# **Recent Developments at ESO Bruno Leibundgut**



## **European Southern Observatory**

- Mission
  - Develop and operate world-class observing facilities for astronomical research
  - Organise collaborations in astronomy
- Intergovernmental treaty-level organization
  - Currently 14 member states
  - Some interest by other countries
- Observatories in Chile
  - La Silla Paranal: VLT, VLTI, 3.6m, NTT, VISTA, VST
  - Chajnantor: APEX and ALMA partnerships
- HQ in Garching and Office in Santiago



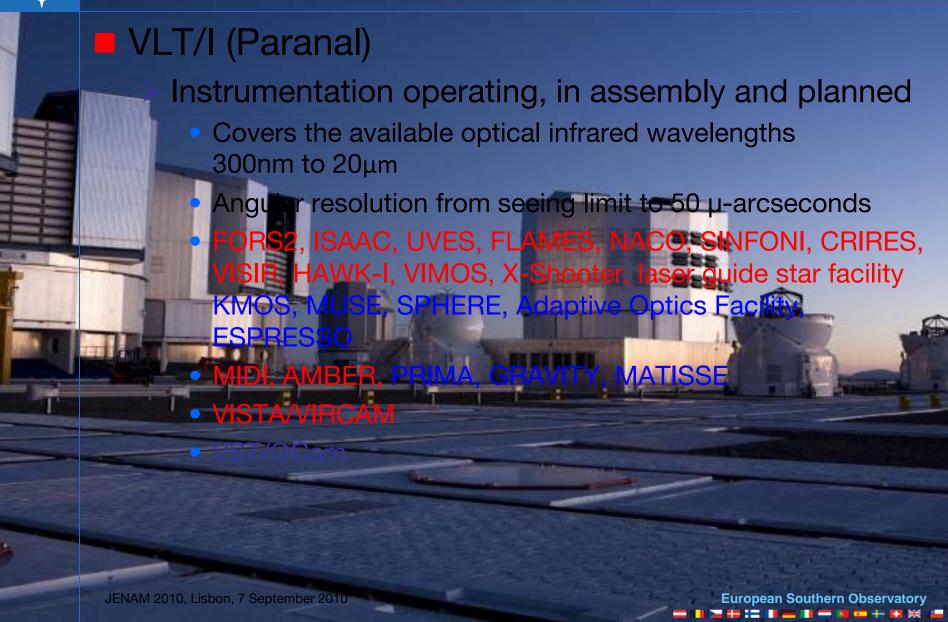
## ESO's sites







## La Silla Paranal





## La Silla Paranal





#### **ALMA**

- Science requirements
  - Detect CO and [CII] in Milky
    - Way galaxy at z=3 in < 24 hr
  - Dust emission, gas kinematics in proto-planetary disks
  - Resolution to match Hubble, JWST and 8-10m with AO
  - Complement to Herschel
- Specifications
  - 66 antennas (54x12m, 12x7m)
  - → 14km max baseline (<10mas)</p>
  - > 30-1000 GHz (10-0.3mm), up to 10 receiver bands

z<1.5 simulation 3 days of integration 4'x4' arcmir z<1.5 5AU

850 GHz

**European Southern Observatory** 

(12 days of integration)

JENAM 2010, Lisbon, 7 September 2010



## E-ELT

- Detailed design study
  - Baseline 42m primary mirror
  - Adaptive optics built-in
    - 8 instrument studies and 2 adaptive optics modules studies finished
      - Industry strongly engaged
  - Construction proposal completed and under review
- Project
  - Builds on entire expertise at ESO and in the member states
    - Synergy: JWST/ALMA/SKA



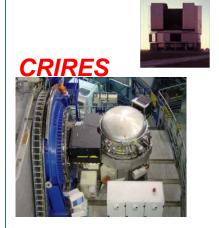
## **Paranal**







## **VLT Instruments**

















FLAMES.







