

ALMA Cycle-3 Capabilities

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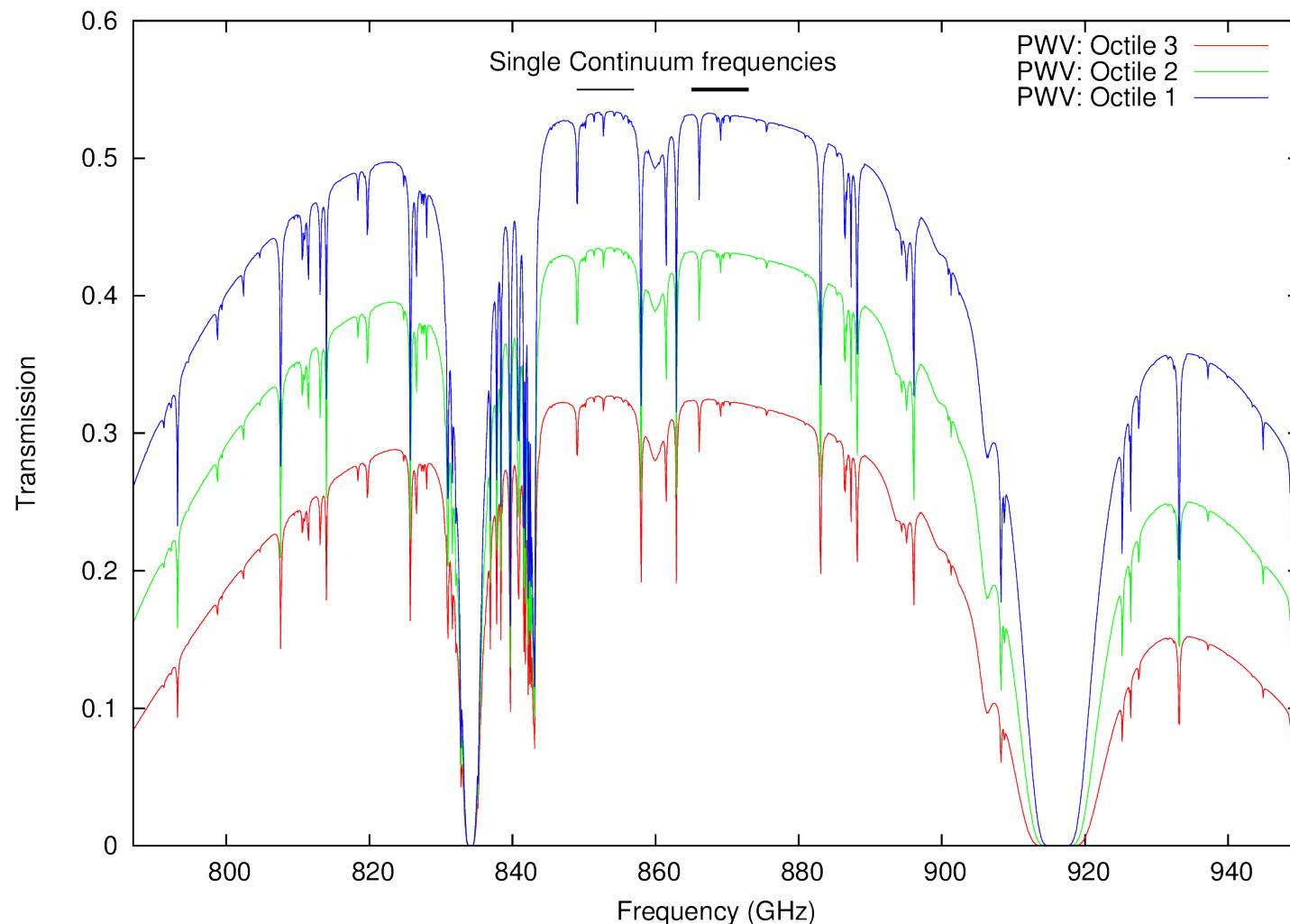
New ALMA capabilities for Cycle 3

- Band 10
- More antennas
- Baselines up to 10 km
- Expanded full-polarization tuning range
- ~~High spectral resolution (FDM) full polarization~~
- ~~TP at Bands 9 and 10~~

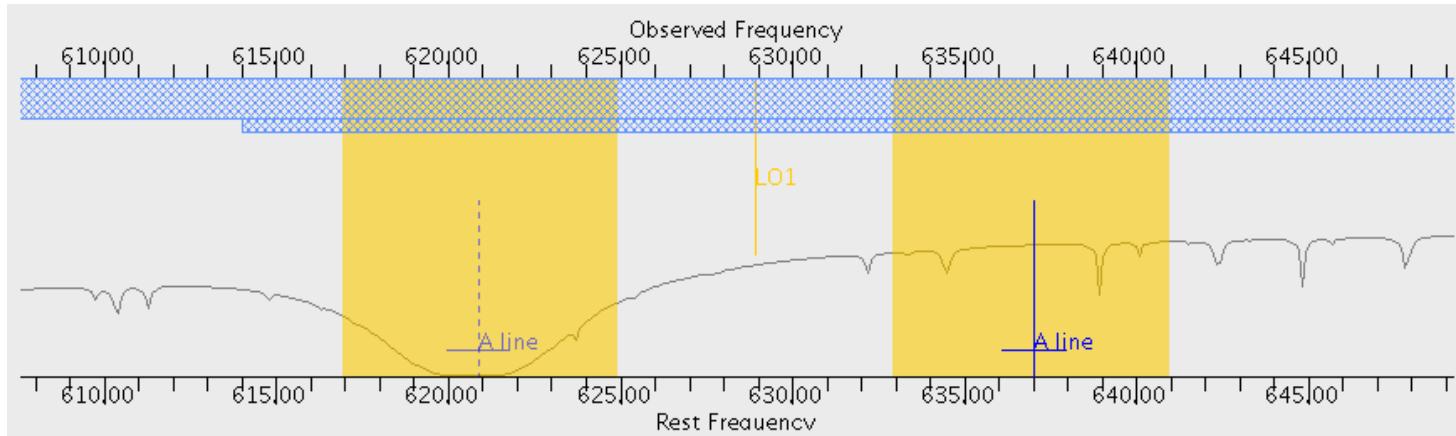
Band 10

- Frequency coverage
 - 787-950 GHz (381-316 μ m)
- Angular resolution
 - 400-30 mas
- Receiver temperature
 - OT uses specification of 230 K (DSB) over 80% of band
- 1st octile of PWV across whole band
 - 0.452 mm
- No mosaics
 - Primary beam is not well characterised

