

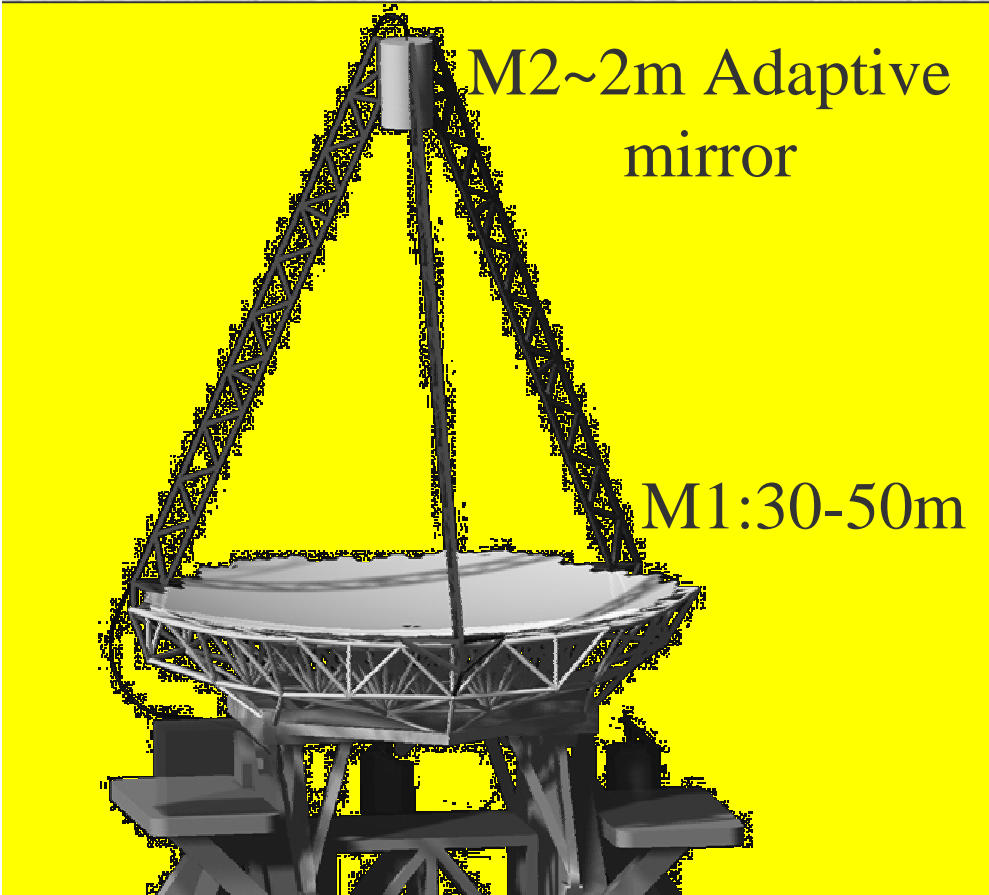


First generation of AO systems for ELTs: a short review of possible systems

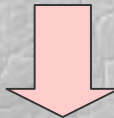
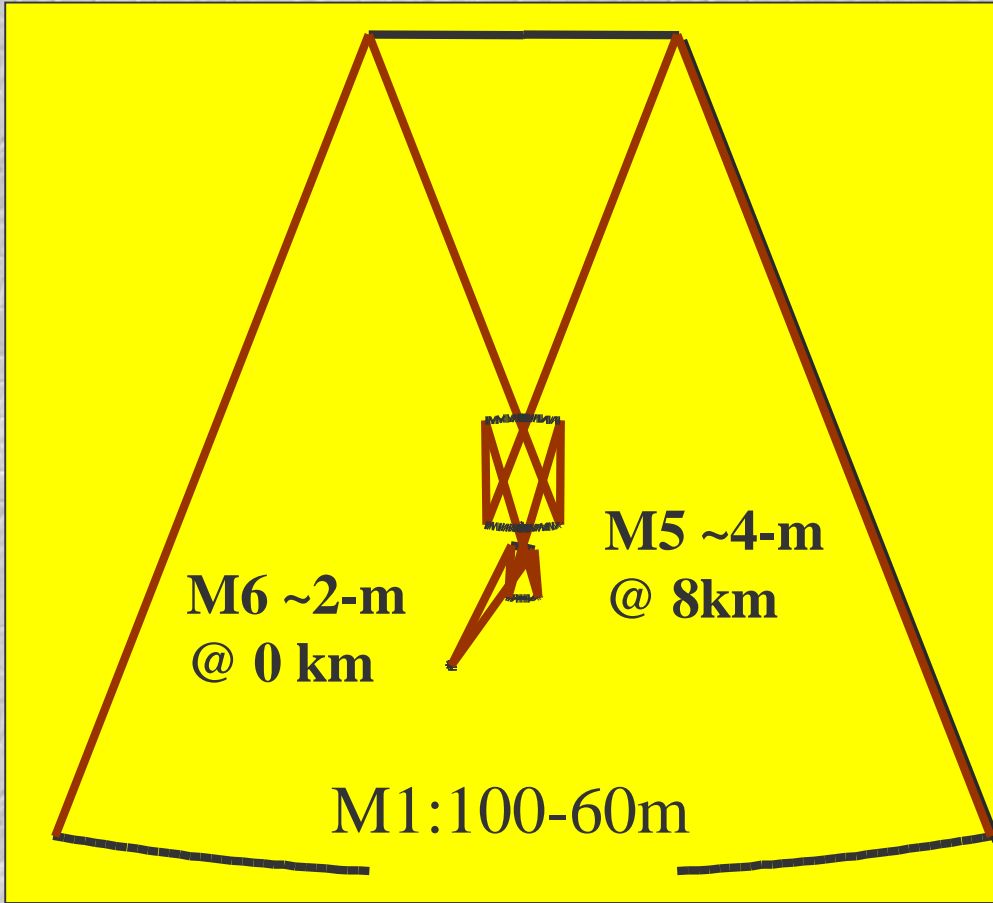
N. Hubin
European Southern Observatory



ELT optical design and Adaptive Optics



+ post focal AO correctors



+ post focal AO correctors



Conventional AO with NGS

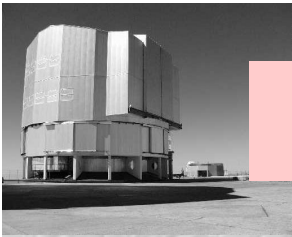


- § Telescope 30-50 m
- § 1 DM with $\sim 2-3$ act/m
- § DM diameter 2m
- § 2800-5000 actuators
- § 1 WFS within θ_0
- § Diffraction limit at $1\mu\text{m}$
- § XAO above L-band

- § Telescope 60-100m
- § 1 DM with $\sim 1/2$ act/m
- § DM diameter 2-4 m
- § 5000-20000 actuators
- § 1 WFS within θ_0
- § Diffraction limit at $1.5\mu\text{m}$
- § XAO > M-band!

Questions:

- Median seeing, seeing variability
- Median correlation time and variability
- Outer scale L_0 and its impact on max. actuator stroke, low order modes
- Median H_0 (in case of mono-conjugate DM)

