_id] [area] [size,fac]
        Make a histogram of the pixel intensities and rebin the input frame
TESTB5/CCD in_cat [out_id] [area] [size,fac]
        Do the statistics of the bias frame is a catalogue.
TESTBA/CCD in_cat [out_id] [meth] [row_ran] [col_ran] [area] [size,fac]
        Do a series of tests of a catalogue of bias frames
TESTC/CCD in_frm [rows] x_pix [columns] y_pix
        Compute the horizontal and vertical charge transfer efficiency.
TESTD/CCD in_cat [out_id] [dec_fac]
        Do a test on a catalogue of dark current frames
TESTF1/CCD in_cat [out_id] [meth] [area] [exp_ran] [option]
        Combine the flat frame in the input catalogue and display
TESTF2/CCD in_frm [out_id] [area] [thresh] [option]
        Find the cold pixels in the combined low count flat.
TESTFA/CCD in_cat [out_id] [meth] [area] [exp_ran] [thresh]
        Do a series of tests on a catalogue of low count flat frames
TESTS/CCD in_frm1 inf_frm2 [out_frm] n_exp [dec_fac]
        Find the shutter error distribution
TESTT1/CCD in_cat [out_id] [area] [option]

```

Display the linearity and transfer curves of pairs of flat frames.
TESTT2/CCD *in_tab* [*out_id*] [*select*] [*tim_int*]
Fit the linearity curves and determine the shutter offset
TESTT3/CCD *in_tab* [*out_id*] [*select*]
Fit the transfer curve and determine the ADU conv. factor and RON
TESTTA/CCD *in_cat* [*out_id*] [*area*] [*tim_int*] [*select*]
Do linearity and transfer tests on a catalogue of flat frames

A.3.3 do

ASSOCIATE/IMA *ost exptype rule_table outtable [flag] [nexp]*
associates to scientific exposures a set of suitable calibration images
CLASSIFY/IMAGE *table descr outcol outchar*
classify images according to one or several rules.
CREATE/CRULE *table rule*
create an classification rule for a given Observation Summary Table
CREATE/OST *file_specs [file_pref] intable outtable flag*
create an Observation Summary Table
GROUP/ROW *table incol outcol*
group the rows of a table by the value of one of its column

A.3.4 echelle

AVERAGE/TABLE *frame table xy_col outcol [size]*
Read pixels in a frame at positions defined by a table.
BACKGROUND/ECHELLE *in out [radx,rady,step] [degree] [smooth] [method]*
estimate interorder scattered light of an echelle spectrum
BACKGROUND/SMOOTH *input output [radx,rady] [niter] [visu]*
estimate interorder scattered light of an echelle spectrum
CALIBRATE/ECHELLE [*defmtd*] [*wlcmtd*]
performs order definition and wavelength calibration
CLEAN/ECHELLE
Clears contexts Echelle and Spec and removes process tables
CONVERT/ECHELLE *input output domain function param option*
resample echelle orders
DEFINE/ECHE [*ordref*] [*width1,thres1,slope*] [*defmtd*] [*defpol*]
define echelle order positions
DEFINE/HOUGH [*ordref*] [*nbord*] [*hwid*] [*hough_par*] [*thresh*] [*degx,deg_y*]
[*hot_thres,step*] [*hough_setup*]
define echelle order positions; automatic detection by Hough transform.
DEFINE/SKY *ima [nsky] [possky] [half_width]*
defines limits of the sky windows

DISPLAY/ECHELLE *image* [*g_flag*]

 Optionally creates a display and graphic windows and scales an image to be displayed.

ERROR/ECHELLE *command keyword*

 Low-level error message generator for the echelle package

EXTR/ECH *input output* [*params*] [*method*]

 extract echelle orders

EXTR/OPT *input output slit,ord1,ord2* [*ron,g,sigma*] [*table*] [*coeffs*]

 weigthed extraction of echelle orders

EXTRACT/ORDER *inp out sl,ang,off meth table coeff* [*ord1,ord2*]

 Extract echelle orders and produces a frame in space pixel-order

EXTRACT/SKY *in out* [*mode*]

 Extracts sky spectrum.

FILTER/ECHELLE *input output*

 filter echelle frame for cosmic ray hits and subtract background

FLAT/ECHELLE [*flat*] [*correct*] [*blaze*]

 subtract background from flat-field image and approximate blaze profile

HELP/ECHELLE *keyword* [*mode*]

 Provides short help on an echelle session keyword

**HOUGH/ECHELLE *input* [*scan*] [*step,nbtr*] [*nbord*] [*flags*] [*hwid*] [*thres*]
[*params*]**

 perform Hough transform and orders detection on a flat-field frame

IDENT/ECHELLE [*wlc*] [*lincat*] [*dc*] [*tol*] [*wlcloop*] [*wlcmtd*]

[*guess*, [*shift*]] [*ccdbin*]

 perform wavelength calibration of echelle spectra

INIT/ECHELLE [*name*]

 initializes echelle parameters

INITIAL/EMMI [*ref*] [*grism*]

 Initializes the Echelle context for a given EMMI configuration

KEYDEL/ECHELLE [*table*]

 Deletes echelle session keywords

LOAD/CALIBRATION

 display wavelength calibration result

LOAD/ECHELLE

 display echelle orders (and optionally background) positions

LOAD/IDENTIFICATION

 display initial identifications.

LOAD/SEARCH

 Loads on display the position of the lines found by the

SEARCH/ECHELLE *command*.

MERGE/ECHELLE *inframe outframe* [*params*] [*method*]

 merge echelle orders

MERGE/OPTIMAL *rebima weight out* [*delta*]

Optimal weighted merging of echelle orders

OFFSET/ECHELLE *[image] [range] [cover] [ordtab] [mode]*
Determines the offset along the slit between the order
coefficients and a given echelle spectrum.

OVERLAP/ECHELLE *rebima order*
Plots the overlap region between adjacent orders n and n+1

PLOT/CALIBRATE *[ord1,ord2]*
plot dispersion relation in echelle reduction

PLOT/ECHELLE *frame [ord1,ord2] [printer]*
plot extracted echelle orders.

PLOT/IDENTIFICATION *frame [ord1,ord2] [printer]*
plot line identifications in echelle reduction

PLOT/RESIDUAL *[ord1,ord2]*
plot dispersion residuals in echelle reduction

PLOT/SPECTRUM *in [start,end]*
plots a rebinned spectrum in wavelength range

PREPARE/BACKGROUND *[step] [init] [back_tab] [order_tab] [descr]*
low-level command; create table back.tbl

PREPARE/WINDOW *catalogue flat-bkg lhcuts*
prepare echelle images for the command AVERAGE/WINDOW

REBIN/ECHELLE *input output sample*
rebin echelle orders into wavelength

REDUCE/ECHELLE *input output [bkcor]*
reduction of echelle spectra.

REGRESSION/ECHELLE *[defpol] [niter] [absres] [kappa]*
fit 2-dim. polynomial to order positions (defpol limited to 5,5)

REPEAT/ECHELLE *[scalx,scaly] [response]*
iterate on the response computation

RESPONSE/ECHELLE *[std] [fluxtab] [response]*
compute instrument response

RIPPLE/ECHELLE *input output [params] [method] [option]*
correct for the blaze function

ROTATE/ECHELLE *cat,ima root-name [mode] [flip_axis] [angle] [o_time]*
rotate (and optionally flip) echelle images

SAVE/ECHELLE *name*
saves current echelle session

SAVINIT/ECHELLE *ima,tab mode*
saves/reads echelle session keywords as descriptor of an image/table

SCAN/ECHELLE *frame [scan-par]*
update echelle keywords SCAN and IMSIZE.

SEARCH/ECHELLE *frame [width2,thres2]*
search for emission lines

SEARCH/ORDER *[ordref] [w,t,s] [ordtab] [defmtd]*
define echelle order positions

```

SELECT/BACKGROUND [all]
    interactive unselection of background reference positions
SET/ECHELLE par=value [...]
    set echelle keywords
SHOW/ECHELLE
    show echelle session
SUBTRACT/BACKGROUND input bkg output [bkgmtd] [bkgvisu]
    compute and subtract background from input frame.
TUTORIAL/ECHELLE
    demonstrates main commands of echelle package
UPDATE/ECHELLE image
    low-level command handling image geometry in world-coordinates
UPDATE/ORDER image [offset]
    Updates order definition coefficients and background table.
VERIFY/ECHELLE file [type]
    check consistency of frame size against predefined values.

```

A.3.5 irac2

