

# Advanced Data Products

## A perspective

Eric Emsellem



# Long term Perspective

Science Discovery Toolkit Search  
Real-time Cosmic event detection system  
Multi-dimensional astroinformatics layer

*Radical transformation  
of our understanding of the cosmos*

**AI-driven astronomical data systems**

- ⇒ real-time and continuous data collecting, processing, interpreting
- ⇒ discovery alerts, suggests new theories



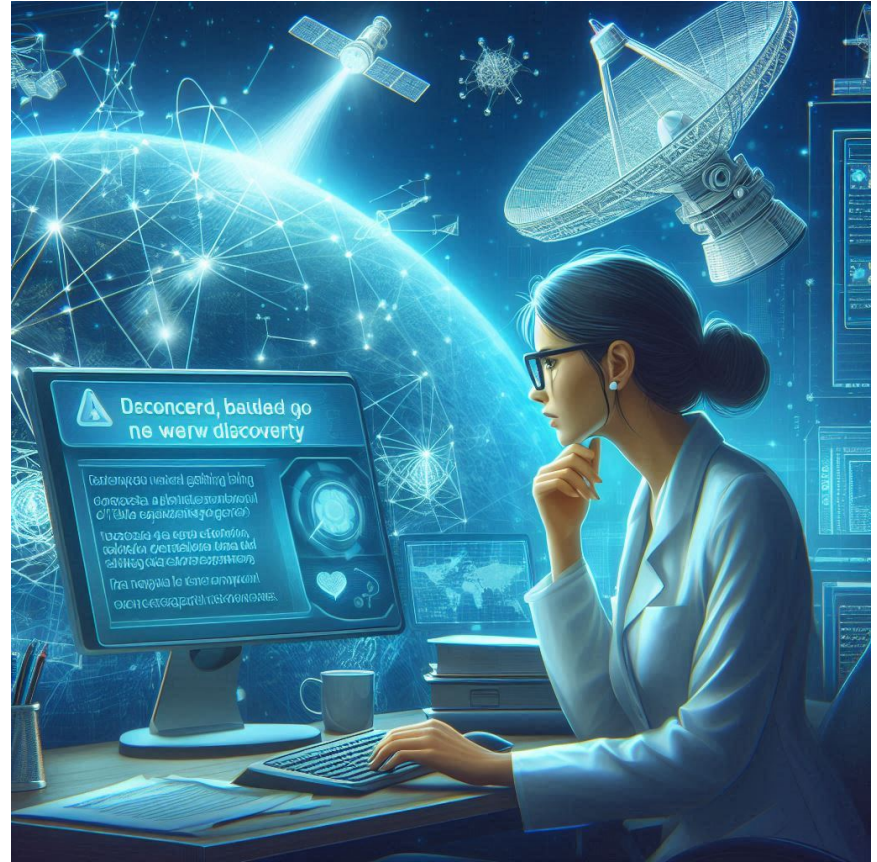
# Long term Perspective

Science Discovery Toolkit Search  
Real-time Cosmic event detection system  
Multi-dimensional astrophysics layer

*Radical transformation  
of our understanding of the cosmos*

**AI-driven astronomical data systems**

- ⇒ real-time and continuous data collecting, processing, interpreting
- ⇒ discovery alerts, suggests new theories



*Thanks ChatGPT but ...*

**AdP :  
what do  
you mean ?**

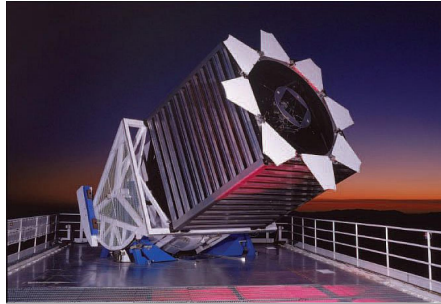
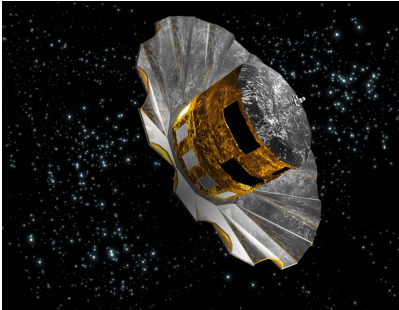
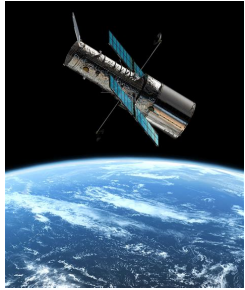


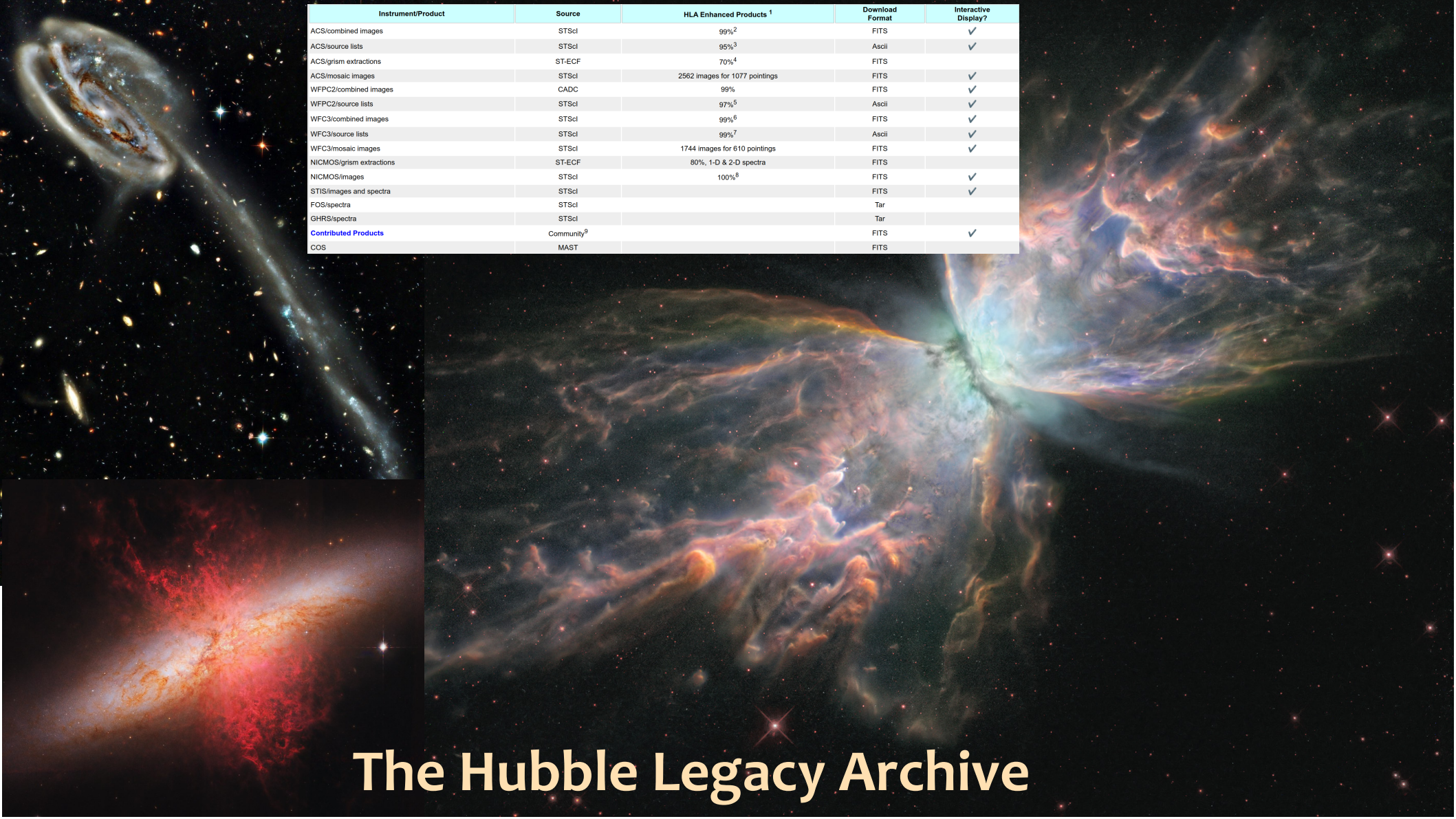
# Advanced Data Products – What do you mean ?

---

## Examples of Data Products

- **The Hubble Legacy Archive** : catalogs and images over 2 decades
- **Gaia Data Releases** : star catalogues, positions, distances, proper motions
- **SDSS Data Releases** : spectra but also redshifts, star formation history
- **CFHT Legacy Survey** : mosaicks of the sky
- **Gemini** : light curve of variable stars

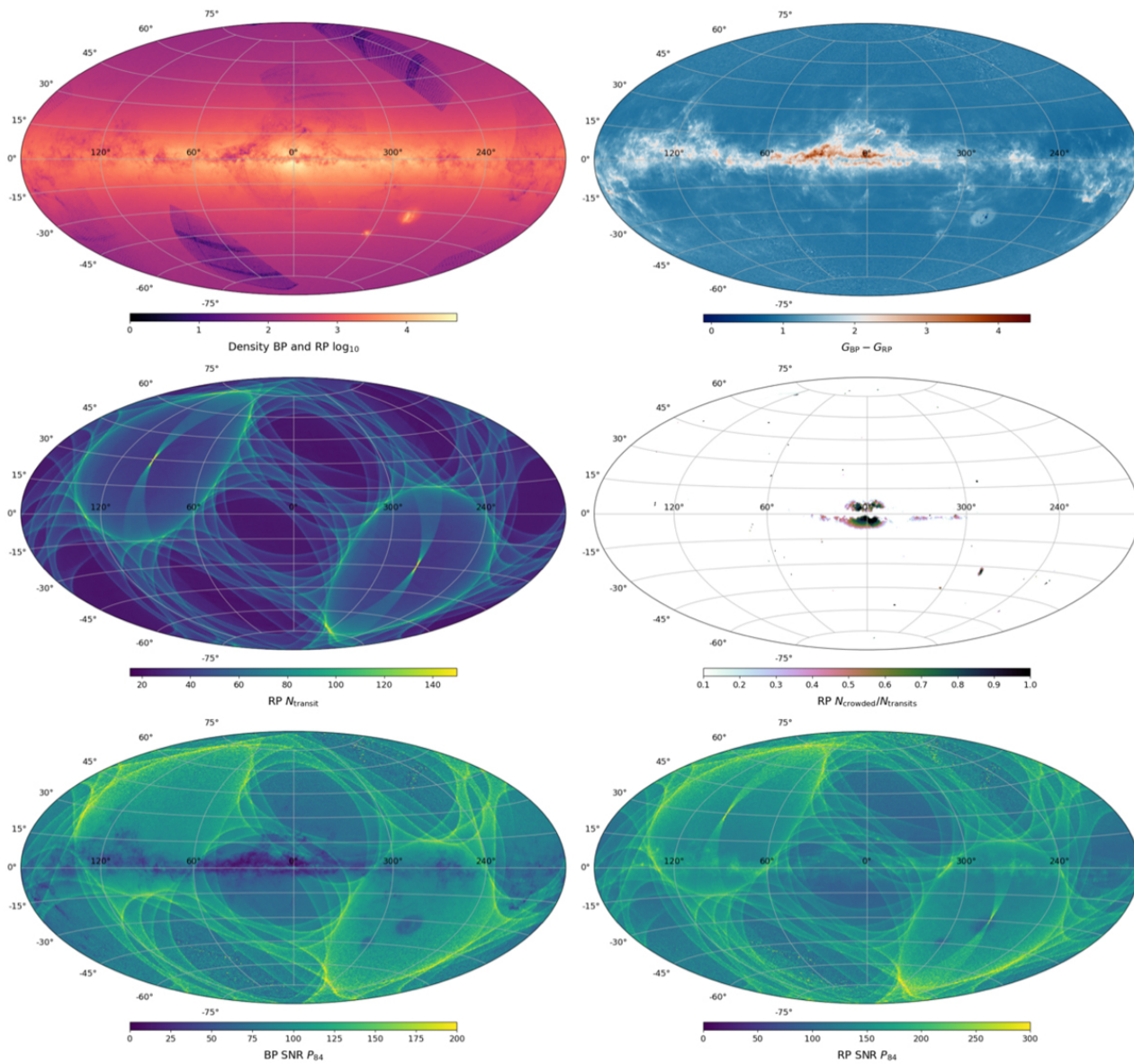




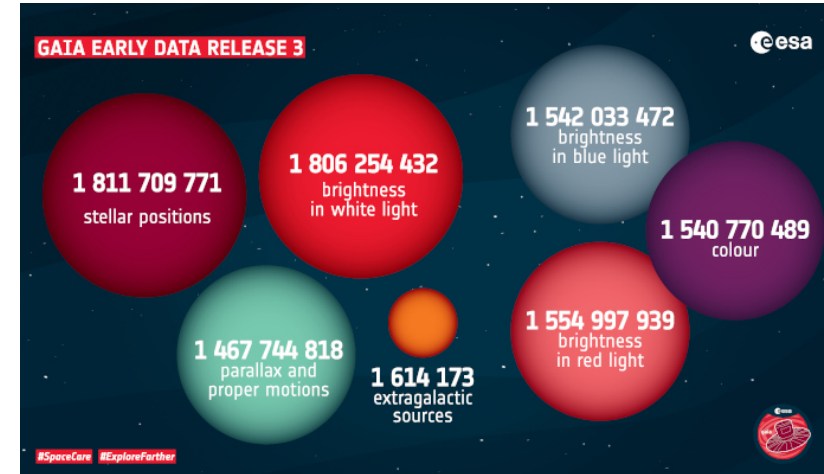
| Instrument/Product          | Source                 | HLA Enhanced Products <sup>1</sup> | Download Format | Interactive Display? |
|-----------------------------|------------------------|------------------------------------|-----------------|----------------------|
| ACS/combined images         | STScI                  | 99% <sup>2</sup>                   | FITS            | ✓                    |
| ACS/source lists            | STScI                  | 95% <sup>3</sup>                   | Ascii           | ✓                    |
| ACS/grism extractions       | ST-ECF                 | 70% <sup>4</sup>                   | FITS            |                      |
| ACS/mosaic images           | STScI                  | 2562 images for 1077 pointings     | FITS            | ✓                    |
| WFPC2/combined images       | CADC                   | 99%                                | FITS            | ✓                    |
| WFPC2/source lists          | STScI                  | 97% <sup>5</sup>                   | Ascii           | ✓                    |
| WFC3/combined images        | STScI                  | 99% <sup>6</sup>                   | FITS            | ✓                    |
| WFC3/source lists           | STScI                  | 99% <sup>7</sup>                   | Ascii           | ✓                    |
| WFC3/mosaic images          | STScI                  | 1744 images for 610 pointings      | FITS            | ✓                    |
| NICMOS/grism extractions    | ST-ECF                 | 80%, 1-D & 2-D spectra             | FITS            |                      |
| NICMOS/images               | STScI                  | 100% <sup>8</sup>                  | FITS            | ✓                    |
| STIS/images and spectra     | STScI                  |                                    | FITS            | ✓                    |
| FOS/spectra                 | STScI                  |                                    | Tar             |                      |
| GHRS/spectra                | STScI                  |                                    | Tar             |                      |
| <b>Contributed Products</b> | Community <sup>9</sup> |                                    | FITS            | ✓                    |
| COS                         | MAST                   |                                    | FITS            |                      |

