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Spatially-resolved Studies of $z \sim 1-4$ Star-forming Galaxies

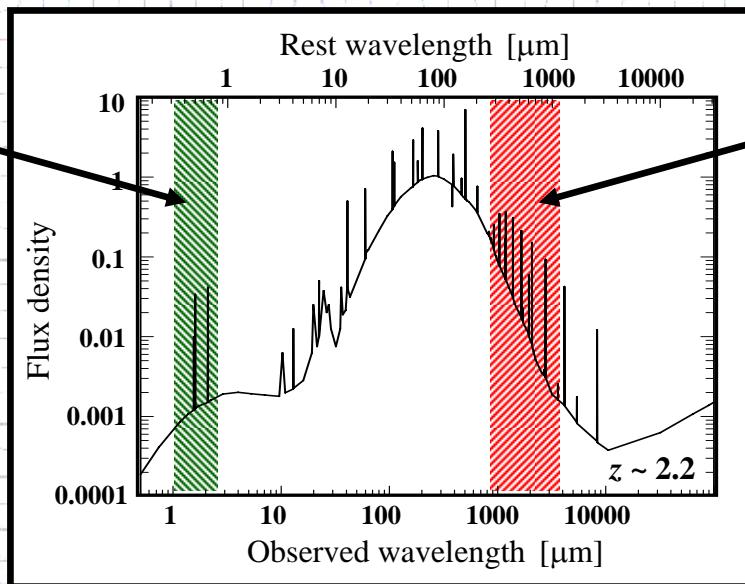
Dynamics, morphologies, physical properties



Rest-frame optical

Stellar & nebular components

SINS survey
in the near-IR with
SINFONI at the VLT



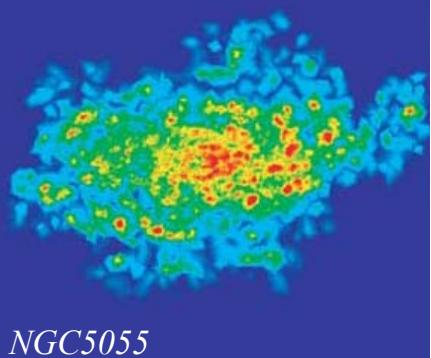
Rest-frame submm

Molecular gas & cold dust

CO survey with the
IRAM/Plateau de Bure
mm-interferometer

Nearby disk galaxy

H α flux



NGC 5055

H α velocity

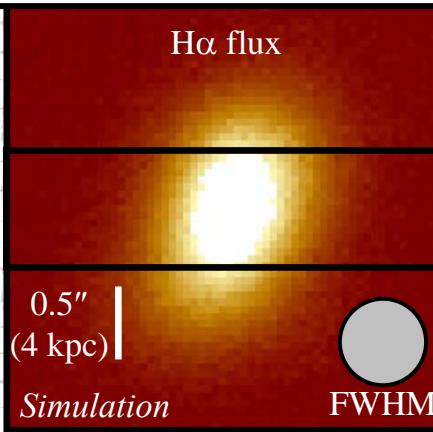


Daigle et al. (2006)

**8-10m (+AO) / HST:
Resolving on
kpc scales**

Disk galaxy at $z \sim 2$: seeing-limited long-slit spectroscopy

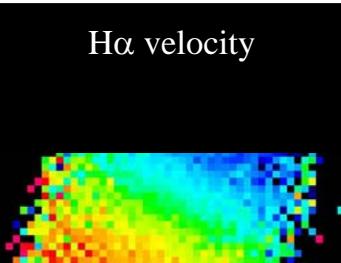
H α flux



$0.5''$
(4 kpc)