## **VLTI Instruments**

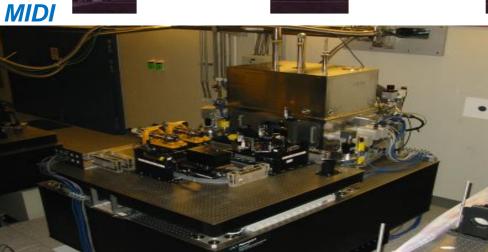












#### **AMBER**



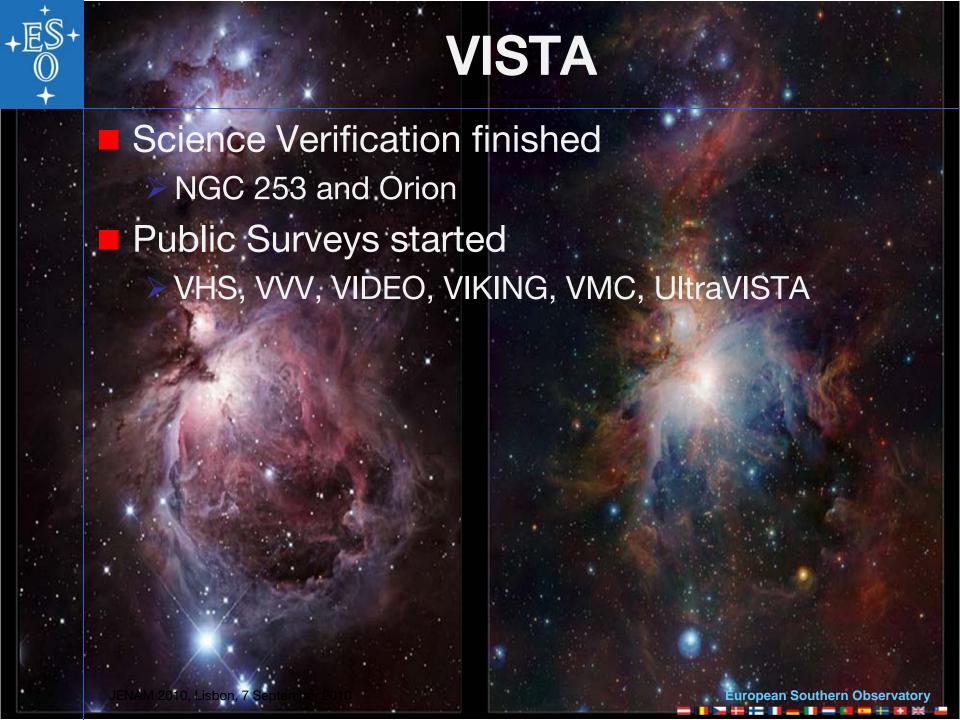


## **VISTA**

- VISTA in operations since April 2010
- Science Verification in late 2009
- Multi-year programme of large public surveys

More in Magda's presentation







## **VST**

- Final assembly and systems tests during the second half of 2010
- Surveys possibly will start in mid-2011
- Review of the public surveys in September 2010
  - > surveys were originally selected in 2005





## Top list of ESO science

- Galactic Centre
  - Supermassive black hole
- Extrasolar planets
  - First images of exo-planets
  - Lightest known planets
  - First direct spectrum of an exo-planet
- Accelerating Universe
  - Spectroscopy of distant supernovae
- Gamma-Ray Bursts/Supernovae
  - Explosion physics
  - Tracers of the distant universe



## Other top science from ESO

- Metal-poor stars
  - Tracing the chemical enrichment
  - Finding the oldest known stars
- Stellar populations in nearby galaxies
  - Measuring stars beyond the Local Group
- Massive galaxies in the distant Universe
  - Puzzles in galaxy formation
- Varying physical constants?
  - Measure the fine-structure constant over time
- Testing the cosmological model
  - Cosmic background temperature