# The Hubble Legacy Archive





# The Gaia Data Releases



NGC 4303



g r i (stars)



g r i (stars) + H $\alpha$  (ionised gas)



[OIII] H $\alpha$  [SII] (ionised gas)



g r i

$\Rightarrow$  Martino Romaniello's Talk

ESO Archive  
Phase 3



NGC 1087



g r i (stars)

g r i (stars) + H $\alpha$  (ionised gas)

NGC 3627



g r i (stars)

g r i (stars) + H $\alpha$  (ionised gas)

NGC 4254



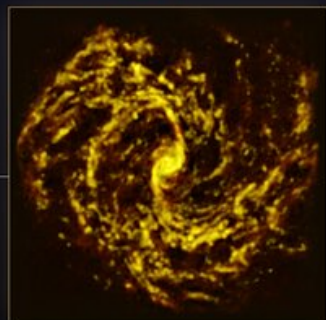
g r i (stars)

g r i (stars) + H $\alpha$  (ionised gas)

MUSE



ALMA



ALMA CO (molecular gas)



g r i (stars) + ALMA CO (molecular gas)



[OIII] H $\alpha$  [SII] (ionised gas)



g r i + [OIII] H $\alpha$  [SII] (ionised gas)



[OIII] H $\alpha$  [SII] (ionised gas)



g r i + [OIII] H $\alpha$  [SII] (ionised gas)



[OIII] H $\alpha$  [SII] (ionised gas)



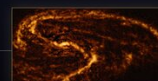
g r i + [OIII] H $\alpha$  [SII] (ionised gas)



ALMA CO (molecular gas)



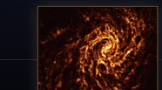
g r i (stars) + ALMA CO (molecular gas)



ALMA CO (molecular gas)



g r i (stars) + ALMA CO (molecular gas)



ALMA CO (molecular gas)



g r i (stars) + ALMA CO (molecular gas)



# Advanced Data Products – Levels @Geophysics

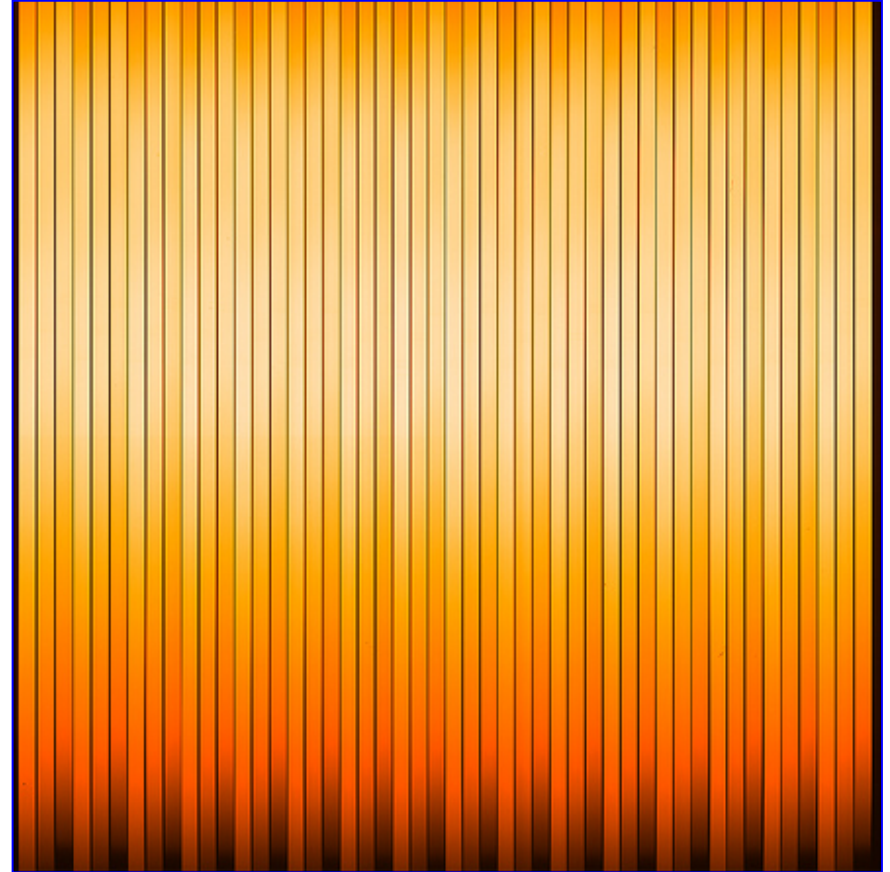
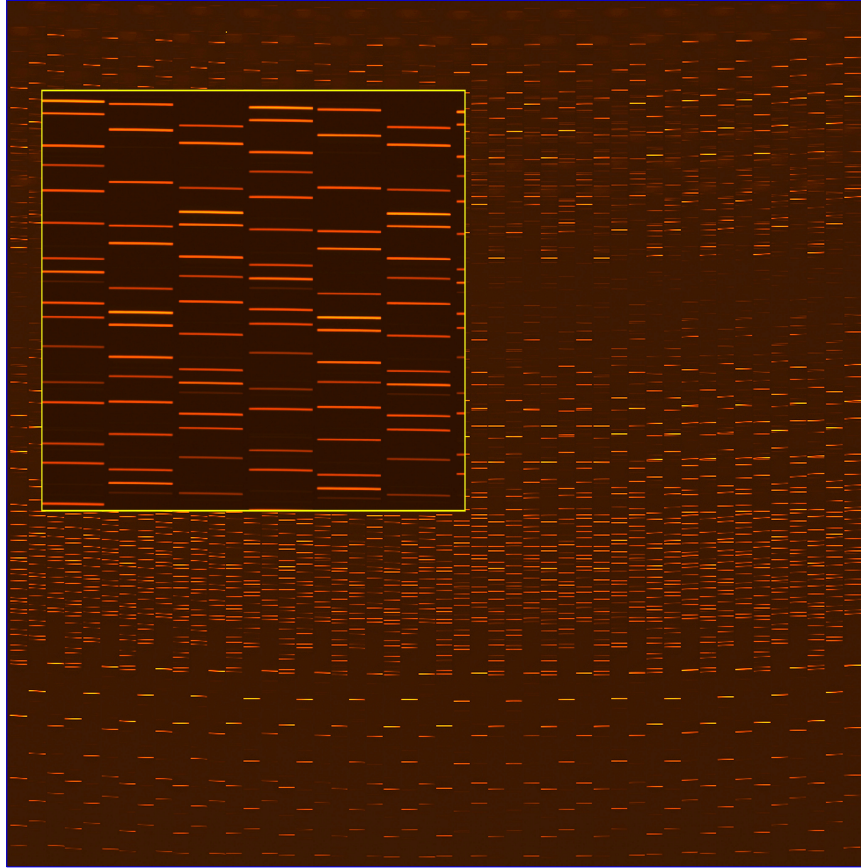
- 
- **Level 0** : Raw data – as measured
  - **Level 1** : processed data, in instrument/sensor units, with main instrumental signature removed (sometimes 1A, 1B, 1C...)
  - **Level 2** : Calibrated data at the same resolution as Level 1 (no loss)
  - **Level 3** : Variables on uniform space/time grids
  - **Level 4** : Model output or results of analysis of lower level data, variables derived from multiple measurements

**Information loss**

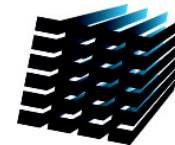


# Level 0

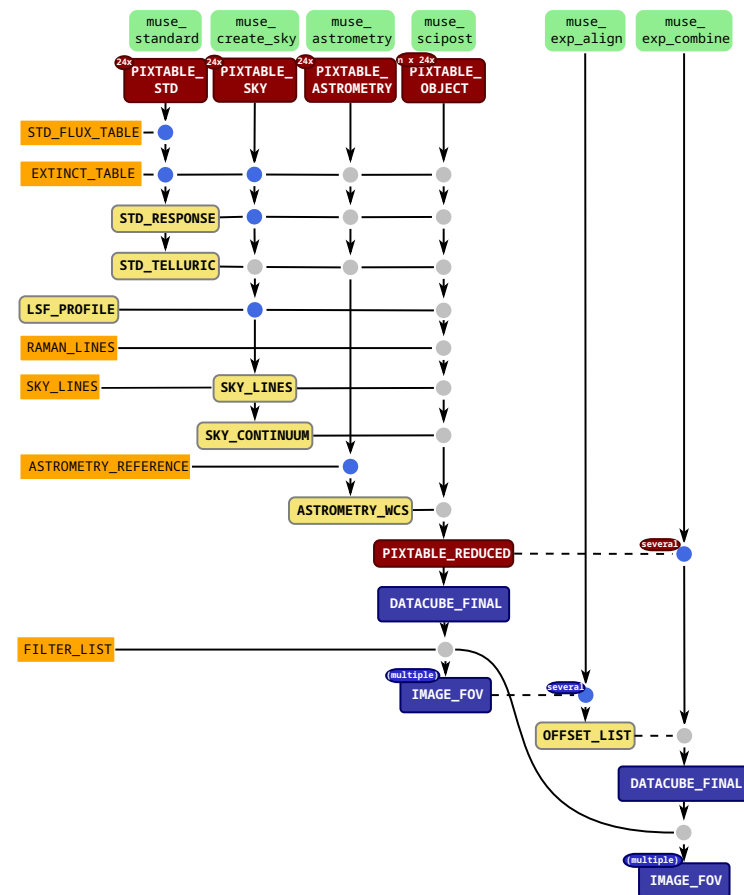
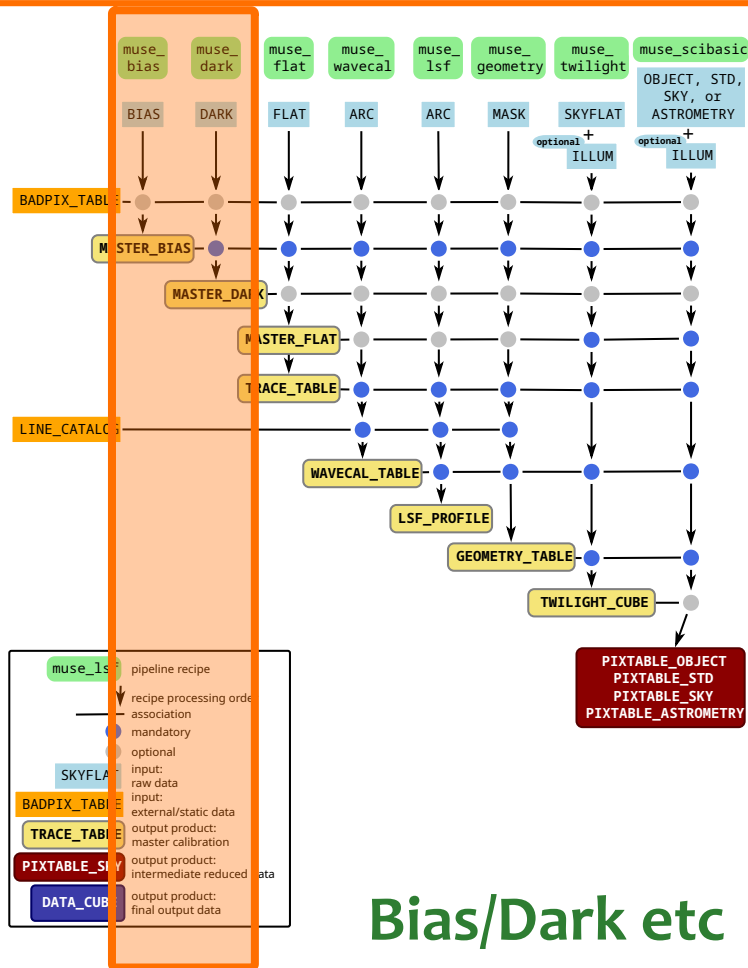
24x