```

ACUTS/IRAC2 [image] [load] [plot] [upper]
    Display an image with cuts mean-3*sig and mean+upper*sig
CMASK/IRAC2 ffield clnffield lthrshold,hthrshold [dispflag]
    create a mask of bad pixels using a flatfield.
DCOMB/IRAC2 [select] [seqname] [accsky] [align] output [trim] [tag]
    Sky subtract and combine dithered images.
FFIELD/IRAC obj_frame ff_frame out_frame
    Flat Field an IRAC frame
FOCUS/IRAC2 seqnum [focout] [create]
    Used to determine the best focus from a focus sequence.
LAST [num]
    Gives very brief information on the most recent exposures
MASK/IRAC2 inframe outframe
    replaces bad pixels by closest good pixel
MKFLAT/IRAC lamp_on lamp_off flat_field
    Make a flat field
OBSLIST/IRAC2 [start] [end]
    Lists a subsection of the IRAC2B OST (Observation Summary Table)
OBSREP start end
    Print out a subsection of the IRAC2B OST (Observation Summary Table)
QL/IRAC2 image1 image2 [outimage]
    Subtracts one IRAC2 image from another taking into account the
    detector integration times.

```

RCOMB/IRAC2 select [align] output
 Combine frames created with the command RCOMB/IRAC2

SEEING/IRAC2
 Determine the seeing, defined as the FWHM of *stellar images*, of IRAC2 images.

SSUB/IRAC *obj_frame* *sky_frame* *out_frame*
 Sky Subtract an IRAC frame

A.3.6 irspec

BADPIX/IRSPEC *in* *out* [*l=load_opt.*] [*dir=clean_opt.*] [*debug=debug_opt.*]
 Clean image of fixed pattern of bad pixels

CALIBRATE/IRSPEC *ima*
 Apply on-line (mechanical) wavelength calibration.

CALIBRATE/IRSPEC *ima_ref* *mode=define*
 Define and store in *ima_ref* precise calibration from sky/lamp lines

CALIBRATE/IRSPEC *ima* *ref=ima_ref*
 Apply precise wavelength calibration to frame *ima* from parameters stored in *ima_ref*

DEFINE/IRSPEC *image* *table* [*mode*] [*threshold*] [*number*] [*load_option*]
 Define fixed pattern of bad pixels and store it into a table.

**FLAT/IRSPEC *in_flat* *in_dark* *out* [*l=load_opt.*] [*t=threshold*]
 [*v=vignetted_value*]**
 Create a normalized flat from an input flat frame.

**FLUX/IRSPEC *in_ima* *response_ima* *out_ima* [*smooth=s1,s2*] [*shift=sh*]
 [*norm=normalize_option*] [*rect=rectify_option*]**
 Flux calibrate a spectrum (either 2D or 1D) using a response frame created using RESPONSE/IRSPEC

**MERGE/IRSPEC *prefix_ima* *i1,i2[,i3]* *out_table* [*excl=#pixels_excluded*]
 [*corr=correct_option*] [*ref=#reference_image*] [*plot=plot_option*]
 [*format=i_format*]**
 Merge 1D spectra into a table forcing connection of overlapping regions

RECTIFY/IRSPEC *in* *out* [*l=load_opt.*] [*tilt=tilt_value*] [*ref=reference_row*]
 Rectify tilted spectral lines.

**RESPONSE/IRSPEC *in_ima* *flux_table* *out_response_ima* [*yrows=y1,y2,y3,y4*]
 [*obs=observation_mode*] [*norm=normalize_option*]
 [*rect=rectify_option*]**
 Create a response frame from a standard star 2D spectrum and a flux table.

**SKYSUB/IRSPEC *ima_obj* *ima_sky* *out* *factor[,shift[,deltax,deltay]]*
 [*sky=sky_table*] [*force=force_sky_to_zero*]
 [*cuts=cuts_values*] [*debug=debug_option*]**
 Perform obj-sky correcting for variation and shift of sky lines.

STANDARD/IRSPEC *in_ascii_file out_table interp_method*
[degree=degree] [step=wavelength_step] [limits=w11,w12]
[units=wavelength_units] [plot=plot_option]
 Create standard star flux table from a "flux ascii file".
SUBTRACT/IRSPEC *in_ima out_ima degree [exclude=area_to_exclude]*
[cont=continuum_image] [load=load_option]
 Fit and subtract, row by row, polynomial to a given image.
TUTORIAL/CALIBR
 demonstration of wavelength calibration commands in IRSPEC
TUTORIAL/IRSPEC
 General tutorial for the package IRSPEC
TUTORIAL/SKYSUB
 Tutorial for sy subtraction with IRSPEC package

A.3.7 long

APPLY/DISPERSION *in out [y] [coef]*
 Apply the dispersion relation to a 1D spectrum and generates a table
BATCH/LONG
 Prepare the Batch Reduction user interface
CALIBRA/FLUX *in out [resp]*
 Correct an image for the instrumental response function
CALIBRATE/LONG *[tol] [deg] [mtd] [guess]*
 Wavelength calibration of 1D and long-slit spectra
CALIBRATE/TWICE
 Performs the wavelength calibration on a selected set of lines.
CLEAN/LONG
 Clear context Long
COMBINE/LONG *catalog output [mtd]*
 Average a catalog of images
EDIT/FLUX *[resp]*
 Edit the instrumental response table
ERASE/LONG
 Interactive rejection of dispersion relation nodes.
ESTIMATE/DISPERSION *wdisp wcent [ystart] [line] [cat]*
 Estimate a linear approximation of the dispersion relation
EXTINCTION/LONG *in out [scale] [table] [col]*
 correct spectra for interstellar or atmospheric extinction
EXTRACT/AVERAGE *in out [obj] [sky] [mtd]*
 extract a long-slit spectrum by averaging rows
EXTRACT/LONG *in out [sky] [obj] [order,niter] [ron,g,sigma]*
 Optimal extraction of a long-slit spectrum
GCOORD/LONG *[number] [outtab]*

Get coordinates from the display window

GRAPH/LONG [*size*] [*position*] [*id*]
Creates a graphic window

HELP/LONG [*keyword*]
provides information about session keywords.

IDENTIFY/LONG [*wlc*] [*ystart*] [*lintab*] [*tol*]
Interactive calibration of lines in an arc spectrum

INITIALIZE/LONG [*session*]
Initialises parameters of context long

INTEGRATE/LONG *std* [*flux*] [*resp*]
Generates an intermediate response table from a standard star spectrum

LINADD/LONG *in* *w,bin* [*y*] [*mtd*] [*mode*] [*line*]
Adds entries to the table line.tbl

LOAD/LONG *image* [*scale_x*, [*scale_y*]]

MAKE/DISPLAY
Creates a display window

NORMALIZE/FLAT *in* *out* [*bias*] [*deg*] [*fit*] [*visu*]
Normalisation of flat-fields

PLOT/CALIBRATE [*mode*]
Plot wavelength calibration identifications.

PLOT/DELTA [*mode*]
Plot the fitted dispersion relation and allow interactive rejection of arc lines.

PLOT/DISTORTION *wave* [*delta*] [*mode*]
Plot the fitted position of arc lines in wavelength/y-coordinate space.

PLOT/FLUX [*fluxtab*]
Plot the flux table

PLOT/IDENT [*wlc*] [*line*] [*x*] [*id*] [*wave*]