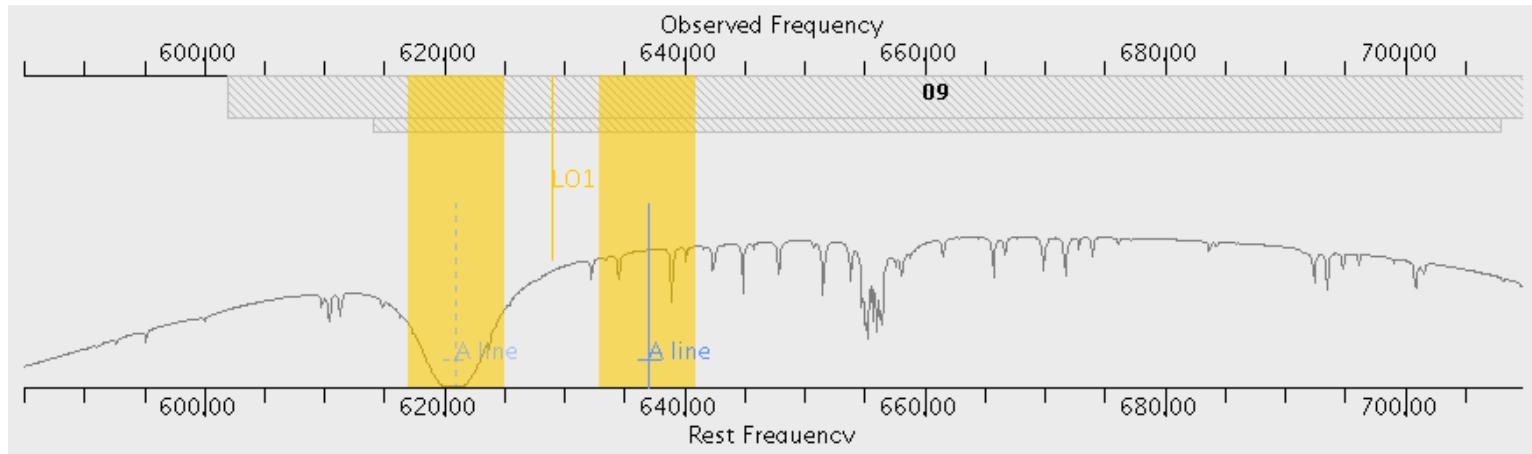
Band-10 Transmission



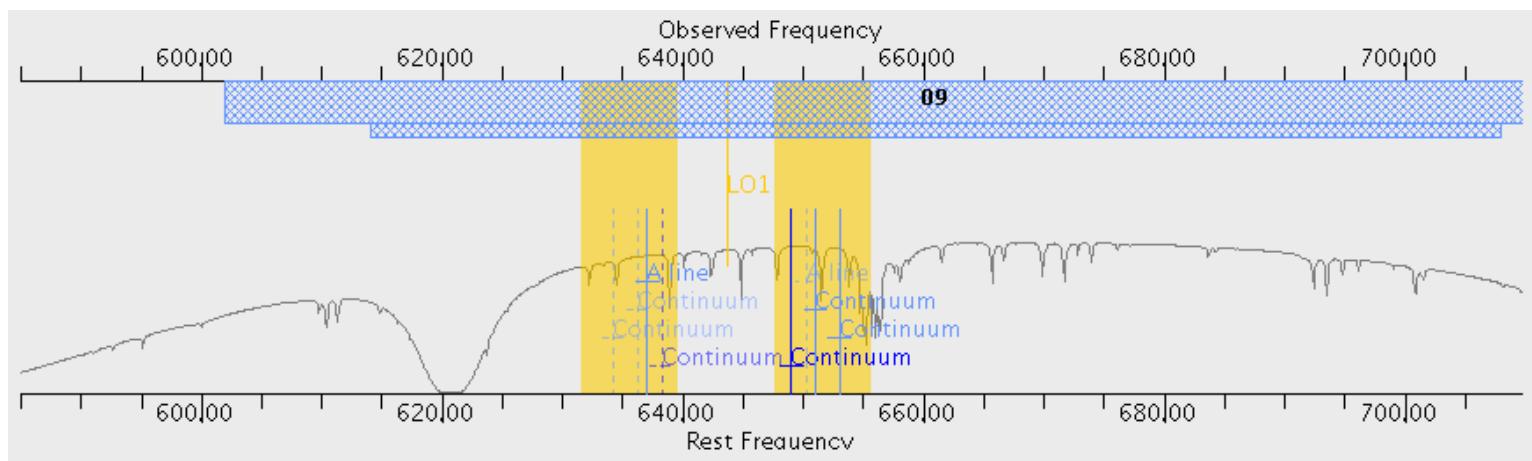
DSB-receiver time estimates



- Noise from “image” sideband contributes to “signal” spw
 - Only an issue for DSB receivers (Bands 9 and 10)
 - (Astronomical signal **is** suppressed)
- Noise is higher where atmospheric opacity is high
 - Atmospheric absorption lines should be avoided
- OT time estimates now include image noise explicitly
 - Visual editor shows location of image spws



Single spw – defaults to Upper Sideband (6.05 min on source)