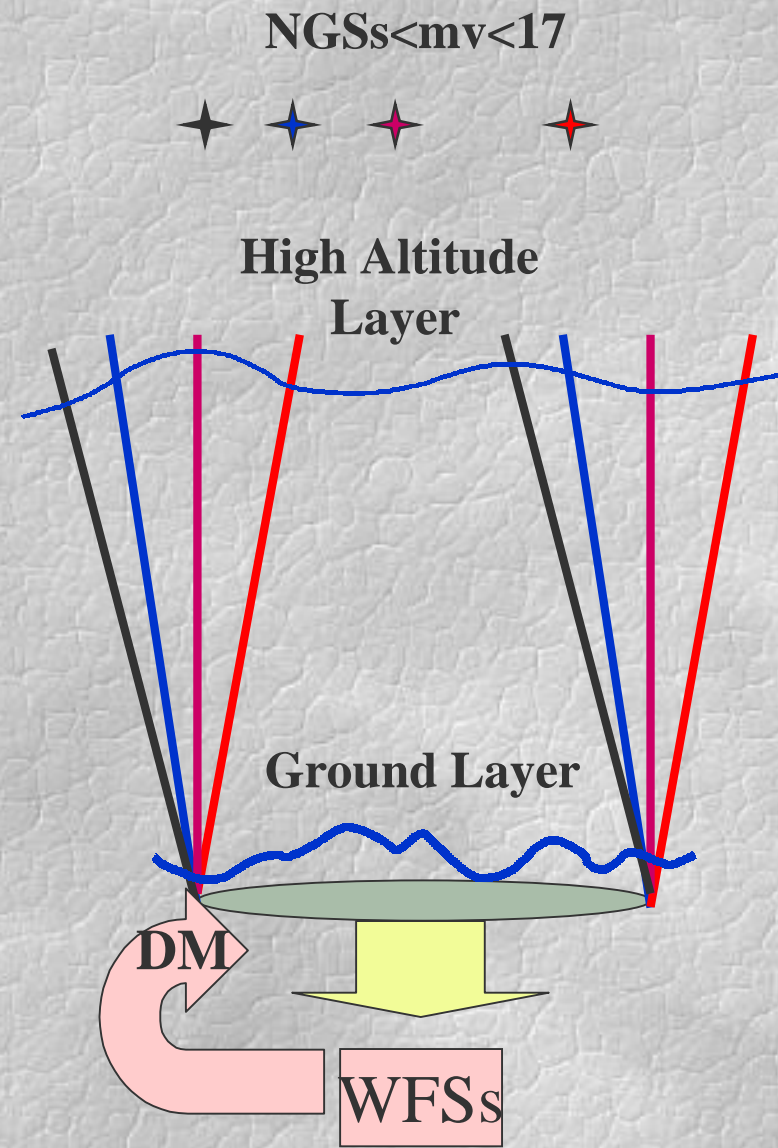


Large FOV Ground layer tomography with N or LGS

- § Tomography with 6-8 NGSs in 6'
- § Multi-WFSs with $\sim mv=15.5$
- § Or tomography with 5-9 LGSs in 5'
- § Corrected FOV 5-6'
- § One 2.5- 5kact. DM corrector at 0km
- § Seeing reducer in K-band (factor 2)

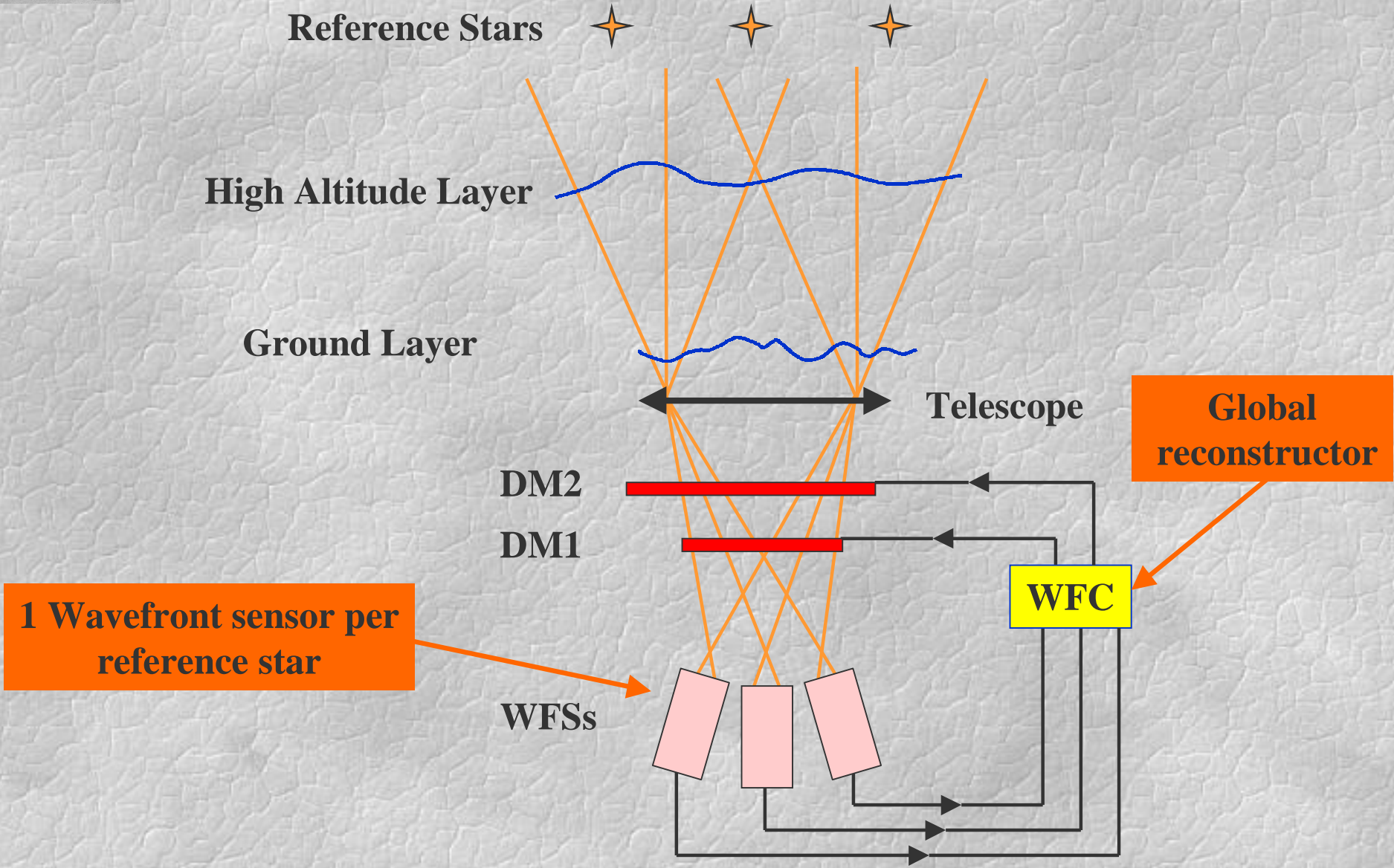
Questions:

