Ice in the Solar System

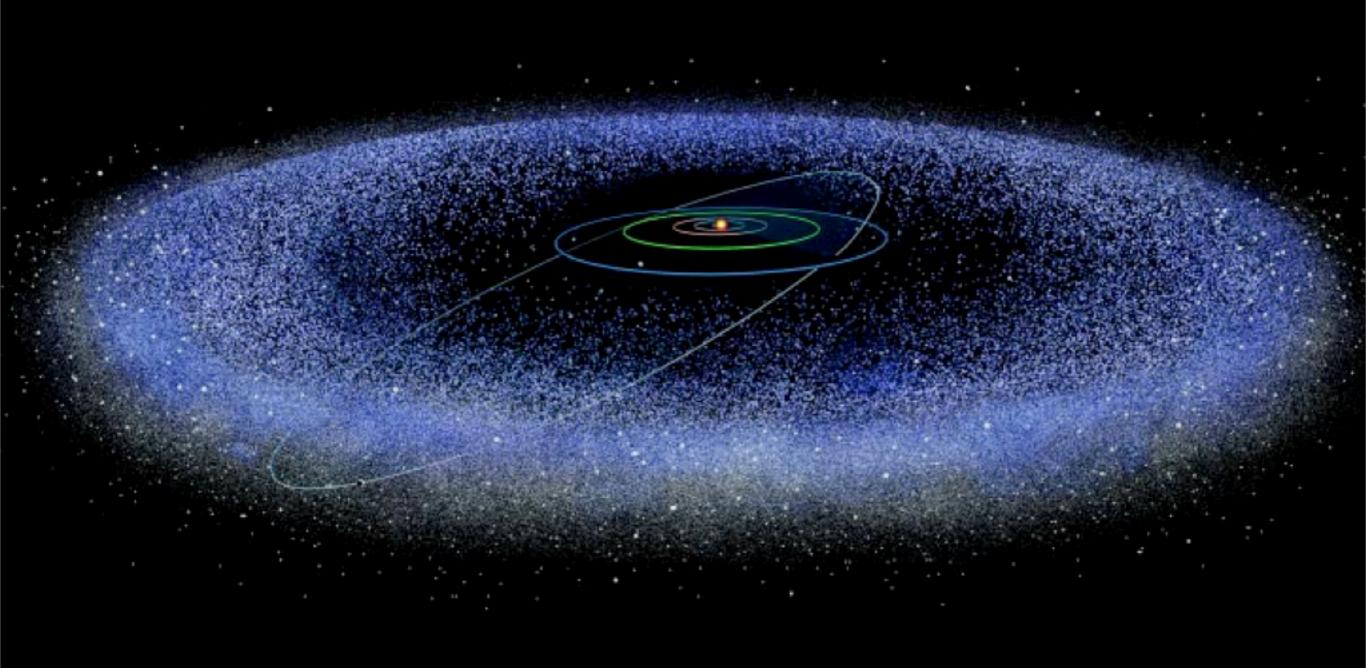
David Jewitt (UCLA)

Objective: To give a broad overview of ice and its importance.

Outline of the Talk:

- I. Introduction & Motivation
- 2. Surface Ice (ice beyond snow-line)
- 3. Interior Ice

Kuiper Belt - our local icebox



few x100 AU diameter

Isothermal Blackbody Temperature

$$T_{BB} = \frac{278}{R^{1/2}}$$

Ice too cold to sublimate R > 6 AU

Outside the Snow Line

Ice sublimates negligibly R > 6 AU

Surface ice, detected through vibrational transitions (fundamental @ 3 µm, overtones & combinations at 2 µm and 1.5 µm) is widespread at larger distances

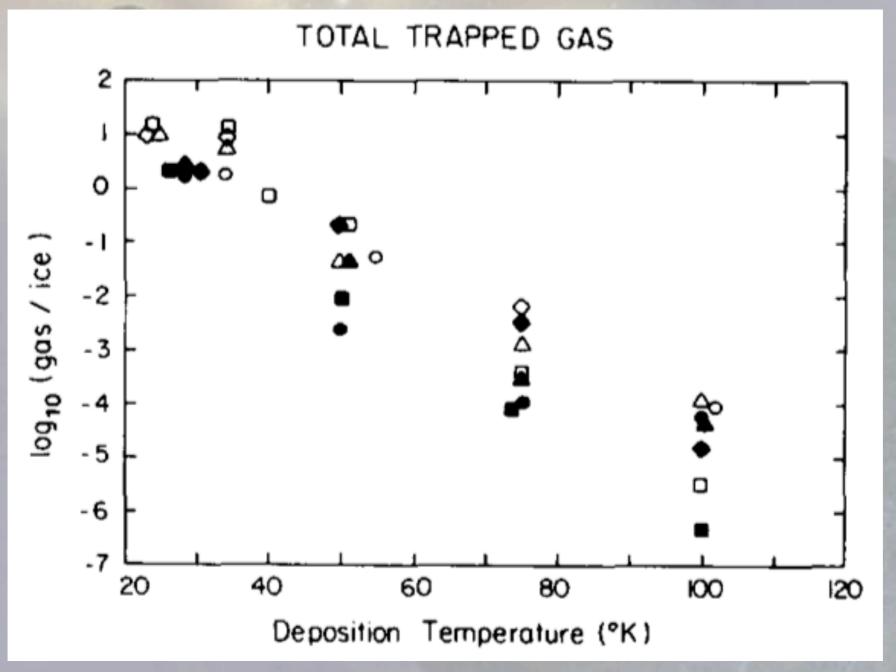
Crystallization Time

$$\tau_{cr} = 3.0 \times 10^{-21} \exp\left[\frac{E_A}{kT}\right]$$

$$\Delta E = 9 \times 10^4 \quad Jkg^{-1}$$

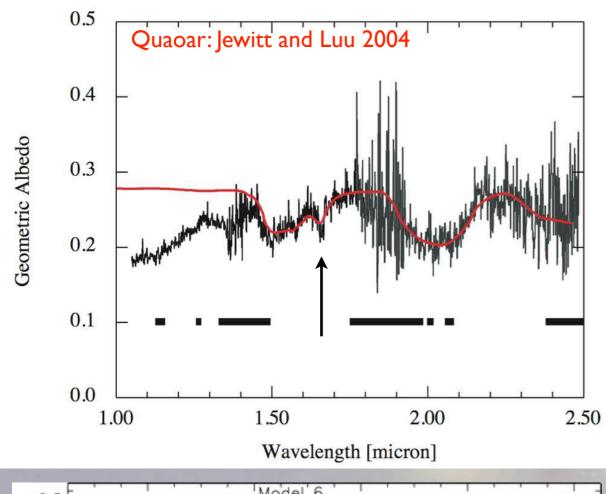
What is the state of the ice?