Plot interactive identifications

PLOT/RESIDUAL [*y*] [*table*]
Plots residual after wavelength calibration

PLOT/RESPONSE [*resp*]
Plots the response correction function

PLOT/SEARCH [*mode*] [*table*]
Plot the results of SEARCH/LONG

PLOT/SPECTRUM *table*
Plots a 1D spectrum *in table format, as supplied by APPLY/DISPERSION*

PREPARE/LONG *in* [*out*] [*limits*]
Extracts sub-images from an image or a catalog.

REBIN/LONG *in* *out* [*start, end, step*] [*mtd*] [*table*]
Rebin a long-slit spectrum using the row-by-row method

RECTIFY/LONG *in* *out* [*reference*] [*nrep*] [*deconvol_flag*] [*line*]
rectify geometrically a distorted 2-D spectrum

REDUCE/INIT *partab*
 Initialises the batch reduction parameters
REDUCE/LONG *input*
 Batch reduction of long-slit spectra.
REDUCE/SAVE *partab*
 Saves the batch reduction parameters
RESPONSE/FILTER *std* [*flux*] [*resp*]
 Generate a response image by filtering based method.
RESPONSE/LONG [*plot*] [*fit*] [*deg*] [*smo*] [*table*] [*image*] [*visu*]
 Converts the response correction from table to image format.
SAVE/LONG *session*
 Saves session keywords
SEARCH/LONG [*in*] [*thres*] [*width*] [*yaver*] [*step*] [*mtd*] [*mode*]
 search for spectral features in a long-slit spectrum
SELECT/LINE
 Select lines identified in all rows of an arc spectrum.
SET/LONG *key=value* [...]
 Assigns a value to long-slit session keywords
SHOW/LONG [*section*]
 Displays values of session keywords.
SKYFIT/LONG *input* *output* [*sky*] [*degree*] [*mode*] [*r,g,t*] [*radius*]
 fit polynomial to spatial flux distribution in windows of every column
TUTORIAL/LONG
 demonstrate commands of the package Long
VERIFY/LONG *file mode*
 Checks conformity of files in the long-slit context
XIDENT/LONG [*wlc*] [*ystart*] [*lintab*] [*tol*]
 Invoke the identification graphical user interface

A.3.8 optopus

CREATE/OPTOPUS *inp_file* [*out_tab*] [*fmt_file*] [*old_equinox*]
 create input table for HOLES/OPTOPUS command
DRILL/OPTOPUS *in_table* [*name*]
 write OPTOPUS *drill command file*
HOLES/OPTOPUS [*inp_tab*] [*out_tab*] [*HH,MM,SS.sss*] [+/-*DD,AM,AS.ss*]
[*ac_flag*] [*p_flag*] [*old_eq,new_eq*]
 determine holes positions on Optopus plate.
MODIFY/OPTOPUS [*table*]
 plot positions of holes on plate and enable rejection of objects.
PLOT/OPTOPUS [*table*] [*label*] [*EW_flip_flag*]
 plot positions of holes on Optopus plate.
PRECESS/OPTOPUS [*inp_tab*] [*new_equinox*]

p recess RA and DEC coordinates in table created by CREATE/OPTO.
REFRACTION/OPTOPUS [*inp_tab*] [*out_tab*] [*year,month,day*] [*exp*] [*lambda1,lambda2*]
 [*start_st_slot,end_st_slot*] [*opt_st*] [*ast_flag*]
 correct for atmospheric refraction X and Y coord. on Optopus plate.
RESTORE/OPTOPUS *table*
 restore previously saved session parameters.
SAVE/OPTOPUS *table*
 save session parameters in descriptors
SET/OPTOPUS *option1[=value1]* *option2[=value2]* ...
 set Optopus context parameters
SHOW/OPTOPUS
 show session parameters.
ZOOM/OPTOPUS [*table*] [*zooming_factor*]
 blow up section of Optopus plate and enable rejection of objects.

A.3.9 pisco

REDUCE/PISCO *catalog table sky calibration* [*mode*]
 perform complete reduction of polarimetric data

A.3.10 spec

CENTER/HISTOGRAM *image*
 Median estimate and scale estimates of an image
COMPUTE/PARAL *ra dec st wave refw*
 Computes parallactic angle and atmospheric differential refraction
CONTINUUM/SPEC *in out* [*radius/meth*] [*type*] [*smooth*] [*degree*]
 Fitting of a spectrum continuum by smoothing splines
CORRELATE/LINE *table_1 table_2* [*pixel*] [*cntr,tol,rg,st*] [*pos,ref,wgt*]
 [*ref_value*] [*outima*]
 Cross-correlation between table columns.
CUMULATE/HISTOGRAM *in out*
 Transforms a histogram image into the cumulated histogram
DEBLEND/LINE *infile* [*fitim*] [*fitpar*] [*method*] [*contin*] [*input*] [*intab*]
 multiple component Gaussian fitting of spectral lines
DISPERS/HOUGH [*wdisp*] [*wcent*] [*fr_specs*] [*line*] [*cat*] [*mode*] [*range*]
 [*vflag*]
 Determination of dispersion relations by HT
EXTINCTION/SPECTRUM *inframe outframe* *scale* [*table*] [*col*]
 correct spectra for interstellar or atmospheric extinction
FILTER/RIPPLE *frame outframe* *period* [*start,end*]
 correct 1-dim. images for periodic ripple (Reticon)

GRAPH/SPEC *[size] [position] [id]*
 Creates a long graphic window adapted for spectroscopy
MERGE/SPECTRUM *spec1 spec2 out [interval] [mode] [vari] [var2]*
 merge two 1D spectra
NORMALIZE/SPECTRUM *inframe outframe [mode] [table] [batch_flag]*
 approximate continuum of 1-dim. spectra for later normalization
OVERPLOT/IDENTIFICATION *[table] [xpos] [ident] [ypos]*
 overplot line identifications
REFRACTION/LONG *inim outim [mode]*
 Differential atmospheric correction for slit spectra
REGRESSION/ROBUST *tab y x1[,x2,..,xn] [file] [out_col] [res_col]*
 Robust multi-variate regression by Least Median of Squares
ROTATE/SPEC *cat [root] [meth] [flip] [angle] [mode]*
 rotate (and optionally flip) a catalog of images
SEARCH/LINE *frame w, t[,nscan] [table] [meth] [type]*
 search for spectral lines
VERIFY/SPEC *file dir keyw [type]*
 low-level spec command checking the existence of calibration tables

A.4 Contributed Commands

A.4.1 astromet

ASTROMETRY/COMPUTE *mes option out trail*
 Convert the coordinates from the measured xy to RA,Dec vice versa
ASTROMETRY/EDIT *std plot*
 Delete/undelete the astrometric standards
ASTROMETRY/POS1
 Interactive procedure for the POS1 astrometry package
ASTROMETRY/TRANS *std mes center pla,cat schmidt,blink tol xterm,yterm std*
 Compute the astrometric transformation parameters of a data

A.4.2 cloud

COMPUTE/ABSORPTION *inframe outframe [cm_table] [ap_table] [psframe]*
 computes a synthetic 1dim. absorption spectrum
COMPUTE/EMISSION *outframe [em_table]*
 computes a synthetic 1dim. emission spectrum
CREATE/PSF *[outframe] fwhm*
 creates a 1-dim. image of a normalized gaussian

A.4.3 daophot