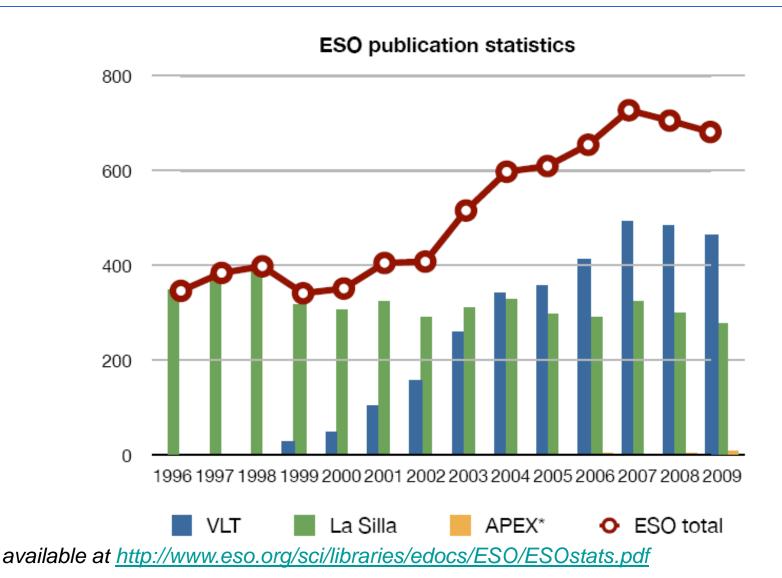


## More top science

- Detecting and imaging the tori around AGN
- Measure the geometric shape of stars
- Determine the size of stars
  - ➤ E.g. Cepheids to calibrate the period-luminosity relation
- Star formation
  - Debris disks, chemistry in circumstellar disks
- Measure the structure of the Milky Way
  - Local spiral arm
  - ➤ Bulge, disk and halo, run-away stars
- Solar System objects
  - Comets, asteroids, weather on Titan



#### **ESO Publication Statistics**



JENAM 2010, Lisbon, 7 September 2010



## Richest planetary system

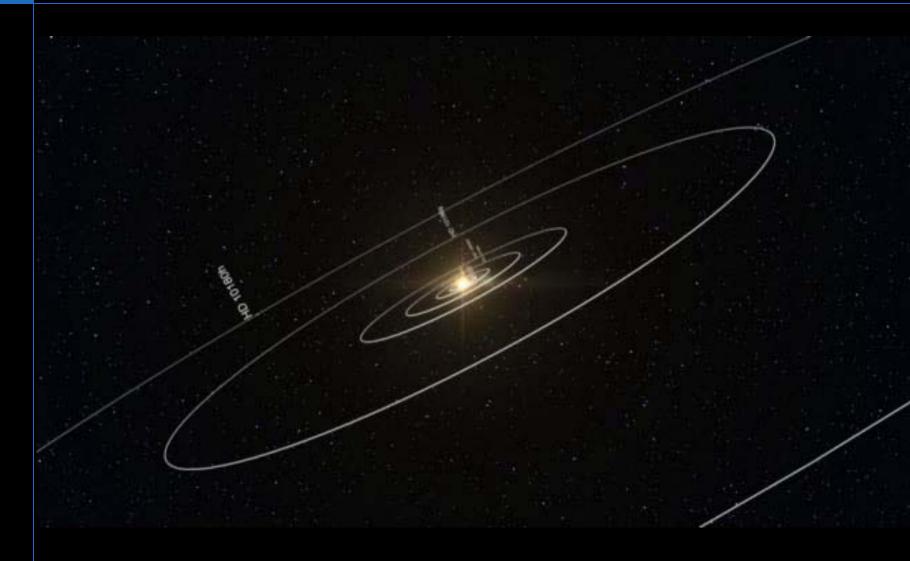
- Five planets with  $12M_{\oplus} < M < 25M_{\oplus}$ 0.06AU < D < 1.4AU
- One candidate with M=1.4M<sub>⊕</sub> at D=0.02AU and another with M=65M<sub>⊕</sub> at 3.4AU
- HARPS detections

HD 40307 GJ 581 HD 69830 HD 115617 HD 10180 HD 181433 46 23 28 Solar System HD 37124 GJ 876 HD 160691 HIP 14810 Ups And HD 74156 HD 125612 Semi-major axis [AU]

