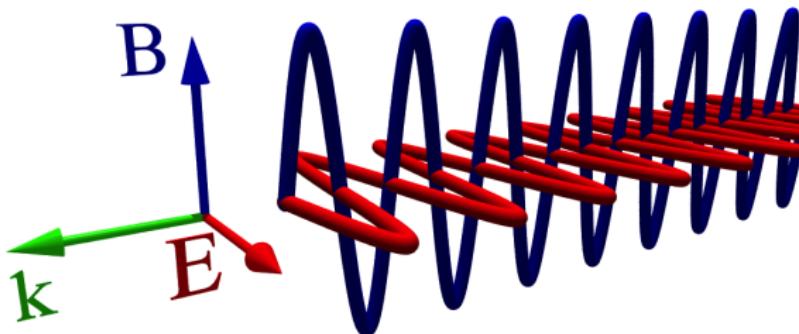


## Fundamentals of Polarized Light

- ① Polarized Light in the Universe
- ② Descriptions of Polarized Light
- ③ Polarizers
- ④ Retarders

# Polarized Light in the Universe

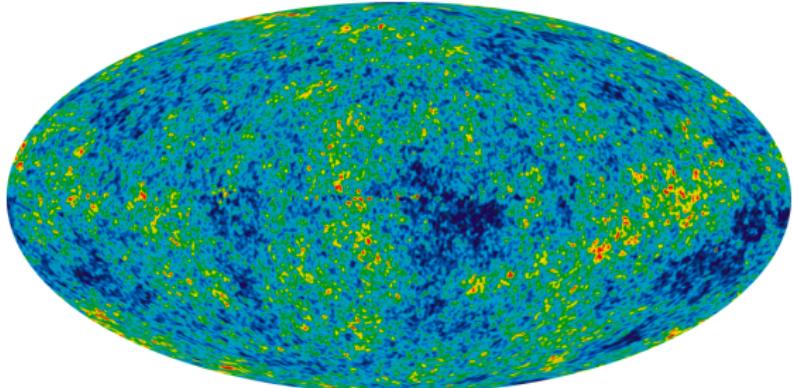


*Polarization* indicates *anisotropy*  $\Rightarrow$  not all directions are equal

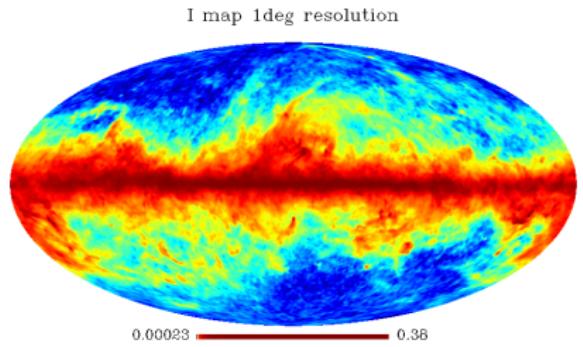
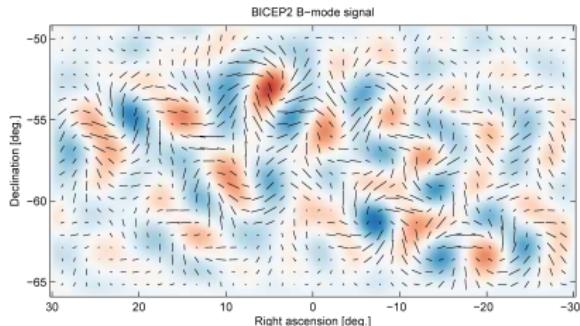
Typical anisotropies introduced by

- geometry (not everything is spherically symmetric)
- temperature gradients
- density gradients
- magnetic fields
- electrical fields

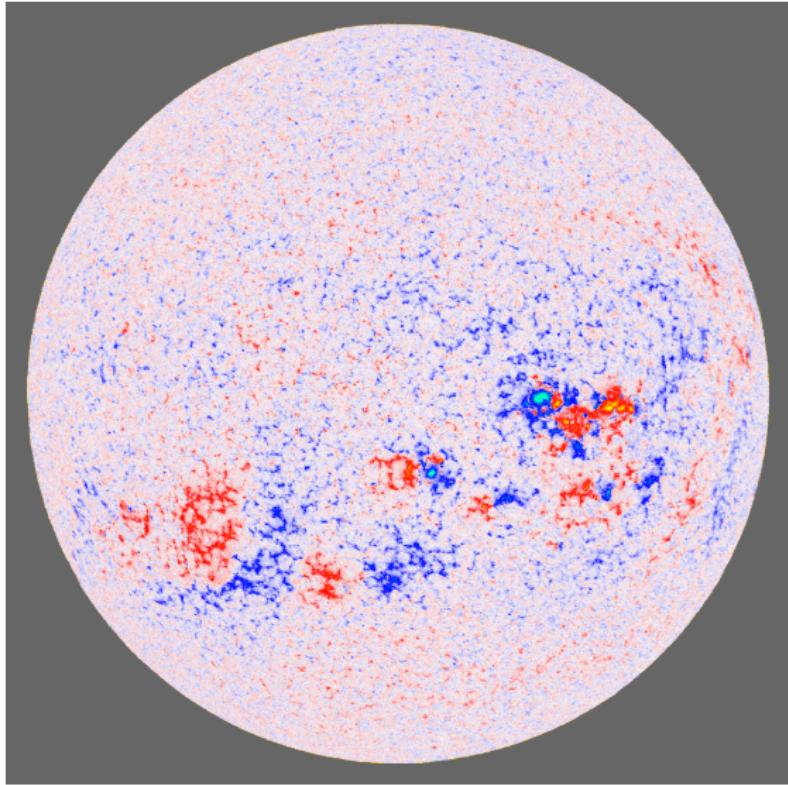
# 13.7 billion year old temperature fluctuations from WMAP