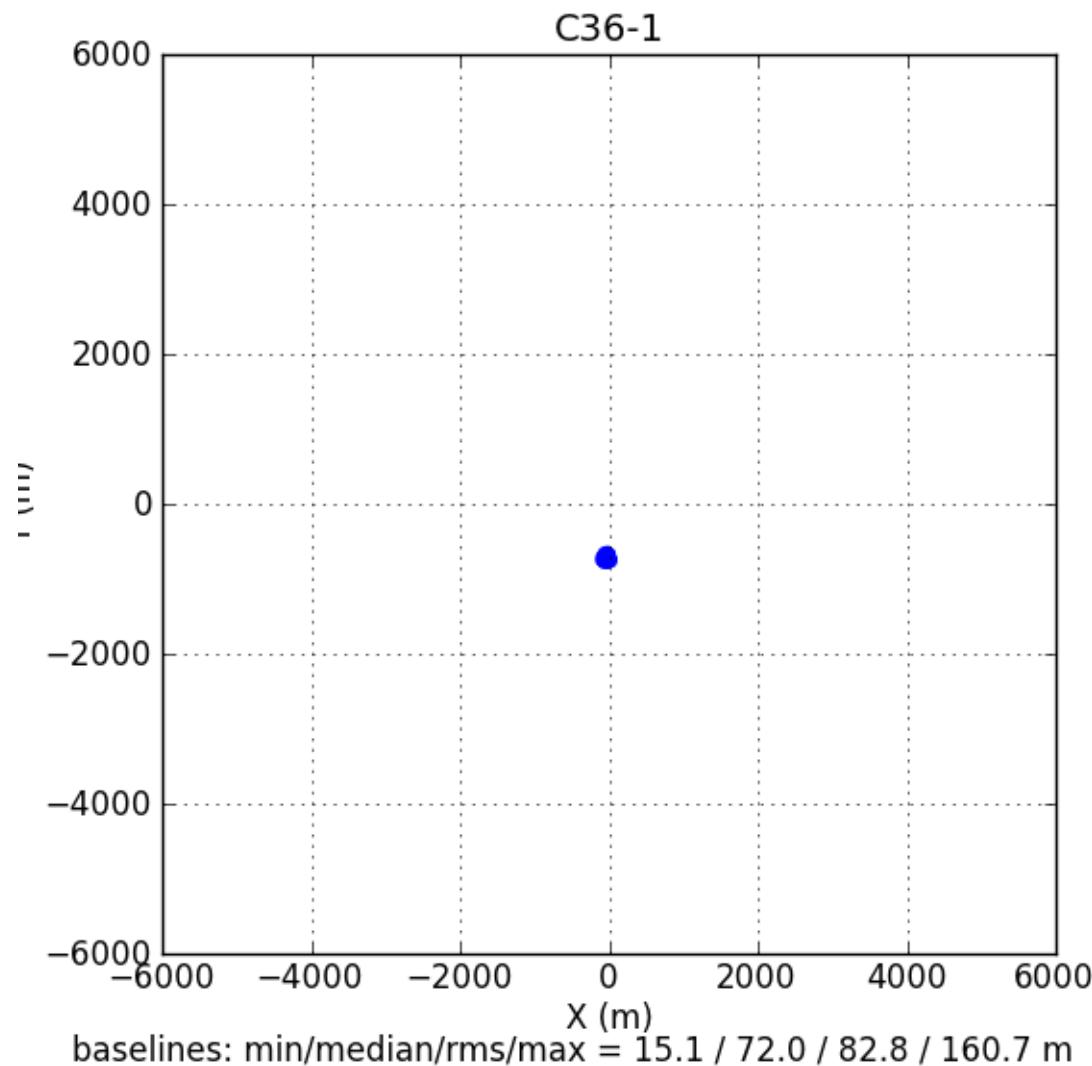
Place continuum spws such that line is forced into LSB (4.03 min on source)