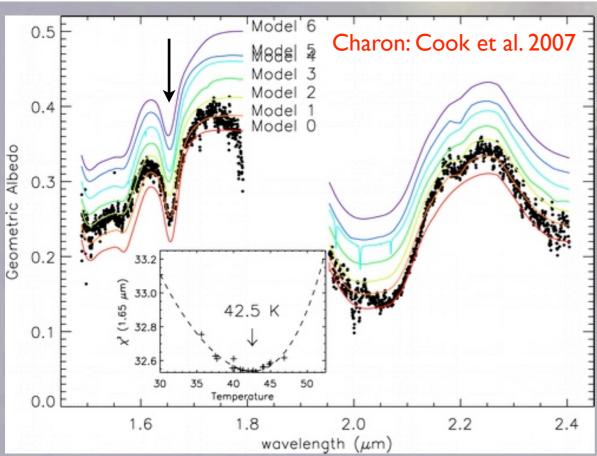
Gas Trapping

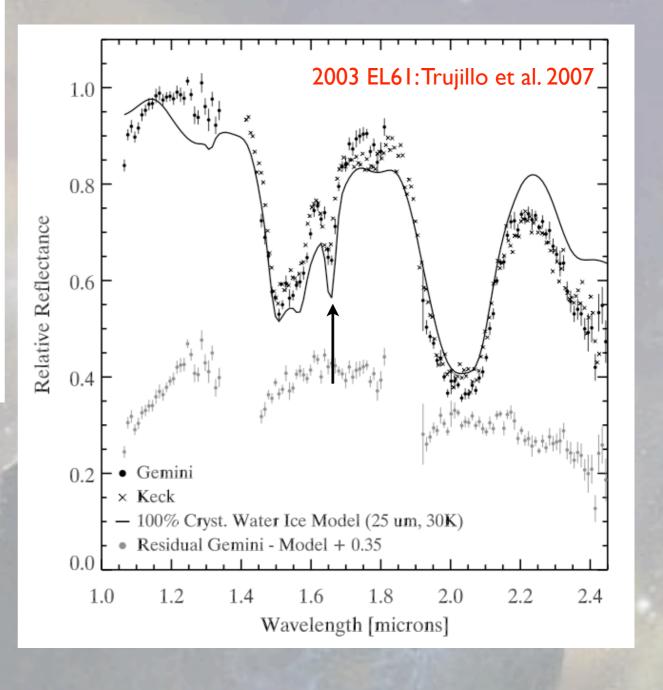


Bar-Nun & Kleinfeld 1989

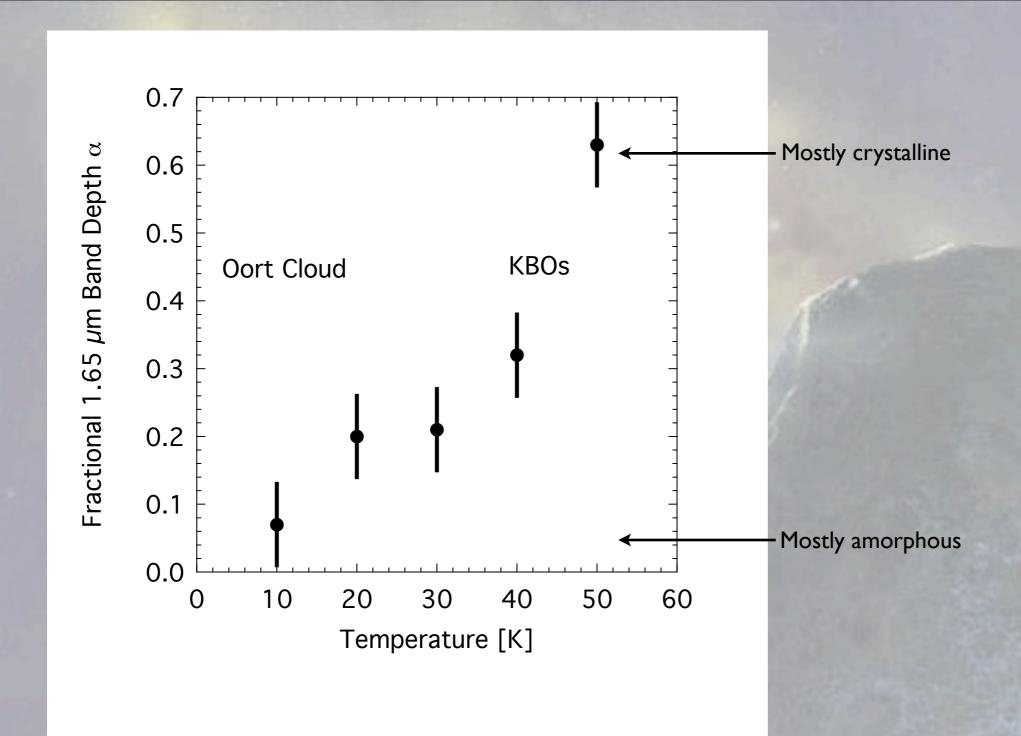
Crystallization releases trapped gas and energy: result = explosive