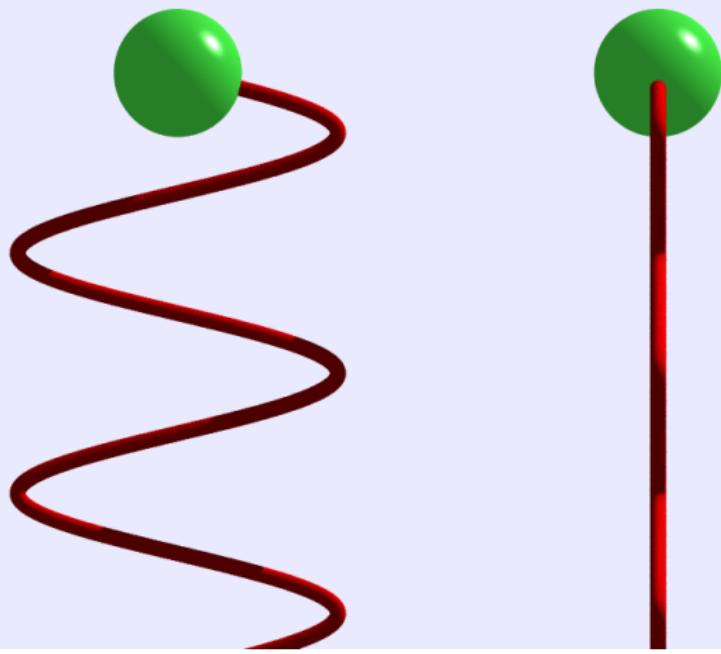
## BICEP2 results and Planck Dust Polarization Map



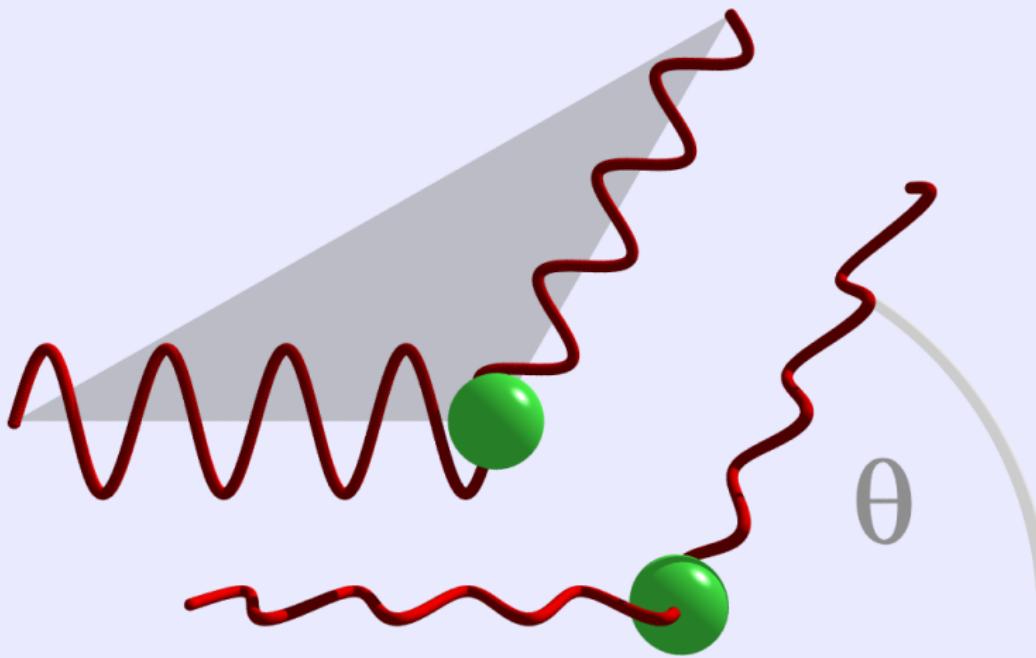
# Solar Magnetic Field Maps from Longitudinal Zeeman Effect



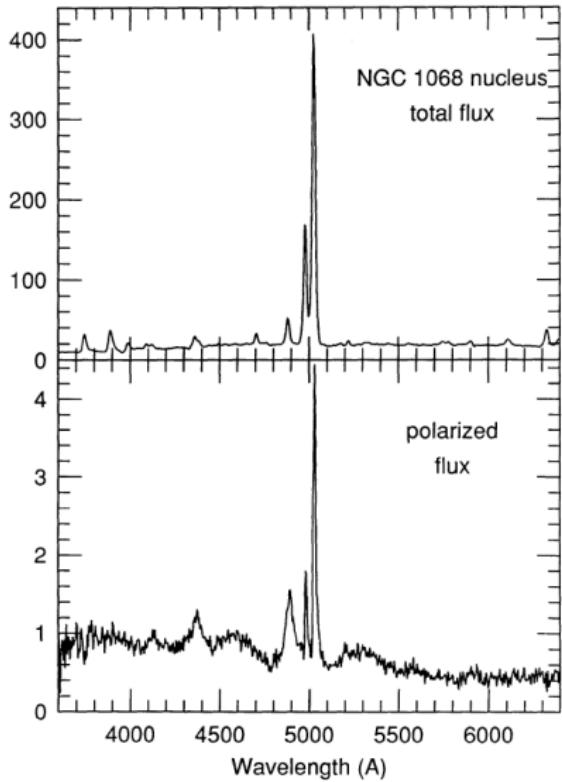
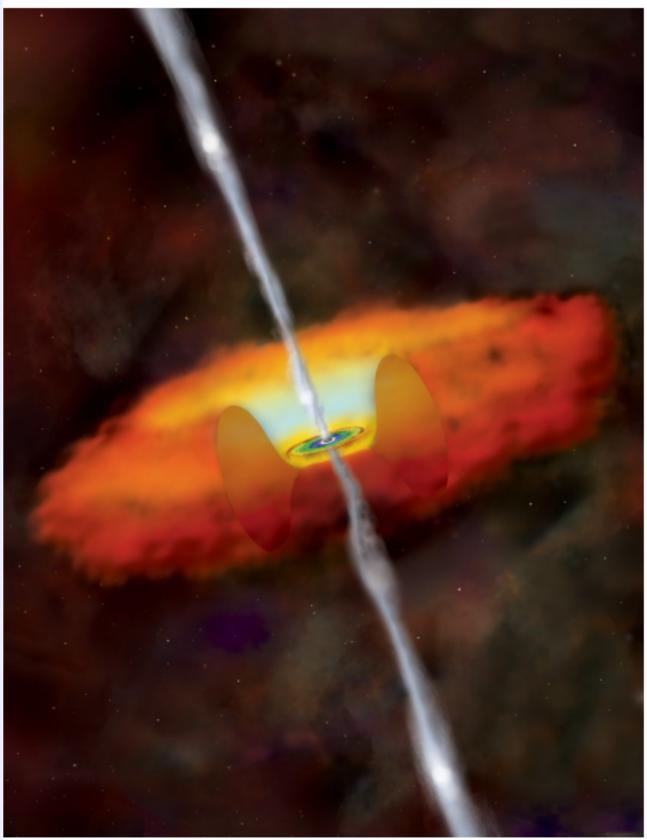
# Scattering Polarization 1