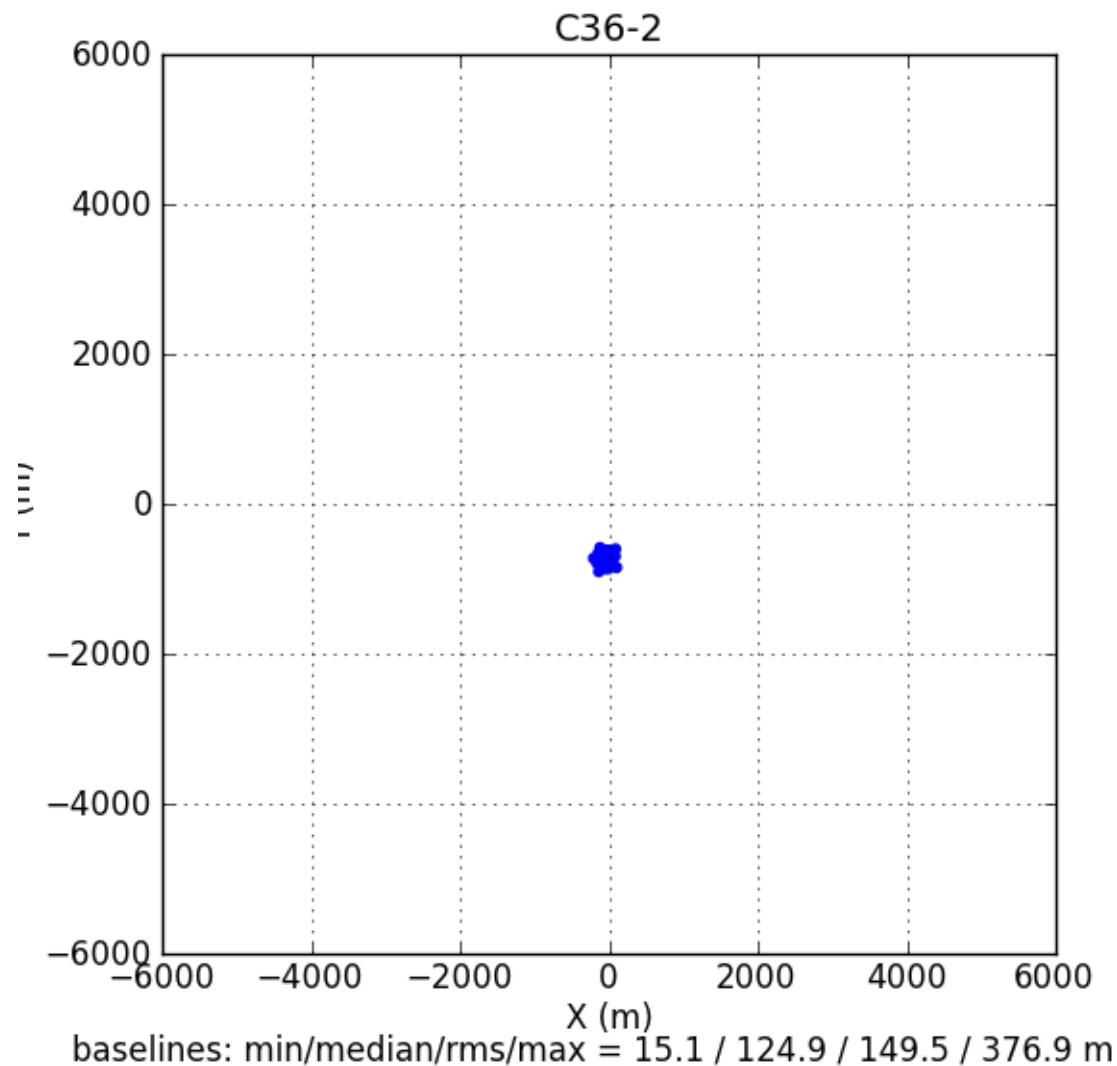
Number of antennas

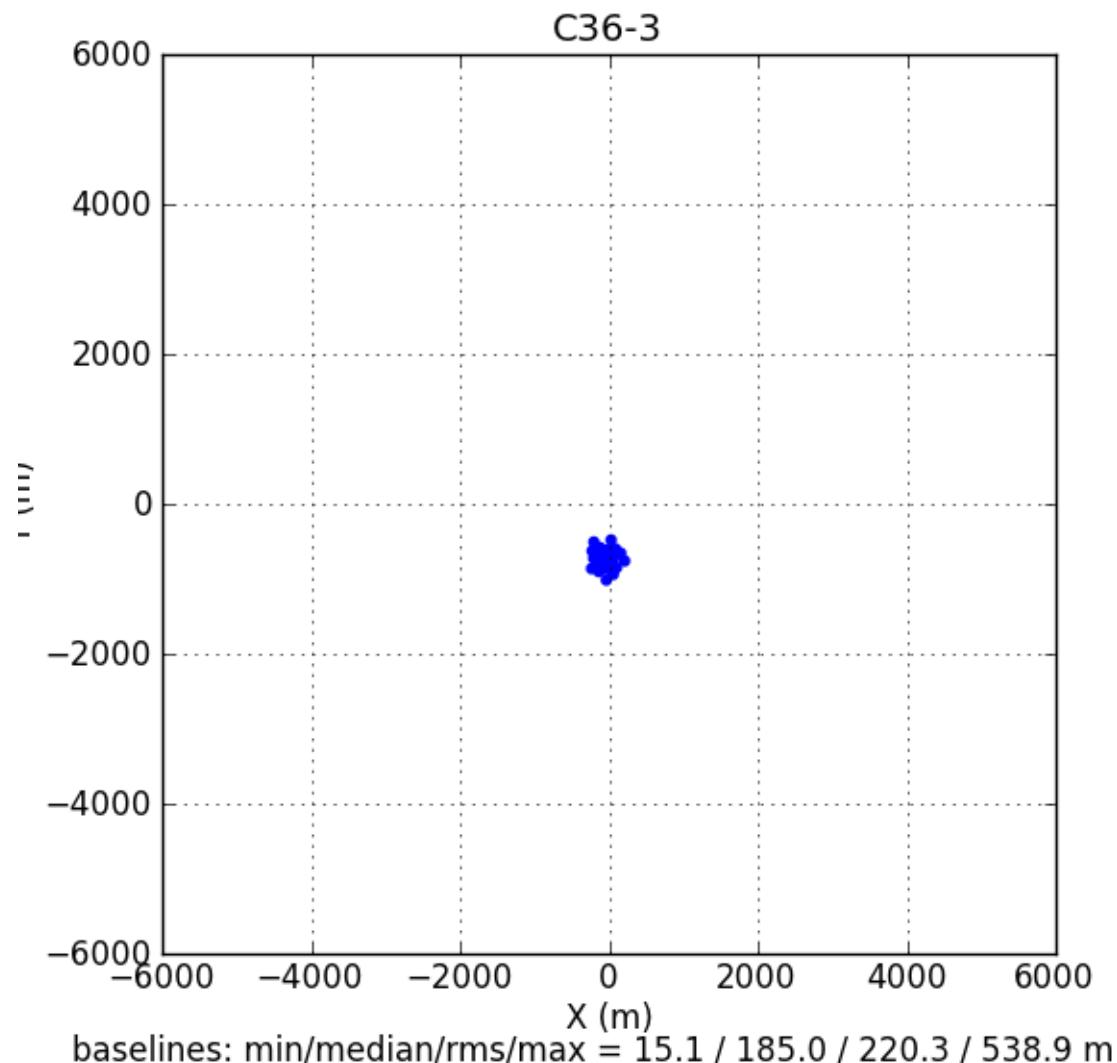
- Expected (minimum) number of antennas
 - 12-m array: 36 [Cycle 2: 34]
 - 7-m array: 10 [9]
 - TP array: 2 [2]
- Modest increase in sensitivity
- Improvement in (u,v) coverage is better
 - 12% more 12-m baselines
 - Better calibration and imaging