```

ALLSTAR/DAOPHOT
    do simultaneous multiple-profile-fitting
DAOMID/DAOPHOT table
    convert a DAOPHOT table into a MIDAS table
DAOPHOT/DAOPHOT
    do precise photometry and astrometry in a 2-dim frame
MIDDAO/DAOPHOT table
    convert a MIDAS table into a DAOPHOT table

```

A.4.4 esolv

```

FROMOD/ESOLV [mode] [colour] [intable] [column]
    retrieve frames from optical disk
MTABLV/ESOLV [table] col1 col2 lfrac col3
    find semidiameter of ellipses at given fraction of light
STATPL/ESOLV table col1 select disp
    computes mean and sd of table file column given in p2
TABFLV/ESOLV [table] ascii_file flag
    lists the contents of special table file
TEXLV/ESOLV [table] tex_file
    prepare Tex file with selected parameters from ESOLV

```

A.4.5 geotest

```

CREATE/ART_IMAGE frame frame dims [starts,steps] [func_type] [coefs]
CREATE/ART_IMAGE frame = ref_frame [func_type] [coefs]
    create artificial image
CREATE/RAMP image [slope] [angle] [dimension]
    generate uniform sloping image, with mean flux per pixel of 100 units
CREATE/SPC1 image [slope] [ampl] [period] [phase] [dim]
    generate sinusoidal, sloping 1-dimensional image
CREATE/SPC2 image [period] [slope] [phase] [dimension]
    generate a discrete 1-dimensional image
CREATE/SPC3 image psf_option centring table boxwidth-or-fwhm
    generate an artificial spectrum with lines
CREATE/WAVE image [amplitude] [period] [dimension]
    generate 2-dimensional sinusoidal background image

```

A.4.6 invent

ANALYSE/INVENTORY *frame in_tab [out_tab] [ver_par] [deb_mode] [out_psf]*
 verify the used table of objects and calculates the image parameters

CLASSIFY/INVENT *table*
 classify the analysed objects into stars, galaxies and spurious objects

SEARCH/INVENTORY *frame table*
 search objects in an image frame and store the parameters

SET/INVENTORY *par1 [par2]*
 display and modify the values of the keywords used by Inventory

SHOW/INVENTORY *par1 [par2]*
 display the values of the keywords used by the Inventory package

A.4.7 mva

CLUSTER/TABLE *intable outable [method]*
 hierarchical clustering

CMDS/TAB *input_table output_table ncols.._output_table*
 multidimensional scaling

CORRES/TAB *input output row/column_analysis ncolumn outable*
 correspondence analysis

EDIST/TAB *input_table output_table*
 standard distances

KNN/TAB *training_table no._of_gp.1_members test_table no._of_NNs*
 discriminant analysis

LDA/TAB *Input_table Output_table*
 Fisher's linear discriminant analysis.

MDA/TAB *input_table output_table eigenvectors*
 discriminant analysis

MST/TABLE *intable outable grid_size*
 create minimal spanning tree for position table

PARTITION/TABLE *intable outable [no_of_class] [alg] [min. card] [s_value]*
 non-hierarchical clustering

PCA/TAB *in_tab out_tab option row/col_anal ncols_table eigenvectors*
 principal components analysis

PLOT/TREE *intable [col_ref]*
 plot output created by minimal spanning tree algorithm

A.4.8 pepsys

CONVERT/PHOT
 Helps you make a new table of observational data

MAKE/HORFORM

Make a blank FORM to fill in with horizon-obstruction data
 MAKE/PHOTOMETER
 generates or checks the instrumental table file for a photometer
 MAKE/PLAN
 generates a photometric observing plan
 MAKE/STARFILE arglist
 Helps you make a new file of program or standard stars
 REDUCE/PHOT
 Reduces tables of observational data

A.4.9 romafot

ADAPT/ROMAFOT int_tab [thres] [fac_int] [fac_sky] [fac_hol] [x_siz,y_siz]
 derive trial values for fitting a new frame
 ADDSTAR/ROMAFOT in_frame out_frame [reg_tab] [cat_tab] [x_dim,y_dim]
 [n_sub]
 create an artificial image with subframes added at random positions
 ADSTAR/ROMAFOT in_frame out_frame [reg_tab] [cat_tab] [x_dim,y_dim]
 [n_sub]
 create an artificial image with subframes added at random positions
 ANALYSE/ROMAFOT frame [cat_tab] [int_tab] [sigma,sat]
 INPUT MODE select all stars within selected subfields;
 OUTPUT MODE check at the results of the fit operation and select
 CBASE/ROMAFOT frame_1 frame_2 [out_tab1] [out_tab2]
 create two tables for coordinate transformation
 CHECK/ROMAFOT cat_tab reg_tab err_mag
 examine number of artificial stars recovered and check the accuracy
 CTRANS/ROMAFOT int_tab [tab_1] [tab_2] [pol_deg]
 find transformation of coordinates and apply to an intermediate table
 DIAPHRAGM/ROMAFOT frame [regi_tab] [rego_tab] ap_rad
 do aperture photometry with fixed diaphragm
 EXAMINE/ROMAFOT int_tab [hmin,hmax]
 examine quality of fitted objects and flag badly fitted ones
 FCLEAN/ROMAFOT cat_tab inti_tab [into_tab]
 selects windows in intermediate table present in catalogue table
 FIND/ROMAFOT frame [cat_tab]
 select objects using the image display
 FIT/ROMAFOT frame [int_tab] [thres,sky] [sig,sat,tol,iter] [meth,[beta]]
 [fit_opt] [mean_opt]
 determine characteristics of stellar images by non-linear fitting
 GROUP/ROMAFOT frame [area] [cat_tab] [int_tab] [thres] [wnd_max]
 [end_rad,sta_rad] [wnd_perc]
 automatic grouping of objects

```

MFIT/ROMAFOT frame [int_tab] [thres,sky] [sig,sat,tol,iter] [meth,[beta]]
  [fit_opt] [mean_opt] [mod_file]
      determine characteristics of stellar images by non-linear fitting
MODEL/ROMAFOT [mod_file]
      compute (sub)pixel values for a model observation
REGISTER/ROMAFOT int_tab reg_tab [wnd_opt] [obj_opt]
      computes and store the absolute quantities in the registration table
RESIDUAL/ROMAFOT in_frame out_frame diff_frame [reg_tab]
      compute reconstructed image and difference with original image
SEARCH/ROMAFOT 07-AUG-1989 RHW
      do the actual search for objects above a certain threshold
SELECT/ROMAFOT frame [int_tab] [wnd_size]
      select objects and store the positions in intermediate table
SKY/ROMAFOT frame [sky_tab] [area] [nrx,nry] [min,max]
      determines intensity histogram and sky background in selected areas

```

A.4.10 surfphot

```

COMPUTE/FCOEFF infram orient rin,rout,rstep outtab
      compute fourier coefficients of azimuthal profiles in spiral galaxies
COMPUTE/GRID angle
      create image and table with distorted rect. and evenly spaced grid
COMPUTE/SKY infram1 infram2 caltab method sky_factor
      compute the sky background and restitute the frame
FILTER/FILL inframe outframe rx,ry thresh
      fill up low-flux (below threshold) pixels with nearby high-flux pixels
FIND/PAIR intab1 intab2 outtab columns [errors] [coo_syst]
      match (pair) two coordinates tables and produce an output table
FIND/POSINC infram x_pos,y_pos rin,rout,rstep
      find the position angle and inclination of a galaxy
FIT/BACKGROUND outframe = inframe(s) [deg,it] [clp1,clpn] [skew] [outbck]
FIT/BACKGROUND outframe = inframe(s) [coef] [outbck]
FIT/BACKGROUND inframe(s) [deg,it] [clp1,clpn] [skew] [outbck]
FIT/BACKGROUND inframe(s) [coef] [outbck]
      compute 2-dim. polynomial fit of the background
FIT/ELL1 inframe outframe l_iso,h_iso x_cen,y_cen max_rad
      fit an ellips with respect to predefined center
FIT/ELL2 inframe pol_opt iso_tol iso_levels [center[ [radius] [sky_level]
      fit an ellips with respect to predefined center
FIT/ELL3 inframe outframe [step] [x,y] [low,high] [min,max] [opt]
      fit ellipses to the isophotes of an object in a 2-dim. frame
FIT/POSINC infram orient rin,rout,rstep region
      fit the position angle and inclination to 2nd and 4th harmonic

```

```

INTEGRATE/ELLIPS frame [ellips.par] [flag]
    integrate pixel intensities within ellipse in 2-dim. image
NORMALIZE/IMAGE infram outfram trunc_vals control_vals
    normalize and truncate a frame
REBIN/DECONVOLVE frame psf result zoom_x,zoom_y n_iter
    rebin image linearly in space and simultaneously deconvolve it with psf
RECTIFY/IMAGE in out table [nrep] [deconvol_flag]
    rectify geometrically a distorted direct image
SUBTRACT/SKY inframe outframe nx,ny
    remove sky by subtraction of histogram-modeled substitute-sky

```

A.4.11 tsa

```

AOV/TSA intab outima start step nsteps [order] [cover]
    compute analysis of variance periodogramme
BAND/TSA intab [maxobs] Evaluate frequency band for time series analysis
COVAR/TSA intab1 intab2 outtab start step nsteps scale
    compute discrete covariance function for unevenly sampled data
DELAY/TSA intab1 intab2 outtab start step nsteps [func,mode] [parm]
    compute chi2-time lag function
INTERPOLATE/TSA intab outtab func parm
    Interpolate an unevenly sampled series using its covariance function
NORMALIZE/TSA intab1 outtab column [mode]
    Normalize mean and variance to 0 and 1
POWER/TSA intab outima start step nsteps
    Compute discrete power spectrum for uneven sampling by slow method
SCARGLE/TSA intab outima start step nsteps
    Compute Scargle periodogramme for unevenly spaced observations
SET/TSA set global keywords for TSA context
SHOW/TSA show global keywords for TSA context
SINEFIT/TSA intab outtab freque order iter
    fit sine (Fourier) series, subtract it from input and return residuals
TABLE/TSA inascii [outtab] [type] [mxcol] convert ASCII table into
MIDAS table
WIDTH/TSA inima [width] [centre] Evaluate line width and profile

```

A.5 Procedure Control Commands

```

BRANCH var comparisons labels
    multi-way branching
CROSSREF label1 label2 label3 label4 label5 label6 label7 label8
    define cross reference labels for the 8 parameters

```

