

Stereo-SCIDAR

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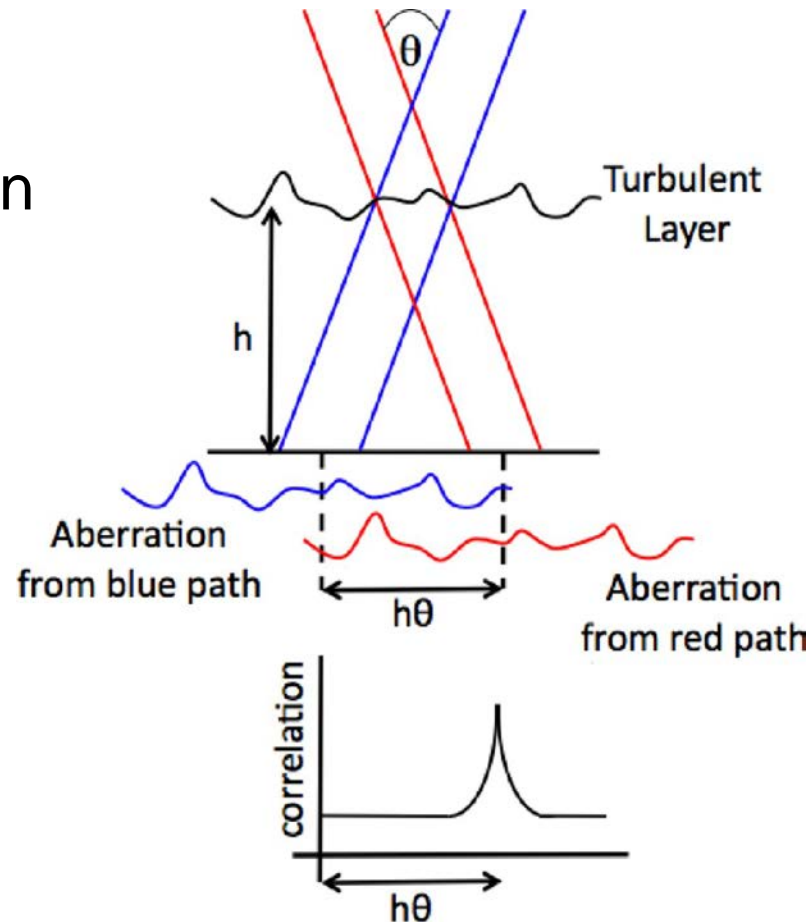
Tim Morris

Optical turbulence profiling regimes:

- **SLODAR** at Paranal: aim is to profile the surface-layer of turbulence (up to $\sim 200\text{m}$) with very high resolution ($\sim 10\text{-}20\text{m}$)
 - **Applications** - understand and estimate UT seeing
 - support GLAO / AOF
- **SCIDAR**: Profile the whole atmosphere up to $\sim 20\text{km}$
 - Lower resolution, but can be $\sim 200\text{m}$ on 2m telescope
 - Very high sensitivity – detect weak turbulent layers
 - Measure layer velocities
 - **Applications** – model and support tomographic AO for large telescopes
- Requires a larger telescope, $> \sim 1\text{m}$ (2m is much better)

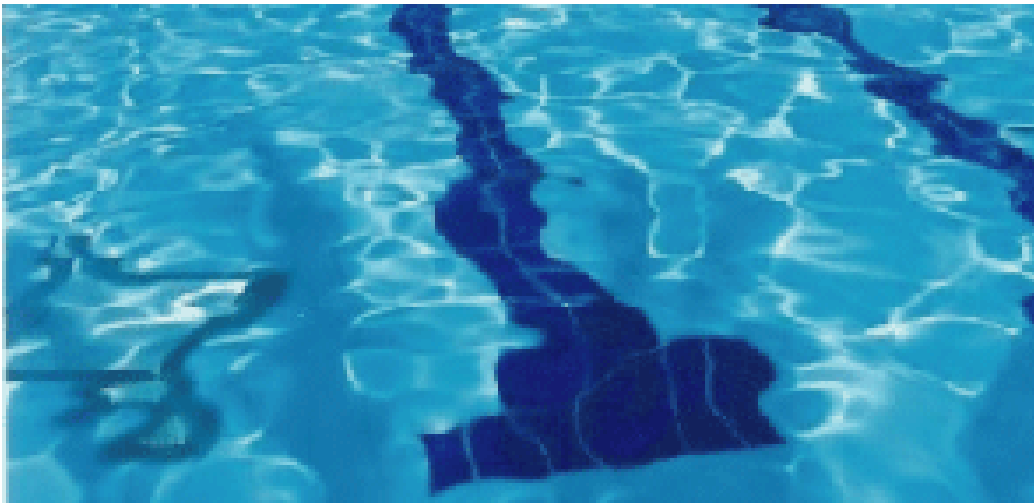
SCIDAR: Optical Turbulence Profile From Double Star Scintillation Patterns

- Triangulation method
- Peaks in cross-correlation pattern determine height and strength of turbulent layers

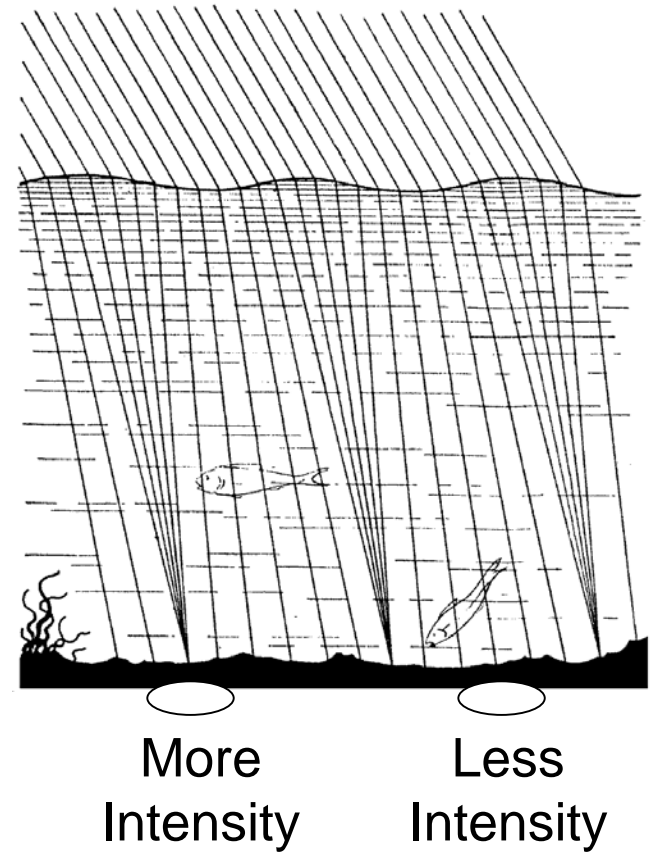


If a turbulent layer at height, h , is illuminated by two stars of angular separation, θ , then two copies of the aberration will be made on the ground separated by a distance $h\theta$.