- $C_n^2(h)$
- Intensity of the ground layer
- Variability of intensity/altitude
- Impact on correction efficiency over large FOV





Global Reconstruction MCAO concept





MCAO with LGSs on 30m ELT



- 3 DMs at 0, 5 and 10 kms
- ~ 8500 actuators total
- 5 sodium LGS
- 3 NGSs
- FOV: 2'
- NIR imager

Questions (for scale 30-100m!)

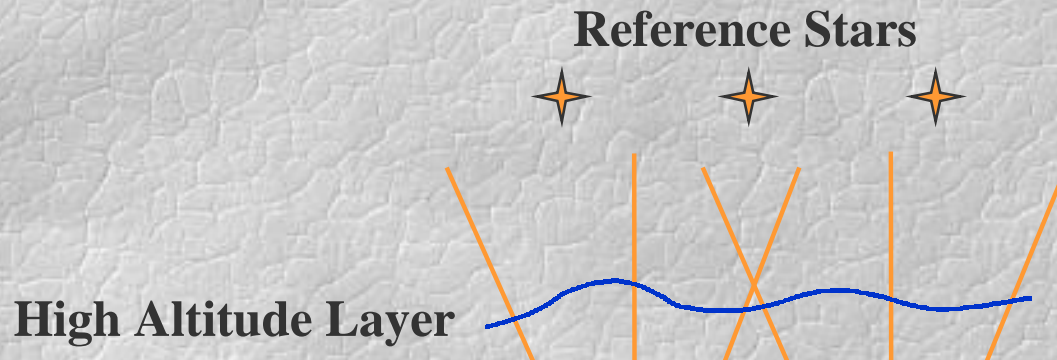
- Best altitudes & variability for DMs: $Cn^2(h)$
- Predictability of WF evolution/layer: WF frozen?
- WF temporal information for each layer
- Impact of L_0 (h)!
- WF spatial & amplitude/freq. for each layer
- Sodium density/altitude/profile variability
Seasonal and mid-short term