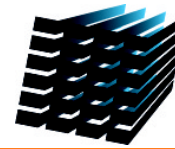
# ⇒ Level 1



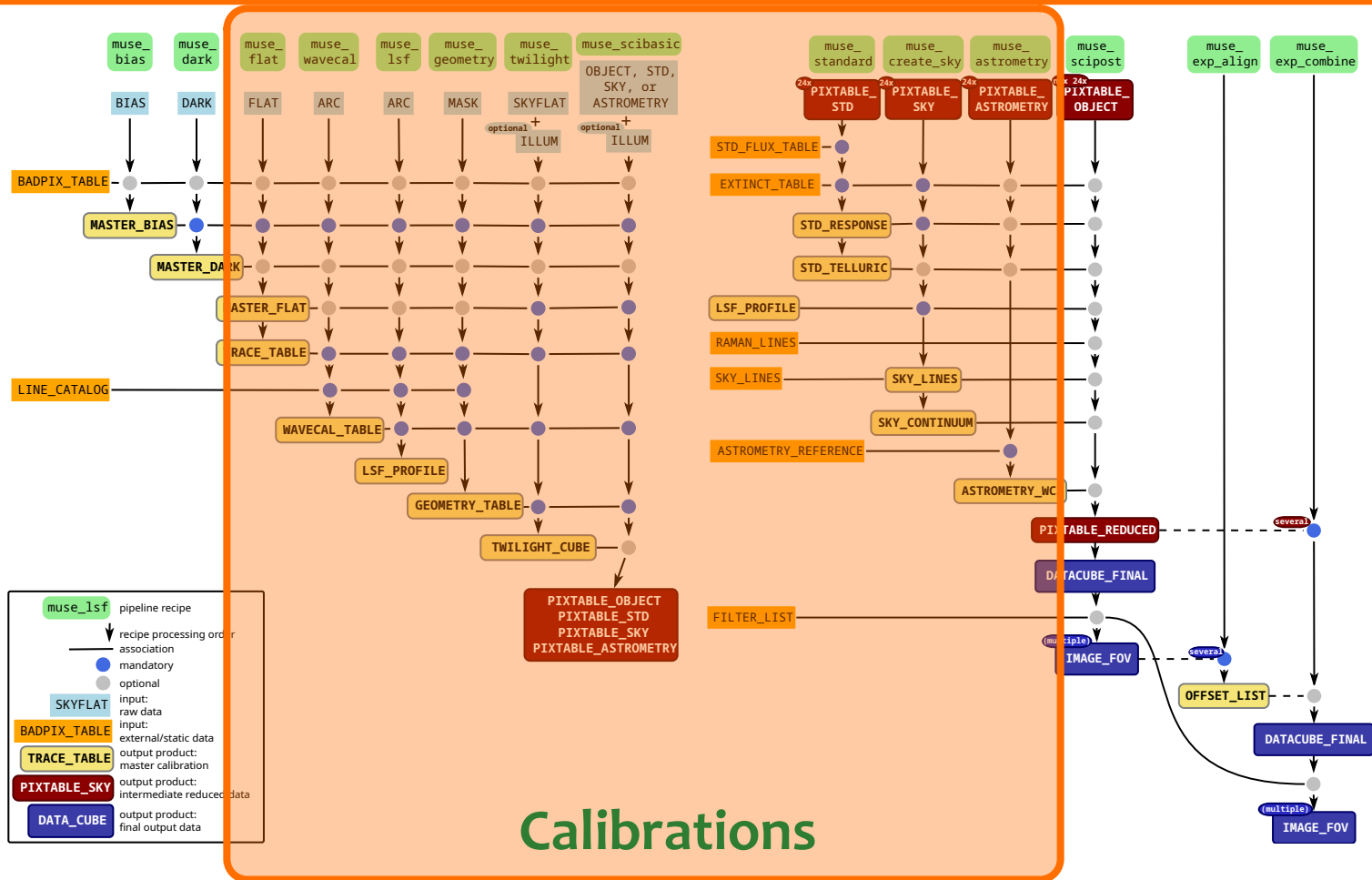
**MUSE**  
multi unit spectroscopic explorer



## ⇒ Level 2

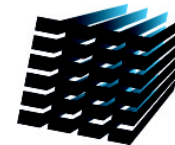


**MUSE**  
multi unit spectroscopic explorer

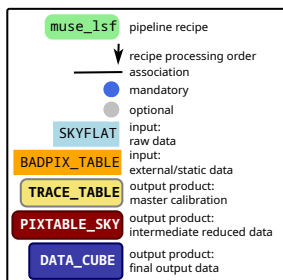
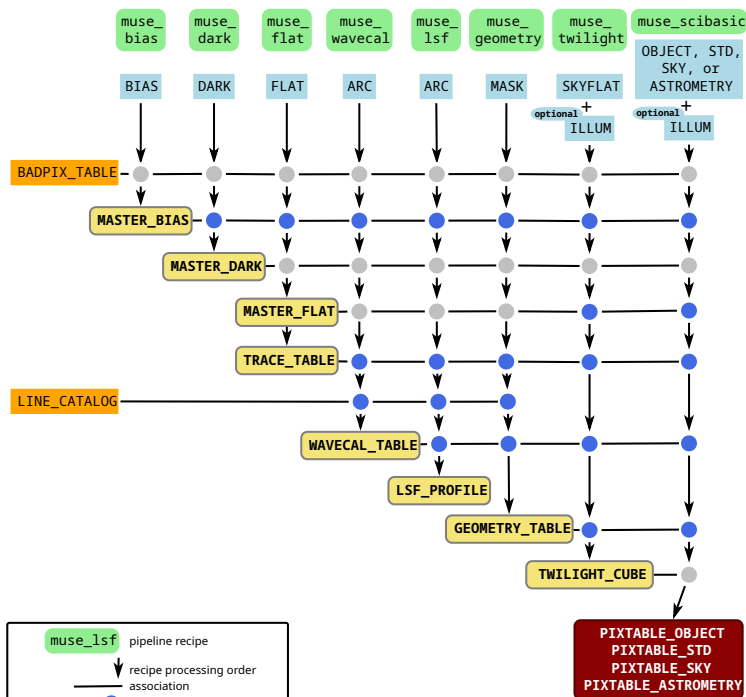




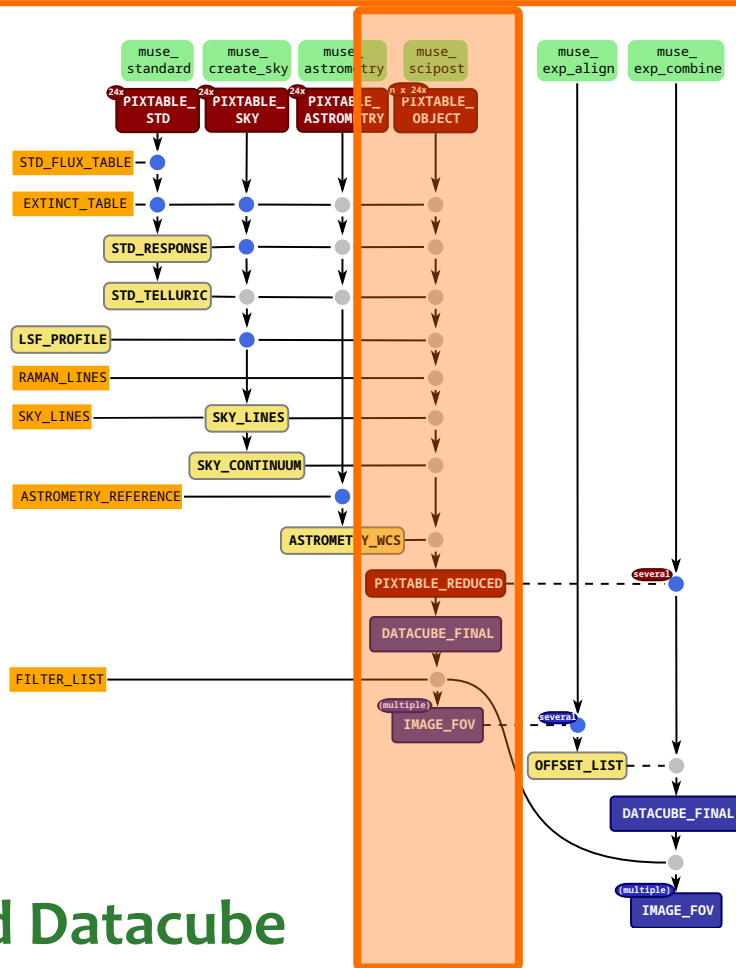
# ⇒ Level 3



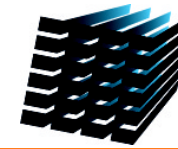
**MUSE**  
multi unit spectroscopic explorer



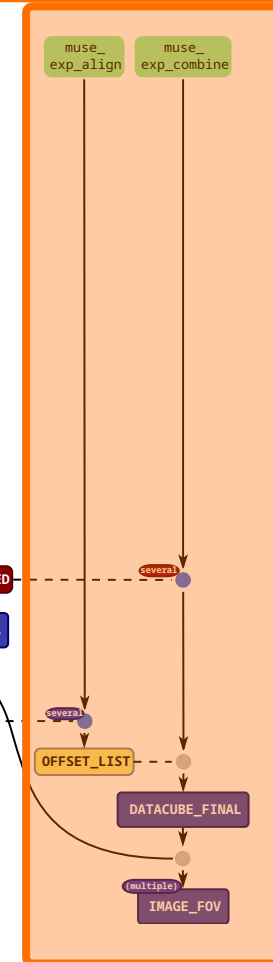
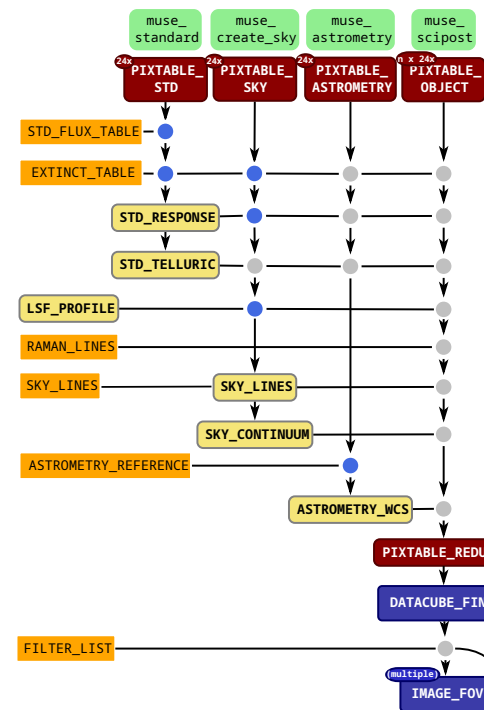
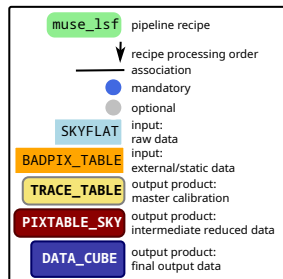
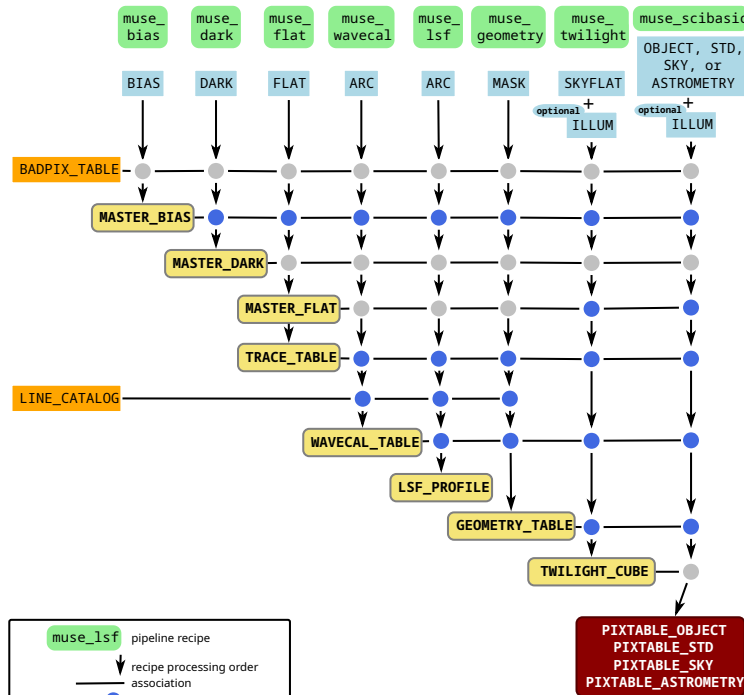
⇒ resampled Datacube



# ⇒ Level 3 / 4



**MUSE**  
multi unit spectroscopic explorer

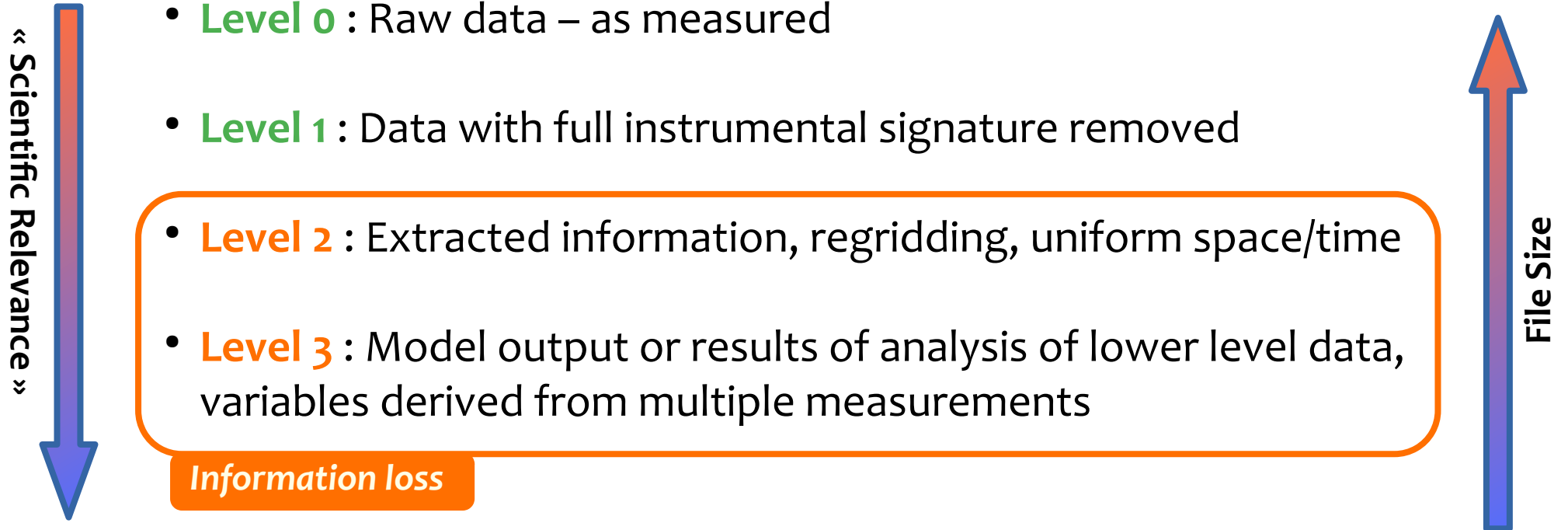


⇒ combined Datacube

(is that « Science-Ready »?)