## Scattering Polarization 2



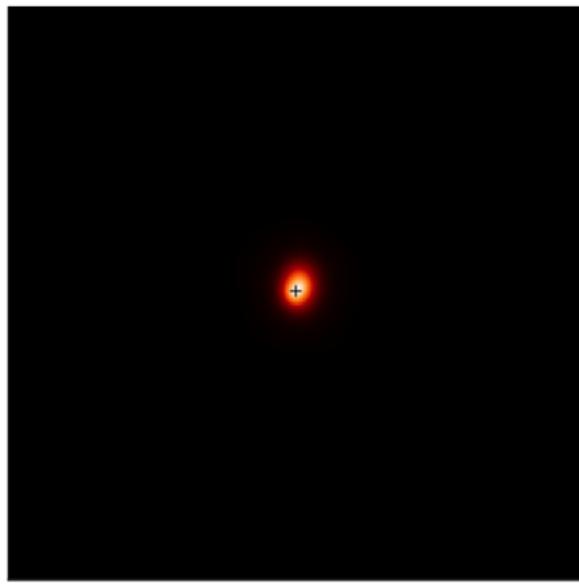
# Unified Model of Active Galactic Nuclei



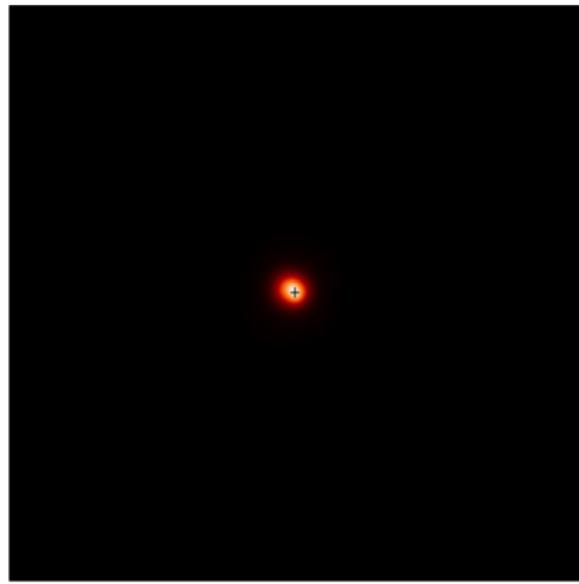
Miller et al., 1991, ApJ 378, 47-64

# The Power of Polarimetry

T Tauri in intensity

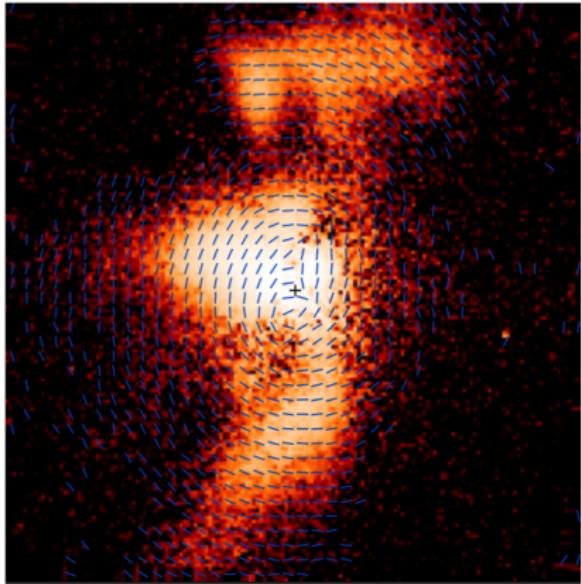


MWC147 in intensity

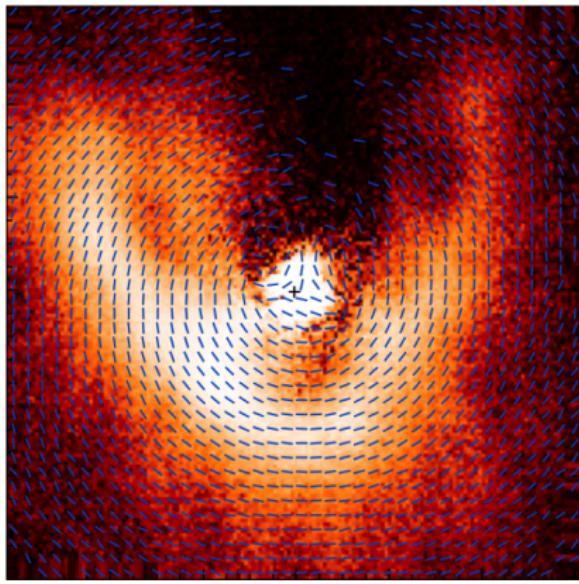


# The Power of Polarimetry

T Tauri in Linear Polarization

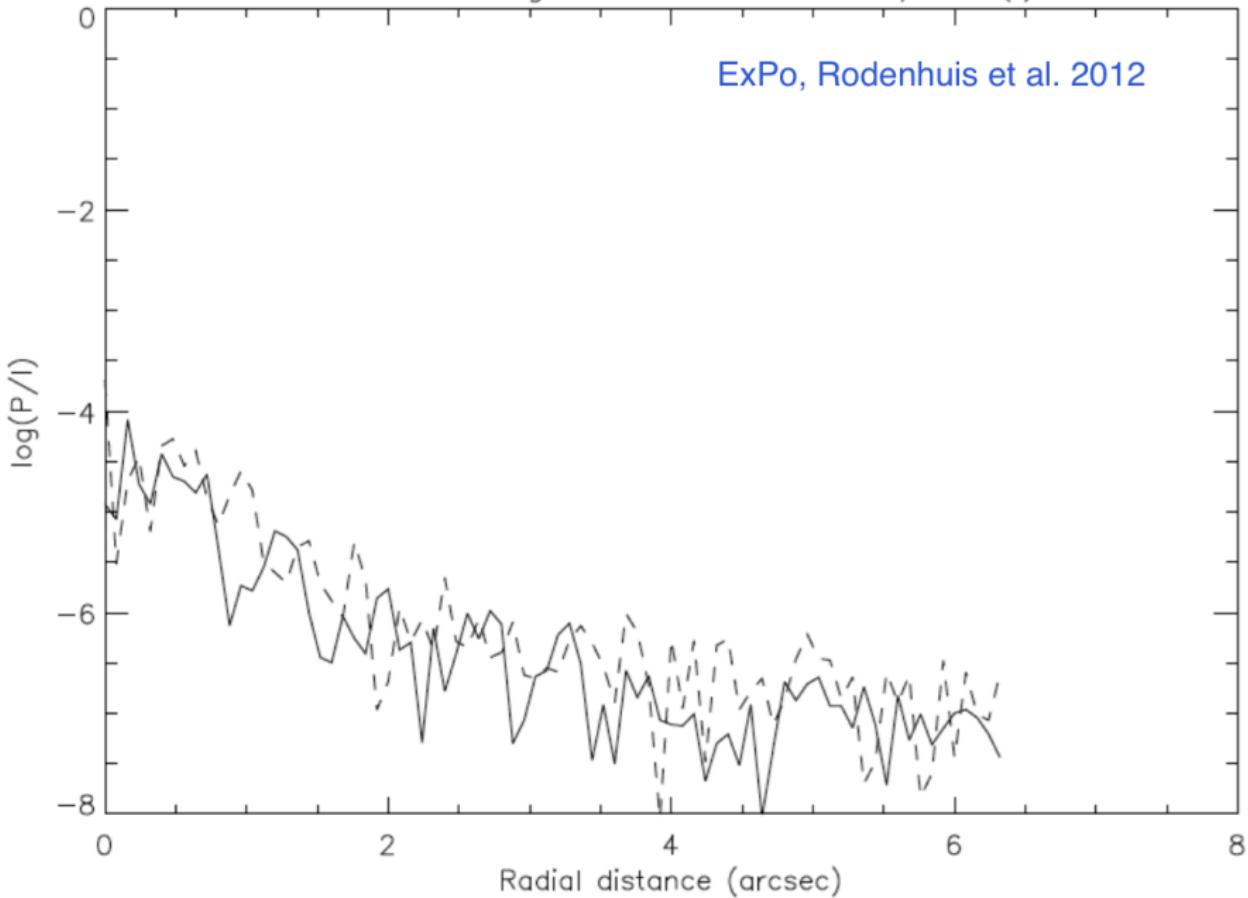


MWC147 in Linear Polarization



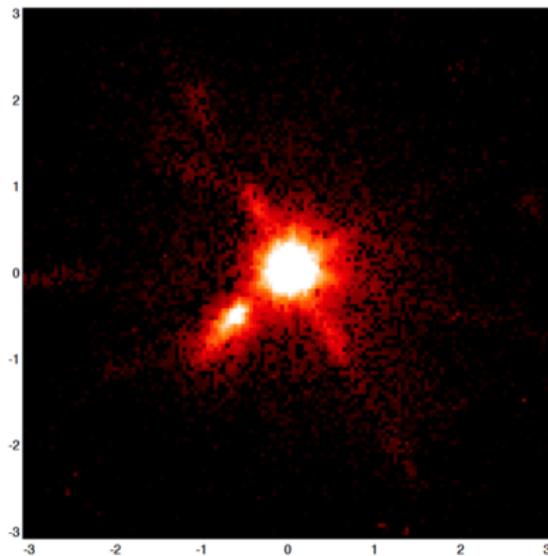
# MWC480 Tangential Contrast $\tan P / \max(I)$