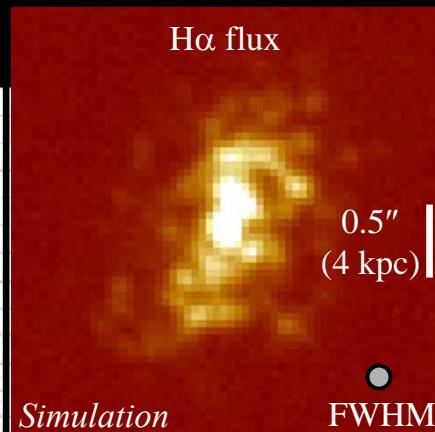
Simulation

H α velocity



Disk galaxy at $z \sim 2$: integral field spectroscopy + AO

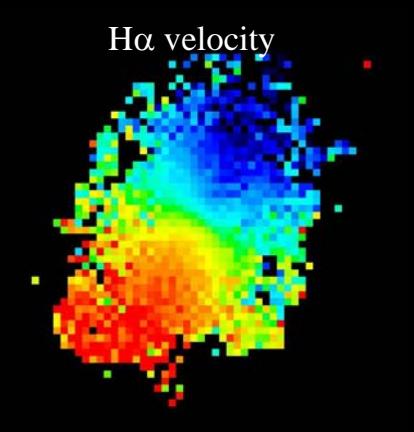
H α flux



$0.5''$
(4 kpc)

Simulation

H α velocity



Förster Schreiber et al. 2006, 2009;

Genzel et al. 2006, 2008; Bouché et al. 2007;

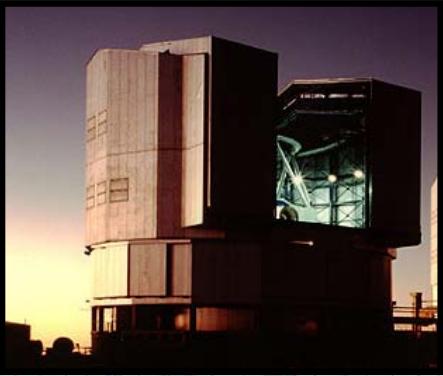
Cresci et al. 2009; Shapiro et al. 2009

Also, e.g., Tecza et al. 2004; Swinbank et al. 2006, 2007;

Nesvadba et al. 2006a,b, 2007, 2008; Law et al. 2007, 2009;

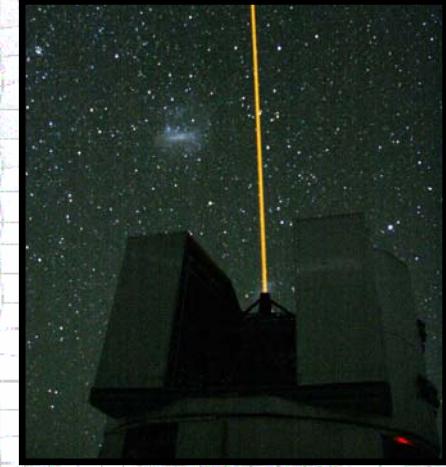
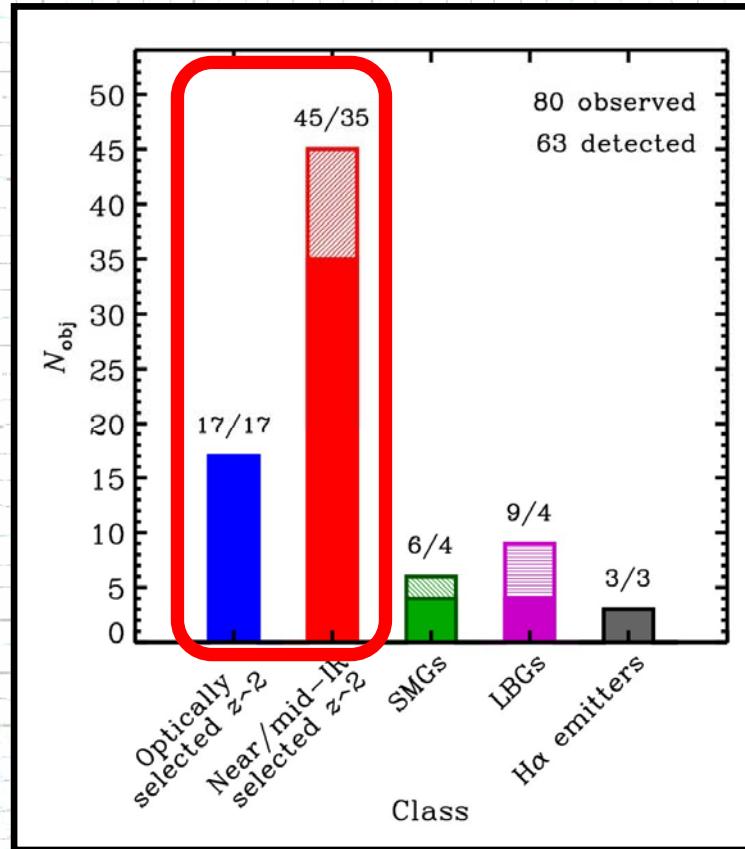
Wright et al. 2007, 2009; van Starkenburg et al. 2008;