```

DEFINE/LOCAL key_def data A lower_levels_flag
  define the maximum no. of parameters for a procedure
DEFINE/PARAMETER Pi default type/option prompt_str low_lim,hi_lim
  define default, type and valid interval for parameter i
DO loopvar = start end [step]
  define a DO loop (as in FORTRAN)
ENTRY proc
  define begin of procedure in a file with different name than the procedure
GOTO label
  branch to command line containing label:
IF par1 op par2 command_string
  execute conditional statement
INQUIRE/KEY key prompt_string
  get terminal input in a MIDAS procedure
LABL:
  define a label, LABL in this example
RETURN par1 par2 par3
  return to calling procedure (or terminal) and optionally pass up to 3
  parameters back

```

A.6 Commands Grouped by Subject

In the following list, general MIDAS commands are given grouped in main application areas. Only the most common commands are listed to make the list easier to use. Commands used for special types of data reduction are given in appropriate chapters in the main part of this manual.

A.6.1 MIDAS System Control

@@	Execute a MIDAS procedure
BYE	Terminate the MIDAS session
CHANGE/DIRECTORY	Change the default (current) directory for MIDAS
CLEAR/CONTEXT	Clear current context level or all levels
COMPUTE/KEYWORD	Compute values of a keyword
CONNECT/BACK_MIDAS	Connect "command syntax" to another MIDAS
CREATE/COMMAND	Create a user command
CREATE/DEFAULTS	Create special defaults for MIDAS commands
DEBUG/PROCEDURE	Run MIDAS procedures in debug mode
DEBUG/MODULE	Run MIDAS modules in debug mode
DELETE/COMMAND	Delete a user defined command

DELETE/DEFAUTLS	Delete special defaults for command
DELETE/IMAGE	Delete an image frame
DELETE/KEYWORD	Delete user defined keyword
DELETE/LOG	Delete log file
DISCONNECT/BACK_MIDAS	Disconnect from a background MIDAS
ECHO/FULL	Show substitutions in program files
ECHO/OFF	Suppress display of input from program files
ECHO/ON	Display input from program files
LOG/OFF	Disable logging
LOG/ON	Enable logging
PLAYBACK/LOG	Playback log file
READ/KEYWORD	Display contents of keywords
RENAME/IMAGE	Rename an image frame
RUN	Execute program inside MIDAS
SAVE/COMMANDS	Save commands from command window in a procedure
SET/CONTEXT	Set new context level
SET/FORMAT	Format for "number-to-string" conversion
SET/MIDAS_SYSTEM	Set different modes and options for MIDAS
WAIT/BACK_MIDAS	Wait until command in background MIDAS terminates
WAIT/SECS	Suspend MIDAS monitor for no_of_secs second
WRITE/COMMANDS	Store commands from a procedure into the command window
WRITE/KEYWORD	Store values into a keyword
WRITE/OUT	Write out text

A.6.2 Help and Information

HELP	Display help info for a command
HELP/...	Display info about various topics
INFO/...	Get information about frames, descriptors and specific setup
PRINT/HELP	Print help information
PRINT/LOG	Print log file
SHOW/COMMANDS	Display MIDAS commands
SHOW/DEFAUTLS	Display all special defaults

A.6.3 Tape Input and Output

INDISK/...	Read data from disk in FITS or ASCII format
INTAPE/FITS	Read data from tape in FITS or IHAP format
OUTTAPE/FITS	Write data to tape in FITS format

A.6.4 Image Directory and Header

ADD/xCAT	Add one or more entries to a catalogue
COPY/DD	Copy descriptors from one file to another
CREATE/xCAT	Create a catalogue

DELETE/...	Delete a frame
DELETE/_DESCRIPTOR	Delete a descriptor
INFO/_DESCRIPTOR	Get type and size of descriptor
READ/_DESCRIPTOR	Read descriptors
RENAME/...	Rename a frame
SORT/xCAT	Sort entries in a catalogue
SUBTRACT/xCAT	Remove an entry from a catalogue
WRITE/_DESCRIPTOR	Write a descriptor
WRITE/DHELP	Write descriptor help

A.6.5 Image Display

BLINK/CHANNEL	Blink between channels
CLEAR/ALPHA	Clear the alpha-numerics memory
CLEAR/CHANNEL	Clear and initialize memory channel
CLEAR/DISPLAY	Reset monitor
CLEAR/LUT	Bypass LUT in screen segment on monitor
CLEAR/SPLIT	Disable split screen
CLEAR/ZOOM	Clear zoom
COPY/CHANNEL	Copy image memory channels
COPY/DISPLAY	Hard copy of image display
CREATE/CURSOR	Create cursor window
CREATE/ZOOM	Create zoom window
CREATE/DISPLAY	Create a display window (using Xwindow)
CUTS/IMAGE	Set display thresholds for image
DELETE/DISPLAY	Delete the display windows
DISPLAY/CHANNEL	Display image loaded into channel
DRAW/...	Draw rectangle and other figures in the overlay plane
EXTRACT/CURSOR	Extract a subframe from the frame currently displayed
EXTRACT/ROTATED	Extract a rotated subimage from displayed image
EXTRACT/TRACE	Extract interactively a line from an image
GET/CURSOR	Coordinates from display device by cursor
GET/IMAGE	Read currently loaded image from channel
INITIALIZE/DISPLAY	Initialize the image display
LABEL/DISPLAY	Write character string on display device
LOAD/CURSOR	Display cursor into display device (DeAnza only)
LOAD/IMAGE	Load image into display device
LOAD/ITT	Load an intensity transfer table
LOAD/LUT	Load a colour lookup table into display unit
LOAD/TABLE	Display table data on image display
MODIFY/LUT	Modify the currently active lookup table
SCROLL/CHANNEL	Scroll image on given channel

SET/CURSOR	Set cursor form and position
SET/DISPLAY	Define colour display control, size of screen etc. (DeAnza only)
SET/LUT	Enable use of colour lookup table
SET/SPLIT	Enable split screen (DeAnza only)
SHOW/CHANNEL	Show information related to channel
VIEW/IMAGE	Explore an image interactively
ZOOM/CHANNEL	Zoom image on display

A.6.6 Graphics Display

ASSIGN/GRAFICS	Define plotter output device and replot
CLEAR/GRAFICS	Clear graphic screen
COPY/GRAFICS	Copy the plot file to the specific graphic device
CREATE/GRAFICS	Create a graphic window (using Xwindow)
CUTS/IMAGE	Set plot thresholds (high and low) for image
DELETE/GRAFICS	Delete the graphic windows
GET/GCURSOR	Coordinates from graphic device by cursor
LABEL/GRAFICS	Plot text in an existing plot
OVERPLOT/ERROR	Overplot table error column
OVERPLOT/HISTOGRAM	Overplot histogram of table column or image
OVERPLOT/ROW	Overplot row/line of image data on previous plot
OVERPLOT/TABLE	Overplot table data on previous plot
PLOT/AXES	Plot a coordinate box with large and small tickmarks and labels
PLOT/CONTOUR	Contour plotting of an image
PLOT/DESCRIPTOR	Plot an entry in a descriptor
PLOT/HISTOGRAM	Plot a histogram of a table column or an image
PLOT/ROW	Plot row/line of an image
PLOT/PERSPECTIVE	Perspective plotting (3-dim.) of an image
PLOT/TABLE	Plot table data
SET/GRAFICS	Set plot characteristics like scaling
SHOW/GRAFICS	Show graphic characteristics

A.6.7 Image Coordinates

CENTER/...	Find center
GET/CURSOR	Coordinates from image display via cursor
GET/GCURSOR	Get coordinates from graphics device by cursor
READ/DESCRIPTOR	List reference coordinates
WRITE/DESCRIPTOR	Write reference coordinates

A.6.8 Coordinate Transformation of Images

ALIGN/IMAGE	Calculate linear transformation between 2 images
EXTRACT/IMAGE	Extract part of image
FLIP/IMAGE	Flip image in x and/or y

GROW/IMAGE	Repeat one scan line to make 2 dim images
INSERT/IMAGE	Insert a subimage into father image
REBIN/II	Logarithmic, exponential, $r^{1/4}$ frequency rebin
REBIN/LINEAR	Pixel rebinning of image
REBIN/ROTATE	Rotate an image any angle
REBIN/SPLINE	Rebin an image with cubic splines
REBIN/WAVE	Rebin 1-D image to linear wavelength
RECTIFY/IMAGE	General geometric correction
ROTATE/CLOCK	Rotate clockwise 90 degrees
TRANSPOSE/CUBE	Rearrange planes of 3-dim data cube
TRANSPOSE/IMAGE	Transpose an image

A.6.9 Image Arithmetic

AVERAGE/AVERAGE	Compute simple average of all pixels in a subimage
AVERAGE/COLUMN	Compute average of image columns
AVERAGE/IMAGE	Calculate the average of images
AVERAGE/ROW	Compute average of image rows
AVERAGE/WINDOW	Compare images, then take the mean
COMPUTE/COLUMN	Perform arithmetic expression on image column
COMPUTE/IMAGE	Compute arithmetic expression of images
COMPUTE/PIXEL	Perform arithmetic operations on images using pixel coordinates
COMPUTE/ROW	Compute arithmetic expression on image scan lines
COMPUTE/..PLANE	Do arithmetic on planes of a data cube

A.6.10 Filtering

CONVOLVE/IMAGE	Convolve image with given point spread function
CREATE/FILTER	Create filter image
DECONVOLVE/IMAGE	Deconvolve image with point spread function
FILTER/GAUSS	Use Gauss filter on image
FILTER/MAX	Apply maximum filter to an image
FILTER/MEDIAN	Median filter image
FILTER/MIN	Apply minimum filter to an image
FILTER/SMOOTH	Smooth an image
FFT/IMAGE	Compute discrete fourier transform of a complex input frame
FFT/INVERSE	Compute inverse discrete fourier transform of a complex input frame

A.6.11 Image Creation and Extraction

COPY/II	Copy image frames
CREATE/IMAGE	Create new image

CREATE/RANDOM	Create a new image from a random distribution
EXTRACT/CURSOR	Extract a subframe from the frame displayed
EXTRACT/IMAGE	Extract part of an image
EXTRACT/LINE	Extract a line from a frame
EXTRACT/ROTATED	Extract a rotated image
EXTRACT/SLIT	Extract subimage defined by fixed slit
EXTRACT/TRACE	Extract line from an image
INDISK/ASCII	Read ASCII file from disk
INDISK/FITS	Read FITS file from disk
INSERT/IMAGE	Insert a subimage into father image

A.6.12 Transformations on Pixel Values

FIT/FLAT_SKY	Correct an image for sky variations
ITF/IMAGE	Transform pixel values in an image