ExPo, Rodenhuis et al. 2012

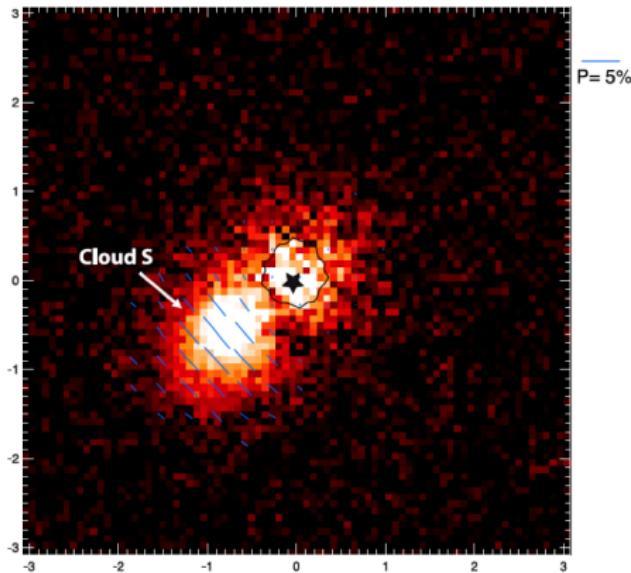


## The Power of Polarimetry 2

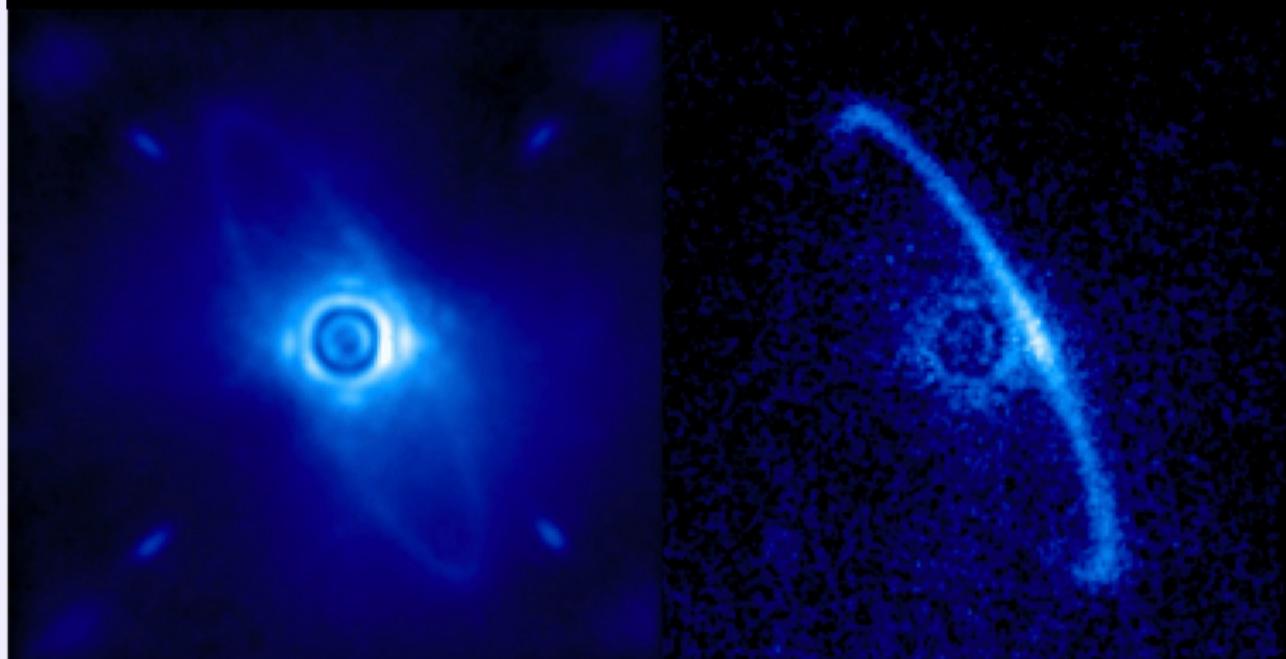
R Cr B with HST in intensity



R Cr B in Linear Polarization

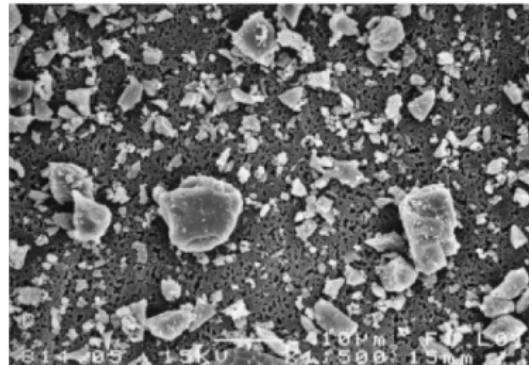


# Characterization

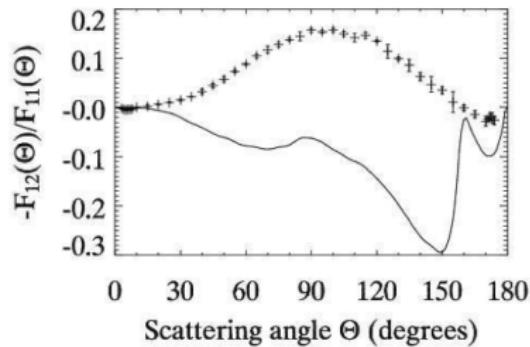
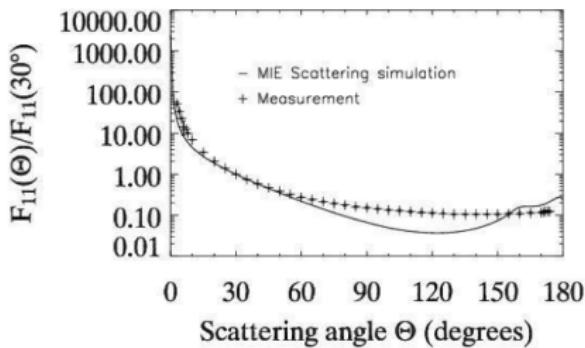