Multi-LGSs



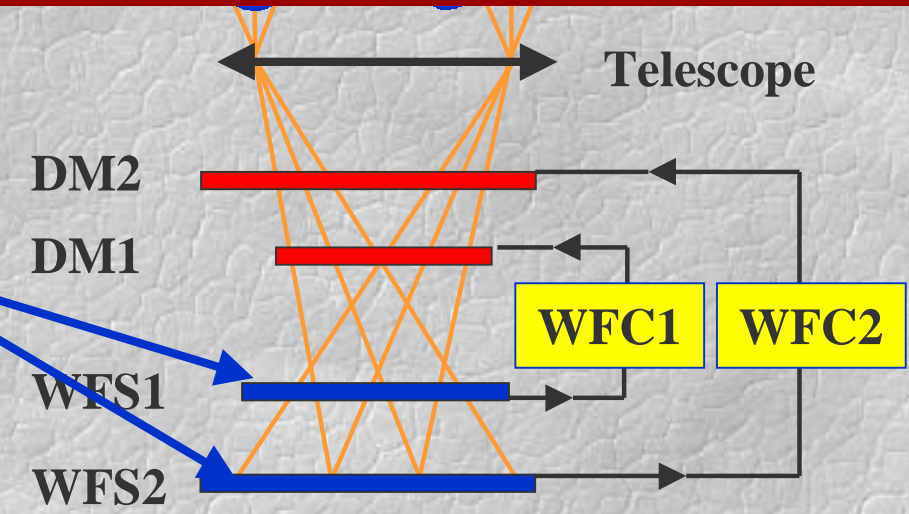


Layer Oriented MCAO concept



**By enlarging the WFS FoV we can collect the light of more stars
NGS MCAO is now possible with good sky coverage**

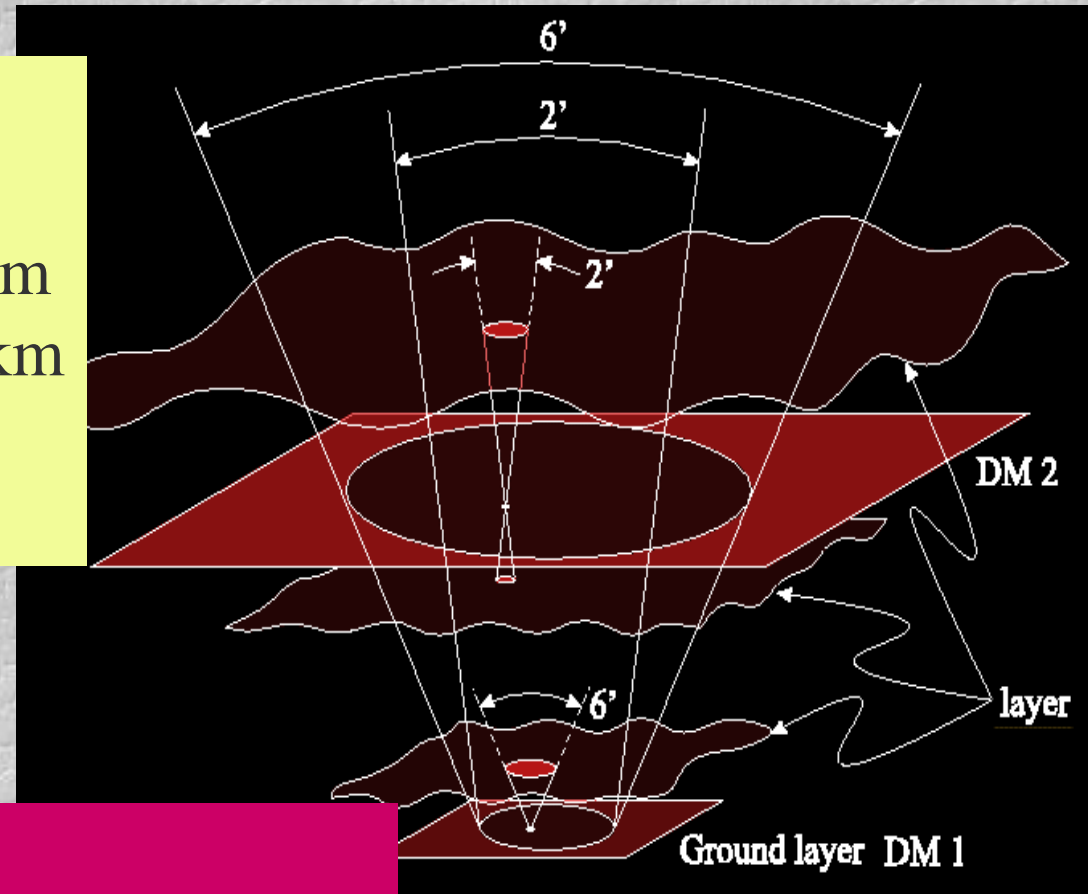
Co-adding of all reference star light !!





Multiple FOV Layer oriented MCAO with NGSs

- 2 DMs at 0 and 8kms
- ~15000 actuators/DM
- up to 12 NGSs in 6' FOV for 0 km
- up to 6 NGS in 2-3' FOV for 8 km
- Corrected FOV: 1-2'
- NIR imager with $Sr(K) < 20-30\%$



Questions (for scale 30-100m!)