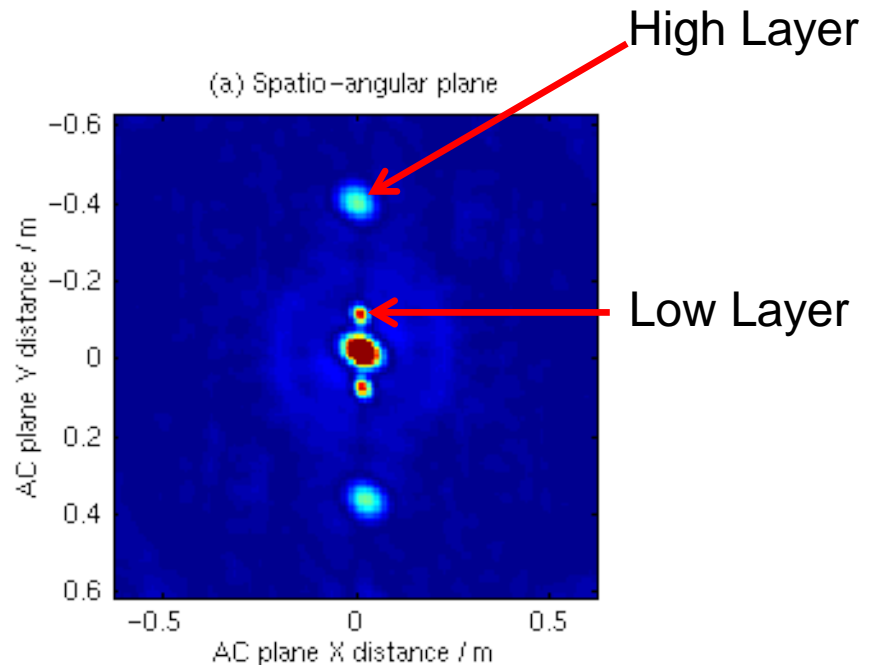
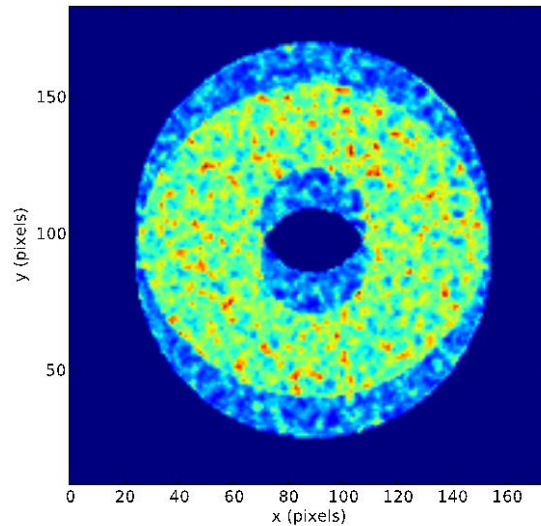
Scintillation ('Flying Shadows')



'Shadow Patterns' in Water



(Conventional) SCIDAR: SCIntillation Detection And Ranging



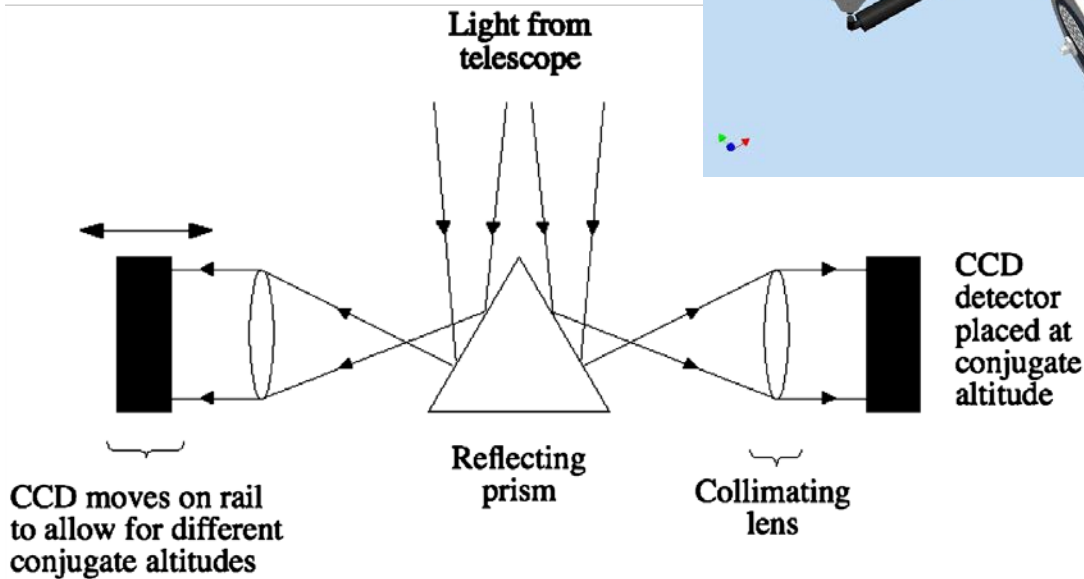
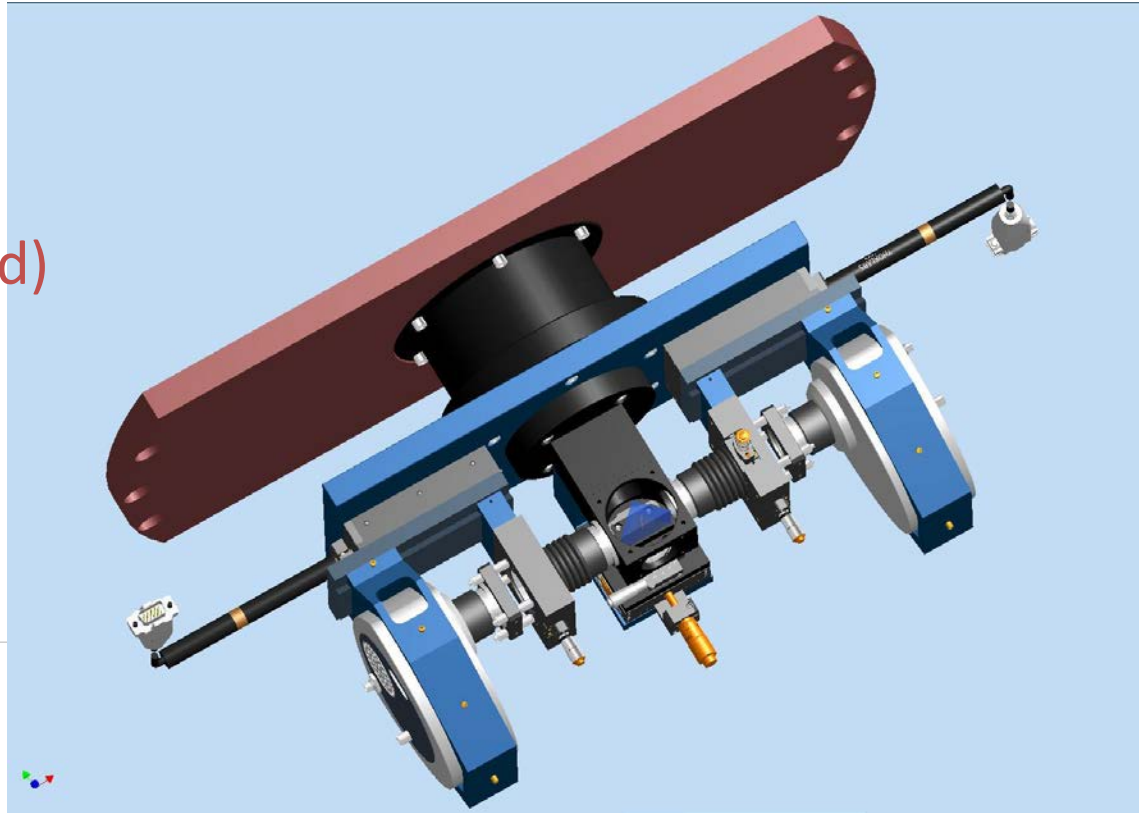
Double Star Scintillation Pattern

