

Background

Non-common path aberrations (NCPA) inside astronomical instruments hinder optical performance. Novel calibration techniques are needed to estimate aberrations and compensate them. Improved performance means enhanced observations. Better data means new exciting science!
Goal: *Could we use ML to estimate the aberrations in our instrument?*

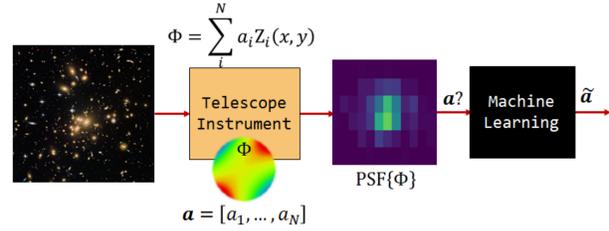


Figure 1 – Machine Learning in the context of NCPA calibration

Methodology

Single-Network approach: MLP regressor trained on PSF images to identify underlying aberrations $\bar{a} = [a_1, \dots, a_N]$

Main caveat: *Curse of dimensionality.* Size of training set blows up with N number of aberrations we want to calibrate

Proposed solution: multi-network approach where we split the task among several networks, each trained to recognize a small subset of aberrations (keeps training sets manageable), see Fig. 2

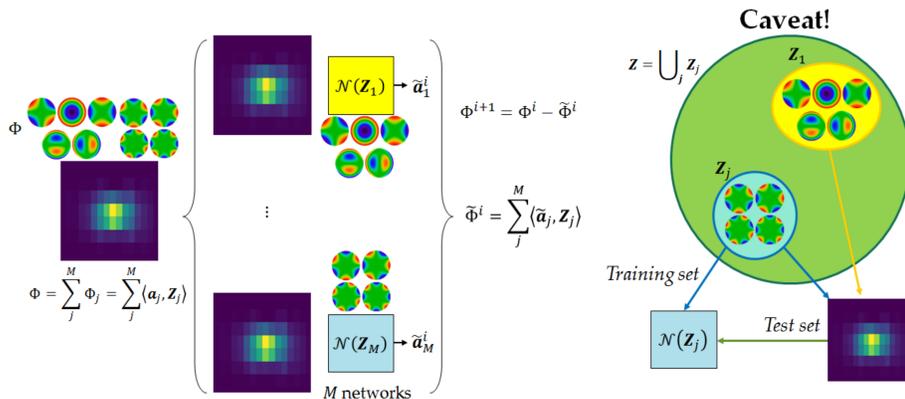


Figure 2 – Multi-Network approach. Each network is trained to recognize a small set of aberrations. Problem: distribution mismatch training-testing!

Problem: each network sees images with features from aberrations it doesn't know about! **Feature contamination** degrades performance
What can we do to clean the images?

Denoising AutoEncoders [1]: trained to minimize reconstruction error between corrupted input and clean output. We teach them how to remove 'noise' from other aberrations.

Our calibration architecture: parallel DAE + MLP

DAE remove the spurious features to help the MLP do their task. But how do we use the data? Two options: **reconstructed:** train MLP on clean images $\mathbf{r} = g(f(\tilde{\mathbf{x}}))$; vs. **encoded:** train MLP on encoded data $\mathbf{h} = f(\tilde{\mathbf{x}})$.

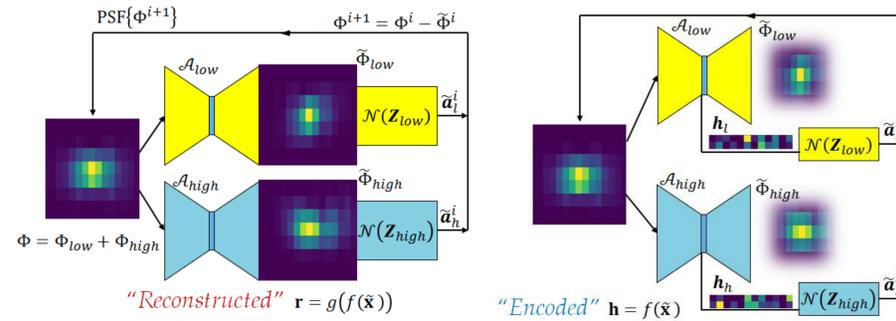


Figure 3 – Two architectures, depending on how we use the DAE data to train the MLP.

Clean is not enough. Taking denoising one step further

Encoded: denoising is *not* the goal; it is a training criterion for learning to extract useful features to guide the learning of other tasks [2].

Learning to denoise = learning to represent the aberrations

Training MLP on **encoded** outperforms **reconstructed** approach (Fig. 4).
Advantages: faster training (encoded data has lower dimensionality), better performance "robust" features learned by the AE help calibration.