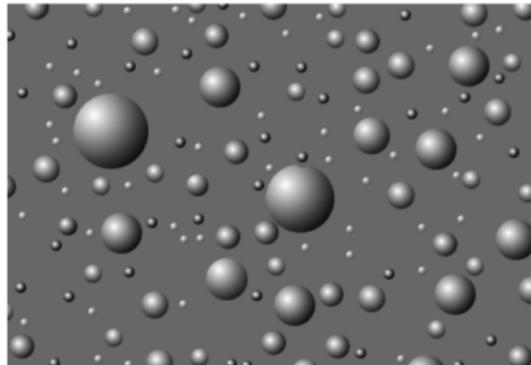


GPI First Light HR4796A (Image credit: Processing by Marshall Perrin, STScI)

real dust

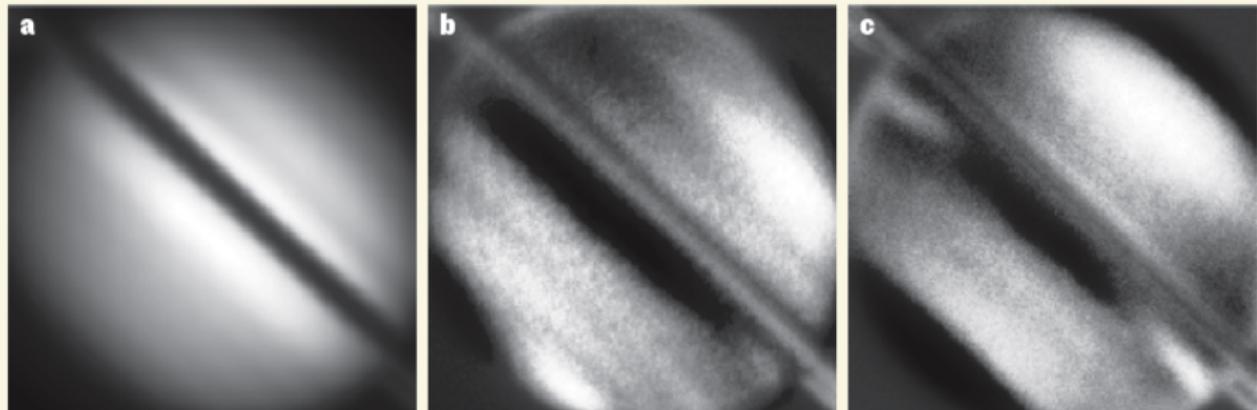


theoretical dust



Laan et al (2007)

## Planetary Scattered Light

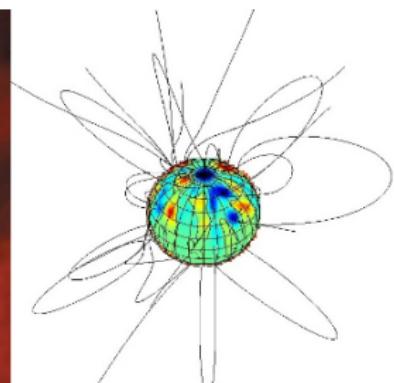
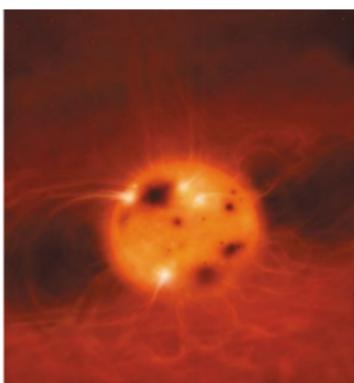
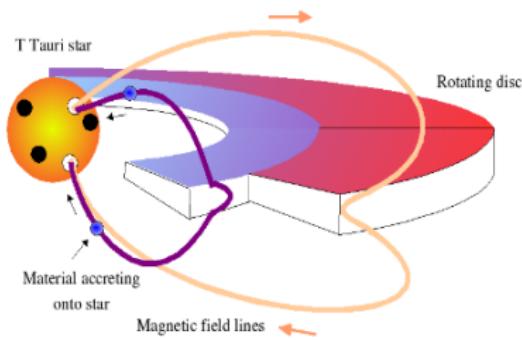