Lovis et al. 2010



## HD10180





## The ESO exo-planet machinery

HARPS at 3.6m telescope

best radial velocity machine at a 4m telescope (supported by UVES on VLT)

extremely stable spectrograph

#### NACO

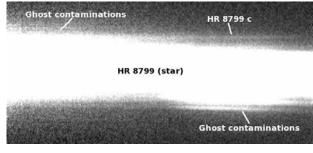
adaptive optics supported imaging and spectroscopy

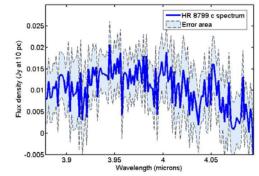
#### VLTI

highest spatial resolution for followup observations of known systems

#### NACO/SINFONI/FORS2

> transit measurements, atmospheres of exo-planets



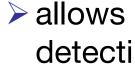


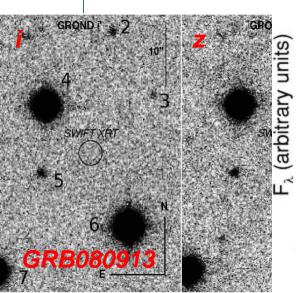
Janson et al. 2010

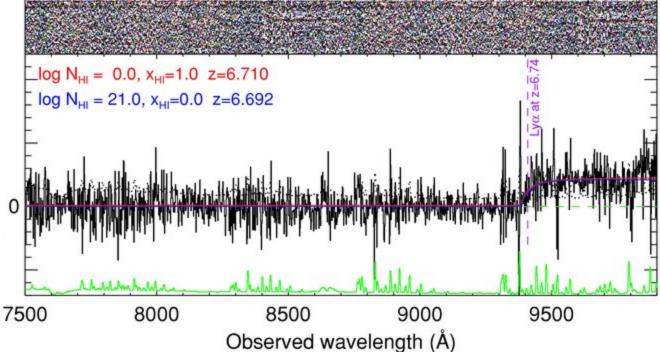


## **Gamma-Ray Bursts**

- Most distant stellar objects ever observed
  - redshifts 6.7 and 8.2 (tentative)
  - ➤ lookback time of nearly 12.5 billion years (or 95% of the age of the universe)
- VLT equipped with rapid response mode



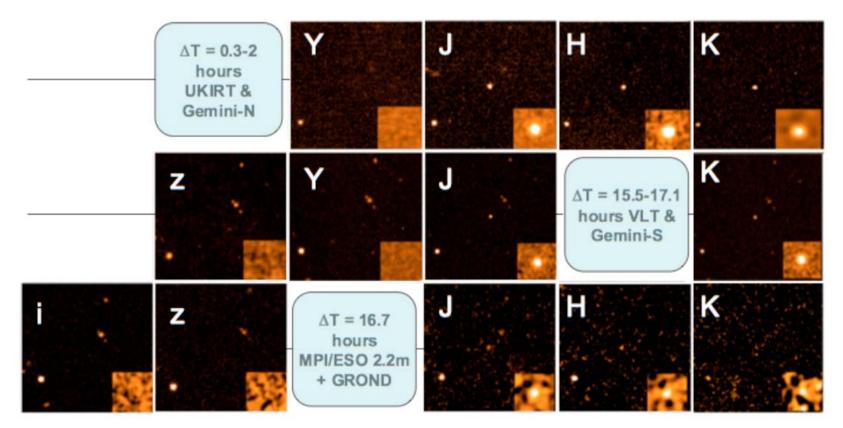






## Most distant stellar object yet observed – GRB 090423

- Optical drop-out, bright in the near-infrared
- Rapid decline

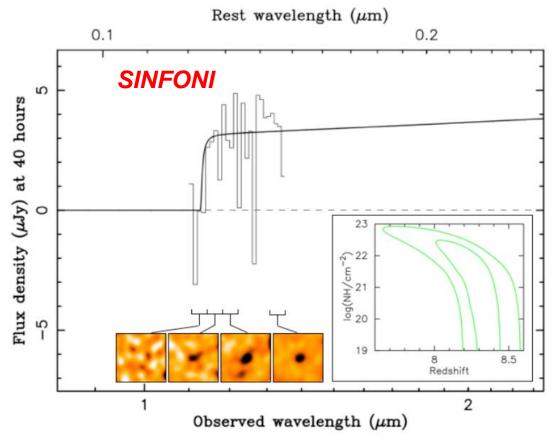


Tanvir et al. 2009



## GRB 090423

- Spectroscopy 17 hours after outburst
- Lyman break indicates a redshift of z≈8.2