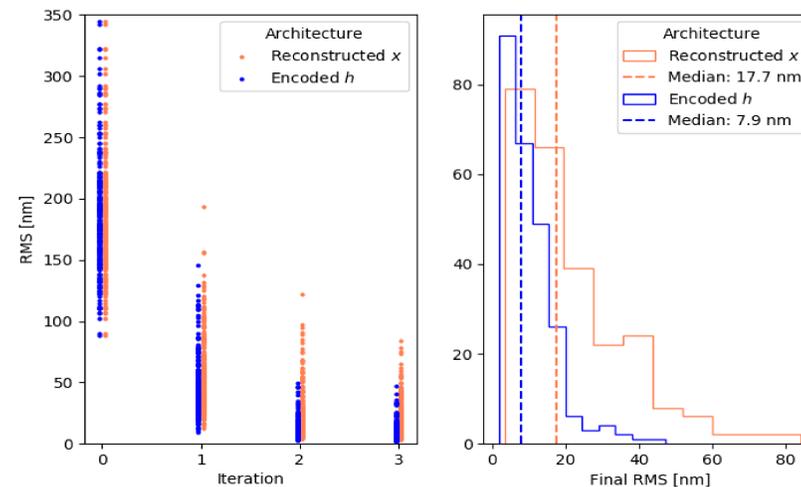


Figure 4 – Calibration results for the two architectures. MLP trained on **encoded** data leads to lower residual errors than when trained on **reconstructed** PSF images.

What is the autoencoder really *learning*?

For \mathcal{A}_{high} , \mathbf{h} doesn't depend on the removed features. It learns to encode the information needed to recover Φ_{high} and disregards Φ_{low} (Fig. 5)

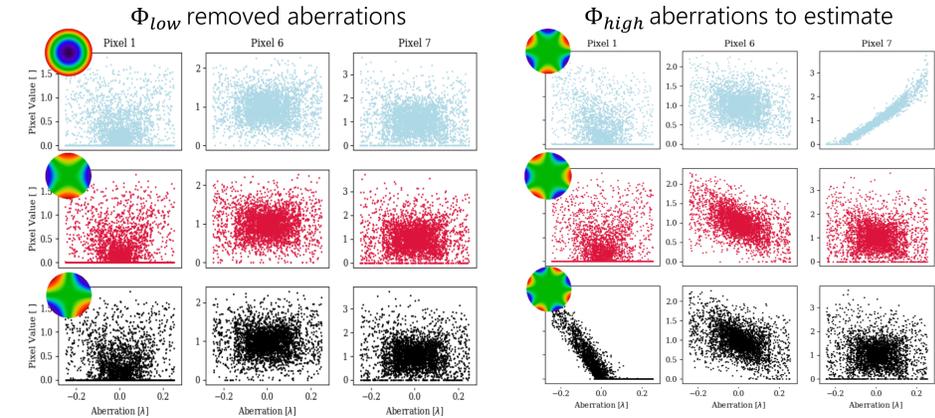


Figure 5 – \mathcal{A}_{high} (trained to remove Φ_{low}). The encoded features as a function of aberration intensity.

PCA analysis of encoded features (Fig. 6) reveals \mathcal{A}_{high} learns a new data representation of the aberrations Φ_{high} . **Advantages:** sparse, robust against feature contamination, leads to better learning of MLP.

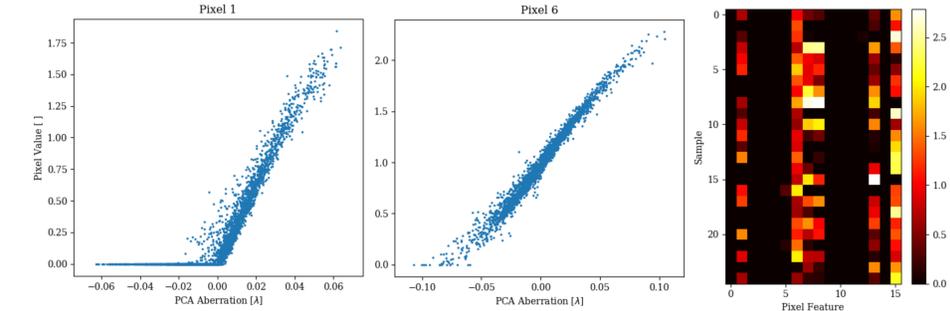


Figure 6 – Encoded data projected onto PCA Φ_{high} aberration space

Conclusions

- ❖ Novel approach to NCPA calibration via ML
- ❖ Use of Autoencoders to mitigate feature contamination
- ❖ Denoising criterion guides learning of new data representation
- ❖ Using that data representation leads to better performance on calibration task.

References

- [1] Vincent et al. 2008 - Extracting and Composing Robust Features with Denoising Autoencoders
- [2] Vincent et al. 2010 - Staked Denoising Autoencoders. Learning useful representations...