# Advanced Data Products – Levels (update)

---





# An enhanced approach : from Level 0 to Level 1... to Level N

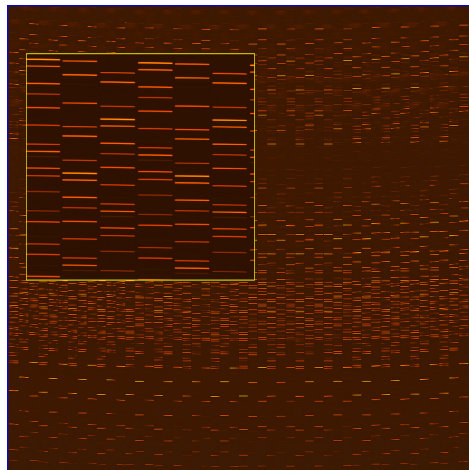
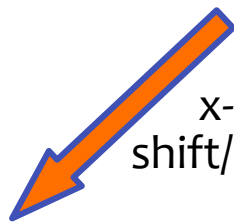


Table  
creation



| x        | y     | $\lambda$ | value   | flag | variance  |
|----------|-------|-----------|---------|------|-----------|
| -75.004  | 11.55 | 8408.7    | 87.72   | 0    | 19785.9   |
| -75.003  | 11.55 | 8410.0    | 10.68   | 0    | 19655.3   |
| -75.002  | 11.55 | 8411.2    | 43.55   | 0    | 20630.6   |
| -75.000  | 11.55 | 8412.5    | 85.31   | 5    | 21929.7   |
| -149.982 | 7.55  | 4212.9    | 1281.37 | 0    | 1321482.0 |
| -149.983 | 7.55  | 4214.2    | 482.40  | 0    | 166733.0  |
| -149.995 | 7.55  | 4215.4    | 394.27  | 0    | 110684.0  |
| -149.997 | 7.55  | 4216.8    | 967.98  | 0    | 740456.0  |

x-y  
shift/scale



| x        | y      | $\lambda$ | value   | flag | variance  |
|----------|--------|-----------|---------|------|-----------|
| -95.504  | 17.325 | 8408.7    | 87.72   | 0    | 19785.9   |
| -95.503  | 17.325 | 8410.0    | 10.68   | 0    | 19655.3   |
| -95.502  | 17.325 | 8411.2    | 43.55   | a0   | 20630.6   |
| -95.500  | 17.325 | 8412.5    | 85.31   | 5    | 21929.7   |
| -170.482 | 11.325 | 4212.9    | 1281.37 | 0    | 1321482.0 |
| -170.483 | 11.325 | 4214.2    | 482.40  | 0    | 166733.0  |
| -170.495 | 11.325 | 4215.4    | 394.27  | 0    | 110684.0  |
| -170.497 | 11.325 | 4216.8    | 967.98  | 0    | 740456.0  |

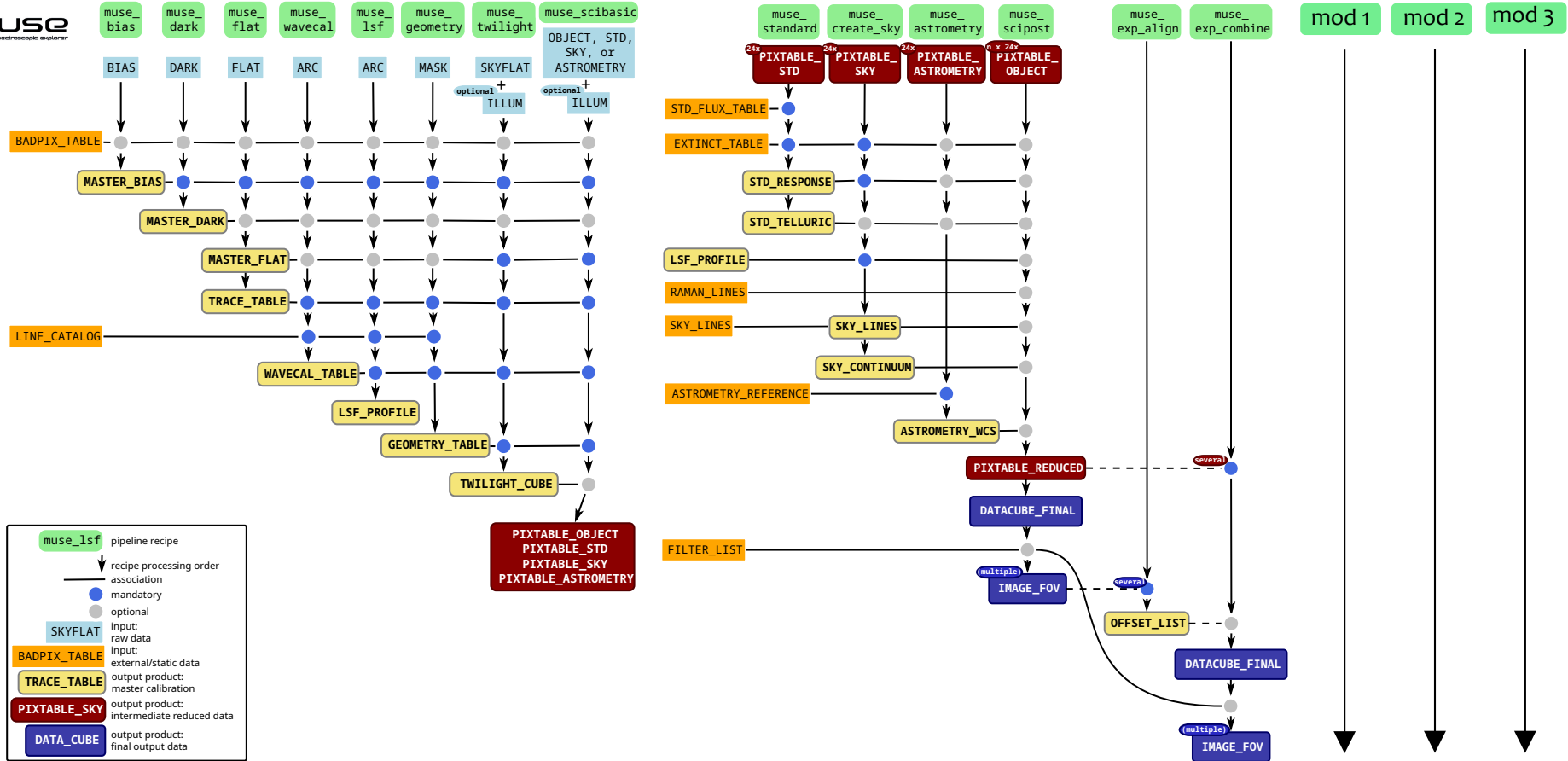
astrometric  
calibration



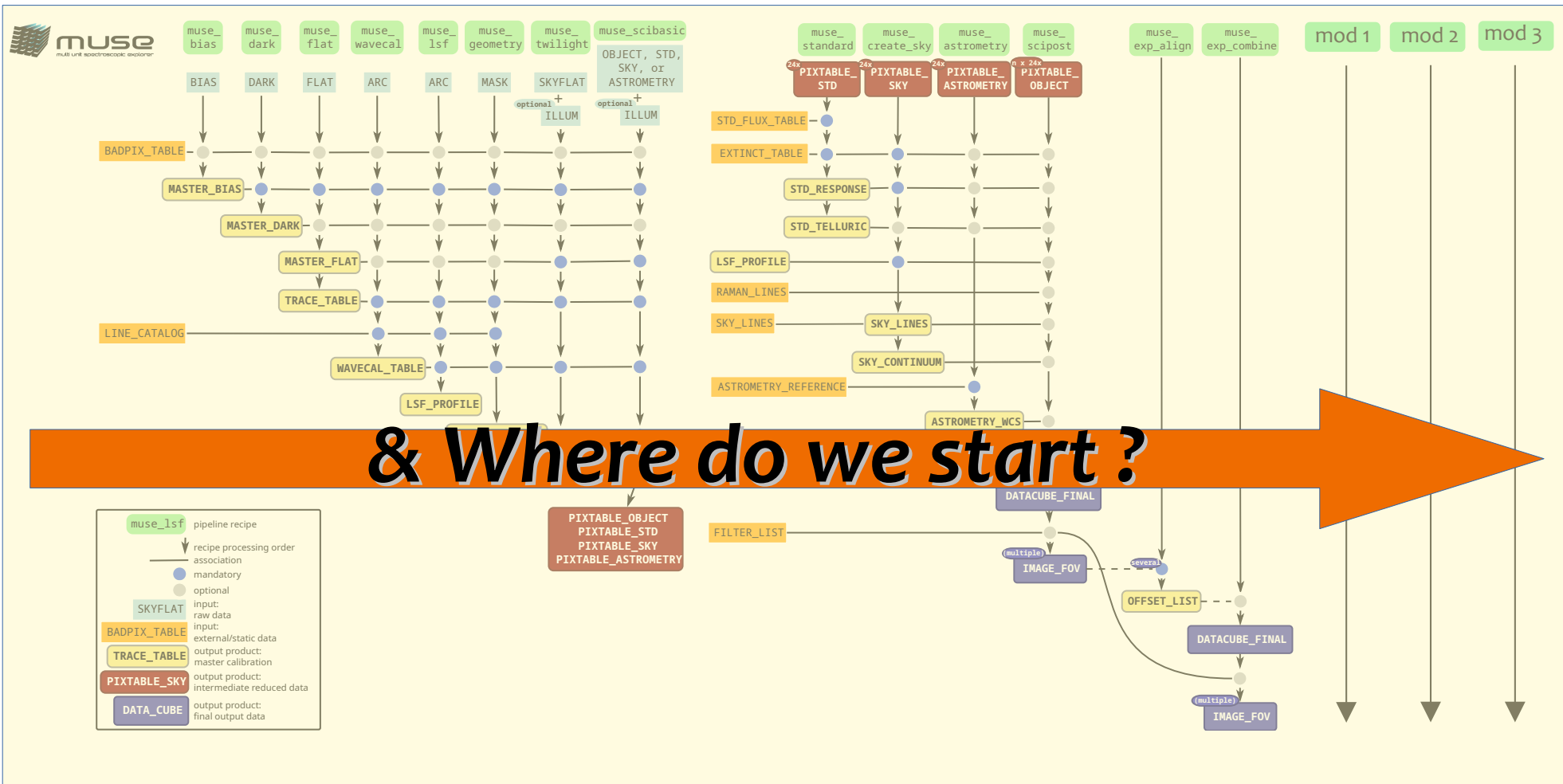
| RA       | DEC   | $\lambda$ | value   | flag | variance  |
|----------|-------|-----------|---------|------|-----------|
| -19.1008 | 3.465 | 8408.7    | 87.72   | 0    | 19785.9   |
| -19.1006 | 3.465 | 8410.0    | 10.68   | 0    | 19655.3   |
| -19.1004 | 3.465 | 8411.2    | 43.55   | 0    | 20630.6   |
| -19.1000 | 3.465 | 8412.5    | 85.31   | 5    | 21929.7   |
| -34.0964 | 2.265 | 4212.9    | 1281.37 | 0    | 1321482.0 |
| -34.0966 | 2.265 | 4214.2    | 482.40  | 0    | 166733.0  |
| -34.0990 | 2.265 | 4215.4    | 394.27  | 0    | 110684.0  |
| -34.0994 | 2.265 | 4216.8    | 967.98  | 0    | 740456.0  |

Change of coordinates  
No change to the values

# Are those levels relevant for science?

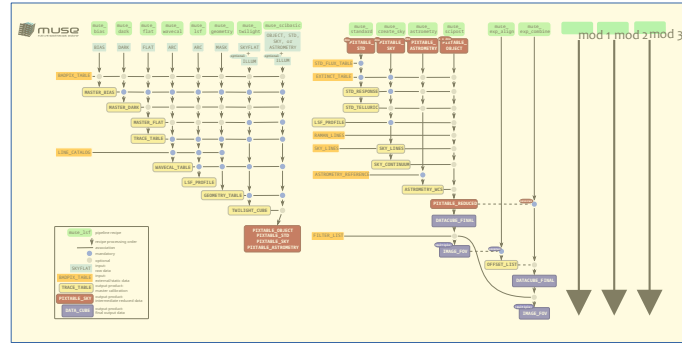


# ADPs ?



# ADPs : where do we stop ?

---



« Observations become scientific when used as evidence of phenomena »

Borgman 2015



**What are  
we trying  
to solve ?**

# Observational science :

Transform a set of observations into  
a **coherent picture**



## Objectives and Gains

- ⇒ **Optimise** and expand exploitation of facilities / data
- ⇒ **Lower** barriers
- ⇒ **Respond** to potential changes in our ecosystem (agile)
- ⇒ Use it as a **leverage for progress**

*[but probably **not** : to shorten the way to papers]*

**An  
evolving  
landscape**

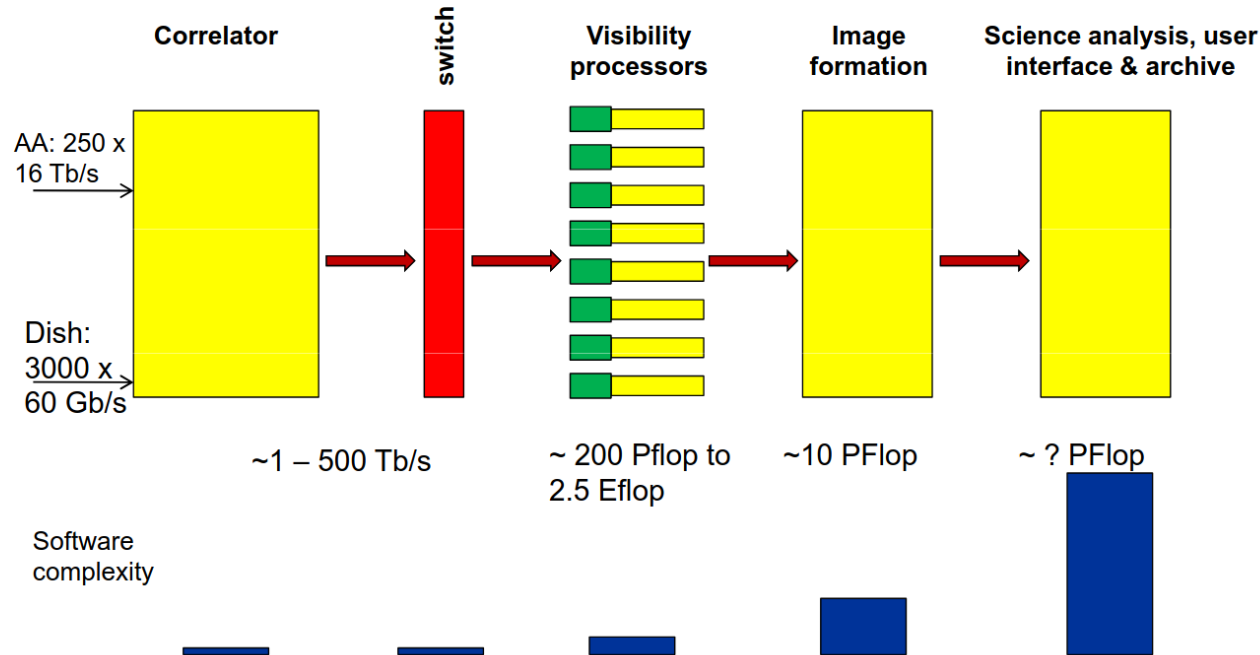
# Era of « Big » & « multi-messenger » Data

- **LSST / SKA** : petabytes or even exabytes during its mission  
(LSST = 20 Tb / night)
- **Analysis speed** : e.g., LIGO that requires processing **within minutes**
- **Richness** : multi-lambda / instruments / formats / dimensions
- **Complexity** : data sources, data formatting, nature of the information, need for forward modelling (and instrument simulators)
- **Accuracy** : LIGO, or ANDES/ELT, errors (e.g., covariance)
- **Cost & Efficiency of Processing** : SKA !