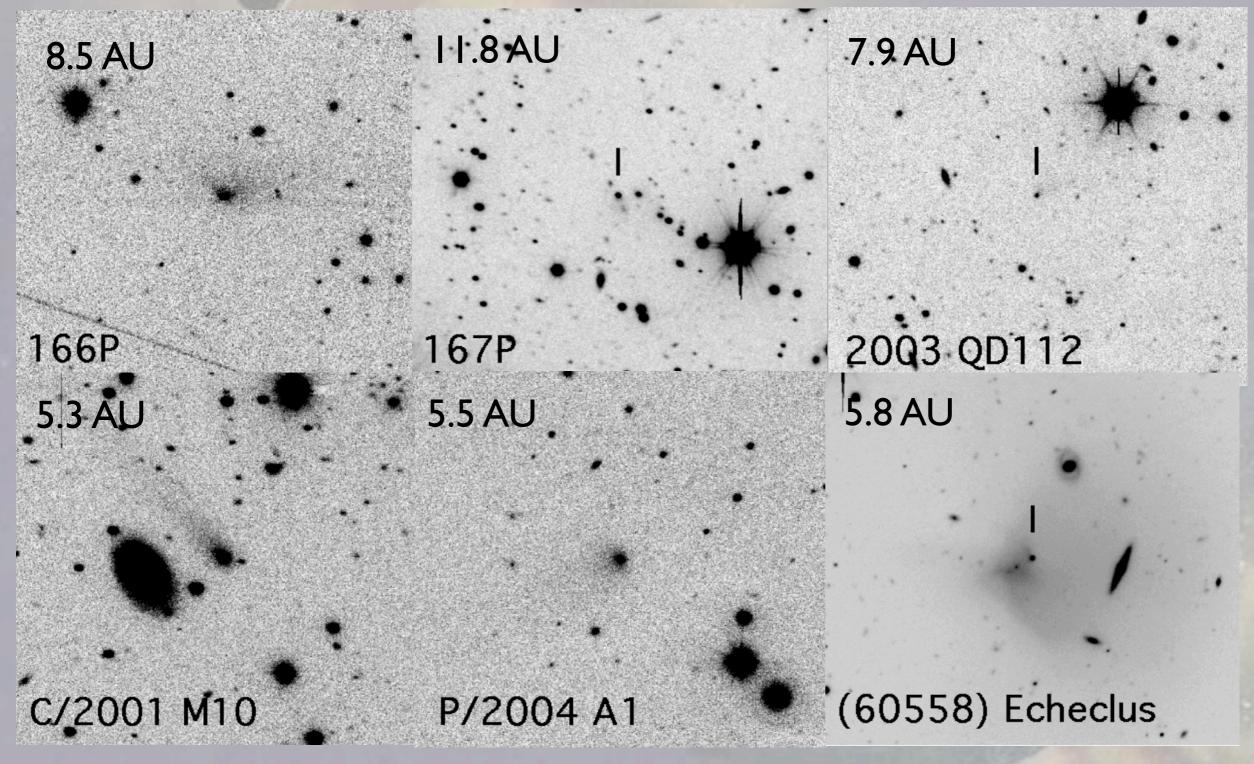


Zheng, Jewitt and Kaiser (2008)

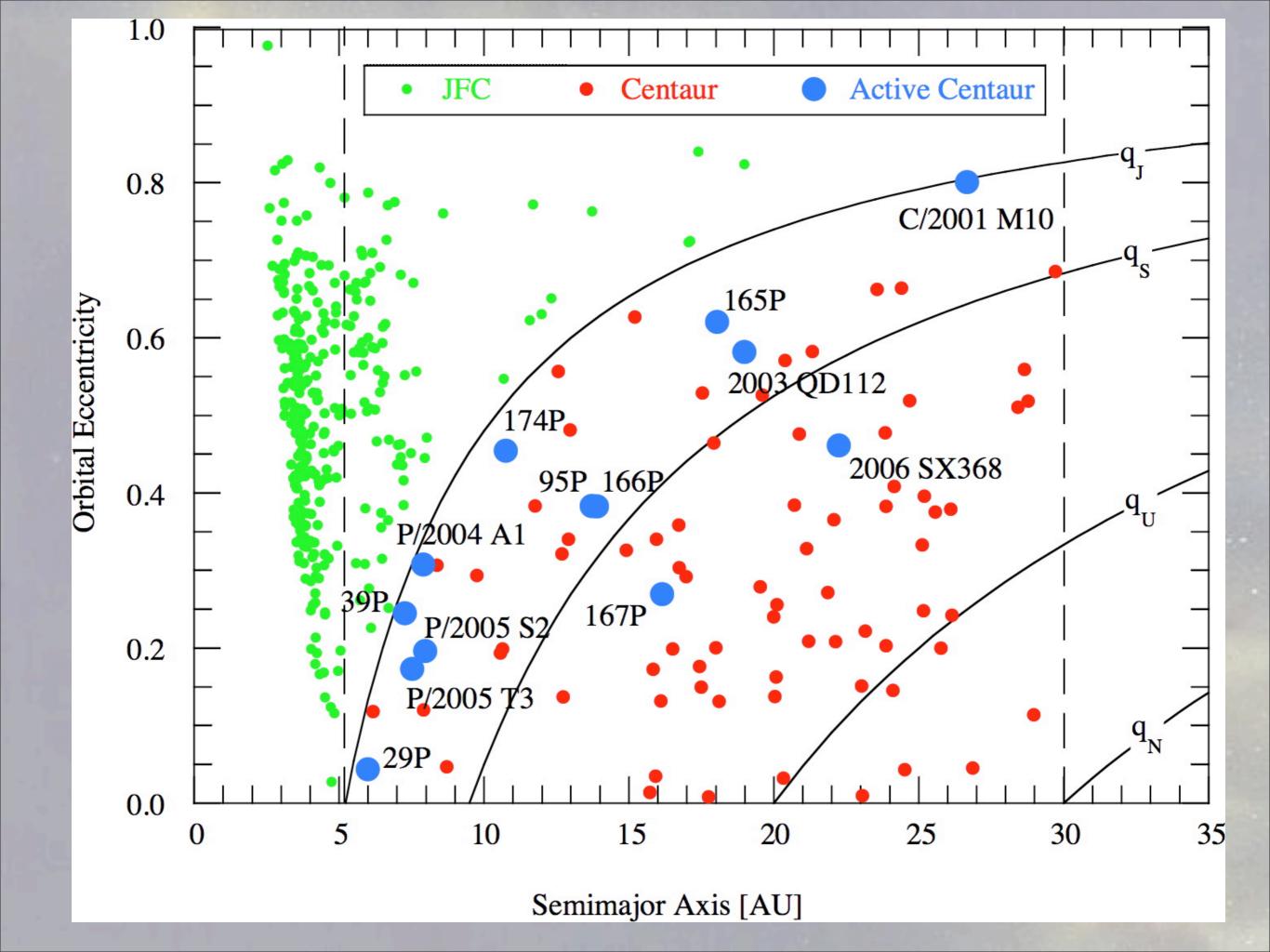


Zheng et al. 2008 (5keV electrons) c.f. Strazzulla et al 1992 (3kev He+ ions)

Evidence from the Active Centaurs



Sample Active Centaurs - Keck and UH 88



$$\tau_{cr} = 3.0 \times 10^{-21} \exp\left[\frac{E_A}{kT}\right]$$

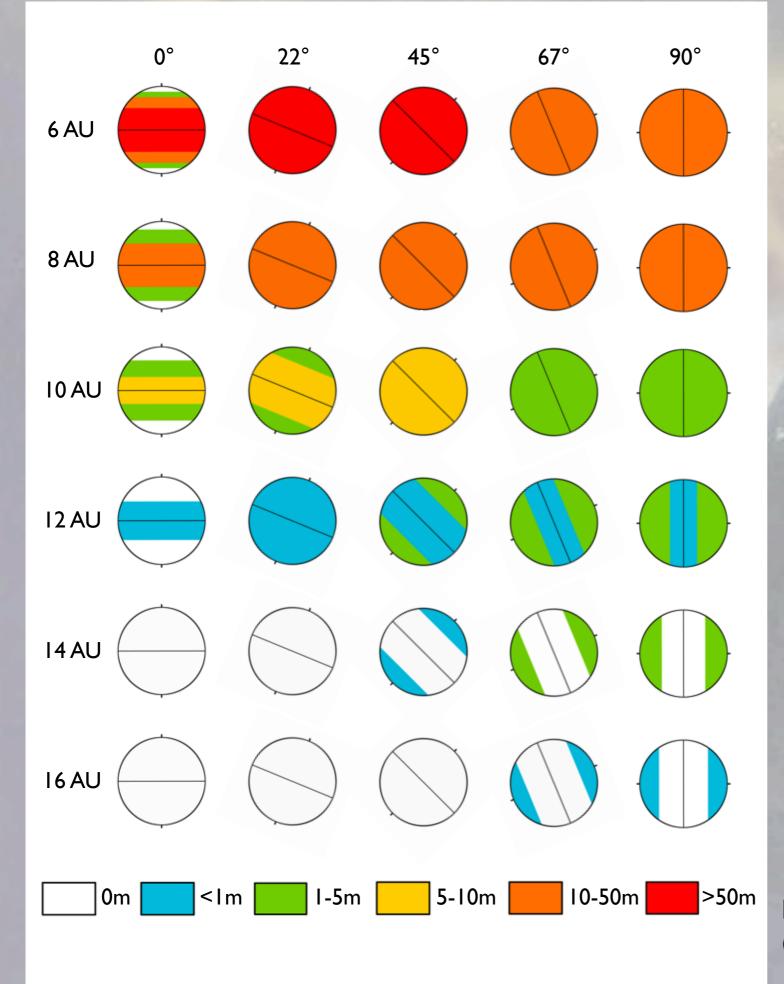
Setting condition for crystallization $T_{cr} \leq T_k$