- solar-system planets show scattering polarization
- much depends on cloud height
- can be used to study exoplanets
- ExPo@WHT, SPHERE@VLT, EPICS@E-ELT

## Other astrophysical applications

- interstellar magnetic field from polarized starlight
- supernova asymmetries
- stellar magnetic fields from Zeeman effect
- galactic magnetic field from Faraday rotation

## Magnetic Fields of TTauri Stars (Courtesy S.V.Jeffers)



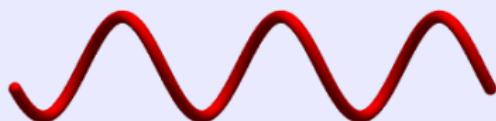
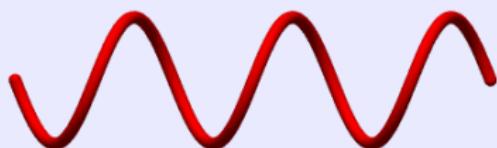
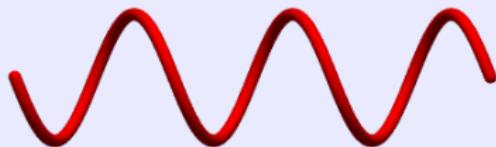
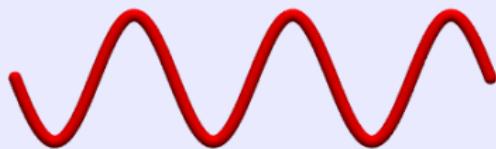
## Summary of Polarization Origin

- Plane Vector Wave ansatz  $\vec{E} = \vec{E}_0 e^{i(\vec{k} \cdot \vec{x} - \omega t)}$
- spatially, temporally constant vector  $\vec{E}_0$  lays in plane perpendicular to propagation direction  $\vec{k}$
- represent  $\vec{E}_0$  in 2-D basis, unit vectors  $\vec{e}_x$  and  $\vec{e}_y$ , both perpendicular to  $\vec{k}$

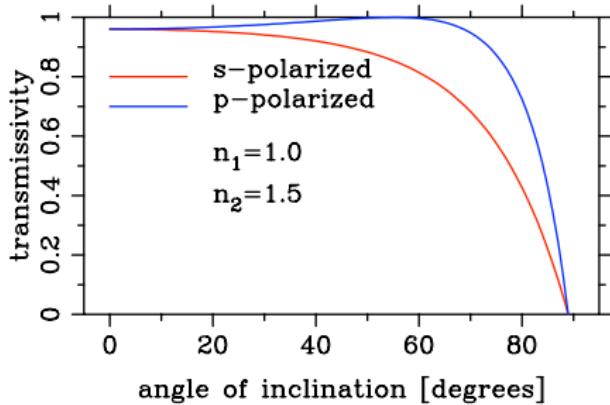
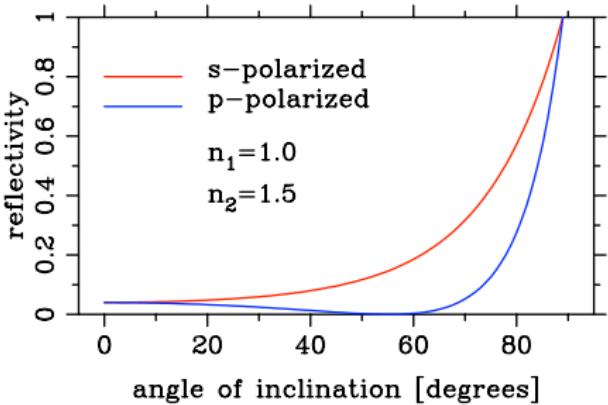
$$\vec{E}_0 = E_x \vec{e}_x + E_y \vec{e}_y.$$

$E_x, E_y$ : arbitrary complex scalars

- damped plane-wave solution with given  $\omega$ ,  $\vec{k}$  has 4 degrees of freedom (two complex scalars)
- additional property is called *polarization*
- many ways to represent these four quantities
- if  $E_x$  and  $E_y$  have identical phases,  $\vec{E}$  oscillates in fixed plane



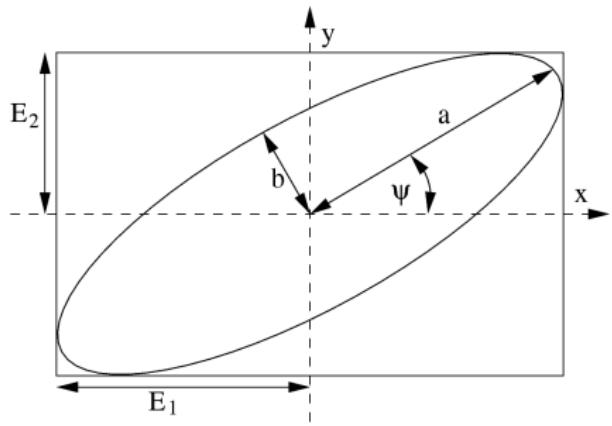
# Polarization at Non-Normal Incidence



- transmitted light is p-polarized (electric field in plane of incidence)
- reflected light is s-polarized (electric field perpendicular to plane of incidence)
- at Brewster angle, reflected light is 100% polarized
- transmitted light is moderately polarized

# Polarization Ellipse

## Polarization Ellipse



## Polarization

- $$\vec{E}(t) = \vec{E}_0 e^{i(\vec{k} \cdot \vec{x} - \omega t)}$$
- $$\vec{E}_0 = |E_x| e^{i\delta_x} \vec{e}_x + |E_y| e^{i\delta_y} \vec{e}_y$$
- wave vector in  $z$ -direction
  - $\vec{e}_x, \vec{e}_y$ : unit vectors in  $x, y$
  - $|E_x|, |E_y|$ : (real) amplitudes
  - $\delta_{x,y}$ : (real) phases