# The Scientific perspective

- Technical / Computing assessment

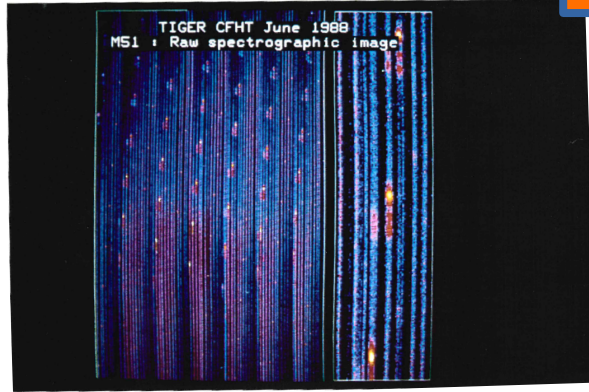
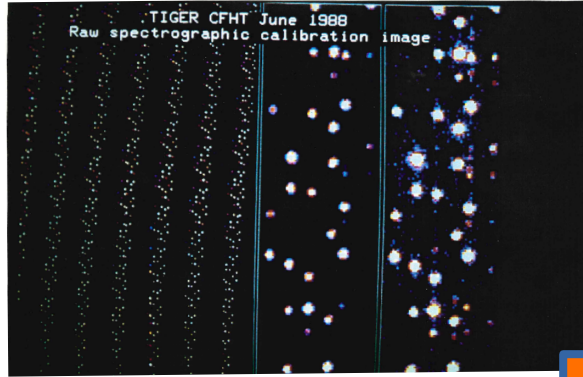
⇒ *what about the end product = scientific discovery ?*