MODIFY/AREA	Remove bad pixel from circular area
MODIFY/CURSOR	Change pixel values in image by cursor
MODIFY/GCURSOR	Change pixel values in image by graphic cursor
MODIFY/PIXEL	Change pixel values in image
REPLACE/IMAGE	Modify pixel values in given intensity interval
REPLACE/POLYGON	Replace pixel values inside a polygon

A.6.13 Numerical Values of Image Pixels

FIND/MINMAX	Display (ands tore) max and min value
FIND/PIXEL	Find first pixel with a value inside or outside the interval
FIT/FLAT_SKY	Fit background image
INTEGRATE/APERTURE	Integrate flux inside aperture
INTEGRATE/LINE	Integrate pixel-values over area in image
MAGNITUDE/CIRCLE	Compute the magnitude of the specified object by integrating over the central area defined by a circular aperture
MAGNITUDE/RECTANGLE	Compute the magnitude of the specified object by integrating over the central area defined by a rectangular aperture
MODIFY/CURSOR	Change pixel values in image by cursor
MODIFY/GCURSOR	Change pixel values in image by graphic cursor
MODIFY/PIXEL	Change pixel values in image
PLOT/HISTOGRAM	Plot histogram of pixel values in image
PRINT/IMAGE	Print an image
READ/IMAGE	List pixel values into image
STATISTICS/IMAGE	Calculate statistics of an image
WRITE/IMAGE	Change pixel values in image (world coordinates)

A.6.14 Spectral Analysis

CALIBRATE/LINE	Calculate coefficients for wavelength calibration
CENTER/...	Compute center of line
CONVERT/TABLE	Make image from table values
EXTINCTION/SPECTRUM	Correct 1-D image for extinction
IDENTIFY/GCURSOR	Identify table entries from graphic display
IDENTIFY/LINE	Equate X positions to wavelengths
INTEGRATE/GCURSOR	Integrate line interactively
MODIFY/GCURSOR	Change data in line interactively
OVERPLOT/IDENT	Overplot line identifications
PLOT/IDENT	Plot line identifications
REBIN/...	Linear or non-linear image rebinning
RESPONSE/SPECTRUM	Make file for flux correction, response curve
SEARCH/LINE	Search calibration lines

A.6.15 Least Squares Fitting

COMPUTE/FIT	Compute fitted image or table
COMPUTE/FUNCTION	Compute function values of image or table
EDIT/FIT	Define function for fitting
FIT/IMAGE	Least squares fitting in image
FIT/TABLE	Least squares fitting in table
PRINT/FIT	Print fitted values
READ/FIT	Read fitted values
SELECT/FUNCTION	Select functions to be fitted
SET/FIT	Control execution of fitting
SHOW/FIT	Display control parameters

A.6.16 Table File Operations

BIN/TABLE	Create a table with averages of col2 in bins of col1
COMPUTE/HISTOGRAM	Compute histogram for a table column
COMPUTE/REGRESSION	Compute column from regression coefficients
COMPUTE/TABLE	Compute arithmetic expression between columns
CONVERT/TABLE	Compute image from table data
COPY/TT	Copy keys from table to table file
CREATE/COLUMN	Create new column in a table file
CREATE/TABLE	Create a table file
DELETE/COLUMN	Delete column from an element in a table file
EDIT/TABLE	Change value of entry in table file
MERGE/TABLE	Merge two table files
NAME/COLUMN	Insert a label name for a column
PRINT/TABLE	Print table
READ/TABLE	List elements of a table file

REGRESSION/POLYNOMIAL	Compute regression between column in table file
SELECT/TABLE	Select a subtable
SHOW/TABLE	List table directory
SORT/TABLE	Order a table file
STATISTICS/TABLE	Computes low order statistics for a column

Appendix B

Acknowledgements

B.1 General

It is of course never possible to adequately acknowledge the many small, but extremely useful, comments which the Data Management Division have received from many colleagues both within ESO and outside. Nevertheless, we would like to express our gratitude to those who have helped to make MIDAS what it is today and hopefully what it will be in the future. In addition, we would like to try to specifically acknowledge certain major contributions.

B.2 Packages and Commands

The fitting routines in MIDAS were developed in close collaboration with O. Richter and later upgraded by Ph. Defert. The INVENTORY programs were developed and written by A. Kruszewski during his extended visits to ESO. The multivariate statistical package has been developed in close collaboration with F. Murtagh. The package for 1-dimensional spectral reductions was designed and tested in collaboration with D. Baade and M. Rosa. The ROMAFOT package for crowded field photometry was developed by R. Buonanno, C. Buscema, C. Corsi, I. Ferraro, and G. Iannicolo at the Osservatorio Astronomico di Roma. The implementation of ROMAFOT in MIDAS was done in collaboration with R. Buonanno. M. Tapia and A. Moneti collaborated in the development of the IRSPEC reduction package. Marguerite Pierre developed modelling commands for interstellar absorption work and contributed to the Long Slit package. P. Stetson created a MIDAS compatible version of DAOPHOT-II.

The digital filter to remove cosmic ray events from single frames was contributed by P. Magain and M. Remy. Several routines used in the COMPUTE command for calculations of airmass, barycentric correction, ST, UT and Julian data were kindly made available by D. Gillet. The FILTER/ADAPTIVE command was kindly provided by G. Richter. The OPTOPUS context was implemented by A. Gemmo. The PISCO context was implemented by M. Schloetelburg and O. Stahl. The astrometry context ASTROMET was made available for MIDAS by O. Hainaut. The original code was written by R. West.

Significant contributions were also added in the application area. A number of important applications were made by ESO La Silla (or in close collaboration with them) such as the new Long Spectral Package and the XAlice graphic user interface for spectral analysis. The IRSPEC reduction was revised by E. Oliva, while an image restoration and co-addition application, based on ideas of L. Lucy, was added by R. Hook (ST-ECF). A Time Series Analysis context, which includes analysis of non-equally spaced data, was made by A. Schwarzenberg-Czerny. The Wavelet package was implemented by J. L. Starck and contains both general routines for wavelet transforms and special applications *e.g.* for deconvolution. Finally, the PEPSYS context was introduced by A.T. Young as the first application in a new context for calibrations of point-source photometry.

B.3 Libraries

B.3.1 AGL

The plotting package in MIDAS is based on the low level routines in the Astronet Graphic Library (AGL) which was developed and is maintained by the Italian ASTRONET. The implementation of the AGL library in MIDAS was done with the help of L. Fini.

B.3.2 IDI

In the early implementations of IDI routines for XWindows the Trieste Observatory (Mauro Pucillo, Paolo Santin, Fabio Pasian) provided a prototype for X10 which has been used for our further developments.

B.4 Manual

This manual has been typeset in \TeX and \LaTeX using an extensive set of macros provided by H.-M. Adorf.

Appendix C

Site Specific Implementation

This appendix describes the site specific hardware setup and implementations used in ESO, Garching.

C.1 Hardware Setup

This section gives short description of the hardware configuration of the general ESO computer facilities in Garching used for image processing. The main installation contains a number of UNIX workstations and Servers (most of which are SUN/SPARC compatible). The logical names of the workstations are `wsn` where n is a running number. The Servers have the prefix `ns` or `mc` depending on their main function as either file servers or main computers. In addition several X-terminal are available. All the systems are interconnected through a Local Area Network using TCP/IP protocols.

C.1.1 UNIX Workstations

A number of SUN/SPARC workstations and X-terminals for general MIDAS use are located in the User room 213. They run UNIX and are configured with the X11 window system. You can login by giving the **Userid/Password** allocated to you. When working in the ESO X11 environment, it is necessary to point the cursor on the X11 OSF/Motif window in question to get access to it. Several MIDAS sessions may be run on a workstation at the same time using the `parallel` option when starting MIDAS with the `inmidas` command. Different MIDAS unit numbers can also be given specified on the command line.

Below we describe the devices currently supported by MIDAS. A complete list of available devices at ESO-Garching the reader is referred to to ESO WWW page <http://http.hq.eso.org/ginfo/computing/devices.html>

C.1.2 Printer and Plotter Queues

To output listings and hard copies of plots a number of devices can be used. These can be used through a number of system queues which spool the output to the appropriate

devices. The majority of these printers are PostScript compatible.

To print the log of the MIDAS session the PRINT/LOG command is used. This command will produce the output on printers, most of which are located in the User room.

Hardcopies of plots are made using the COPY/GRAFH and the PLOT commands (see Chapter 6). The first parameter of the ASSIGN/GRAFH command specifies the plotting queue/device to be used. The help command in MIDAS can be used to get an up to date list of output devices.

C.1.3 X11 Window systems