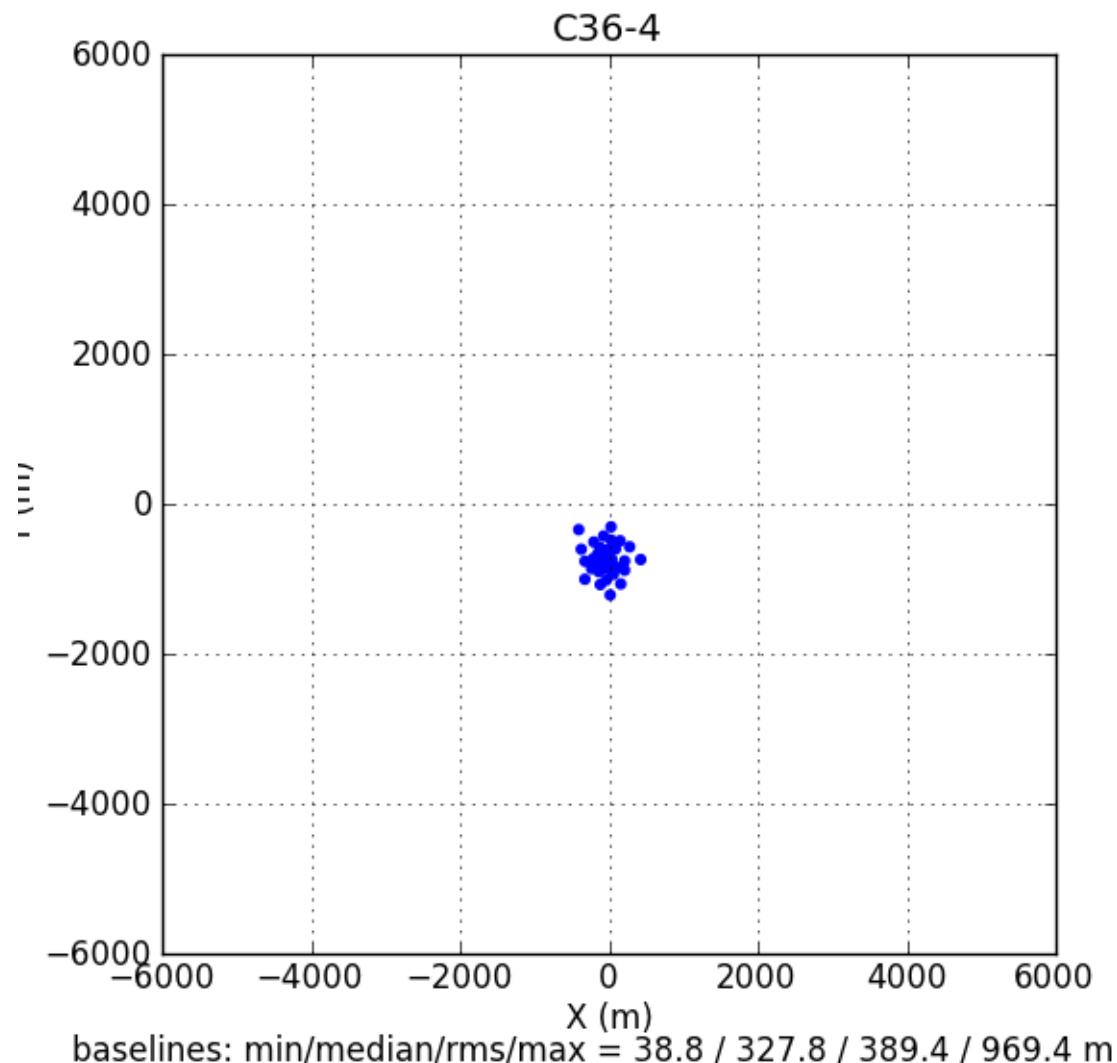
Antenna configurations

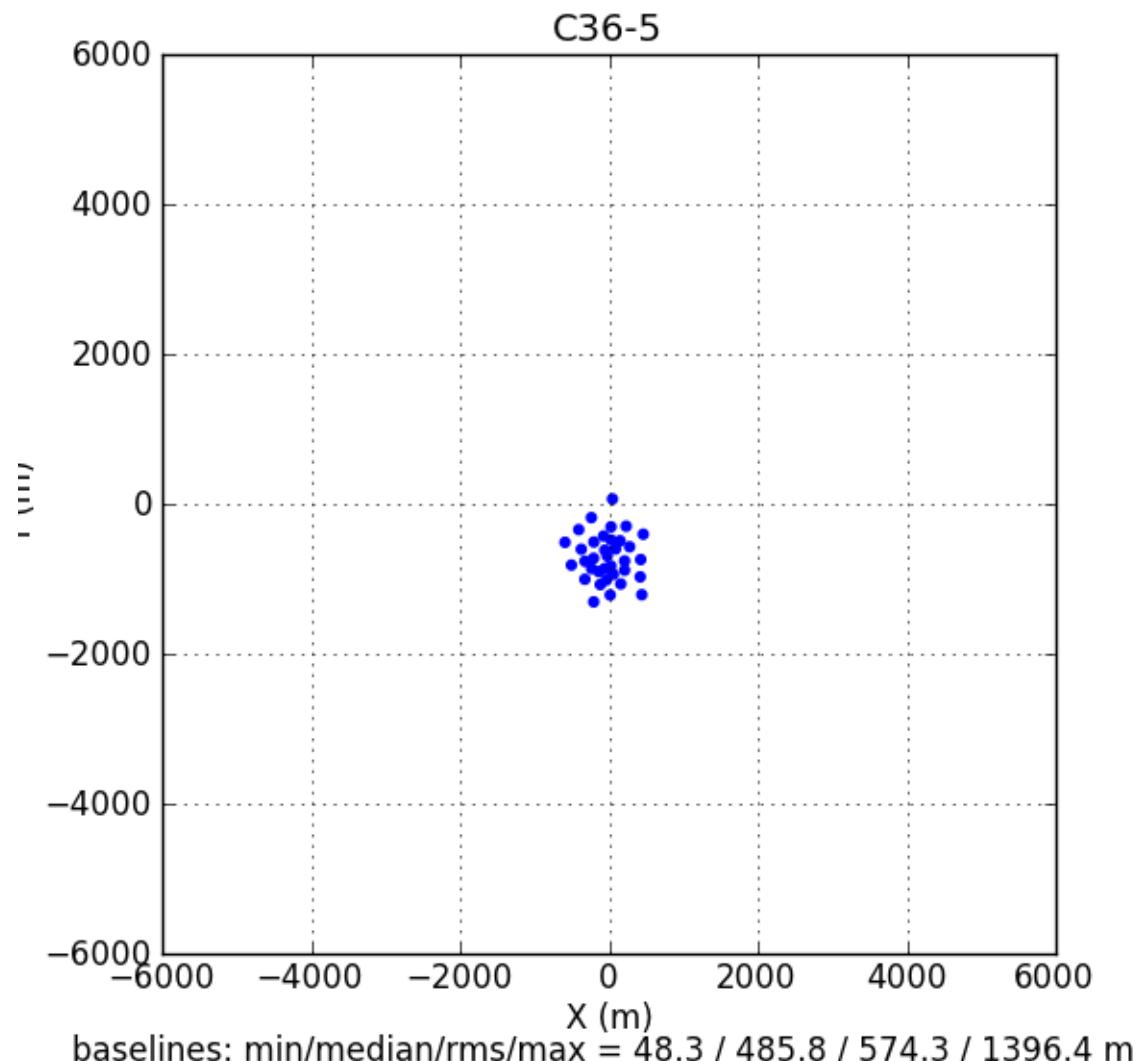
- 8 12-m array configurations in total
 - C36-1: 14.7 | 160.7 [min | max baseline (m)]
 - C36-2: 14.7 | 376.9
 - C36-3: 14.7 | 538.9
 - C36-4: 38.6 | 969.4
 - C36-5: 47.9 | 1396.4 (close to Cycle-2 C34-7)
 - C36-6: 77.3 | 2299.6
 - **C36-7: 248.3 | 6074.2**
 - **C36-8: 346.5 | 9743.7**
- 1 ACA configuration
 - 8.7 | 32.1