gives....
$$\frac{3}{2} \ln R_{au} = -47.26 + \frac{E_A \sqrt{R_{au}}}{kT_1}$$

Solutions for coldest (hottest) isothermal (subsolar) equilibria are

$$7 \leq R_{au} \leq 14 \,\mathrm{AU}$$

- this is about what we observe



Centaur activity is observed only on objects where crystallization time ≤ orbit period

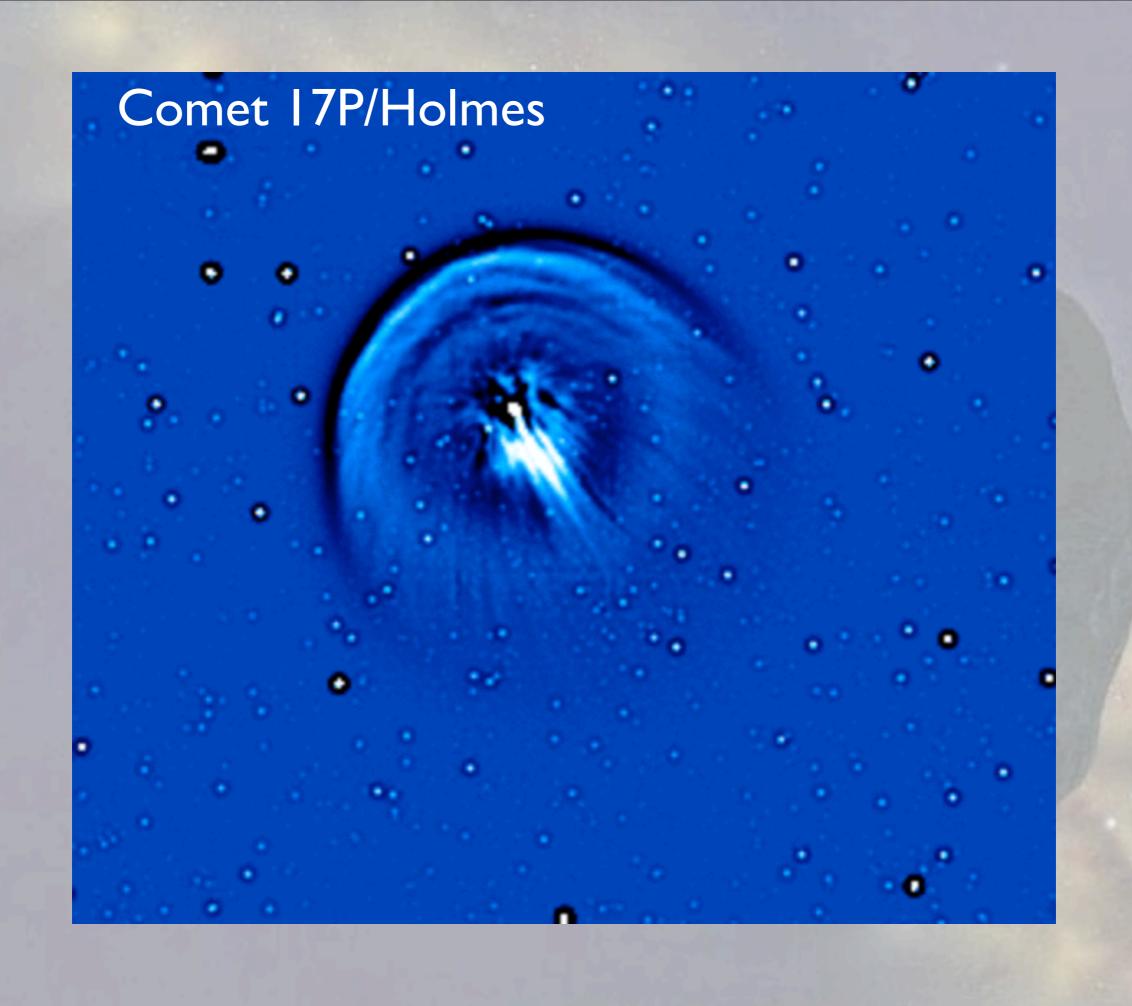
consistent with amorphous -> crystalline transition as activity driver.

Detailed thermal model by Aurelie Guilbert (UCLA)