## Chajnantor

#### APEX

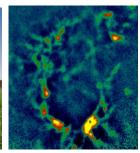
- ➤ 12m sub-millimeter antenna, operated by ESO @ Sequitor
- MPG (50%), Sweden (23%) and ESO (27%)

#### ALMA

- > 66 antennas at 5050m
- Operations support at 2950m
- Global partnership with North America
   East Asia & Chile











## Chajnantor

- Three facility and three `PI` instruments on APEX
- Watch out for ALMA
  - > early science in 2011
  - be prepared



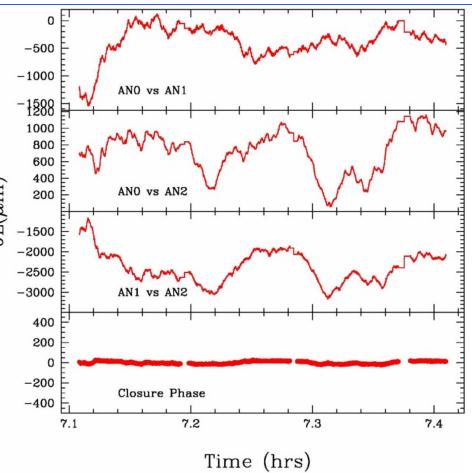






## Closure phase





Commissioning and Science Verification started on 22 Jan 2010





## **ALMA Main Deliverables**

IPT /	Europe	North America	East Asia
Subproject	_		
Site	ALMA Road to AOS	Roads at AOS	Financial Contribution
Development	OSF Technical Facilities	AOS Technical	to Costs
	OSF Residencia	Facilities	
	Santiago Central Office		
	ALMA Power Supply		
	192 Antenna Foundations		
Antennas	25 Antennas (12m)	25 Antennas (12m)	4 Antennas (12m)
	Two Antenna Transporters		12 Antennas (7m)
Front End	Band 7 (70)	Band 3	Band 4
	Band 9 (70)	Band 6	Band 8
	Cryostats (70)	Monitor and Control	Band 10
	Water Vapour Radiom. (53)	26 Integrated Receivers	Integrated Receivers
	Calibration System (70)		
	26 Integrated Receivers		
Back End	Photomixers (hundreds)	Optical Transmitters	
	Digitizers	Local Oscillators	
	Digitizer Clocks	Digital Formatters	
	Multiplexers	Digital De-Formatters	
Correlator	Tunable Filter Boards	Correlator for 64 Ant	Correlator for 16 Ant
Computing	Software	Software	Software

ESO deliverables correspond to ~460 M€



## **ALMA**

#### Progress

- ➤ Nearly all European deliverables on track
- Closure phase with three antennas at the AOS
- Santiago Central Office building completed JAO moved in two weeks ago
- Multi-fuel turbine being procured
- > First four European antennas mechanically integrated
  - two more in assembly

#### Concern

> Antenna delivery schedule: under close scrutiny



## **AEM Antenna #1**





## **AEM Antenna #2**





## European antennas





## **ALMA Early Science**

- 16 antennas with four frequency bands
- Baselines up to 1 km
- Up to 1/3 of the time used for this