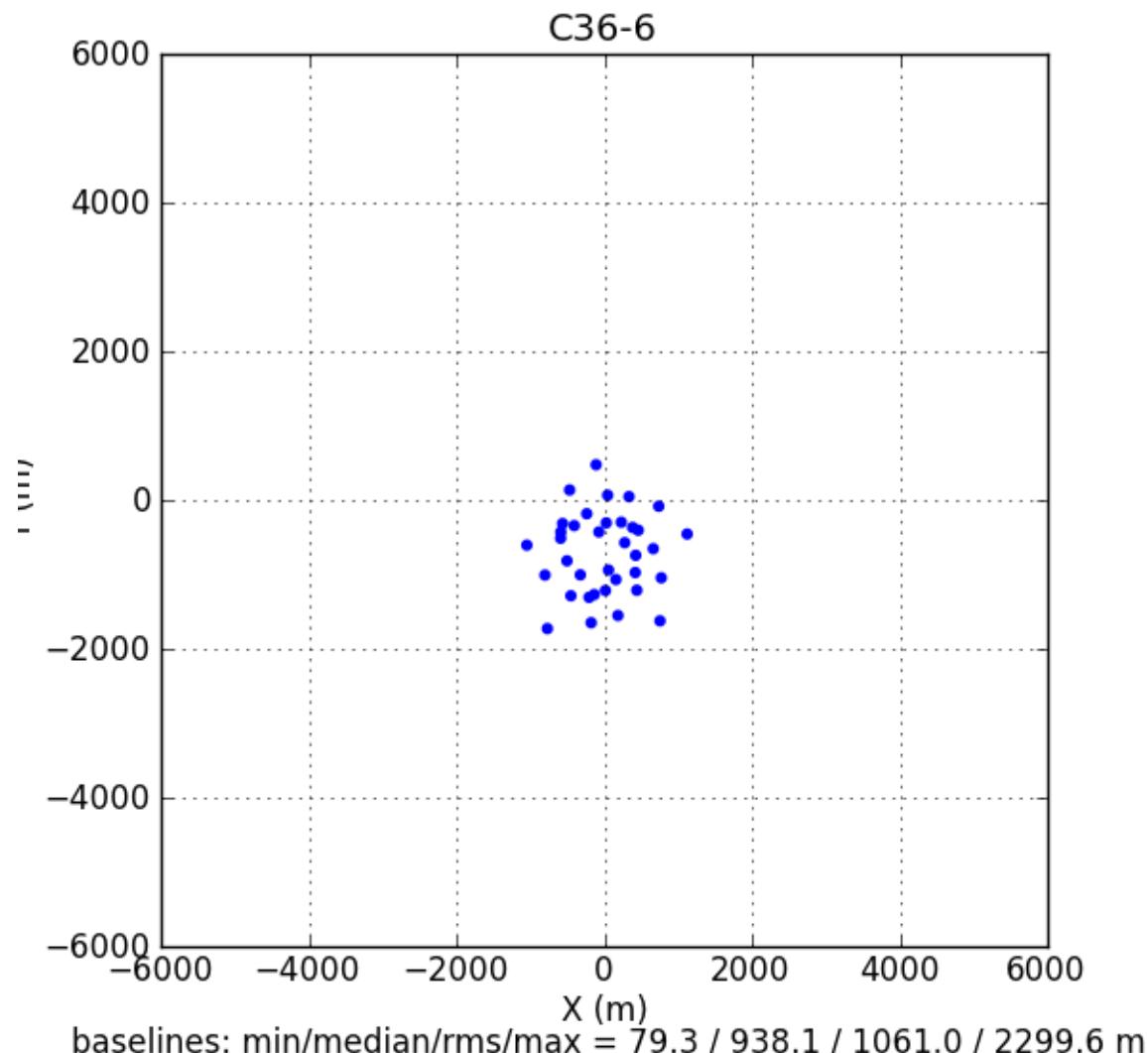


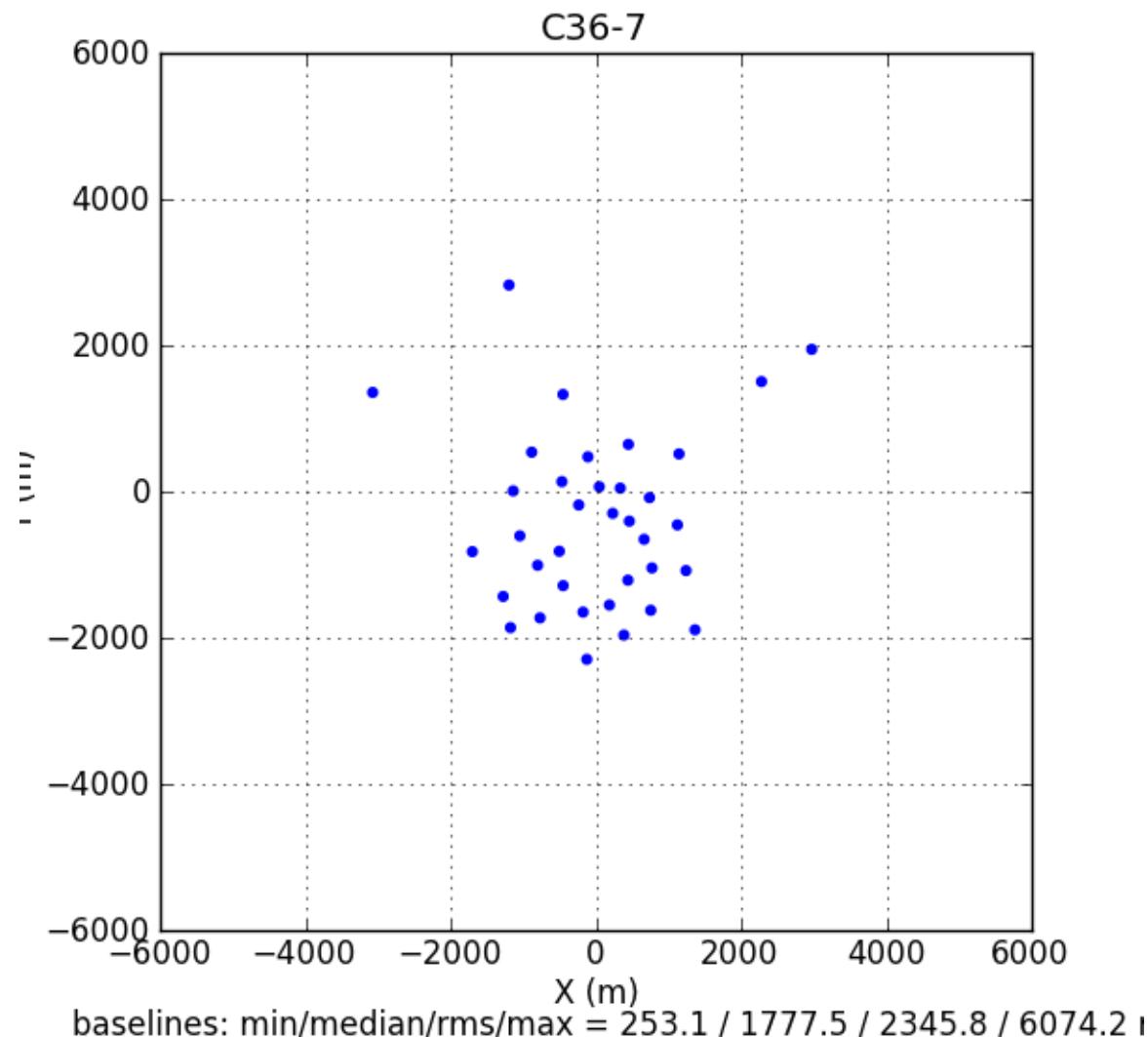


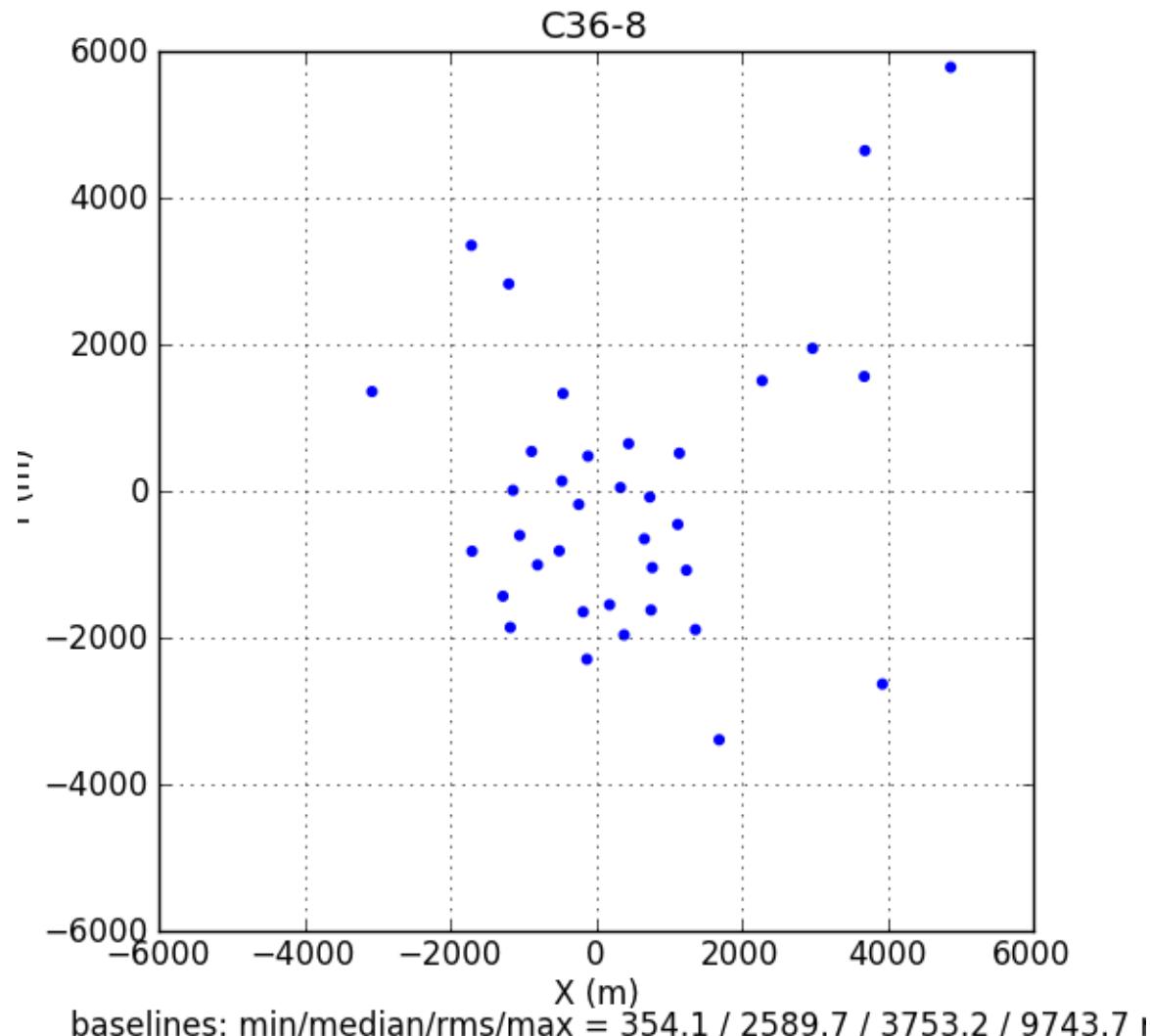












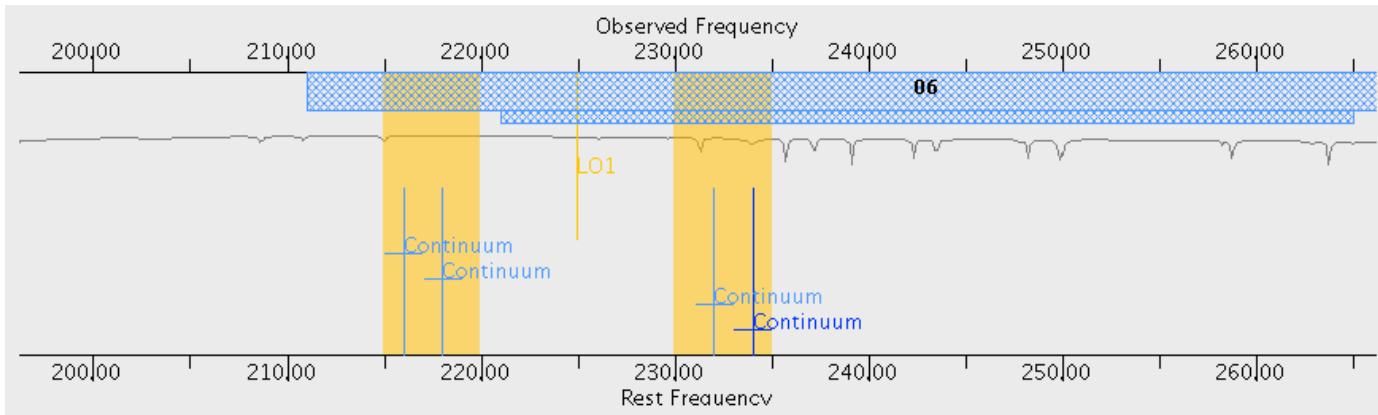
“Long”-baseline configurations

- C36-7 and -8
- Available configurations will be band-dependent
 - B3, 4 and 6: ≤ 10 km
 - B7: ≤ 6 km
 - B8, 9 and 10: ≤ 2.5 km
- More frequent phase referencing
 - Approximately every minute
 - Time estimates are therefore relatively high