Image display windows can be created and deleted from MIDAS using the commands CREATE/DISPLAY and DELETE/DISPLAY, respectively. All MIDAS display commands will work on these window displays, some have to be implemented in software which makes them slow *e.g.* ZOOM and SCROLL. The cursor is implemented through the mouse on which the left button is **Enter** and the middle button (or right button on a 2 button mouse) is **Exit**. When the cursor is positioned on a feature it is also possible to use the **Return** key to read the cursor location. Some commands use a second cursor which is implemented through the arrow key on the keyboard.

C.1.4 Film Hardcopy

This Section describes the Film hardcopy devices available at ESO, Garching. Only a system for recording 35mm Colour Slides is offered.

Film print queues

They are implemented as standard UNIX print queues and can be accessed with the UNIX lpr command. Files written to a queue **must** be in the correct format (see ESO WWW page <http://http.hq.eso.org/garching-info/computing/devices.htm>)

In the case of Colour Slide the MIDAS command COPY/DISPLAY SLIDE can be used. This command will both create the PostScript image file and submit it to the slide queue.

Processing of Film Hardcopies

Depending on the number of entries, images in the colour queue will be checked and activated once per week. Development of the films may take upto two weeks after which they will be send to the user. Thus, a turn-a-round time of approximately one month must be expected when using film hardcopy.

C.1.5 Tape I/O

Several tape drives (1/2 inch, QIC, 8mm, DAT) are available on the UNIX system. They can be accessed from any UNIX workstation on the network. The step by step procedure is as follows:

1. Login on a UNIX workstation.

2. Run MIDAS with **inmidas**.
3. Mount your tape on a drive connected (see below). If you want to write on the tape remember to enable writing.
4. Use the **intape** and **outtape** commands to access the tape unit using the logical device name *e.g.* **tape0**, or the physical device name *e.g.* **/dev/nrst0** or **ws1:/dev/nrst0** for a tape unit on remote host **ws1**. Since the recording density for UNIX tape units is defined by their name, the device name will overrule the density given on the command line! Some 1/2 inch drives require the density to be set manually on the drive as well. During reading the tape drive will itself sense the tape density used. Thus, the generic tape name can normally be given when reading.
5. Dismount your tape from the drive so that others can use it.
6. Logout from MIDAS with **bye** and from the workstation.

Concerning the usage of tape media please note the following:

1. Tape drives can be selected from any host in the same domain. To get access to drives in other domains you need to have another account with the same name in the remote domain.
2. Tape devices have to be accessed via theirs system names;
3. The temporary *register* accounts used by visitors belong to domain **eso**.
4. New tapes are from factory write enable.
5. Tapes can be obtained from the operator, room 220/1, 2nd floor.
6. The MIDAS error **Permision denied** occurs when the given tape is write protected (check remarks below) or when it is allocated to another user (use **deallocate** command)
7. Concerning 1/2 inch tapes:
 - For each 1/2 tape drive can be accessed in two densities: high (6250 bpi) and medium (1600 bpi). The mode that is used to access the data depends on the device name.
 - 1/2 inch magtape drives have to be set to enable remote density selection: with the drive OFF-LINE press the button DENSITY until LED "REMOTE DEN" comes on. The device name used with the command "INTAPE/MIDAS" will actually set the density.
 - 1/2 inch magtape drives can *read* tapes of any density if the drive is set with "REMOTE DEN" on. That is, for reading there is no difference between, *e.g.* **tape2**, **tape2h**, and **tape2m**.

- 1/2 inch tapes are write enable when they are mounted with the write enable ring and the LED "WRITE ENBL" comes on.

8. Concerning 8mm tapes:

- 8mm-High drives can also read 8mm-Low density tapes, but not vice versa.
- The 8mm tape is write enable when the red switch in front of the tape is in "REC" position.

9. Concerning QIC tapes:

- QIC-150 drives can only write on high density tapes DC600XTD or DC6150 however they can read QIC-24 low density tapes DC300XL or DC600A.
- QIC-24 drives use only low density tapes DC300XL or DC600A.
- QIC tapes are write protected when the round switch in the tape points to the "SAVE" position; write enable otherwise.

10. Concerning DAT tapes:

- DDS drives use Sony 60/90/120 min.
- DAT tapes are write enable when the switch in front of the tape closes the hole.
- An up to date list of available tape drives is posted on the bulletin board in the user room.

C.2 Operating Systems

The main operating system used for image processing in ESO is UNIX. On the SUN workstation the X11 OSF/Motif window system is used to provide virtual terminals or displays. For those who are not familiar with UNIX systems this section tries to provide a very basic introduction. For more information, the interested user is referred to the various UNIX or SUN publications. A few copies of these manuals and other documentation are available through system management (room 220/2).

C.2.1 Login Procedures

The X11 window system will display a login window which has user identification and password files. If the screen is black you may have to move the mouse or press a key on the keyboard to exit the screen save mode. First enter your `userid` and then your password each terminated with the `return` key. This will start an automatic login procedure (see OSF/Motif manual for customising the window manager). In case you don't have an account on the systems you should contact the System Manager (room 220/2).

C.3 Data Format Compatibility

The internal binary data formats of the VAX and SUN systems are different which makes it impossible to share data files (*e.g.* image or table). The SUN systems store data starting with the most significant bits (*i.e.* big endian) and use IEEE floating point format while VAX's are byte swapped and have a proprietary floating point format.

In order to exchange data files between these systems it is necessary to use a machine independent format *e.g.* FITS. For this reason and because internal MIDAS data structures may change, it is **strongly** recommended to save data in FITS format.

Appendix D

Release Notes

D.1 Current Status

This appendix contains on the following pages the Release Notes for the **95NOV** release of MIDAS. A listing of the MIDAS NEWS-file which gives an overview of the modifications and improvements of the system for the present release has been added.

D.2 Installation

The **95NOV** release of ESO-MIDAS can run on most UNIX systems, and is most likely to run under DEC/VMS 4.7 and DEC/OPEN-VMS 6.0 (or higher). If you are going to install ESO-MIDAS on your system you have retrieved a copy of this release through ftp. However, as have been announced, as of 1995 new applications are not tested automatically on VMS/Open VMS systems.

The monitor and low level interfaces are coded in C which means the portable MIDAS requires a C compiler for installation. Many applications are written in FORTRAN-77 and can therefore only be installed if either a FORTRAN-77 compiler or a FORTRAN-to-C conversion package (*e.g.* **f2c** for UNIX) is available. The installation of MIDAS has been certified with the public domain GNU ANSI-C compiler which can be obtained from the Free Software Foundation, and verified with the quality control tool Purify from Pure-Software.

The instructions for installation are given in separate documents for either VMS or UNIX systems which are included in the distribution kit. Read the appropriate one carefully and proceed as described (the procedure may have been modified compared with previous installations!). Basic knowledge of your local operating system (*e.g.* VMS or UNIX) is assumed and required.

One area may require special attention during the first installation. It relates to the image display interfaces that, for MIDAS, must conform to the standard IDI interfaces. IDI's are provided only for the X-Window system, Version 11. For other display devices a set of IDI routines has to be written.

For a few commands the NAG mathematical library is required. If you do not have

this library MIDAS can still be installed, however, some commands may not be available. In future versions we will try to reduce our dependency on NAG. A list of programs using NAG and the routines are given in Section D.5.

Starting with the 89NOV release a separate installation of the lower level AGL library is no longer required. The library, i.e. those parts needed for the MIDAS plot facilities, is fully integrated within the MIDAS directory structure and will be generated as every other MIDAS subroutine library. For a full installation of the AGL library refer to the Astronet Documentation Facility, Trieste.

D.3 Software Modifications