Image **Auto**-Correlation (Time-Averaged)

- **Conventional SCIDAR:** The scintillation patterns for the two stars are superposed on the detector
- Each turbulent layer produces **3 peaks** in the auto-correlation

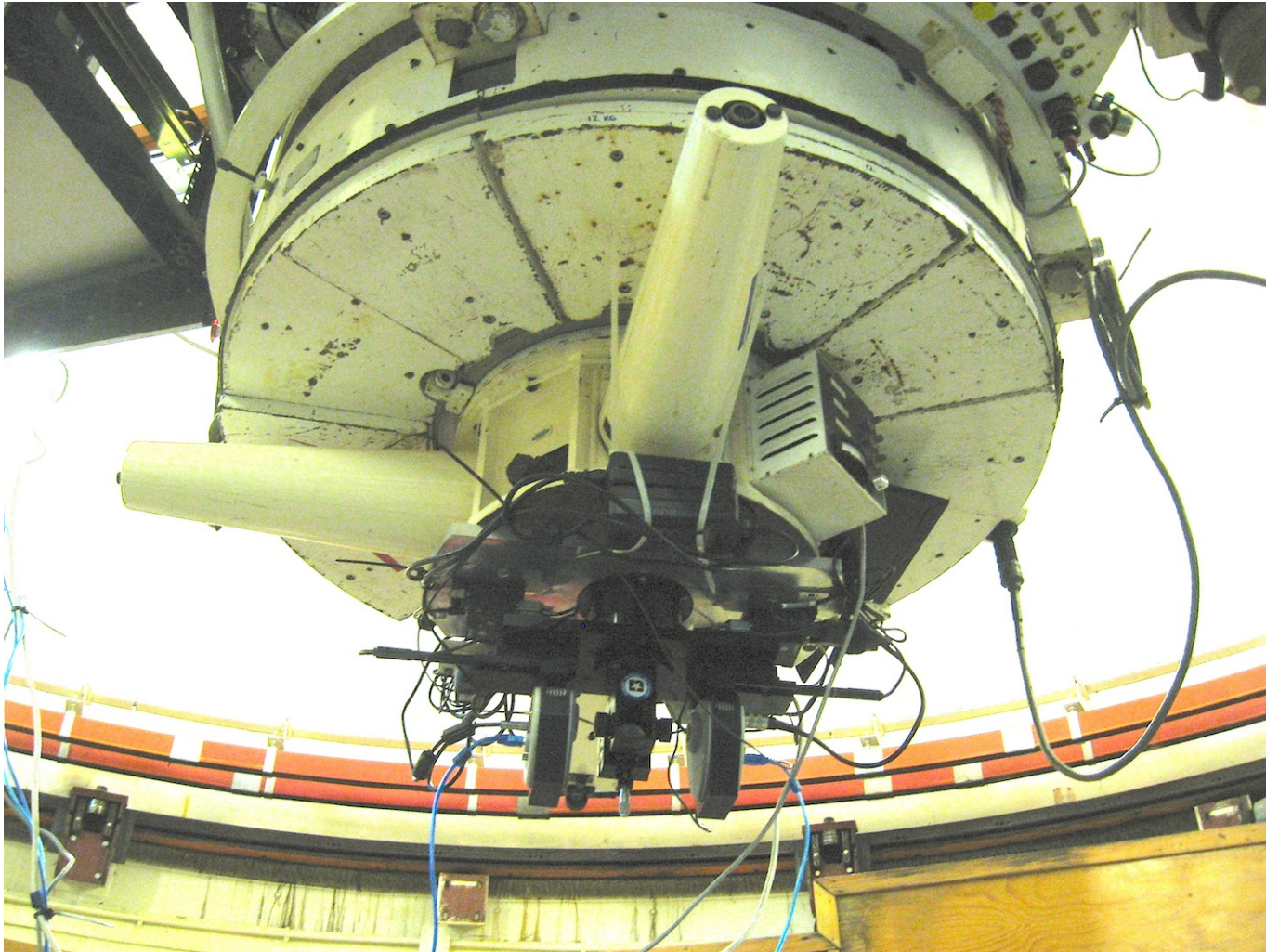
'Stereo – SCIDAR'