Array selection

- Parameters for selection
 - Requested angular resolution ($\Delta\theta$)
 - Largest angular scale in source (LAS)
 - Max recoverable scale of array (MRS, set by min baseline)
- Procedure
 - Choose smallest 12-m configuration that achieves $\Delta\theta$
 - Until MRS > LAS, add
 - A smaller 12-m configuration (if first was C36-6, -5 or -4)
 - The 7-m array
 - The TP array
- Special restriction for “long”-baseline configuration
 - No other configurations allowed

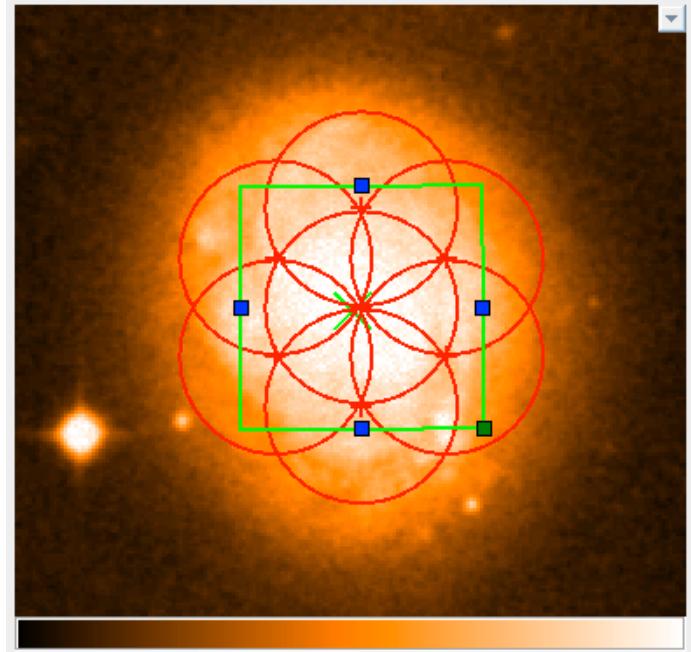
Expanded full-polarization tuning range



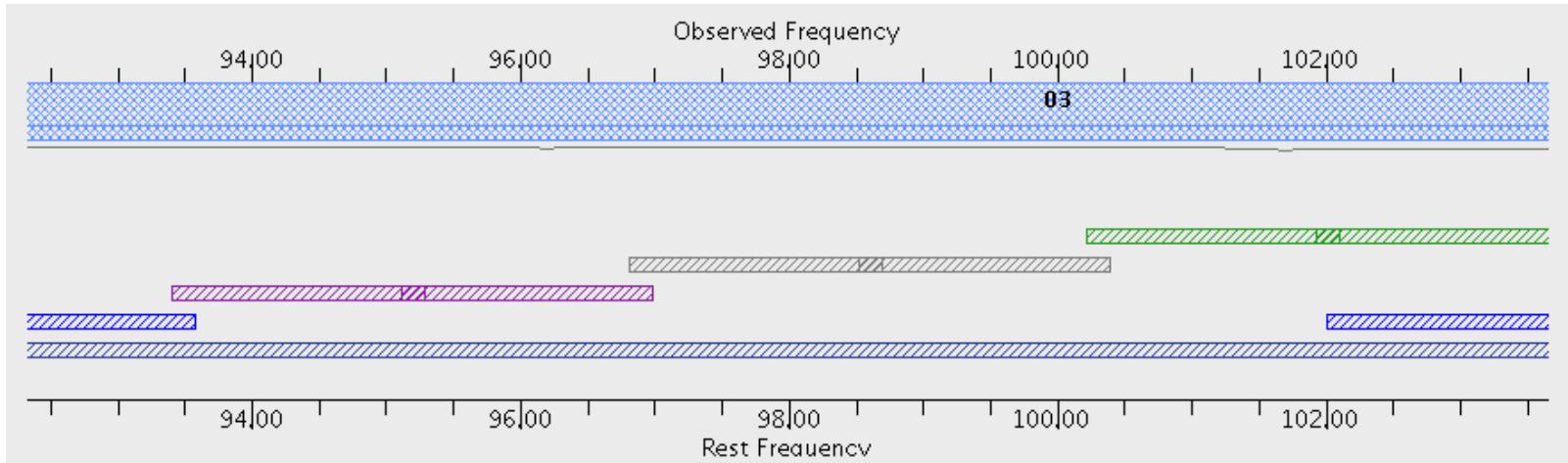
- Cycle 3
 - “Single continuum” setups at arbitrary frequencies
- Continuing restrictions
 - Still only Bands 3, 6 and 7
 - On-axis only (single pointings)
 - Only **linear** polarization officially supported
 - Only one tuning per SG
 - 3-hour minimum time estimate will remain

Antenna beamsize

- HPBW of antenna
 - Also known as ‘primary beam’
 - Sets Field of View
- New definition: $1.13 \lambda / D$
 - OT had previously assumed $1.2 \lambda / D$
- The number of mosaic pointings will not change
 - Mosaic spacing remains at $\lambda / (\sqrt{3} D)$
 - This number is appropriate for triangularly-spaced mosaics



Spectral Scan overlaps



- Quality of edge channels can be poor
 - Ringing
 - Bad T_{sys} measurements
 - This is in addition to the TDM channels that are usually flagged
- An overlap will therefore be enforced in Cycle 3
 - Only a few TDM channels
 - Exact overlap depends on bandwidth of spw

“Non-standard” modes

- Broadly defined as projects that can't be pipelined
 - These can make up 25% of available **12-m** time
- Complete list
 - Bands 8, 9 & 10
 - Narrow aggregate bandwidth (<1 GHz)
 - Full polarization
 - Spectral Scans
 - Long baselines (C36-7 or -8)
 - User-defined calibrations
- Will be indicated to PI on cover sheet

Additional new OT features

- Multiple rectangular fields per Science Goal
- Import and export of field source pointings
- TJ node overhauled
- Improved time constrained interface
- Improved time estimate dialogue
- Various improvements to ALMA Sensitivity Calculator