Exothermic phase transition: $\Delta E = 9 \times 10^4 \text{ J/kg}$



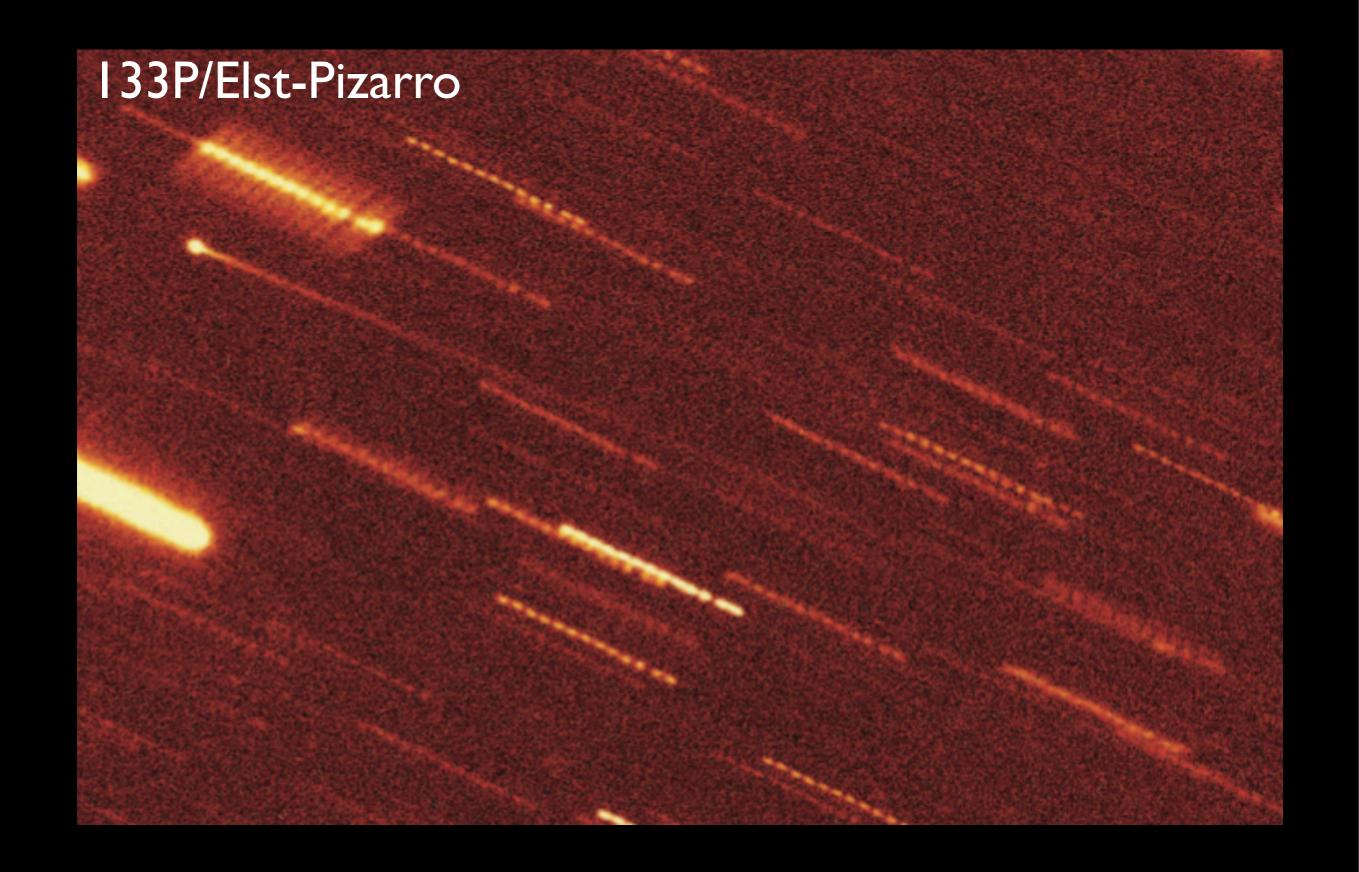
Crystallization might trigger a runaway

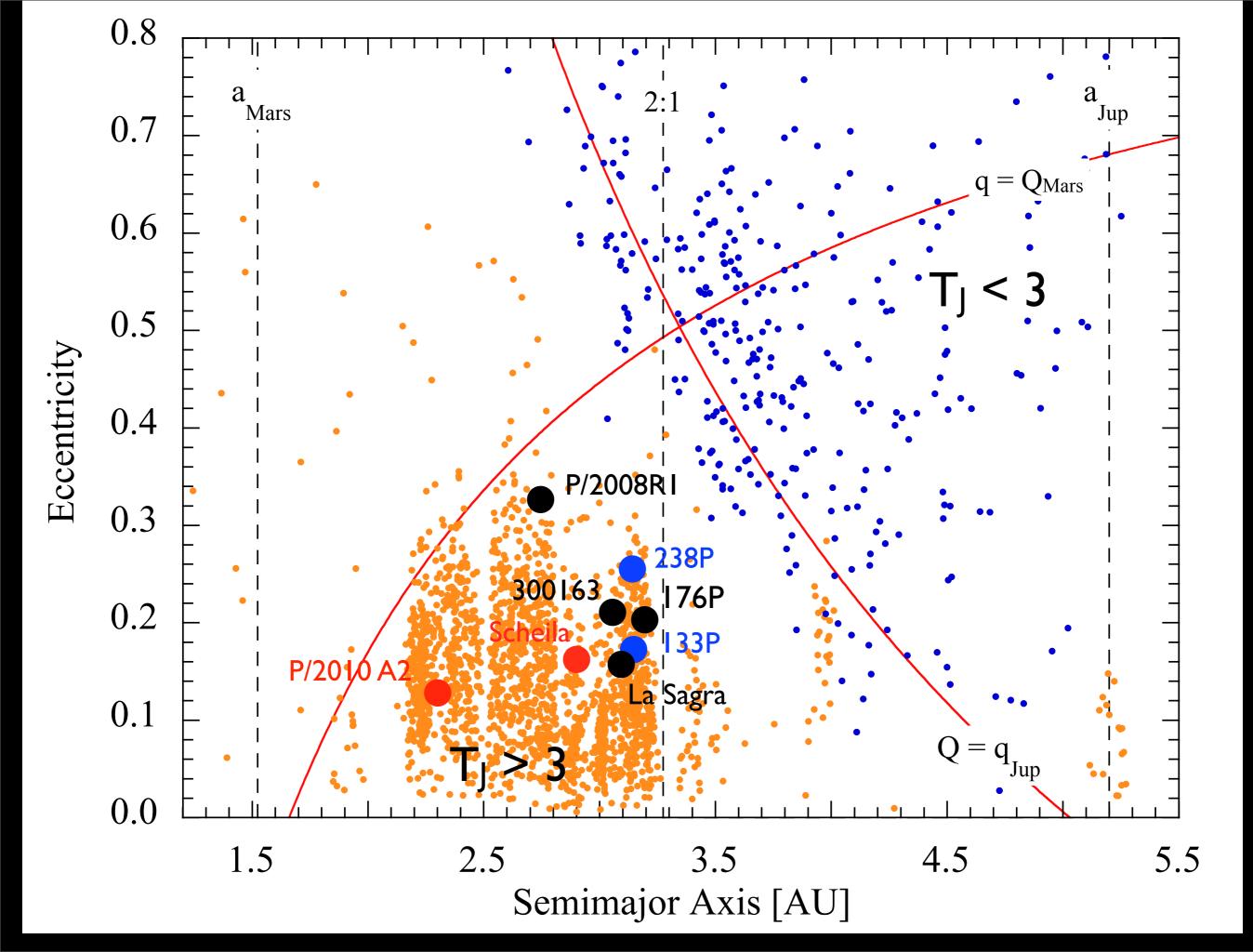


Mini-Summary

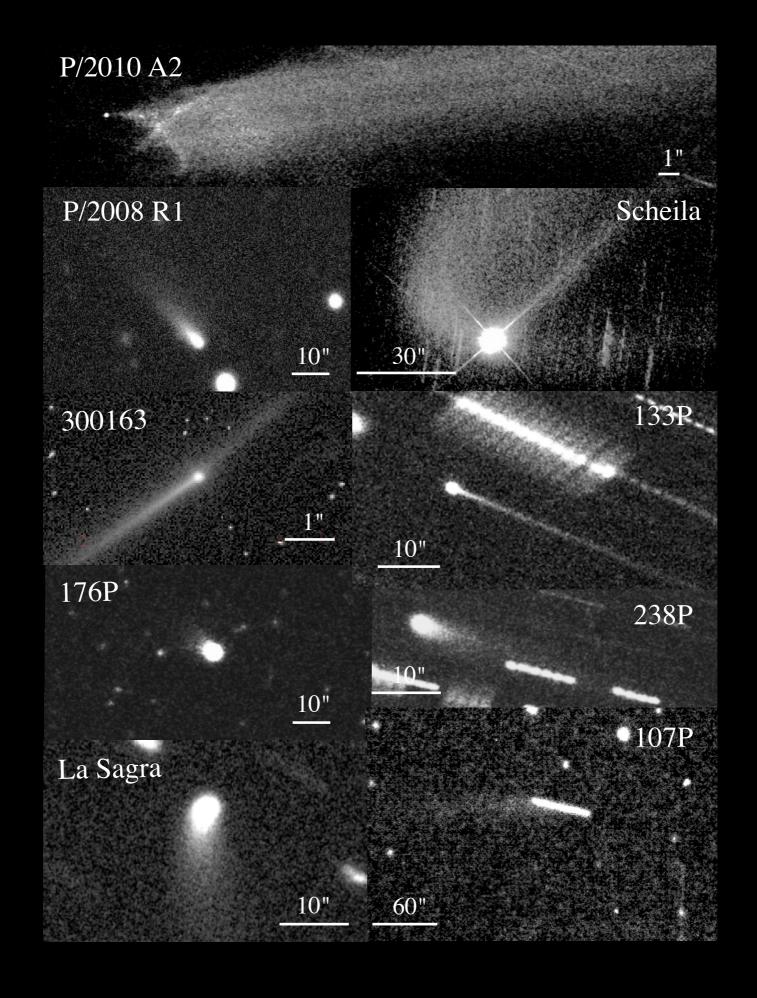
- Ice in the Kuiper Belt (& Oort Cloud) is cold enough to be amorphous. Surface ice, though, is crystalline. Indirect evidence suggests that Kuiper belt objects are *internally* amorphous.
- Crystallinity affects volatile gas carrying ability of the ice.