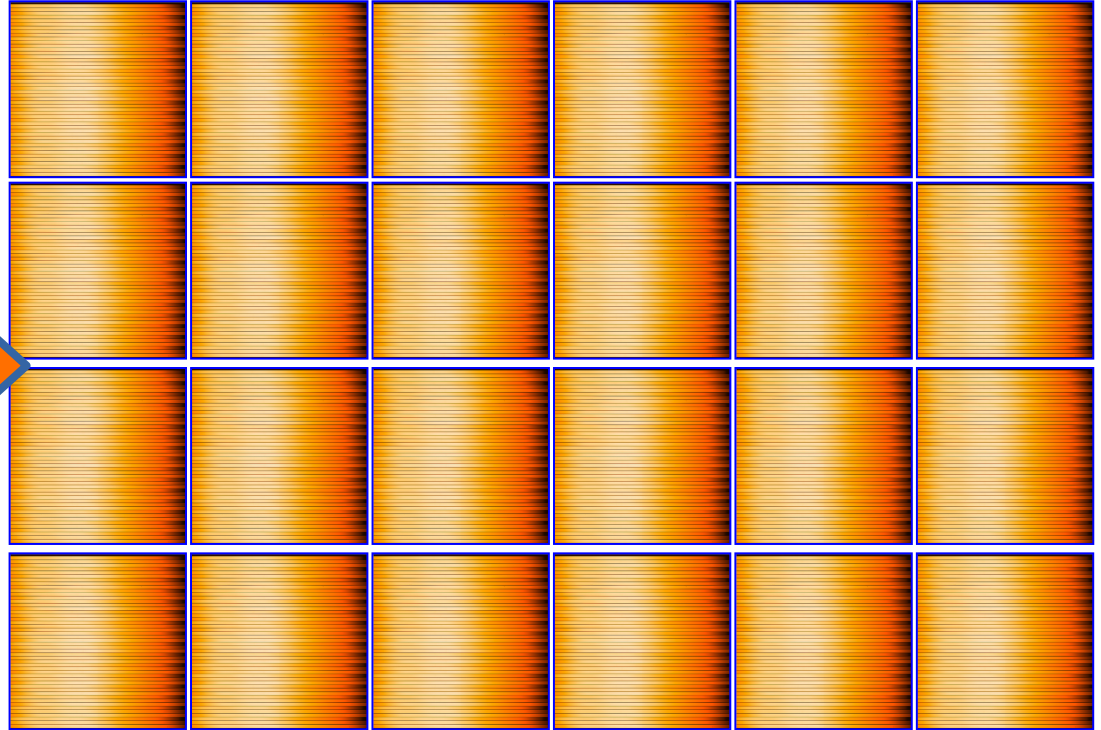
# Emerging from the Dark Ages

50 x 400



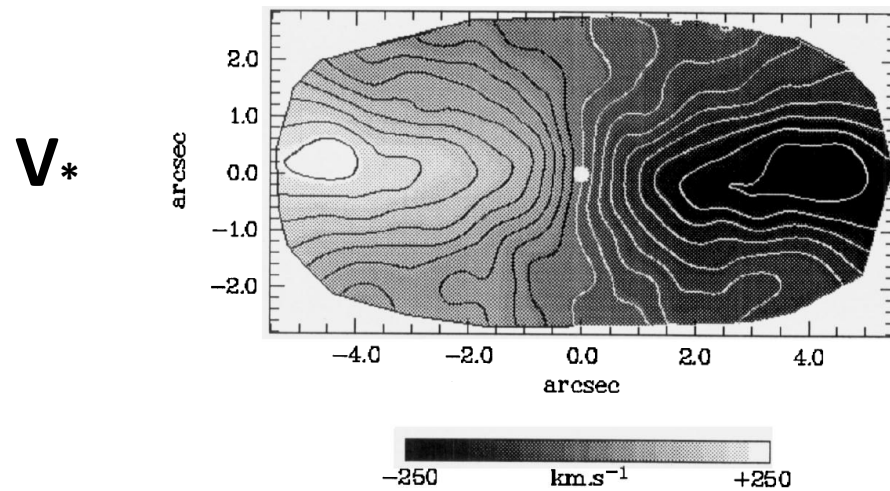
1988

90,000 x 4000



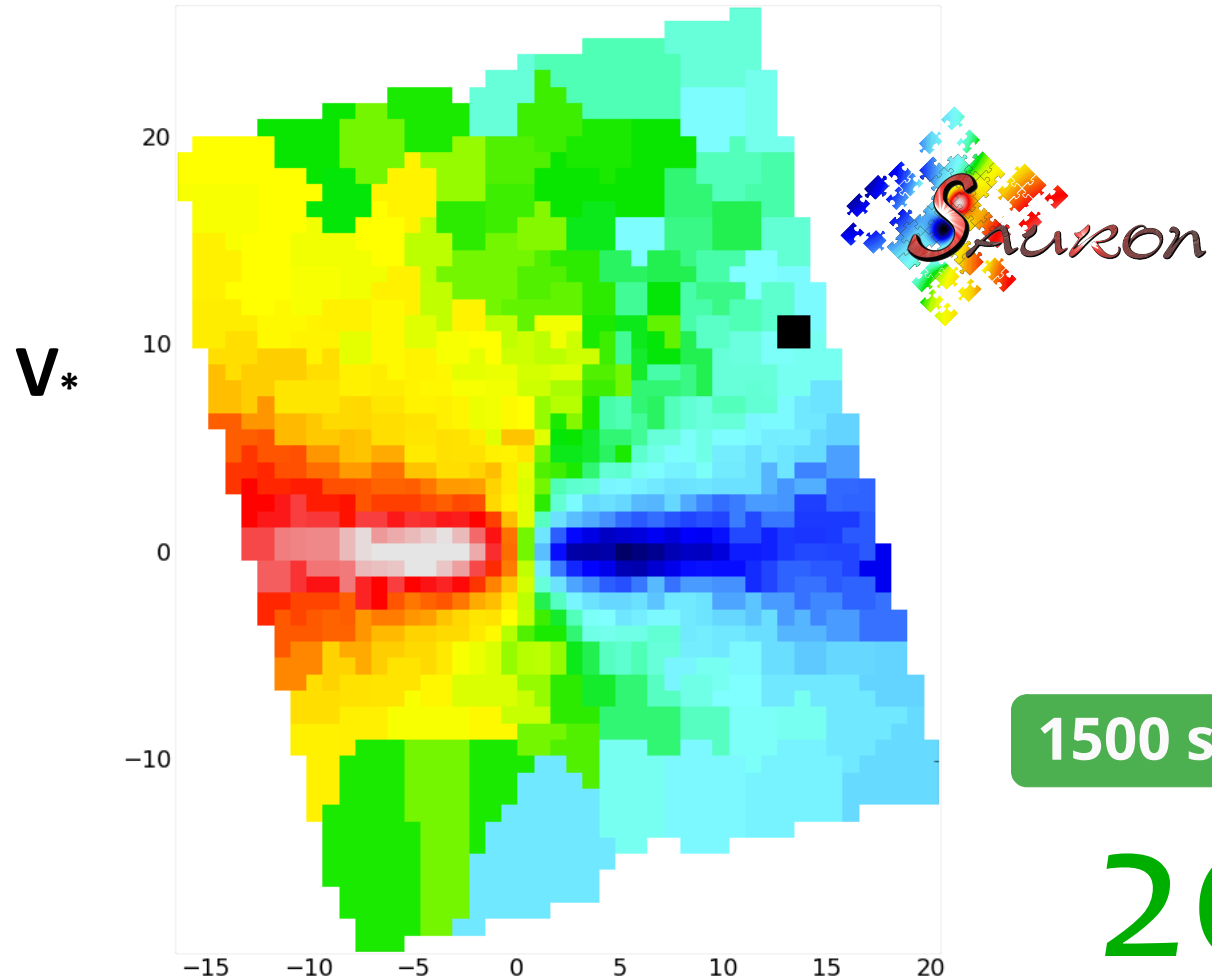
2014

# TIGER

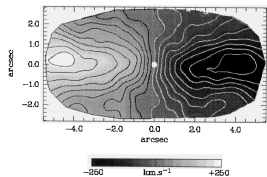


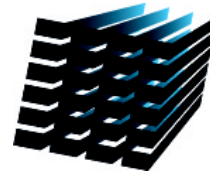
1996

400 spectra / 2 hours

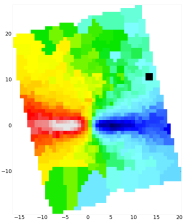


1996



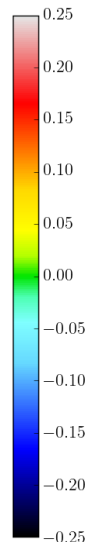
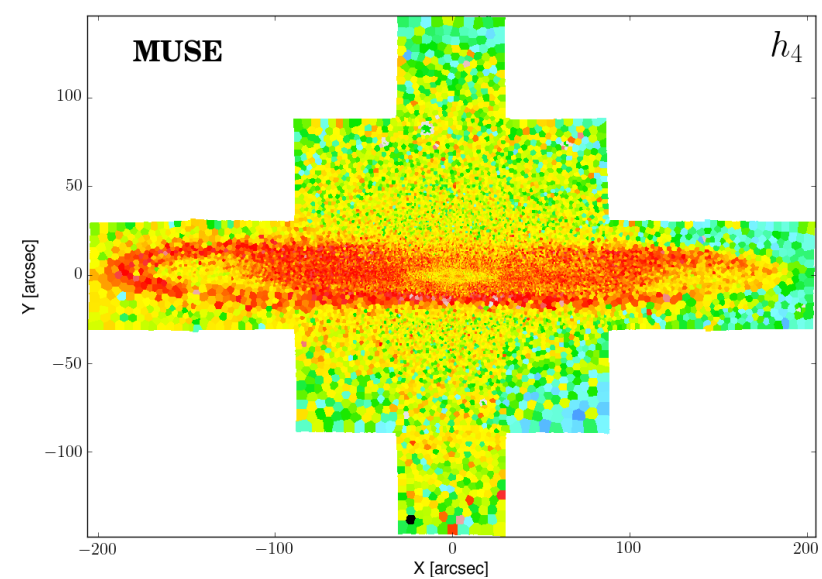
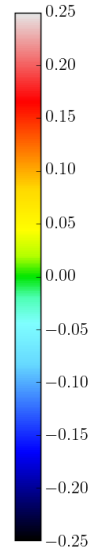
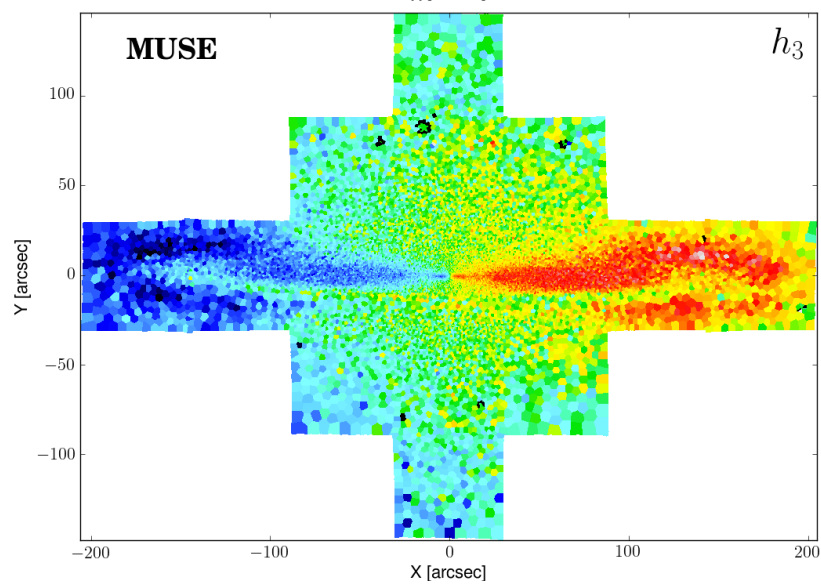
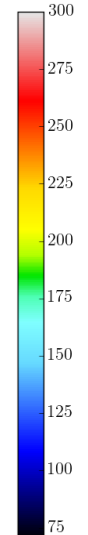
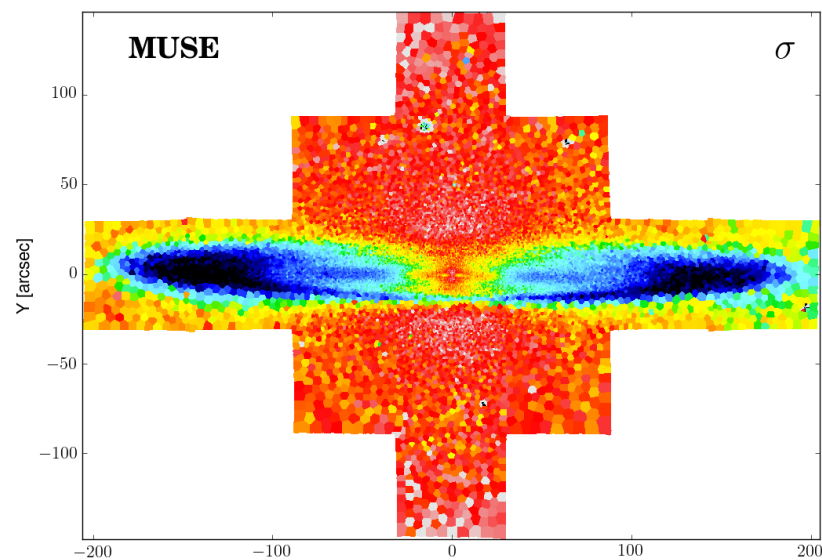
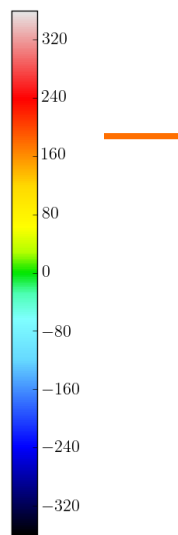
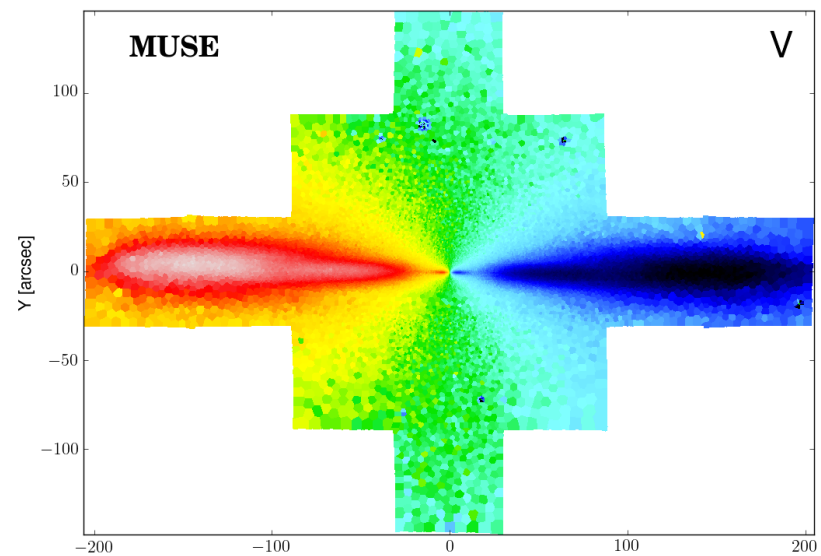


**MUSE**  
multi unit spectroscopic explorer



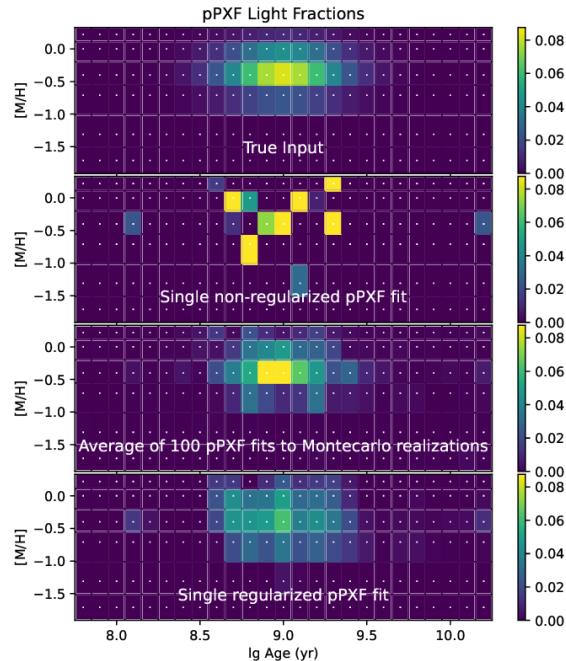
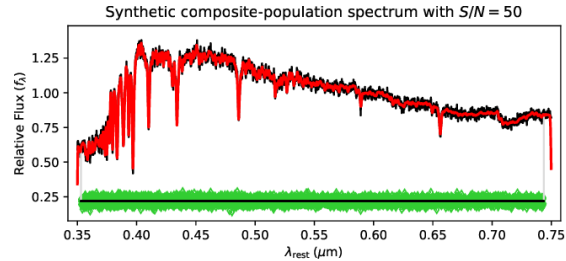
1.4  $10^6$  spectra, < 4h

2014



# What it entails

## Star formation history (Cappellari, 2022)



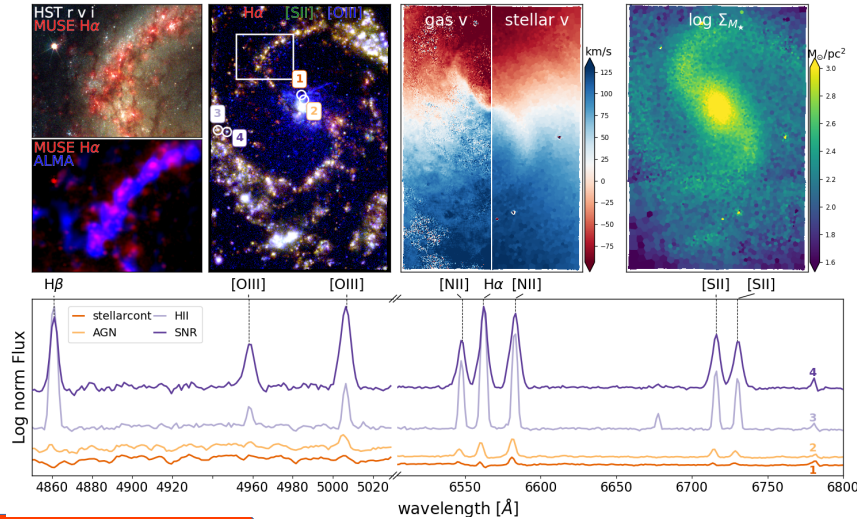
- Input spectrum and variance (& covariance)
- A complex code with algorithmic branches
- Many Input parameters :
  - Polynomial degrees, windows, potential emission lines, regularisation matrix
- Extension law
- Full stellar library
  - More spectra
  - More assumptions (Simple Stellar Population)



# Data do not work in isolation

- Data: catalogs, maps, key astrophysical parameters

ESO Archive  
Phase 3 data products



Emsellem et al. 2022

- Simulated Universe and predictive models

- Simulations & models go hand in hand with data
  - e.g., applied libraries, theoretical assumptions, numerical models and simulations
- (e.g., *MAST High-Level Science Products, HLSPs* ⇒ see Scott Fleming's talk)

# A technical paradigm change

---

- Technological developments
  - hardware and associated frameworks & infrastructures
  - Often software + hardware hand in hand (Supercomputers, GPUs)
- **Algorithmic** improvements : machine learning/AI, optimisation algorithms
- Automated Discovery (some of the fastest algorithmic improvements)
  - Automated classification, anomaly detection
  - Predictive models
  - Optimisation of telescope time

⇒ New partnerships with AI scientists

⇒ Importance of scrutinising its ethical impact
- New development/**communication** platforms and **networks**

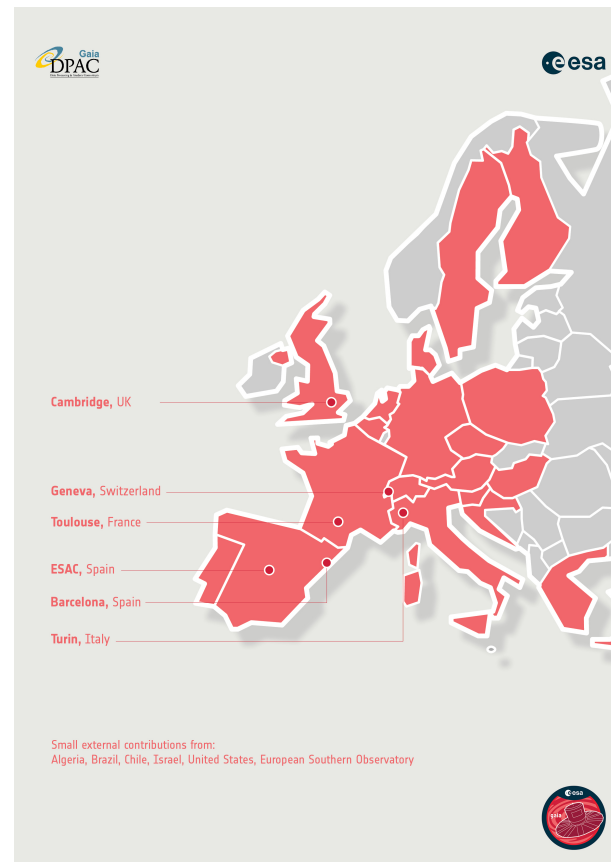
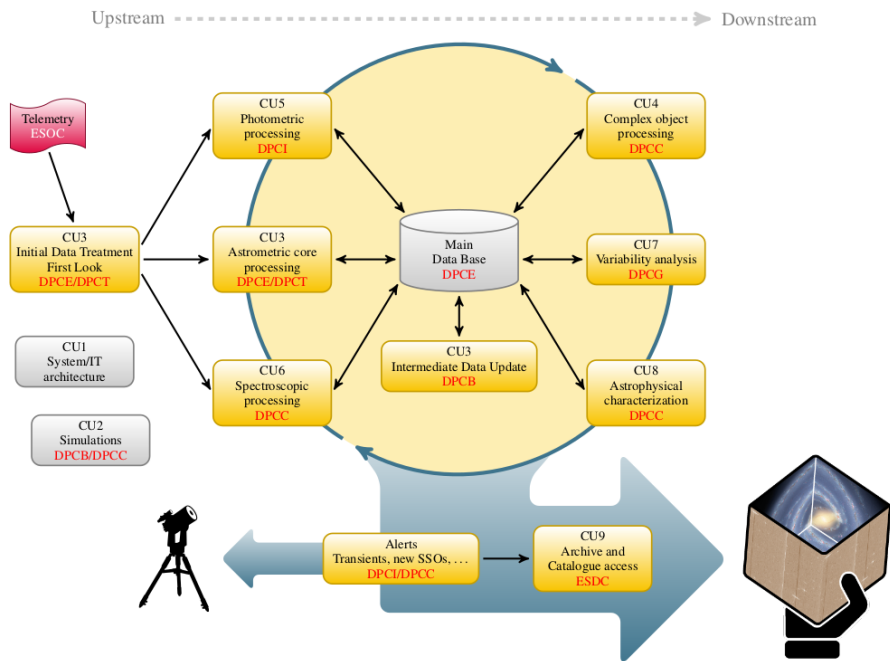


# An Era of collaborations and Interoperability

- Networks of observatories, institutes, organisations

## ⇒ Requires Standards

- Data formats
- Collaborative tools and protocols
- Portals, Platforms, Clouds



# A societal paradigm change

- **The next generation (?)**

- Those who will use SKA, the ELTs  
⇒ are barely coming out of high school



- **Public engagement**

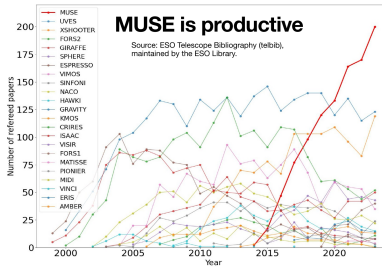
- Citizen science
- What about citizen developments ?



- **Public opinion**

- Everyone has an opinion, and they are visible, impactful
- Harder to discriminate levels of relevance

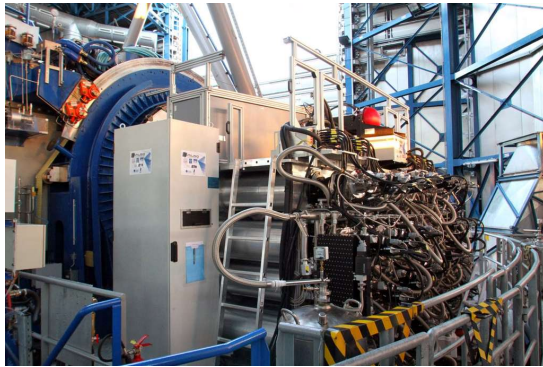
# Why is MUSE at the VLT so “successful” ?



**BTW :** what does this mean ?  
Which metric should we use ?  
[popularity, production, impact ?]

## ● Design and development approach

- Broad community service
- But restricted set of modes
- Robust pipeline : based on simulations, collaborative updates
- **3-ways conversation : ESO/Consortium/Community**
- **Early testing phase of information extraction, simulations**
- **Parallel development of ADP modules**



⇒ Very complex instrument, but « easy » to use

⇒ Relatively easy to transform raw data into a coherent scientific picture

# Risks and Threats



# The boundary limit

---

Metric  
of  
Cost



Metric  
Of  
Impact

- Limitations come **from the community** (size, expertise, coordination)
  - *not so much from the data* [we are data rich]
- If data cannot be used easily (**high cost!** Or low efficiency)
  - ⇒ the community will switch to other datasets
- The origin of the data does not matter, e.g.,
  - ⇒ ESO could turn into the **best data provider** and *have low impact*

# Risk assessment

---

## Big Data / Multi-source

- Being late : data is there, tools are not, no one is ready
- Accessibility becomes a myth
  - « who » can address data ?
  - « who » can use the tools ?