Stark et al. 2008; Épinat et al. 2009; Mannucci et al. 2009



Near-IR
integral field spectroscopy
with SINFONI (+AO)
at the VLT

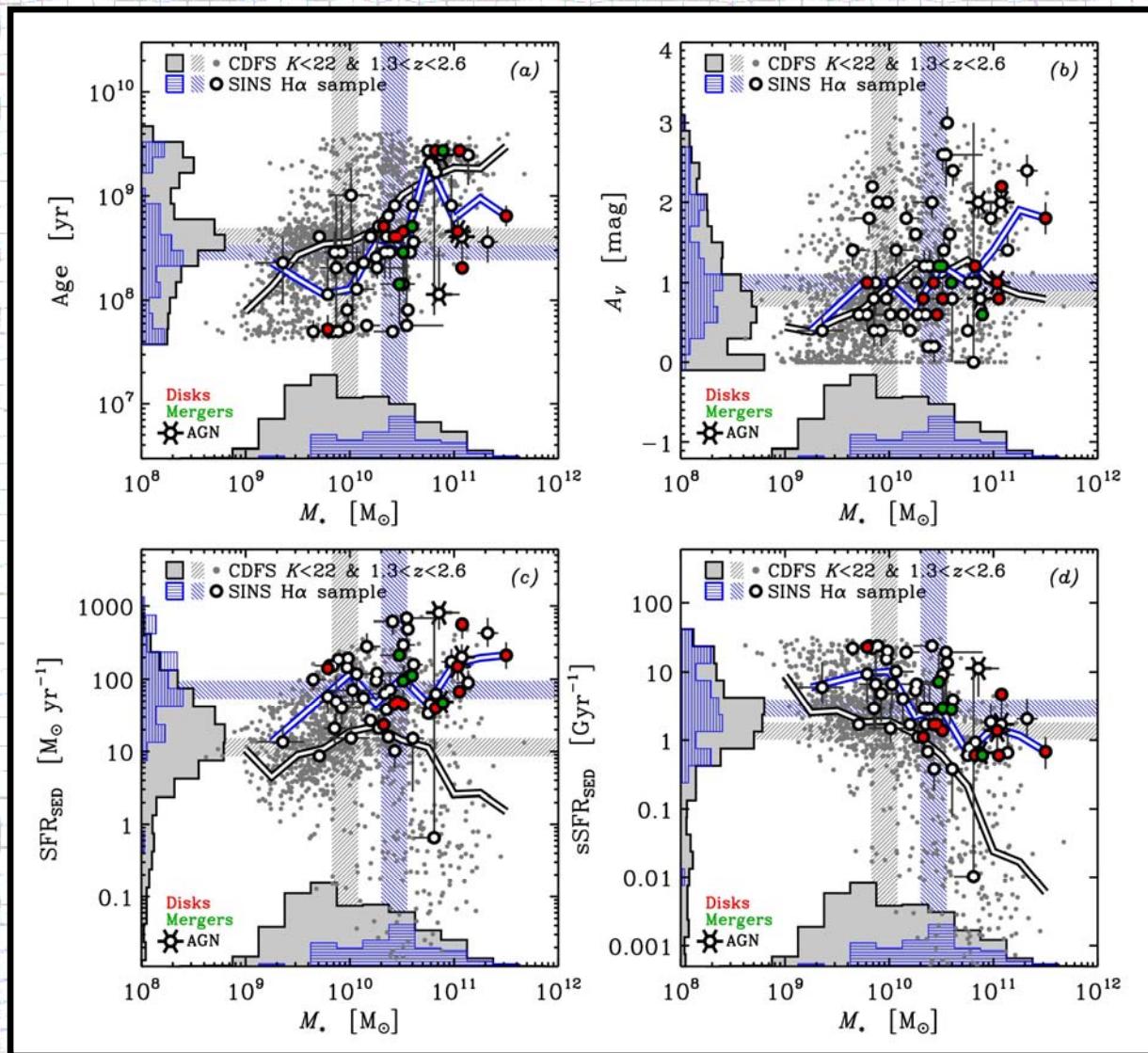
SINS Survey

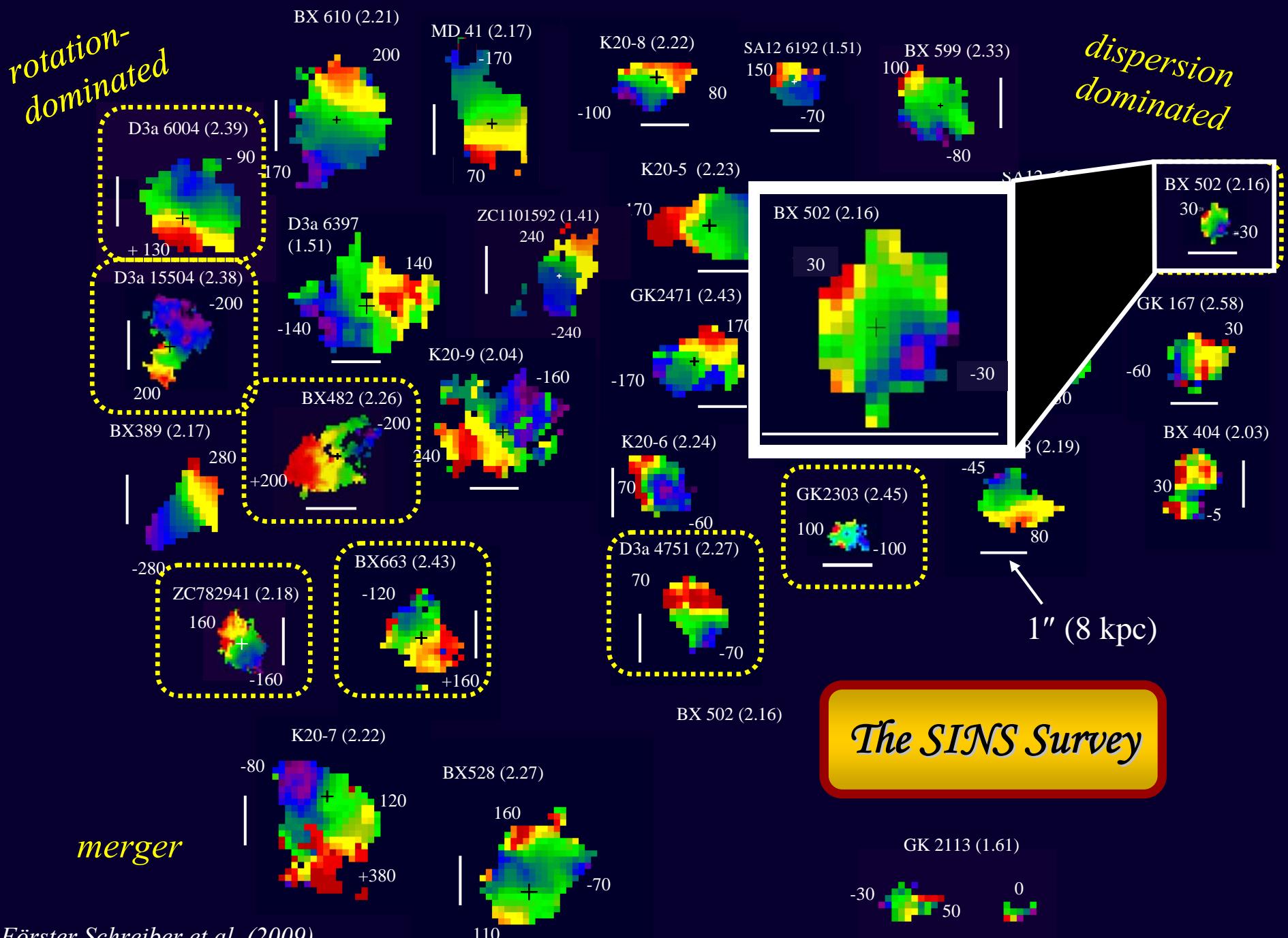