- Use separate (synchronized) cameras for each star

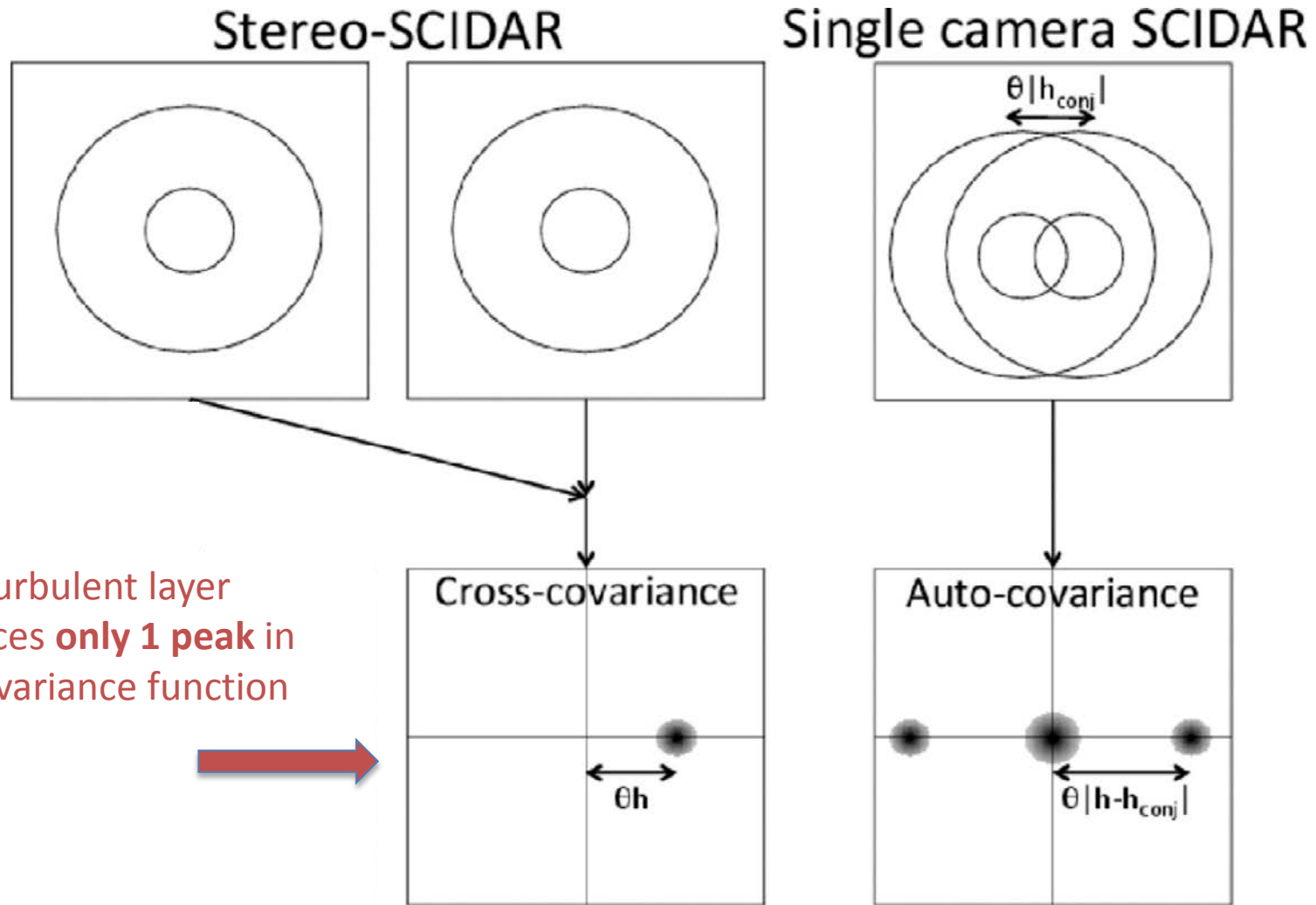


Stereo –SCIDAR on JKT, La Palma

Supports Canary AO demonstrator on WHT



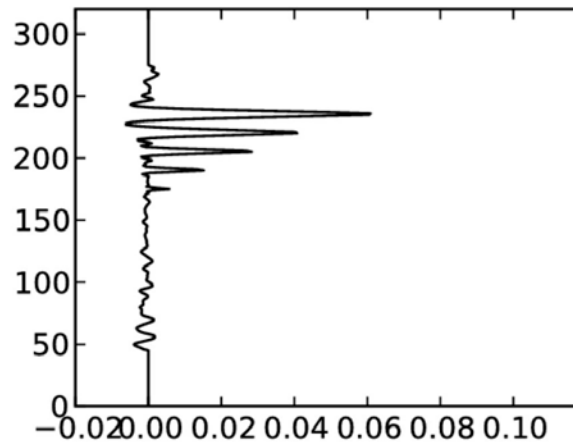
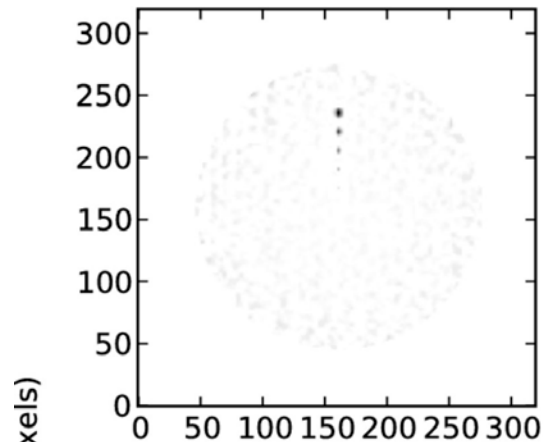
Covariance Response Functions – Conventional and Stereo SCIDAR



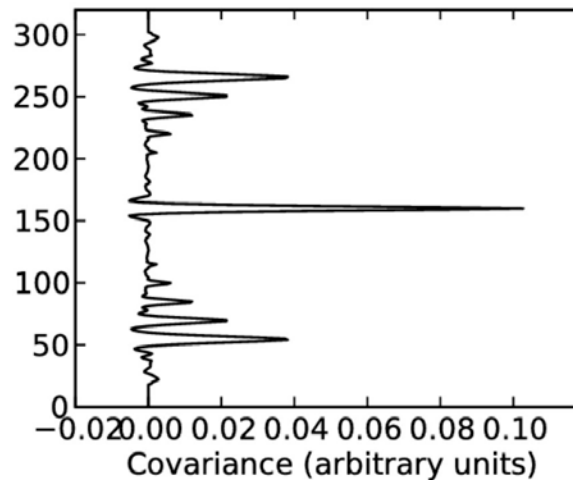
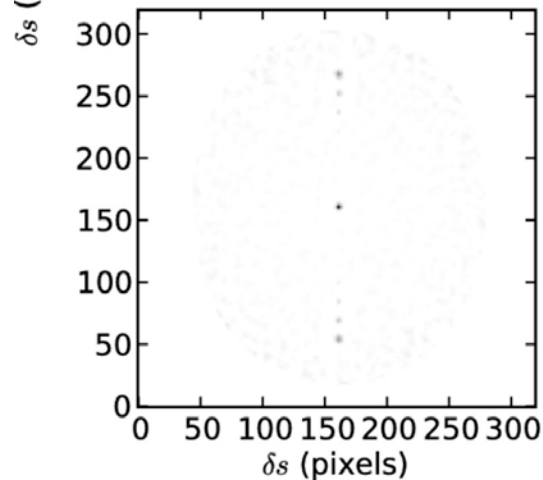
Each turbulent layer produces **only 1 peak** in the covariance function