⇒ **We lose part of our community**

- **& Low efficiency** : less than optimal data exploitation

& Bad use of money and efforts (low optimisation)

⇒ **Be Open : Engage from the start of the design phase**

⇒ **On site and distributed services (nodes, platforms)**



# Risk assessment

## Tech paradigm change

- Running after trains (Python! Platforms ?)
- **Loss of expertise, attractiveness and visibility**
- **Low efficiency** : less than optimal data exploitation  
& Bad use of money and efforts (low optimisation)



⇒ Keep an engaged perspective = active participation

⇒ Keep an attractive ecosystem (careers, projects)

⇒ Be open

# Risk assessment

---

## Societal paradigm change

- Research approach **not in sync** with new generation of researchers
  - Lack of a community identity
  - **Losing « costumers »** (both scientists & the public)
    - Not fully serving our community
    - Far from optimal Investment
    - Drift and disconnect
- ⇒ Monitor changes, embed yourself, adopt modern standards
- ⇒ Promote career paths, hire junior staff

# Risk assessment

---

## Collaborations / Interoperability

- Wrong standards, or ... no Standards
- **Culture gap**, conversation between stakeholders too costly
- Disconnect ⇒ not lowering the barriers, just **offsetting** them
- Loss of visibility, credibility
- **Ivory tower effect** (we are the best)
  - ⇒ engage, expose yourself, embed yourself
  - ⇒ Be Open [see the coming HOWTO slide]

# Black box effect

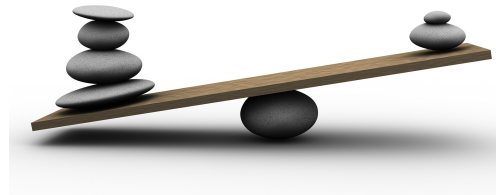
- **Telescope → remote observing**
    - Losing expertise, experience, understanding
  - **Tools and AdPs**
    - Losing expertise, experience, understanding
- ⇒ *Should we care ?*



**Lost value**

**VS**

**Added value**



⇒ Importance of being aware, doing it consciously

⇒ Embed methods and configurations in metadata ⇒ reproducibility



# Challenges

# Challenges

---

- Mid/Long-term maintenance
- Mid/Long-term preservation : of data & software + configurations
- Mid/Long-term sustainability : accessible, usable
  - New models may be needed
- Relevance of service/products w.r.t. science, and our communities
- Funding and resources (in difficult times)



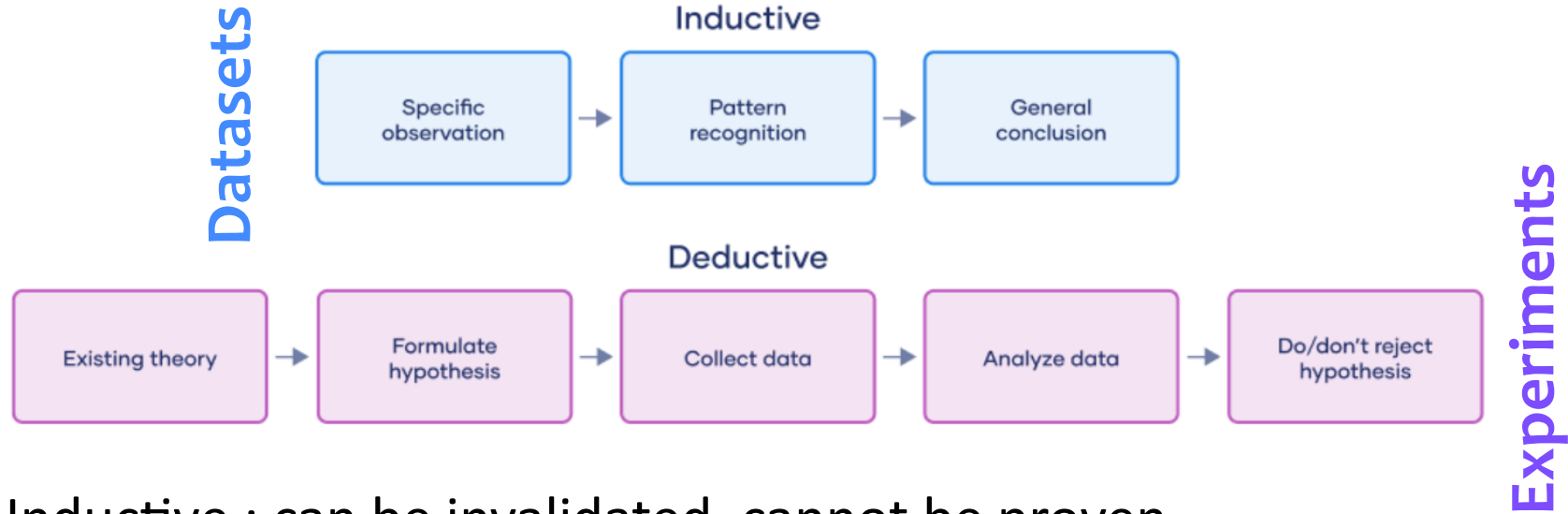
⇒ **Tendency to be satisfied with the current « working » model**

⇒ **Tendency to focus development within on one silo (my own)**

[while inter-connected multi-source data is the new paradigm]

# Which Advanced Data Products ?

## Inductive vs Deductive Research



- Inductive : can be invalidated, cannot be proven  
⇒ **Risk is** to develop useless ADPs
- Deductive : needs to be based on validated input  
⇒ **Risk is** to develop too specific/constrained ADPs

# Opportunities

# Which Advanced Data Products ? No rigid answer

## ● Guided developments

- Sense of purpose : science-driven, problem solving
- Injecting knowledge : e.g., instrument, methods, science overview

⇒ **The observatories/consortia have unique expertise**

## ● Collaborative and open development

- Open framework : allowing entries from the start
- Flexible framework : lowering the barriers for contributions
- Expert reviews : can be distributed
- Observatories can be gate keepers and guides

⇒ **Let them be « The light houses »**

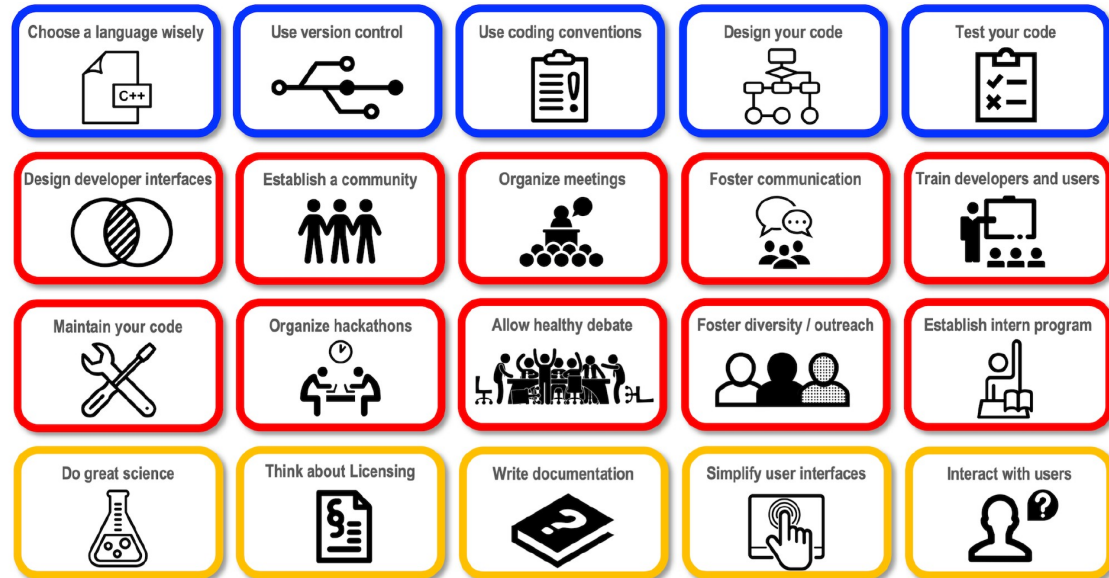
# HOWTO = collaborative developments

## Standards on all sides

- Synergies from facilities
  - VO is great ! But not enough.
- Standards for the development
- Modularity
- Tests : version, unit, data
- Code Review requests

## Community interactions

- Development Meetings, knowledge spread
- Hackathons, Data Challenges
- Training



Koehler Leman et al. 2020 ; Rosetta Commons

# Skills, resources, people

---

- Not reinventing the wheel
- **Embed observatory staff in relevant forums**
  - Make them visible actors
  - Let them contribute, guide, expose
    - And credit them for that
- Creating bridges between experts
  - Need for embedded professionals
    - The model of collaborative University departments does not work well

« **Software/tools cannot be sustained without a community around it** »



# Crediting system

---

- **Recognition and reward system needs a revolution**
  - Promote existing talents, pro-actively sponsor career paths
- **Incentivise**
  - Creative thinking, diversity of contributions
  - By making such contributions part of our astronomy port-folio
- Some practical things we can do (*Christopher Steven Marcum*)
  - « Jettison journal titles in CVs and letters of references »
  - « Substitute impact factors with narrative statement »
    - E.g. : free narrative CV templates
  - « Provide parity to non-publication output for values-based assessments »
    - Diversity of contributions

# Preservation

---

- Of Data = not anymore a bit threat (mid-term)  
⇒ **but still an unknown (long-term)**
- What about Software ?

## Black Box Effect

- *Describe the methods, algorithms, configuration files*
- Metadata associated with data

⇒ **Reproducibility (different than reusability)**

⇒ *The journey is more important than the destination*  
⇒ **The method is more important than the product**

[ But there is no universal answer that applies to all  
Sometimes = **better to reobserve, recompute, store ?** ]

# From hardware, to software, to AdPs

# 1980's

Deliver an instrument → **working hardware**

# 1990's

Space-born facilities = software is key  
Ground-based facilities = « after the fact »

# 2010's

## New paradigm

Instrument  $\equiv$  hardware + reduction software + operational scheme

## Time for

## *An Updated paradigm*

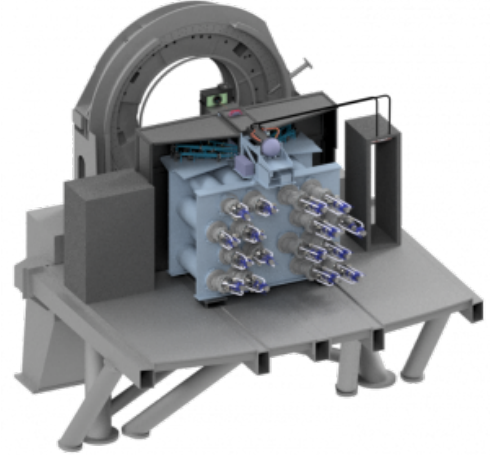
# 2020's

Instrument     $\equiv$  hardware + software + operational scheme  
                   $\equiv$  ecosystem fostering science results

# Opportunities

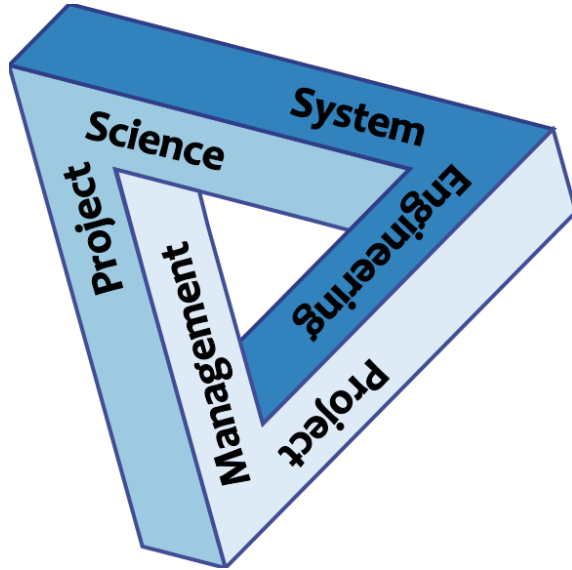
---

- MUSE lessons learned
  - E.g., ⇒ **BlueMUSE @ VLT : PhaseA (not approved)**
    - How do we incentivise ?
    - How do we connect ?
    - How do we sustain ?
  - How do we shape a Data Analysis System for a given facility ?
    - Risk to freeze a product, make it irrelevant
    - Risk not to deliver
- ⇒ **Give a sense of direction, keeping relevance, adaptability**



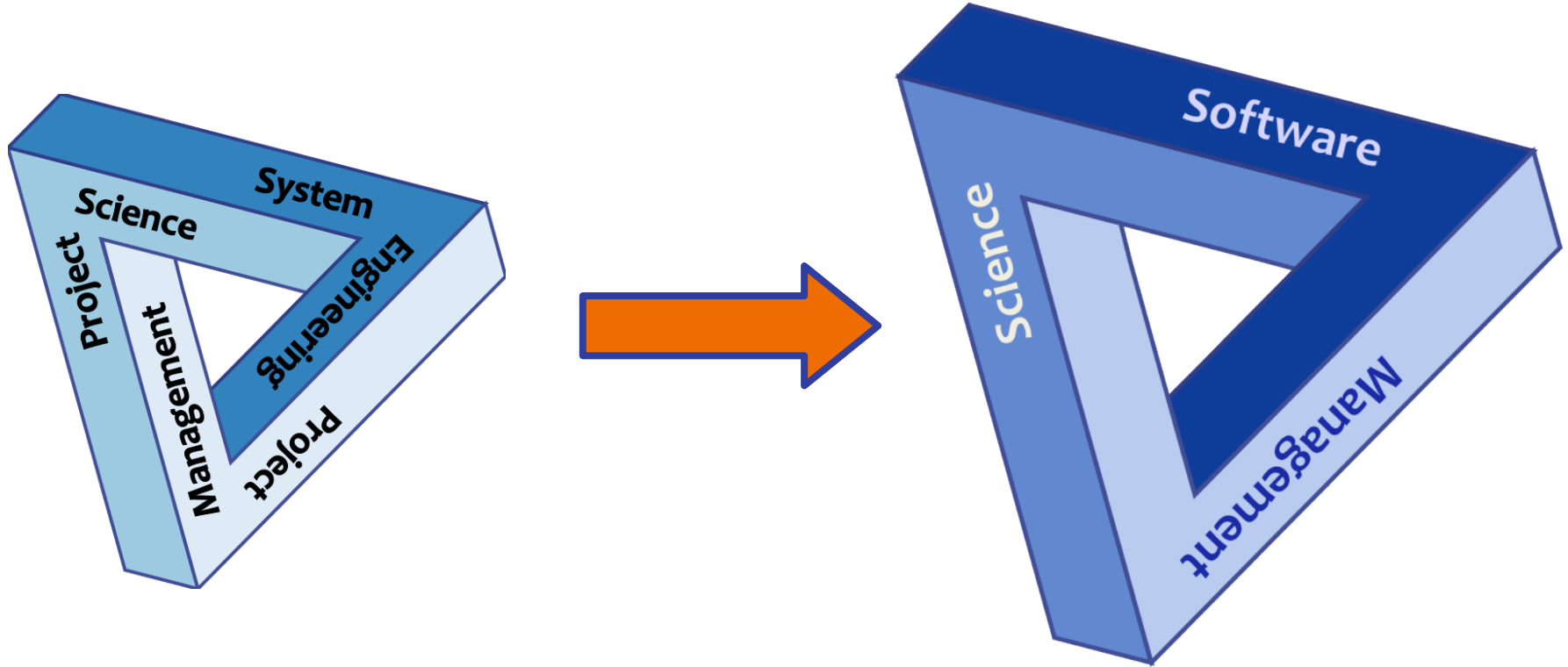
# Management / System Engineering / Science