# Learn more about ALMA this afternoon in the special session

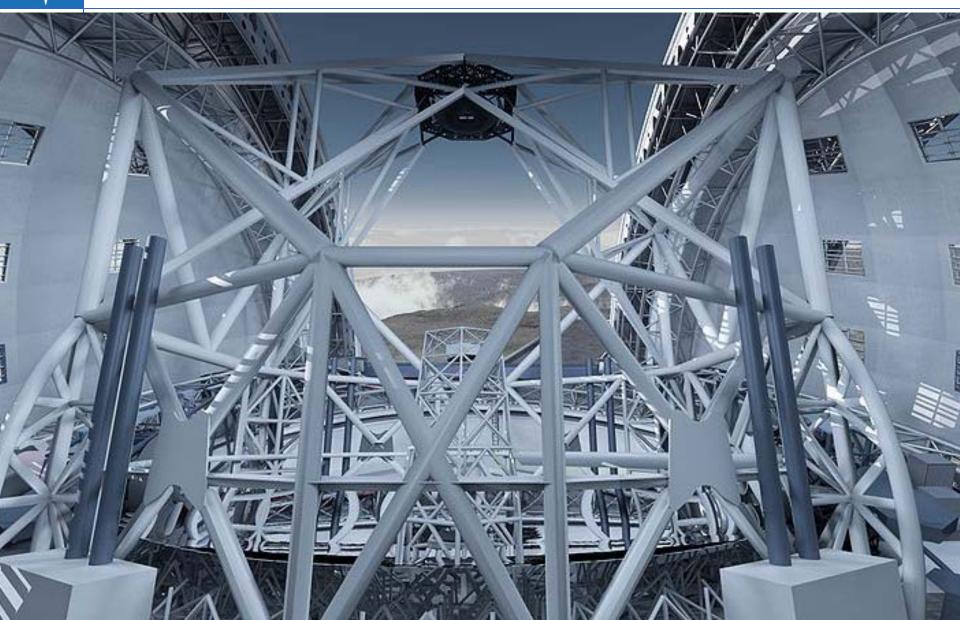
ALMA Band	Frequency Range (GHz)	Wavelength range (mm)
3	84-116	3.6-2.6
6	211-275	1.4-1.1
7	275-373	1.1-0.8
9	602-720	0.5-0.4

## **ALMA 2013**





## E-ELT





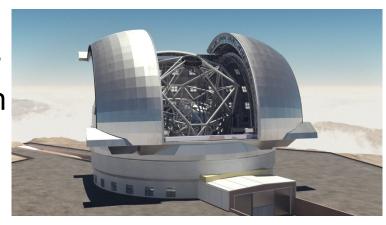
#### **E-ELT Design**

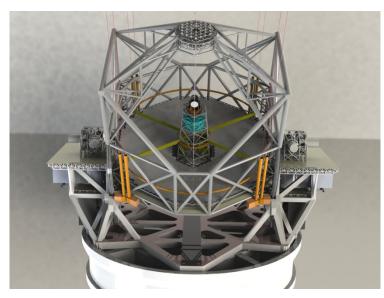
- Detailed design study
  - ➤ M1: ~1000 1.45m segments
  - ➤ Innovative five-mirror design⇒ adaptive telescope
  - Suite of instruments
  - Industry strongly involved
  - > Study cost: ~60 M€
  - Precursor work: ~40 M€, in part from EU programs

#### Status

- Construction proposal completed in Aug 2010
- External review Sep 2010

  JENAM 2010, Lisbon, 7 September 2010

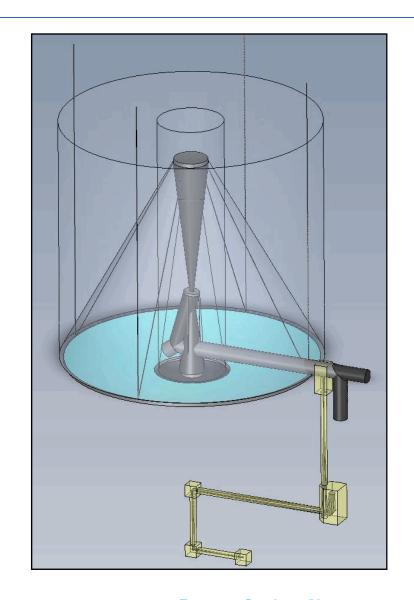






#### Telescope

- Novel 5 mirror concept
   (3 mirror anastigmat + 2 flats)
- Diffraction limited over full 10' field of view
- Flat, almost telecentric focal plane
- 'Zoom' capability
   (Nasmyth, gravity invariant, coudé foci)
- Laser "friendly" (very low aberrations even at zenith)