Covariance Response Functions: Stereo-SCIDAR and Conventional SCIDAR



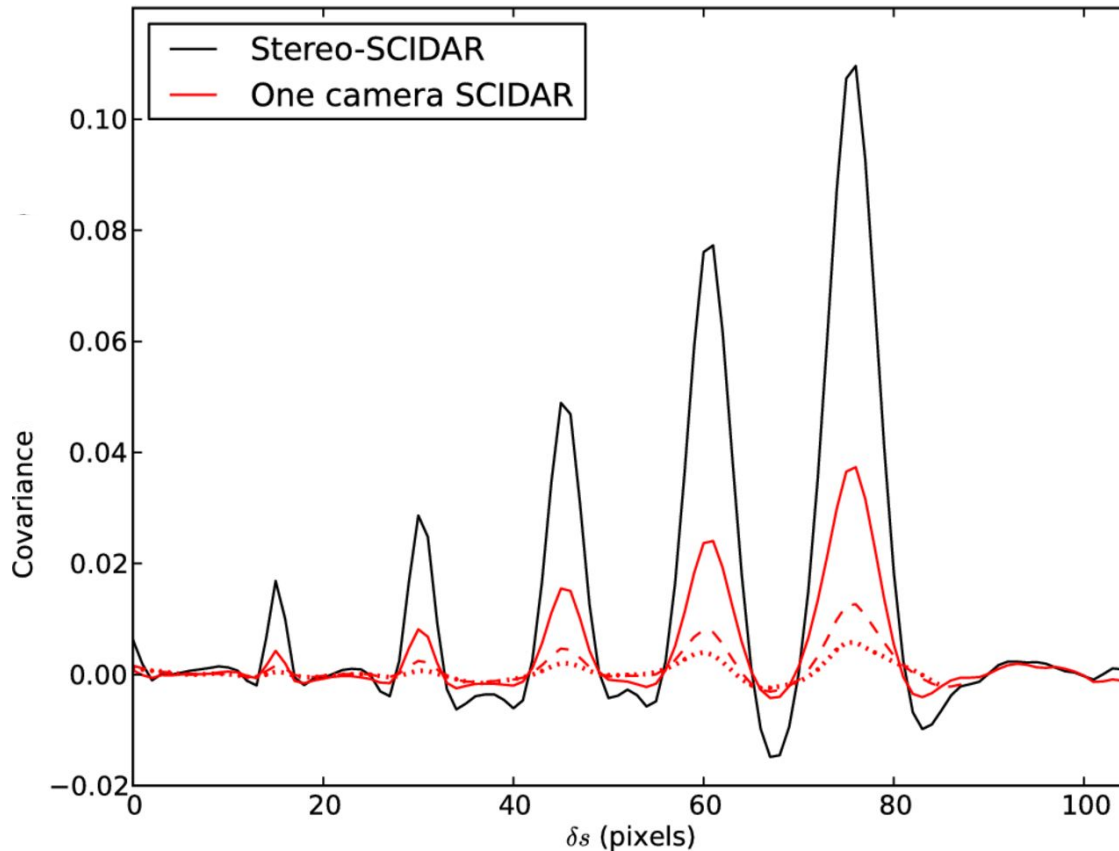
Stereo-SCIDAR



Conventional
SCIDAR

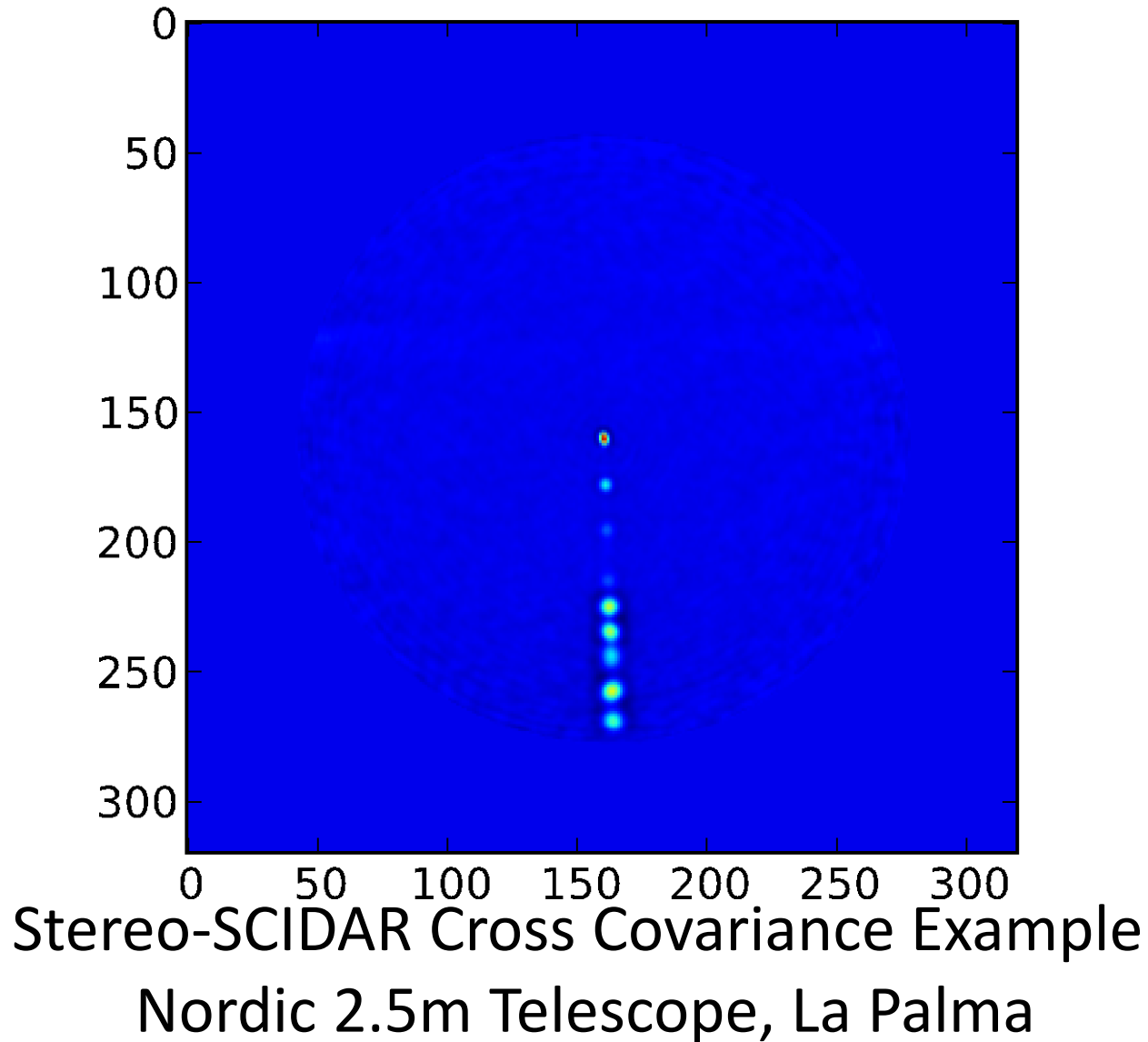
6 layers of equal strength, 2 km intervals between 0 and 10 km