---



# Management / Software Engineering / Science

---



# A Tiered approach

## Lessons learned

- Gaia  $\Rightarrow$  DPAC
- ALMA  $\Rightarrow$  ARCs
- ... SKA  $\Rightarrow$  SRCs

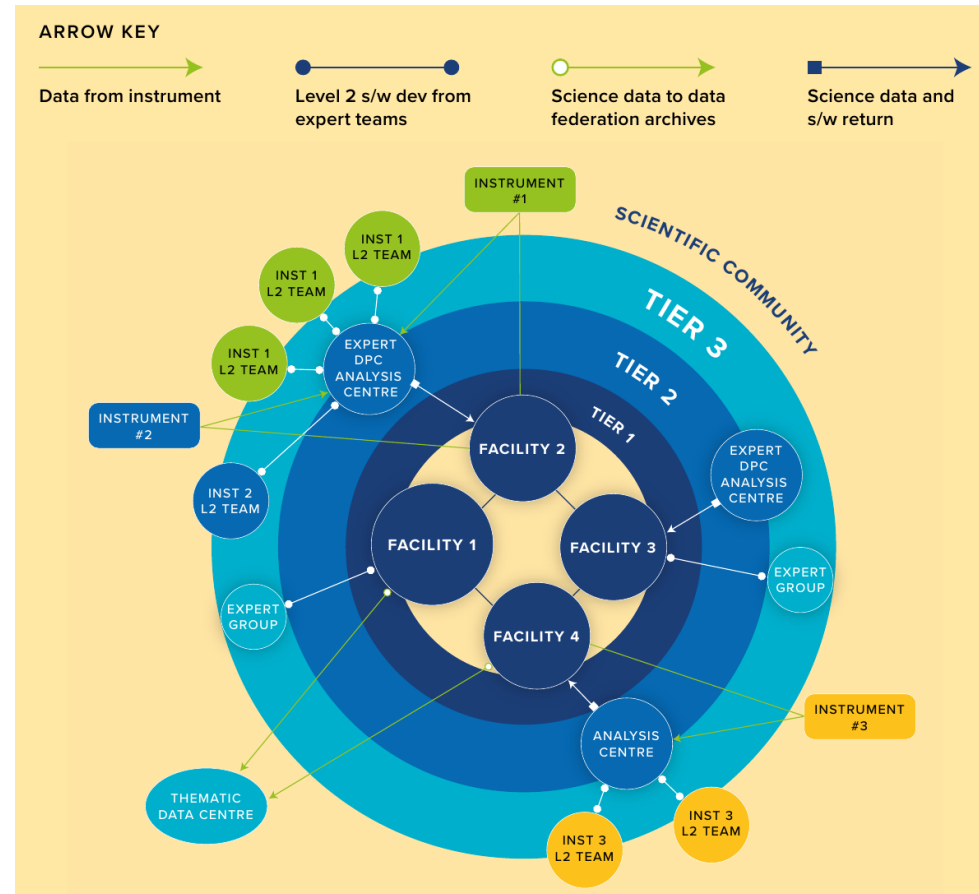
## Identified centers of expertise

- Mission specific tools
- More agile development

## Coordination of a guided effort

- Integrated communication between the builders, the coordinators & the scientific community

$\Rightarrow$  improved common standard

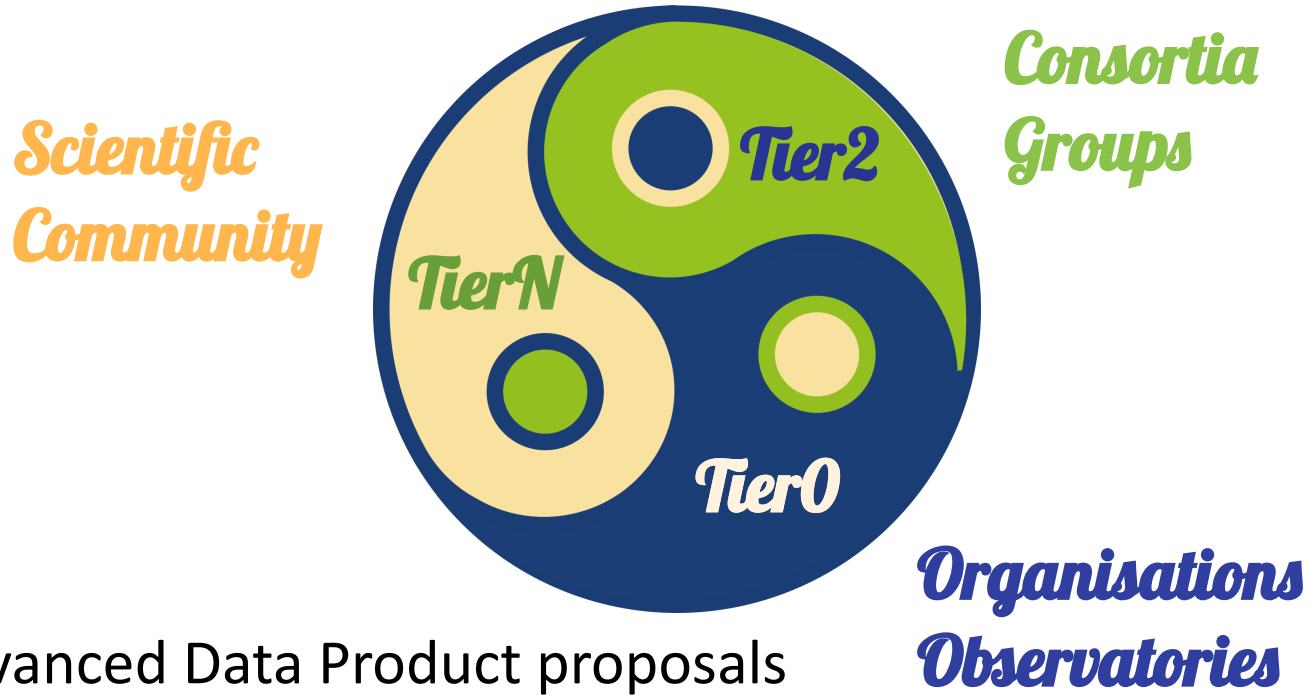


Emsellem et al. (Fig. 1, Panel Report A)



# Engage with the community

---



- Advanced Data Product proposals
  - Direct connections

**Curation** ⇒ [Code and Science Case] Review requests

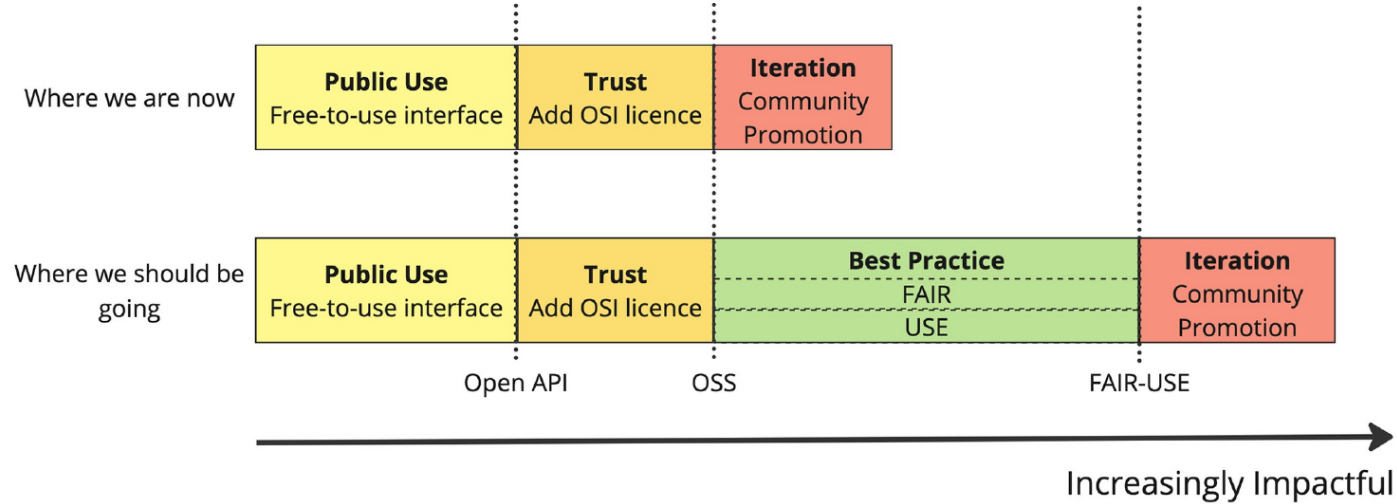
- Distributed / shared peer review

# Lowering the barriers

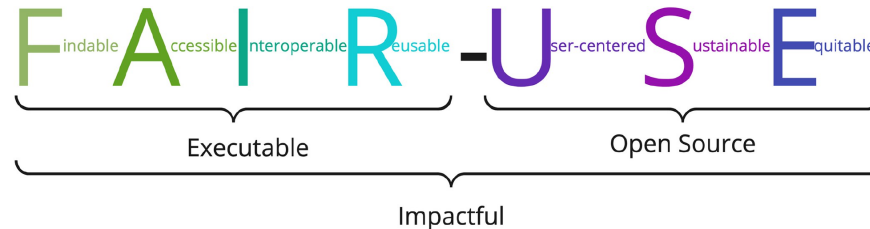
---

- Challenge of *inclusiveness & equitable access*
  - The **Open Science** umbrella
- Provide **Data Products & the Tools** to use them
  - Break the « in-house » view of tool development
  - Break the « we cannot do it, let the community do it » view
  - Let's provide a framework for development and guide it
    - Meaning : make sure it exists and is relevant
- **Provide the infrastructure**
  - Integrate datasets into international **data-sharing platforms**
  - Advertise it and provide training, tutorials
    - ⇒ will the community adopt it ?

# Best practices and more



Sonabend et al. 2024

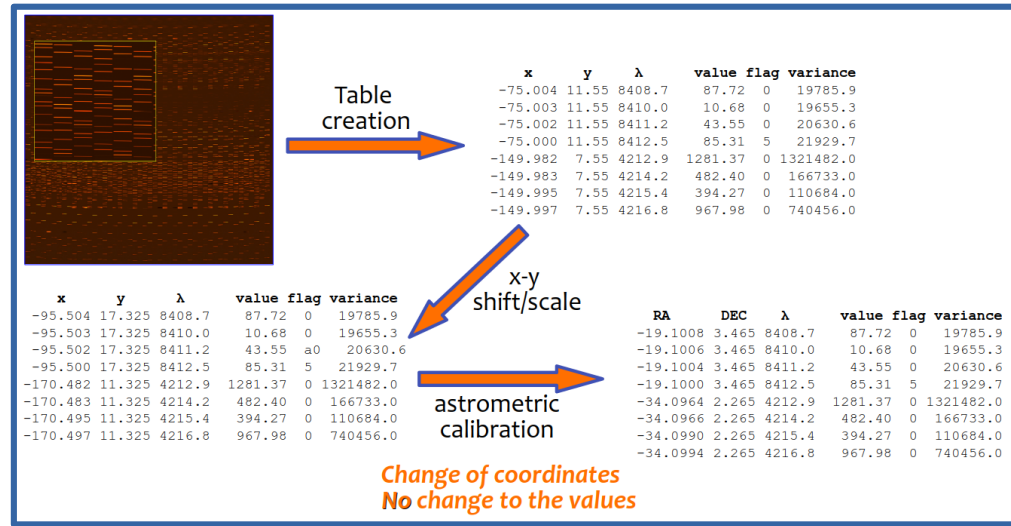


**All stakeholders = involved from the start**

⇒ to ensure the product genuinely meets the key user needs

# One more = the enhanced (spectro)perfectionism

⇒ See the ADPs as an integral part of the data journey



⇒ ADPs : may be a misleading wording

[advanced w.r.t. what?]

# Wrapping up

---

- ADPs as a way to
  - **Optimise the scientific outcome** of our facilities
  - Add value (legacy), **foster collaborations, create an identity**
- A conversation, lowering barriers
  - **Embed observatory**, organisation staff in the community
  - Consider a **Tiered** approach as a supporting, collaborative framework
  - Expose and promote tools
  - **Develop** ADPs and their methods **with** the community **from the start**
- **Upgrade/Change our crediting system**
  - Acknowledge the value of ADP development
  - Create concrete and rewarding career paths
  - Create long-term funding opportunities

# Wrapping up

---

- We need standards, we need testing frameworks
- We need expertise (resources) : Obs. as unique knowledge base
- ADPs (as pipelines) should not be developed in isolation  
(organisation may not wish to turn into data or service providers)



# Take-home messages

---

ADPs acknowledged as  
an **integral part of**  
a mission design and development

**Make**  
the development and preservation of ADPs  
a part of our **astronomical port-folio**