



Summary Information Sheet

The SUN

The Sun is our nearest star. It is a huge, luminous ball of gas like all the other stars. It consists mostly of hydrogen and helium, with tiny amounts of other elements





The Corona is the outer envelope of the Sun's atmosphere. It is extremely hot with temperatures up to 2 million degrees

The Chromosphere is a transparent layer above the photosphere. It extends up to 2000 km with temperatures around 10,000 degrees

Sunspots -

The Photosphere is the visible 'surface' of the Sun. It is about 300 km thick. Here most of the Sun's activitiy takes place, e.g., sunspots

The solar cycle:sunspots and other forms of solar activity vary with an average period of 11 years The Radiative Zone Here energy is transported outwards by radiation. It covers about 70% of the Sun's diameter

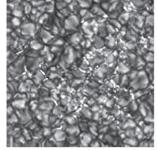
> -The Core In the centre of the Sun the energy is produced by fusion processes through which hydrogen nuclei are fused to produce helium nuclei

The Convective Zone It extends roughly over 30% of the Sun's diameter. Here energy is mainly transported upwards by convective streams of gas

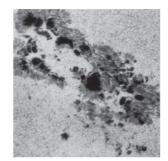
The spectrum of the Sun not only shows the rainbow colours: It also displays dark lines named absorption lines or Fraunhofer lines



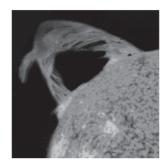
Spectrum of the Sun



Granulation



Sunspot



Eruption



The Sun's corona during a solar eclipse

Physical Data

Property

Distance from the Sun Rotation period Equatorial radius Mass Density

Sun

27 days 695,000 km $2 \times 10^{30} \, \text{kg}$ 1400 kg/m³

Earth

150 million km 23 hrs 56 min 6378 km $5.97 \times 10^{24} \text{kg}$ 5520 kg/m³

Jupiter

779 million km 9 hrs 55 min 71500 km $1.899 \times 10^{27} \text{kg}$ 1330 kg/m^3

Concept: B. Mackowiak

For comparison