## Polarization Description

- 2 complex scalars not the most useful description
- at given  $\vec{x}$ , time evolution of  $\vec{E}$  described by *polarization ellipse*
- ellipse described by axes  $a, b$ , orientation  $\psi$

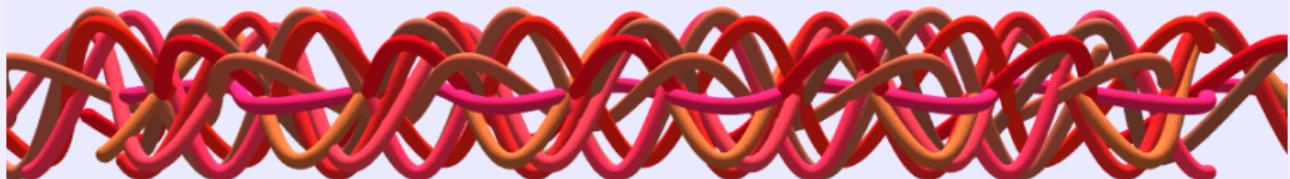
## Stokes Vector

- formalism to describe polarization of quasi-monochromatic light
- directly related to measurable intensities
- Stokes vector fulfills these requirements

$$\vec{I} = \begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \begin{pmatrix} E_x E_x^* + E_y E_y^* \\ E_x E_x^* - E_y E_y^* \\ E_x E_y^* + E_y E_x^* \\ i(E_x E_y^* - E_y E_x^*) \end{pmatrix} = \begin{pmatrix} |E_x|^2 + |E_y|^2 \\ |E_x|^2 - |E_y|^2 \\ 2|E_x||E_y|\cos\delta \\ 2|E_x||E_y|\sin\delta \end{pmatrix}$$

Jones vector elements  $E_{x,y}$ , real amplitudes  $|E_{x,y}|$ , phase difference  $\delta = \delta_y - \delta_x$

- $I^2 \geq Q^2 + U^2 + V^2$
- can describe unpolarized ( $Q = U = V = 0$ ) light



## Stokes Vector Interpretation

$$\vec{I} = \begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \begin{pmatrix} \text{intensity} \\ \text{linear } 0^\circ - \text{linear } 90^\circ \\ \text{linear } 45^\circ - \text{linear } 135^\circ \\ \text{circular left - right} \end{pmatrix}$$

- *degree of polarization*

$$P = \frac{\sqrt{Q^2 + U^2 + V^2}}{I}$$

1 for fully polarized light, 0 for unpolarized light

- summing of Stokes vectors = *incoherent* adding of quasi-monochromatic light waves

## Linear Polarization

- horizontal:  $\begin{pmatrix} 1 \\ -1 \\ 0 \\ 0 \end{pmatrix}$
- vertical:  $\begin{pmatrix} 1 \\ 1 \\ 0 \\ 0 \end{pmatrix}$
- $45^\circ$ :  $\begin{pmatrix} 1 \\ 0 \\ 1 \\ 0 \end{pmatrix}$

## Circular Polarization

- left:  $\begin{pmatrix} 1 \\ 0 \\ 0 \\ 1 \end{pmatrix}$
- right:  $\begin{pmatrix} 1 \\ 0 \\ 0 \\ -1 \end{pmatrix}$

## Mueller Matrices

- $4 \times 4$  real Mueller matrices describe (linear) transformation between Stokes vectors when passing through or reflecting from media

$$\vec{I}' = M \vec{I},$$

$$M = \begin{pmatrix} M_{11} & M_{12} & M_{13} & M_{14} \\ M_{21} & M_{22} & M_{23} & M_{24} \\ M_{31} & M_{32} & M_{33} & M_{34} \\ M_{41} & M_{42} & M_{43} & M_{44} \end{pmatrix}$$

- $N$  optical elements, combined Mueller matrix is

$$M' = M_N M_{N-1} \cdots M_2 M_1$$



- polarizer: optical element that produces polarized light from unpolarized input light
- linear, circular, or in general elliptical polarizer, depending on type of transmitted polarization
- linear polarizers by far the most common
- large variety of polarizers

### Vertical Linear Polarizer

$$M_{\text{pol}}(\theta) = \frac{1}{2} \begin{pmatrix} 1 & 1 & 0 & 0 \\ 1 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

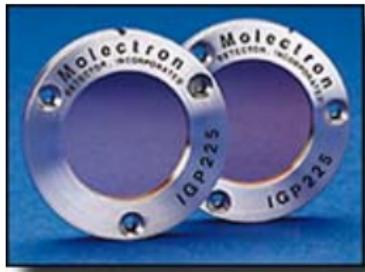
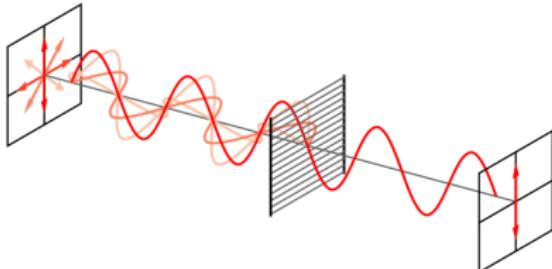
### Horizontal Linear Polarizer

$$M_{\text{pol}}(\theta) = \frac{1}{2} \begin{pmatrix} 1 & -1 & 0 & 0 \\ -1 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

### Mueller Matrix for Ideal Linear Polarizer at Angle $\theta$

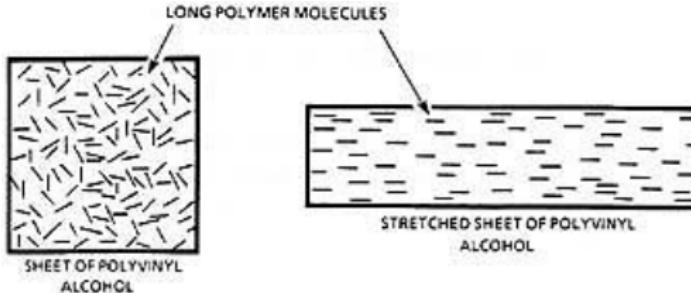
$$M_{\text{pol}}(\theta) = \frac{1}{2} \begin{pmatrix} 1 & \cos 2\theta & \sin 2\theta & 0 \\ \cos 2\theta & \cos^2 2\theta & \sin 2\theta \cos 2\theta & 0 \\ \sin 2\theta & \sin 2\theta \cos 2\theta & \sin^2 2\theta & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