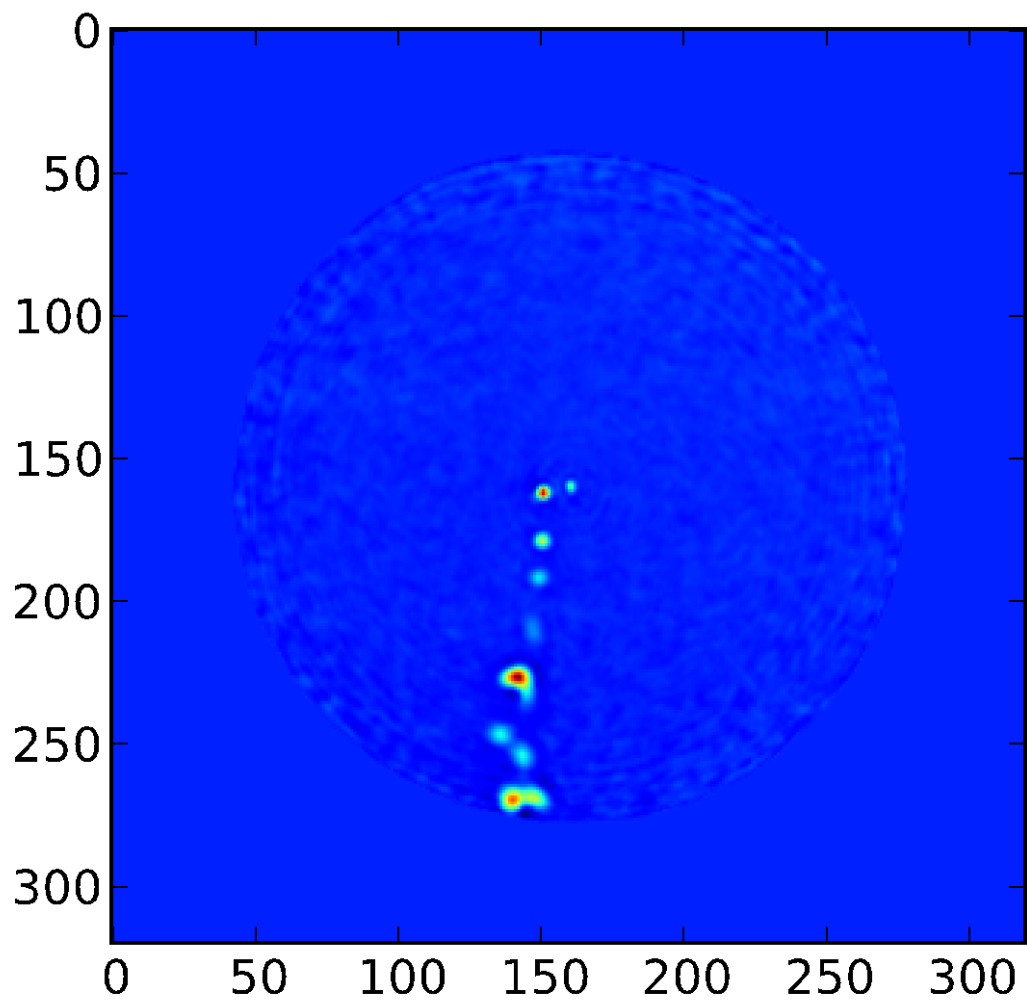
Covariance Response Functions: Stereo-SCIDAR and Conventional SCIDAR