Inside the Snow Line





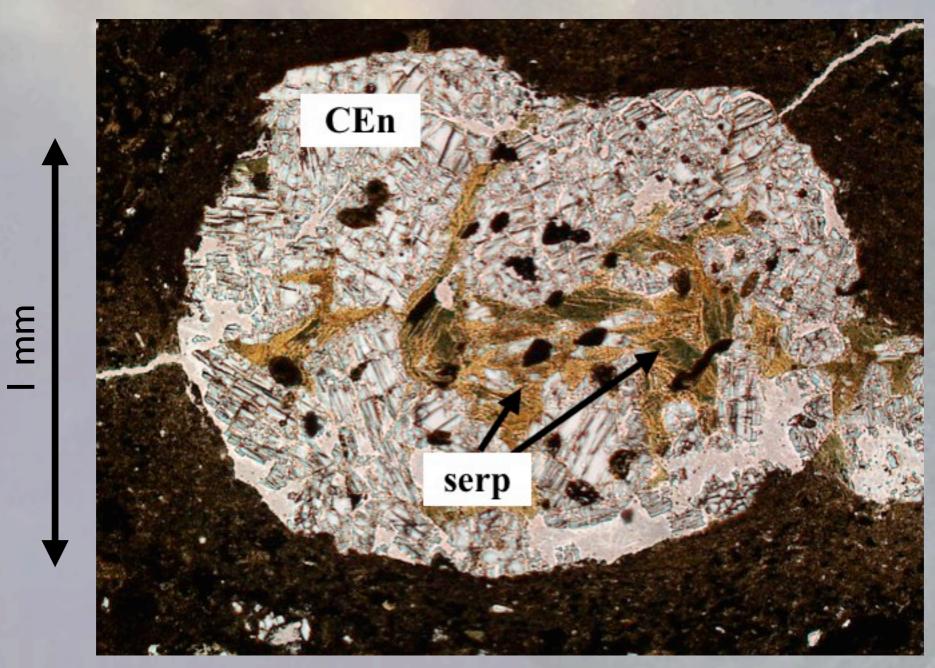
Main-belt Comets



Conduction cooling Time

$$au_c \sim rac{a^2}{\kappa}$$

Hydrated Minerals in Murchison Chondrule



Murchison CM2 Chondrite

Brearley, via Chizmadia

Phyllosilicates, clays, carbonates, serpentine, salt...all found

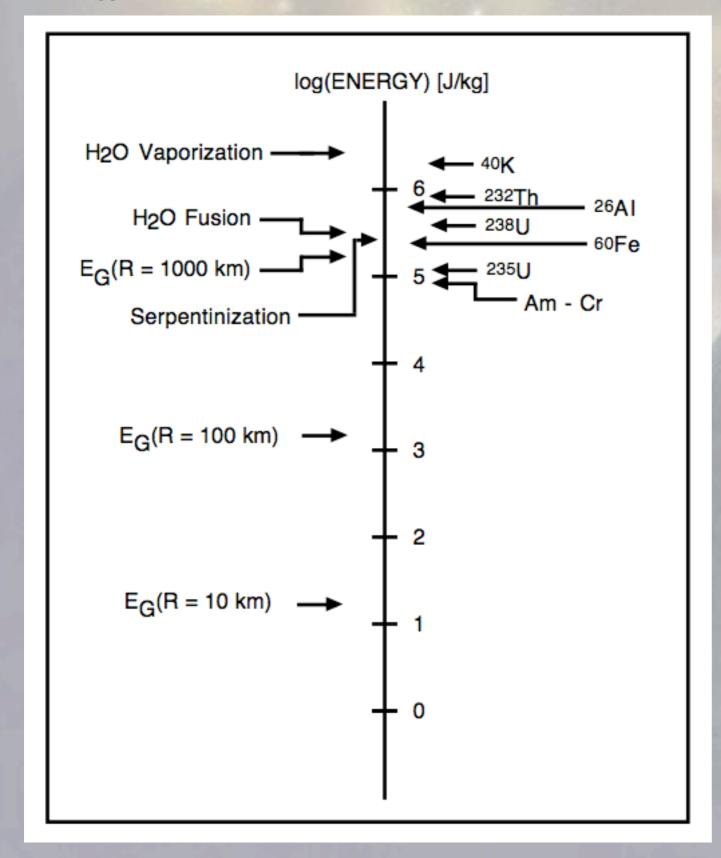
Serpentinization (hydration)

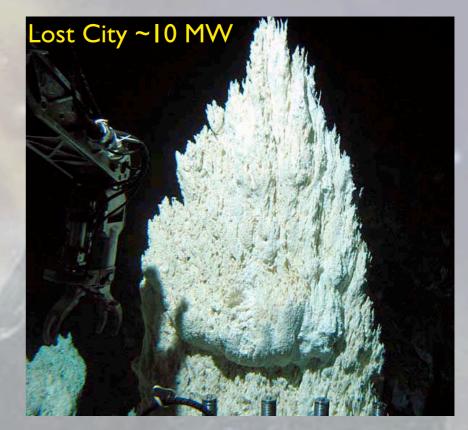
$$Mg_2SiO_4 + MgSiO_3 + 2H_2O \leftrightarrows Mg_3Si_2O_5(OH)_4$$