## Wire Grid Polarizers



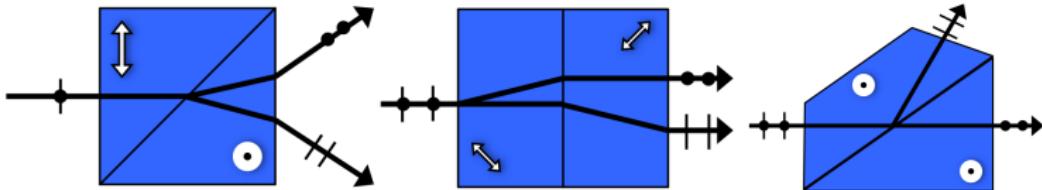
- parallel conducting wires, spacing  $d \lesssim \lambda$  act as polarizer
- polarization perpendicular to wires is transmitted
- rule of thumb:
  - $d < \lambda/2 \Rightarrow$  strong polarization
  - $d \gg \lambda \Rightarrow$  high transmission of both polarization states (weak polarization)
- mostly used in infrared

## Polaroid-type Polarizers



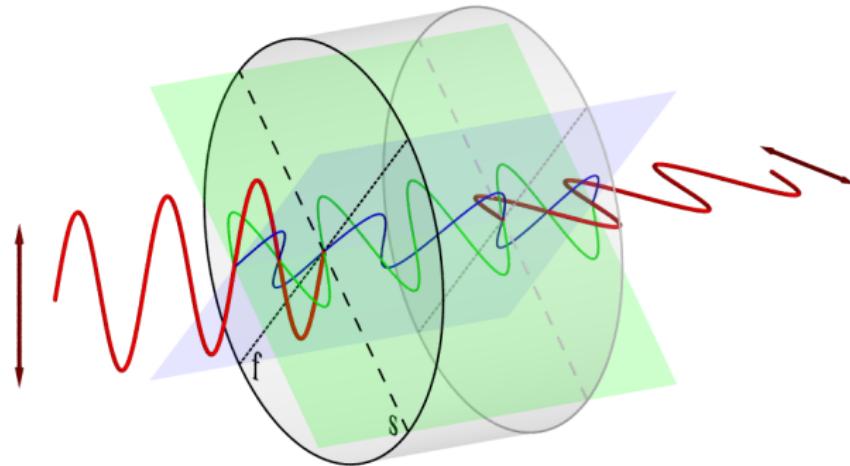
- developed by Edwin Land in 1938 ⇒ Polaroid
- sheet polarizers: stretched polyvinyl alcohol (PVA) sheet, laminated to sheet of cellulose acetate butyrate, treated with iodine
- PVA-iodine complex analogous to short, conducting wire
- cheap, can be manufactured in large sizes

## Crystal-Based Polarizers



- crystals are basis of highest-quality polarizers
- precise arrangement of atom/molecules and anisotropy of index of refraction separate incoming beam into two beams with precisely orthogonal linear polarization states
- work well over large wavelength range
- many different configurations
- calcite most often used in crystal-based polarizers because of large birefringence, low absorption in visible
- many other suitable materials

## Retarders or Wave Plates



- retards (delays) phase of one electric field component with respect to orthogonal component
- anisotropic material (crystal) has index of refraction that depends on polarization direction

## Retarder Properties

- does not change intensity or degree of polarization
- characterized by two Stokes vectors that are not changed  $\Rightarrow$  eigenvectors of retarder
- depending on polarization described by eigenvectors, retarder is
  - linear retarder
  - circular retarder
  - elliptical retarder
- linear retarders by far the most common type of retarder

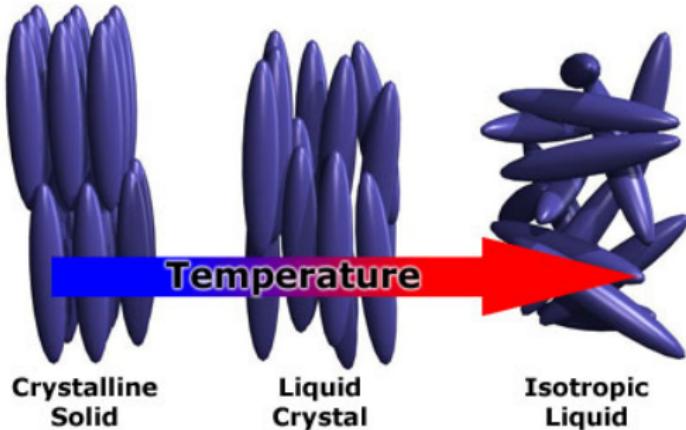
## Mueller Matrix for Linear Retarder

$$M_r = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \cos \delta & -\sin \delta \\ 0 & 0 & \sin \delta & \cos \delta \end{pmatrix}$$

## Variable Retarders

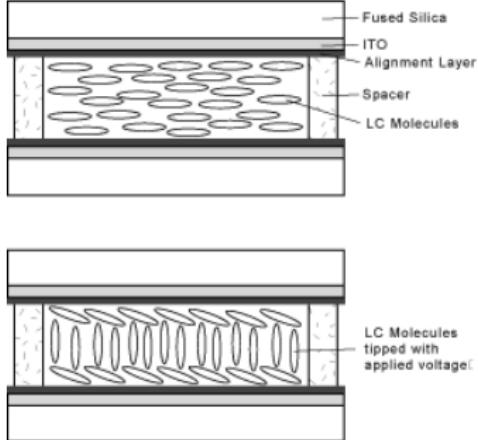
- sensitive polarimeters require retarders whose properties (retardance, fast axis orientation) can be varied quickly (*modulated*)
- retardance changes (change of birefringence):
  - liquid crystals
  - Faraday, Kerr, Pockels cells
  - piezo-elastic modulators (PEM)
- fast axis orientation changes (change of c-axis direction):
  - rotating fixed retarder
  - ferro-electric liquid crystals (FLC)

# Liquid Crystals



- liquid crystals: fluids with elongated molecules
- at high temperatures: liquid crystal is isotropic
- at lower temperature: molecules become ordered in orientation and sometimes also space in one or more dimensions
- liquid crystals can line up parallel or perpendicular to external electrical field

# Liquid Crystal Retarders



- dielectric constant anisotropy often large  $\Rightarrow$  very responsive to changes in applied electric field
- birefringence  $\delta n$  can be very large (larger than typical crystal birefringence)
- liquid crystal layer only a few  $\mu\text{m}$  thick
- birefringence shows strong temperature dependence