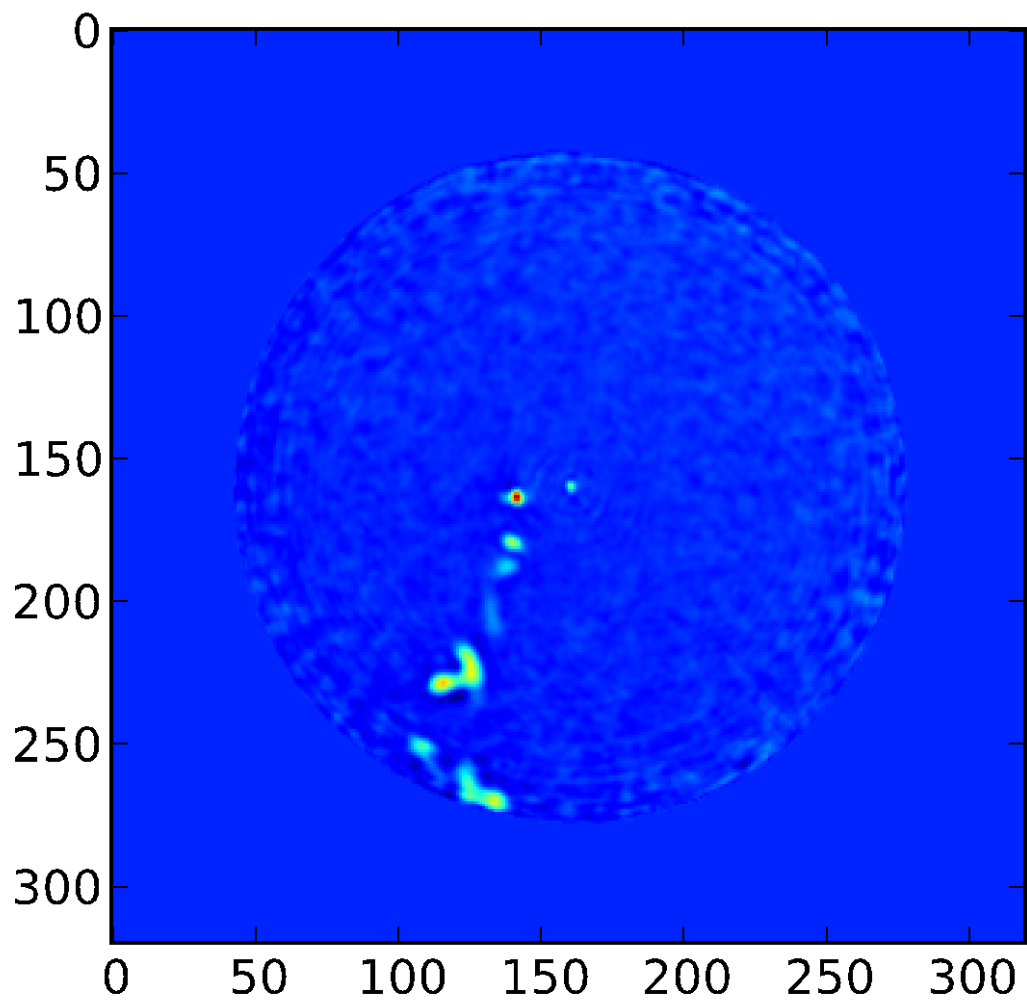


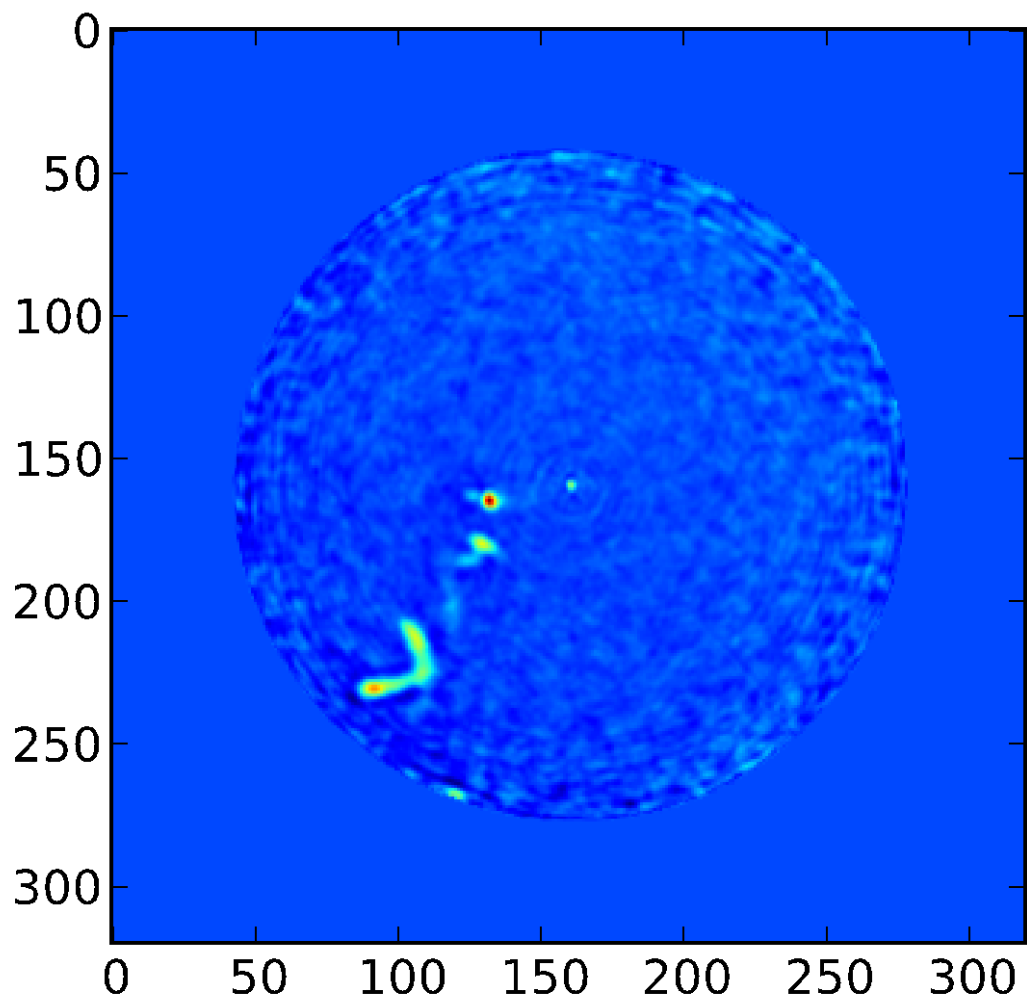
Solid lines: Target stars have the equal magnitude
Broken line: 2 magnitude difference ($m_1 = 4, m_2 = 6$)
Dotted line: 3 magnitude difference ($m_1 = 4, m_2 = 7$).

Turbulence (Wind) Velocity Measurements



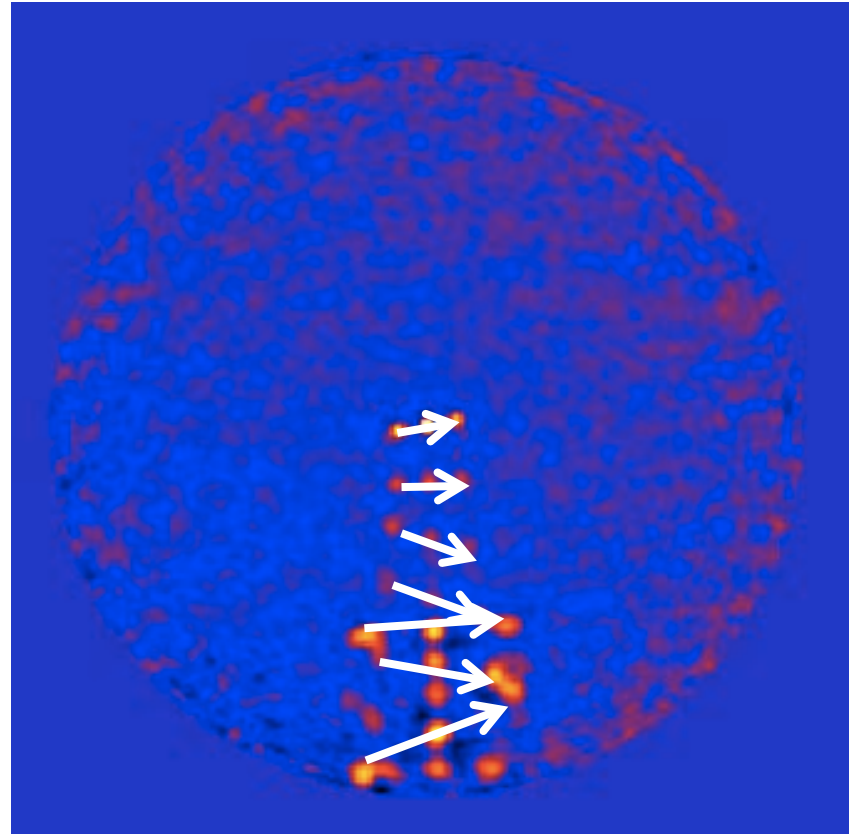






Wind Velocity Profile

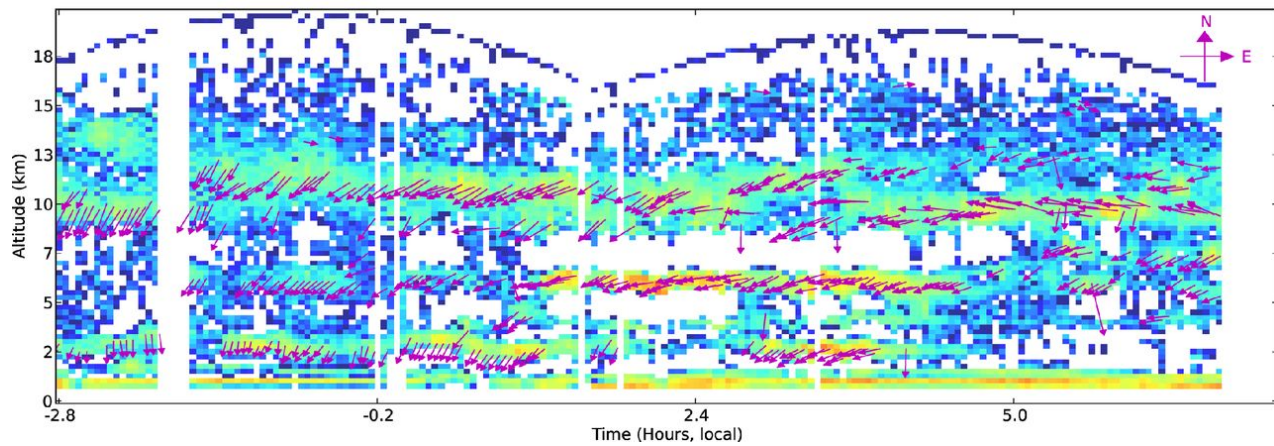
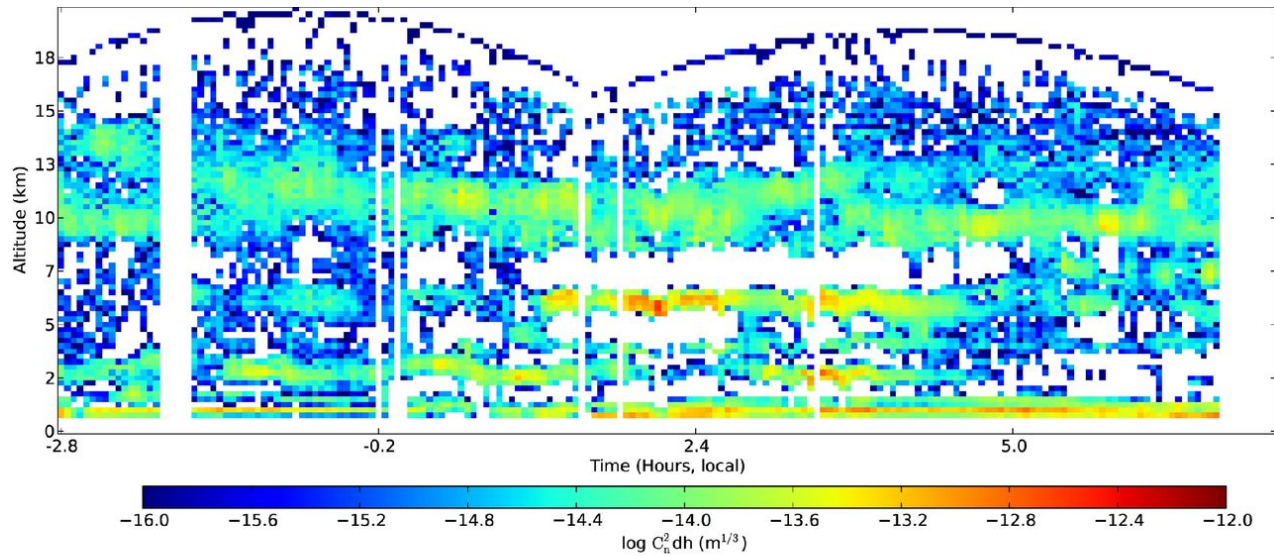
- Cross correlations only contain one set of peaks.
- Geometric wind profile algorithm rather than wavelet analysis
- Stationary central peak corresponds to dome seeing (dt +ve and -ve to avoid confusion with other peaks)