#### **Instrument Studies**

ACRONYM (P.I.)	INSTRUMENT TYPE
EAGLE	Wide Field, Multi IFU NIR Spectrograph with MOAO
EPICS	Planet Imager and Spectrograph with XAO
MICADO	Diffraction-limited NIR Camera- AO assisted
HARMONI	Single Field, Wide Band Spectrograph - AO assisted
CODEX	High Spectral Resolution, High Stability Visual Spectrograph
METIS	Mid Infrared Imager & Spectrograph -AO assisted
OPTIMOS	Wide Field , Visual, MOS (fibre or slit-based)- AO assisted?
SIMPLE	High Spectral Resolution NIR Spectrograph –AO assisted
	POST-FOCAL AO MODULES
MAORY	Multi Conjugate AO module (high Strehl, field up to 2')
ATLAS	Laser Tomography AO Module (high Strehl, narrow field)

STC recommendation on the first two instruments

- imager and integral-field spectrograph to be pursued first
- make sure AO is developed at the same time

JENAM 2010, Lisbon, 7 September 2010





#### **Site: Cerro Armazones**

- Selection by ESO Council 26-04-2010
  - Paranal property will be extended through a generous donation by the Chilean government
- Cerro Armazones
  - Best overall atmospheric quality
  - Highest science return
  - Lowest cost for construction and for operations
  - ➤ Optimal scientific synergy with VLT, ALMA, LSST, SKA
- Overall program
  - Strengthening of Paranal with another unique capability, ensuring its long-term scientific future
  - World-leading ground-based observatory