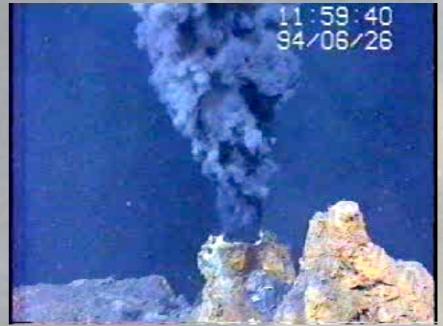
Exothermic $\Delta E = 3x10^5 \text{ J/kg}$

Also releases H₂, reduces C-compounds, making CH₄: the start of organic chemistry

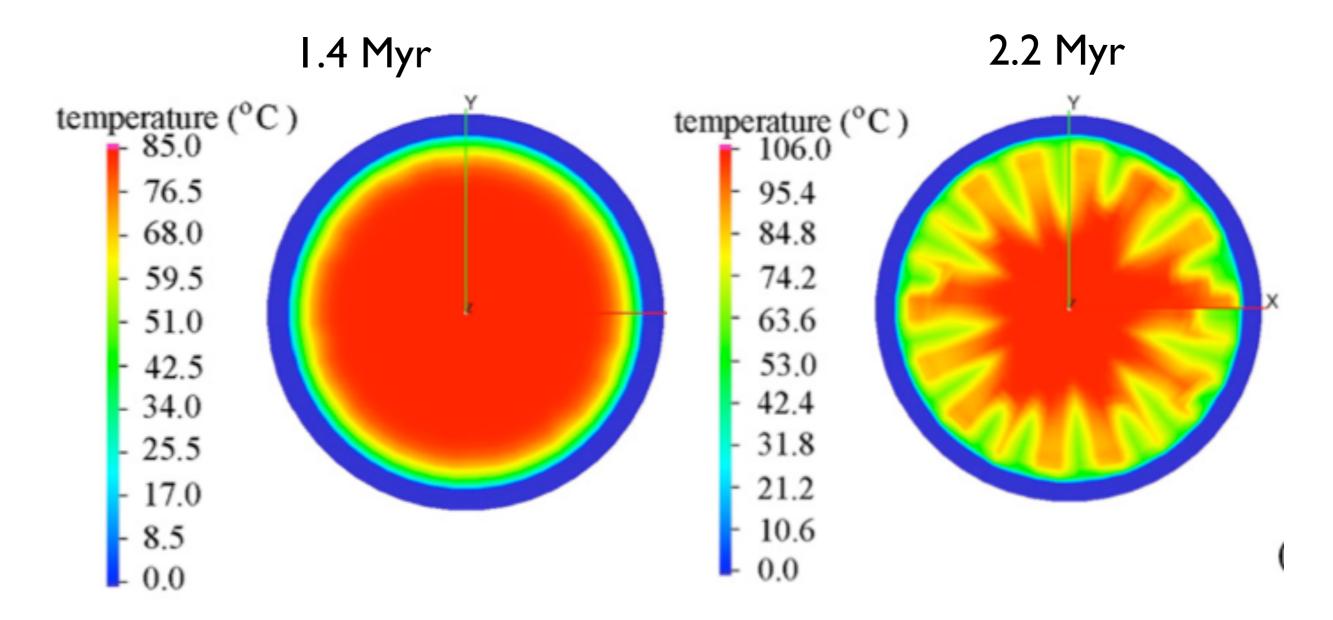
Energy Production







 $a = 50 \text{ km}, {}^{26}\text{Al}(t=0) = 5e-8$

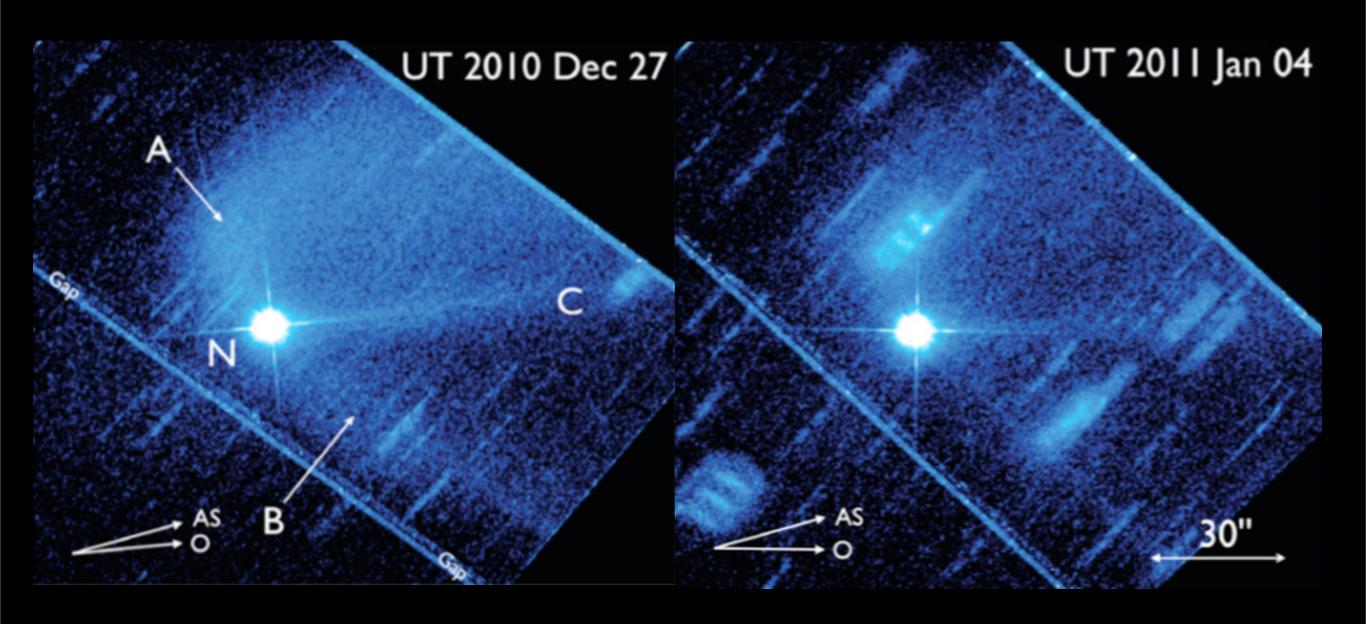


Palguta et al. (2010)