- Best altitudes & variability for DMs: $Cn^2(h)$
- Predictability of WF evolution/layer: WF frozen?
- WF temporal frequency for each layer
- WF spatial & amplitude/freq. for each layer



XAO with on-axis NGS

- 30 m telescope
- High order conventional AO
- 2 arcsec FOV
- One DM at 0km
- 10-50 kact.
- XAO in K-I bands
- Coronagraph
- WFE < 100nm rms!!

- 60-100m telescope
- High order AO
- Few arcsec FOV
- Two DMs at 0km
- 100-200 kact.
- XAO in the K-band; ~10% in V
- Coronagraph-Spectroscopy
- WFE < 130nm rms!!

Questions (for scale 30-100m!)

- Scintillation and impact on AO correction
- WF spatial & amplitude frequency

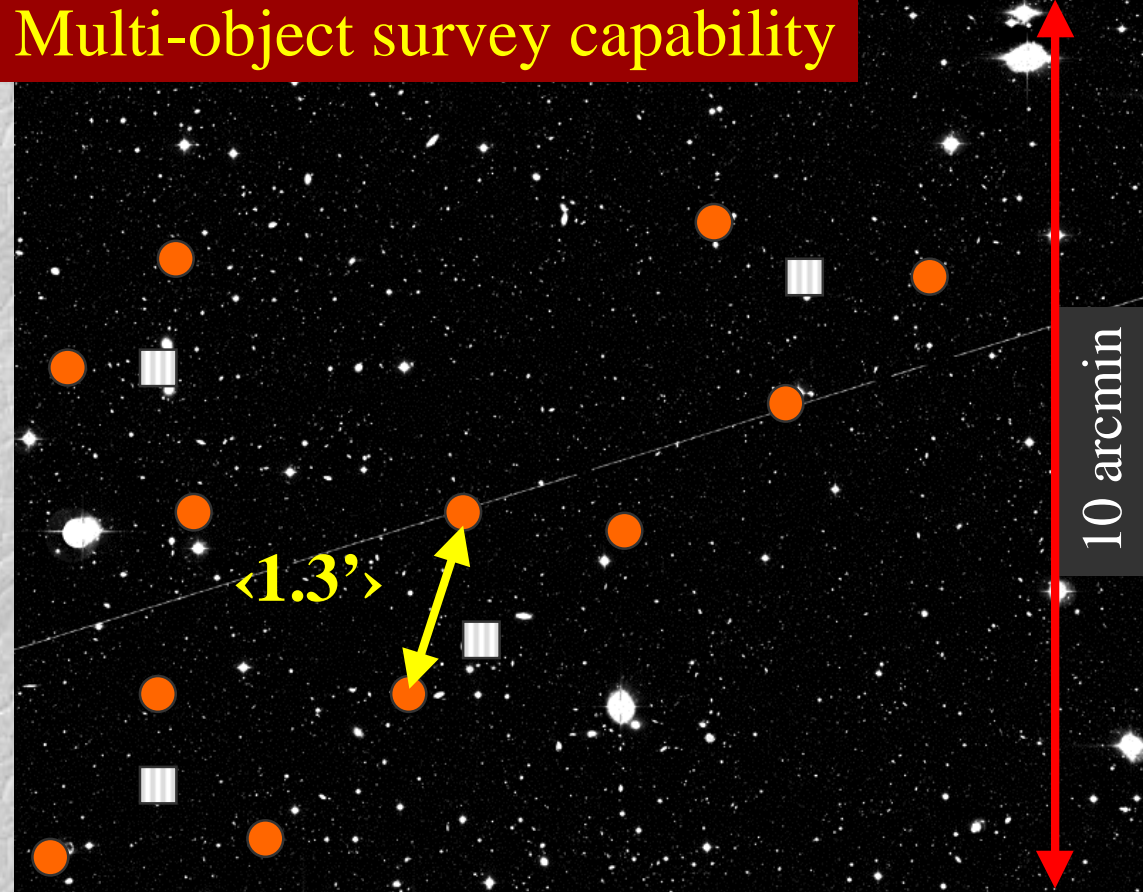


Very Large FOV Multi- seeing reducer



- Search FoV=10-12'
- Corrected FOV: few''
- Multi-AO buttons (~2kact.)
- Coupled with ground layer DM ?
- Multi- DM/IFUs
- Partial correction AO
- Spatial sampling: 0.12''
- Gain 1-2 mag. @ 1.4 μm
- R=3000-20000
- Multiplex gain= 60