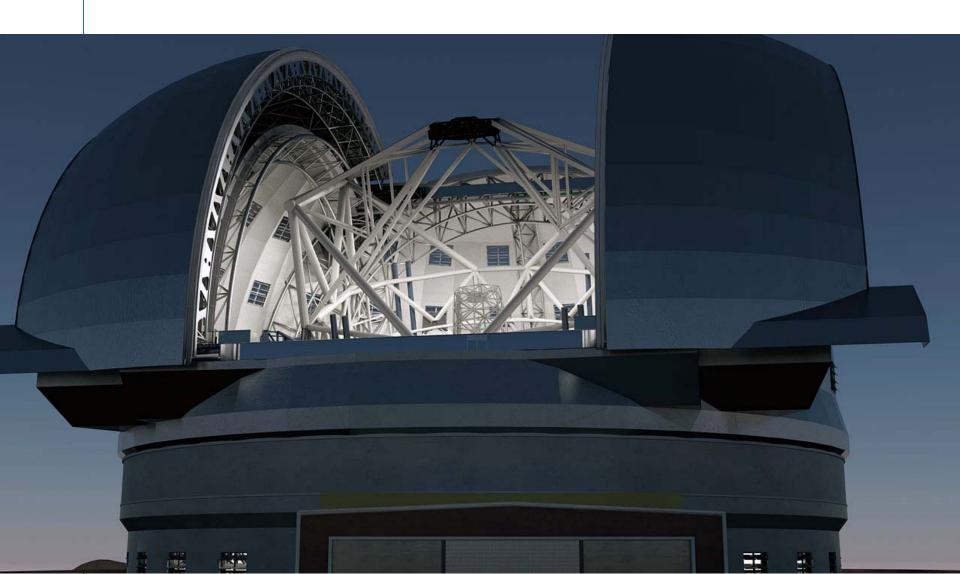


## **Armazones and Paranal**





# Time to turn of the lights and start observing





## **Proposing for ESO time**

Deadline for P87 proposals:30 September 2010



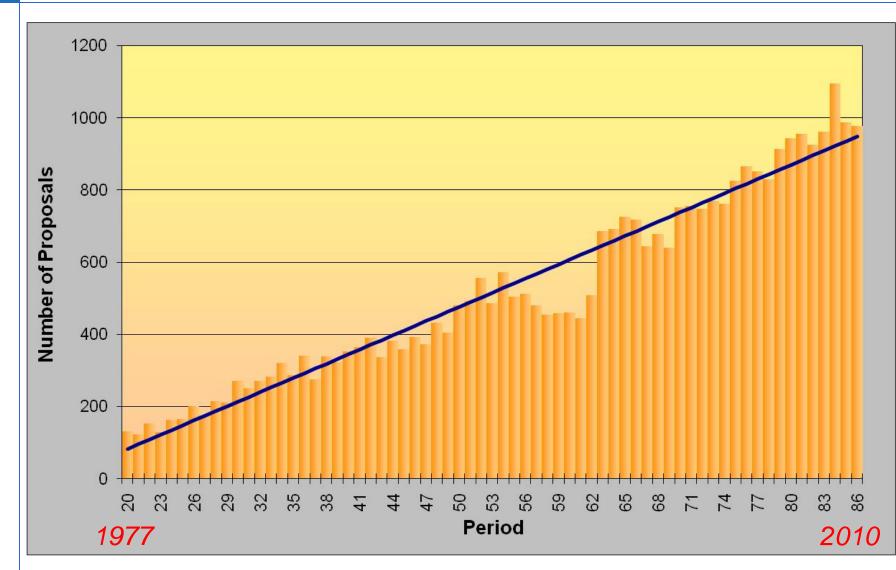


 $ESO\ Call\ for\ Proposals-P87$ 

Proposal Deadline: 30 September 2010, 12:00 noon CEST



## **Proposal submission**





## **OPC Working Group**

- Composed by current OPC members, representatives of other observatories and ESO
- Chair: Elias Brinks (Hertfordshire)
- Evaluation of the current OPC process
  - > efficiency in handling 1000 proposals twice per year
  - keep the load on panel members manageable
  - > keep the process at ESO manageable
- Investigated 'tweaks' and alternatives
  - > included some radical solutions
- Current process fairly good
  - any changes need to be accepted and trusted by the community

European Southern Observatory



## Public Spectroscopic Surveys

- Call for Letters of Intent issued
  - deadline: 15 October 2010
  - http://www.eso.org/sci/observing/proposals/Call for Public Spectroscopic Surveys 2010.pdf
- Pre-selection by a Public Spectroscopic Survey Panel
  - analogue to the public imaging surveys
- Proposals to be submitted for P88 (April 2011) for a start in October 2011
- no restriction on instrument or telescope



## Wide-Field Spectrograph Study

- Call for Letter of Interest issued
  - Deadline: 1 October 2010
  - http://www.eso.org/sci/announcements/MOS final combined.pdf
- Conceptual Design of a multi-object spectrographic facility
- no restriction on telescopes
  - change of UT top ends excluded
- Invitation for full proposal by end of October based on the received Letters of Interest
  - closing date for the proposals: 1 March 2011



#### **ESO Archive**

- The ESO data archive
  - > is a rich source of excellent data
  - > abstracts of previous proposals available
  - data public one year after they have been delivered to the PI
  - > great way to compete with your competitor, if they got observing time
  - > easy retrieval and selection of calibration data



## ESO's goals for coming years

- Best science from La Silla Paranal Observatory
  - Second generation instruments (VLT/VLTI)
  - > new instrument call
  - ➤ Key surveys with VST and VISTA
  - Public surveys (imaging and spectroscopy)
  - > Long-term programs for unique science on La Silla
  - Prepare for ALMA science with APEX
- Deliver ALMA on time and in budget
- Secure funding for construction and operations of the world-leading E-ELT





#### How others see us

#### US Decadal Survey:

- ➤ The 14-nation ESO consortium is on track to become the undisputed leader in ground-based OIR astronomy with its planned construction of the 42m European Extremely Large Telescope (E-ELT) ...
- ➤ By concentrating most of its resources into a single international partnership, Europe has minimized duplication of capability between facilities, created a major international research center, and established a funding line for construction that is intended to lead from ALMA, to E-ELT ...