Wind profile 25/09/12, NOT 2.5 m, 9 Layers
Sum of Cross correlations for three time steps

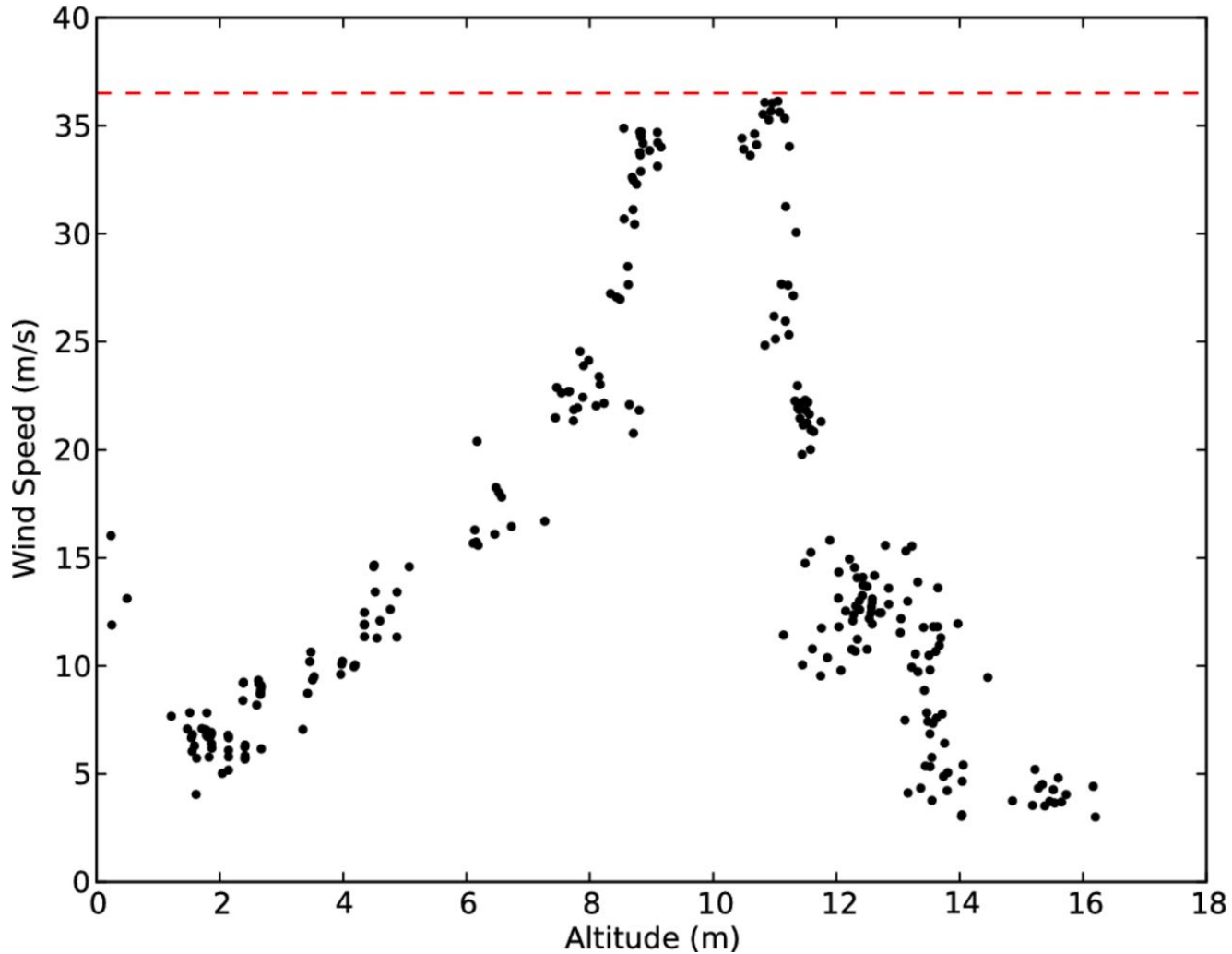
Example S-SCIDAR Turbulence Profile Sequence

JKT, La Palma, 2013 September 15.



Example Distribution of Wind Speed with Altitude

Stereo-SCIDAR on JKT, La Palma, 2013 September 13



Stereo-SCIDAR Advantages

- **SNR improved** for turbulence profiling by factor of 2-20, c.f. conventional SCIDAR (target dependent):
 - Detect weaker layers
 - More accurate determination of layer strengths
 - Improve resolution in altitude:
 - ~ 200m resolution possible on 2m telescope
- **Turbulence velocity** measurements are very sensitive, robust and greatly simplified.
- **Dynamic re-conjugation** is possible – may improve layer altitude accuracy even more (but strong ground-layer is a problem)
- **More target** double stars available (can tolerate bigger magnitude difference) ...

Stereo-SCIDAR: Application at Paranal ?

- Possible Auxiliary Telescope instrument
- Precedent - IAC 'Cute-SCIDAR' deployed on AT (2007)
- Aim for minimal support requirements, 'point-and-shoot' instrument

- Could operate on ~6 nights per month (when ATs are not in use for interferometry)
- Acquire statistical turbulence profile data relevant to Paranal and Armazones:
 - Turbulence strength and velocity up to ~20km
 - High vertical resolution, high sensitivity