The MIDAS release **95NOV** is the 10th official release of portable MIDAS. A number of new applications have been added to this version. Given below is a overview of the changes and improvements compared with the previous release:

1. The Problem Report form of the XHelp GUI has been simplified to make it easier to send us e-mail. After one year of working with the GNATS problem report database, new PR categories have been created, in particular for GUIs and for each context.
2. The Echelle package has been refurbished to allow calibration in world coordinates. The 94NOV version of the package is still available under the package name echellec. Documentation for the EMMI calibration is provided with the command INITIALIZE/EMMI. A MOTIF GUI for the new version of the Echelle package is provided and can be activated with `create/gui echelle`.
3. Several new commands have been included in the Spec package related to spectral analysis:
 - **CONTINUUM/SPEC** is a new command for the determination of spectra continuum using smoothing spline interpolation. This command comes in complement of **NORMALIZE/SPEC**, based on standard polynomials.
 - **DEBLEND/LINE** allows to fit up to 6 multiple-gaussian components in spectral lines (contributed by Goettingen University)
 - **REFRACTION/LONG** provides the correction for differential atmospheric dispersion (contributed by P.W.A. Goerdt, Goettingen University).
 - **COMPUTE/PARALLACTIC** allows to compute the parallactic angle (contributed by A. Smette, Groningen Observatory).
4. The MOS context has been developed by the Observatory of Heidelberg in the context of the FORS data reduction software development. The context includes a tutorial (**TUTORIAL/MOS**).
5. Different search paths for Midas procedures, data and executables are now possible.
6. A first set of 3-dim functions has been implemented.

7. An upgrade of the CCD context (now called CCDRED) that now also includes tools for mosaicing ccds.
8. A separate context CCDTEST containing an updated version of the CCD testing commands.
9. A new context IRAC2 containing on- and off-line commands to support processing of the IRAC2 data.

For a more complete list of all updates/additions use the MIDAS command HELP [NEWS] after having installed this release. A hardcopy of it can be obtained via PRINT/HELP [NEWS].

D.4 Manual Updates

Several sections of the MIDAS User's Manual have been updated in this release. A list of the parts to be replaced is given below :

Volume A, Titlepage		to be replaced
Volume A, Chapter 1	Introduction	to be replaced
Volume A, Chapter 3	Monitor and Syntax	to be replaced
Volume A, Appendix A	Command Summary	to be replaced
Volume A, Appendix B	Acknowledgements	to be replaced
Volume A, Appendix C	Site Specific Implementation	to be replaced
Volume A, Appendix D	Release Notes	to be replaced
Volume B, Titlepage		to be replaced
Volume B, Chapter 1	Introduction	to be replaced
Volume B, Chapter 3	CCD Reductions	to be replaced
Volume B, Appendix A	Command Summary	to be replaced
Volume B, Appendix D	Echelle Reduction	to be replaced
Volume B, Appendix J	IRAC2	to be added
Volume B, Appendix K	CCD Test Procedures	to be added
Volume C, Complete Volume	Detailed Command Summary	to be replaced

D.5 Use of NAG Library

The NAG mathematical library is still used in a few MIDAS commands. A list of these programs and routines are given below:

Program	Package	NAG Routines
fitimag	Fit	e04fdf, e04fcf, e04gcf, e04hev, e04gbf, e04gef, e04gdf, e04ycf, e04jaf, e04hbf, e04jbf, e04kaf, e04hcf, e04kbf, e04kcf, e04kdf
genran	General	g05cbf, g05ddf, g05daf, g05dbf, g05def, g05edf, g05eyf, g05ecf, g05dff
echripp1	Echelle	e04gdf

A set of dummy routines are provided for sites that do not have a NAG library. This implementation enables sites to use the FIT package even without the NAG library (in this case, only the Newton-Raphson method is supported).

Appendix D

Release Notes

D.1 Current Status

This appendix contains on the following pages the Release Notes for the **94NOV** release of MIDAS. A listing of the MIDAS NEWS-file which gives an overview of the modifications and improvements of the system for the present release has been added.

D.2 Installation

The **94NOV** release of ESO-MIDAS can run on most UNIX systems, as well as under DEC/VMS 4.7 and DEC/OPEN-VMS 6.0 (or higher). If you are going to install ESO-MIDAS on your system you have retrieved a copy of this release through ftp.

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The instructions for installation are given in separate documents for either VMS or UNIX systems which are included in the distribution kit. Read the appropriate one carefully and proceed as described (the procedure may have been modified compared with previous installations!). Basic knowledge of your local operating system (*e.g.* VMS or UNIX) is assumed and required.

One area may require special attention during the first installation. It relates to the image display interfaces that, for MIDAS, must conform to the standard IDI interfaces. IDI's are provided only for the X-Window system, Version 11. For other display devices a set of IDI routines has to be written.

For a few commands the NAG mathematical library is required. If you do not have this library MIDAS can still be installed, however, some commands may not be available. In future versions we will try to reduce our dependency on NAG. A list of programs using

NAG and the routines are given in Section D.5.

Starting with the 89NOV release a separate installation of the lower level AGL library is no longer required. The library, i.e. those parts needed for the MIDAS plot facilities, is fully integrated within the MIDAS directory structure and will be generated as every other MIDAS subroutine library. For a full installation of the AGL library refer to the Astronet Documentation Facility, Trieste.

D.3 Software Modifications

The MIDAS release 94NOV is the 9th official release of portable MIDAS. A number of new applications have been added to this version. Given below is a overview of the changes and improvements compared with the previous release:

1. Readline: a new line editor from GNU only on UNIX systems. It enhances the line editing capabilities of MIDAS, like a history stack of commands, Emacs or Vi editing functions, command and filename completion functions, and a communication channel to the MIDAS GUI Xhelp for a help-on-line function
2. INTAPE/FITS with data decompression on-the-fly only on UNIX systems. Compressed FITS data, by default those with extensions .z and .Z, are decompressed automatically on-the-fly (pipeline) when accessed.
3. The generic-tape driver. Based in the semi-standard MTIO system interface to tapes, this new driver handles most of the tape devices without any other particular configuration than the system device-name.
4. A new client/server model for access to remote tapes with machine-independent data structures.
5. A refurbishment of the graphics software that resulted in notable improvements in performance. Graphical representation of three dimensional tables is now also supported.
6. Graphical user interfaces for the Data Organizer context (DO) and for the context IRSPEC.
7. A upgrade of the CCD context which now also include test procedures for monitoring the quality of the detectors by a series of standard tests.
8. A new astrometry package called ASTROMET has been included.

For a more complete list of all updates/additions use the MIDAS command HELP [NEWS] after having installed this release. A hardcopy of it can be obtained via PRINT/HELP [NEWS].

D.4 Manual Updates

Several sections of the MIDAS User's Manual have been updated in this release. A list of the parts to be replaced is given below :

Volume A, Titlepage		to be replaced
Volume A, Chapter 1	Introduction	to be replaced
Volume A, Chapter 3	Monitor and Syntax	to be replaced
Volume A, Chapter 5	Table File System	to be replaced
Volume A, Chapter 6	Graphic and Image Display	to be replaced
Volume A, Appendix A	Command Summary	to be replaced
Volume A, Appendix B	Acknowledgements	to be replaced
Volume A, Appendix C	Site Specific Implementation	to be replaced
Volume A, Appendix D	Release Notes	to be replaced
Volume B, Titlepage		to be replaced
Volume B, Chapter 1	Introduction	to be replaced
Volume B, Chapter 3	CCD Reductions	to be replaced
Volume B, Chapter 7	Echelle Spectra	to be replaced
Volume B, Chapter 16	Astrometry	new
Volume B, Appendix A	Command Summary	to be replaced
Volume B, Appendix D	Echelle Reduction	to be replaced
Volume B, Appendix G	Long Slit and 1D Spectra	to be replaced
Volume C, Complete Volume	Detailed Command Summary	to be replaced

D.5 Use of NAG Library

The NAG mathematical library is still used in a few MIDAS commands. A list of these programs and routines are given below:

Program	Package	NAG Routines
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genran	General	g05cbf, g05ddf, g05daf, g05dbf, g05def, g05edf, g05eyf, g05ecf, g05dff
ecripp1	Echelle	e04gdf

A set of dummy routines are provided for sites that do not have a NAG library. This implementation enables sites to use the FIT package even without the NAG library (in this case, only the Newton-Raphson method is supported).