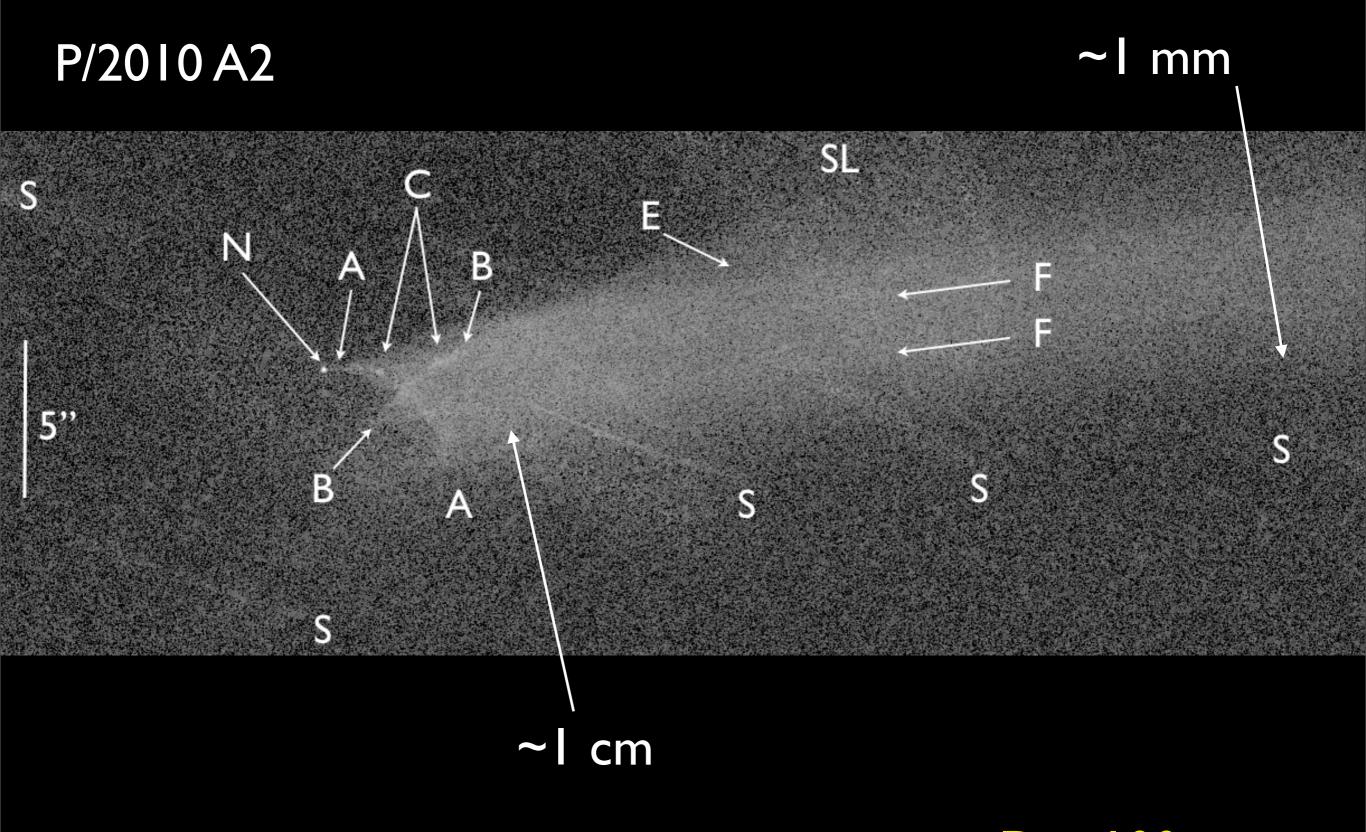
- Convection plumes create non-uniform thermal and mineralogical structure
- Complex hydrothermal systems drive chemistry
- Carbonaceous chondrites contain organics up to amino acids.....

Bonus Discovery

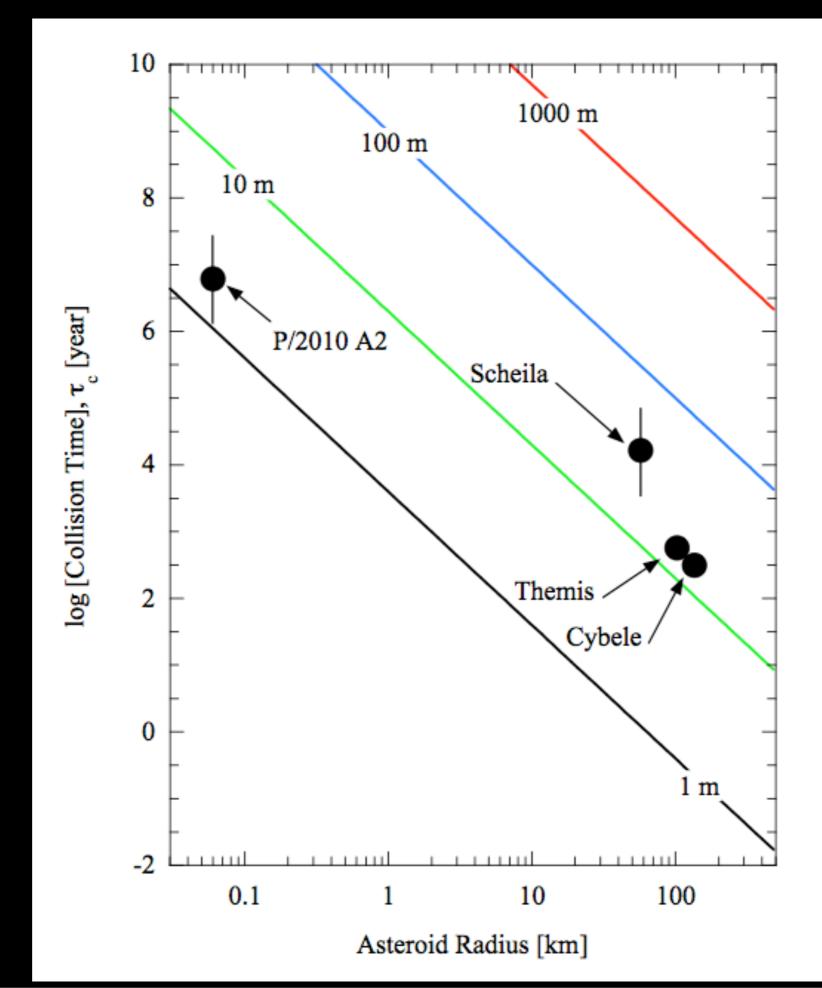
(596) Scheila



D = 113 km



D ~ 100 m



Summary

- 1). Planets and ice are both needed for life as we know it
- 2). Ice in outer solar system should be amorphous. Spectral evidence is lacking. Physical (outgassing) evidence is suggestive. Kuiper belt is probably amorphous.
- 3) Ice in asteroids is unstable unless protected from sublimation. Evidence for ice in asteroids is strong.
- 4) Water on Earth probably derived at least in part from wet asteroids. Biogenic precursors likewise.
- 5) Recently discovered impacts in belt analogs of debris disk production