Multi-object survey capability



Questions:

- Ground layer amplitude & altitude/ very large FOV
- Best second layer for 2nd stage AO corrector

● mini-WFS

▣ mini-DM/IFU

Prime Focus AO System

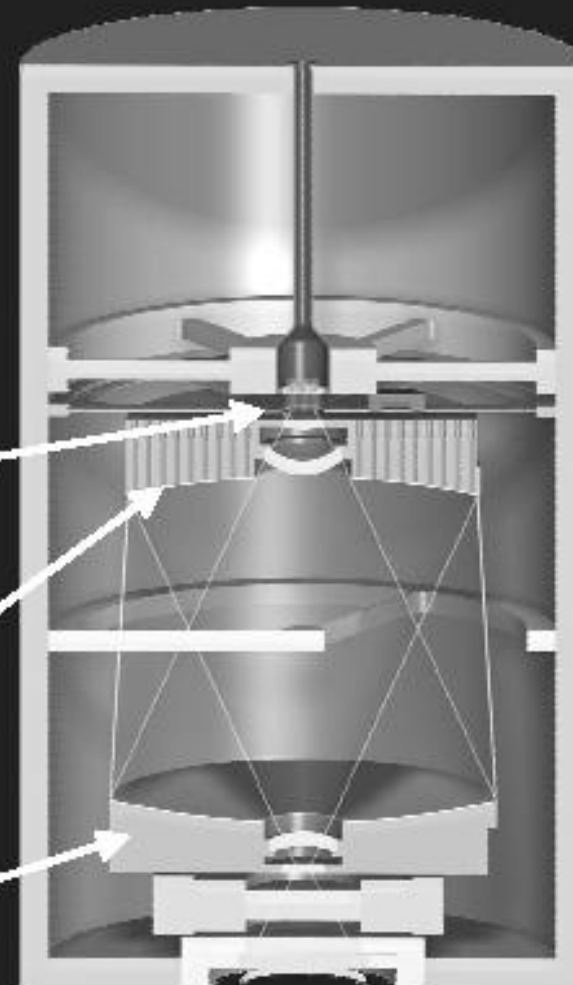
- Corrects M1 warping and ground-level turbulence
 - Achieves moderate improvement over 20-arcmin FOV

Multiple NGS wavefront sensors

Adaptive mirror conjugate to M1

- ~ 1000 actuators

Tip-tilt mirror





Atmospheric parameters requires by AO-ELTs



- § Conventional AO requires r_0 , t_0 , θ_0 , L_0
- § Ground layer tomography requires: $C_n^2(h,t)$ with good resolution at low h ; impacts on corrected FOV of 2-10' !!
- § MCAO requires: $C_n^2(h)$, Spatial and temporal information of WF at scale up to 100m (including frozen WF or not, $L_0(h)$..
- § LGS MCAO requires: Sodium density/altitude/profile variability (Seasonal and mid-short term)
- § XAO requires amplitude of scintillation and effect on WF sensing for high Strehl



Sky coverage issue

Coverage < 0.5

0.5 < coverage < 0.8

Coverage > 0.8

Courtesy: R. Stuik

Atmosphere knowledge & AO for 8-100m telescopes



No of stars	Field Size (')	Limiting Magnitude		
		<14.5	<17.5	<18.5
1	1	0.09	0.43	0.55
	2	0.27	0.76	0.86
	3	0.47	0.91	0.97
	6	0.83	1.00	1.00
3	1	0.00	0.11	0.20
	2	0.02	0.39	0.54
	3	0.09	0.63	0.78
	6	0.45	0.97	1.00
6	1	0.00	0.02	0.07
	2	0.00	0.20	0.32
	3	0.01	0.40	0.54
	6	0.19	0.84	0.96
8	1	0.00	0.01	0.04
	2	0.00	0.14	0.25
	3	0.00	0.31	0.45
	6	0.12	0.74	0.90
12	1	0.00	0.00	0.01
	2	0.00	0.07	0.16
	3	0.00	0.21	0.34
	6	0.05	0.58	0.76