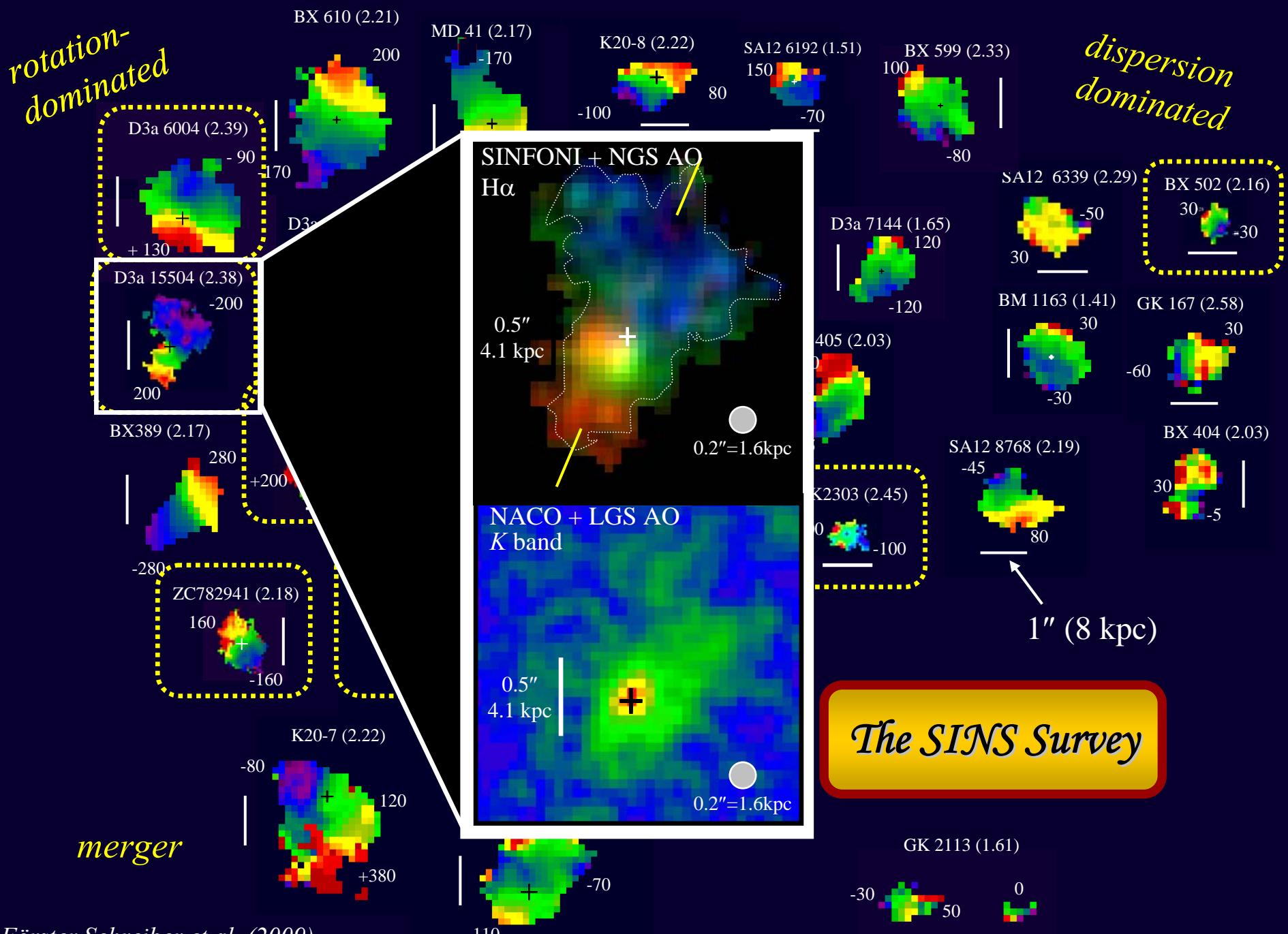
Complemented with
near-IR imaging
with HST/NICMOS-NIC2
and VLT/NACO+AO

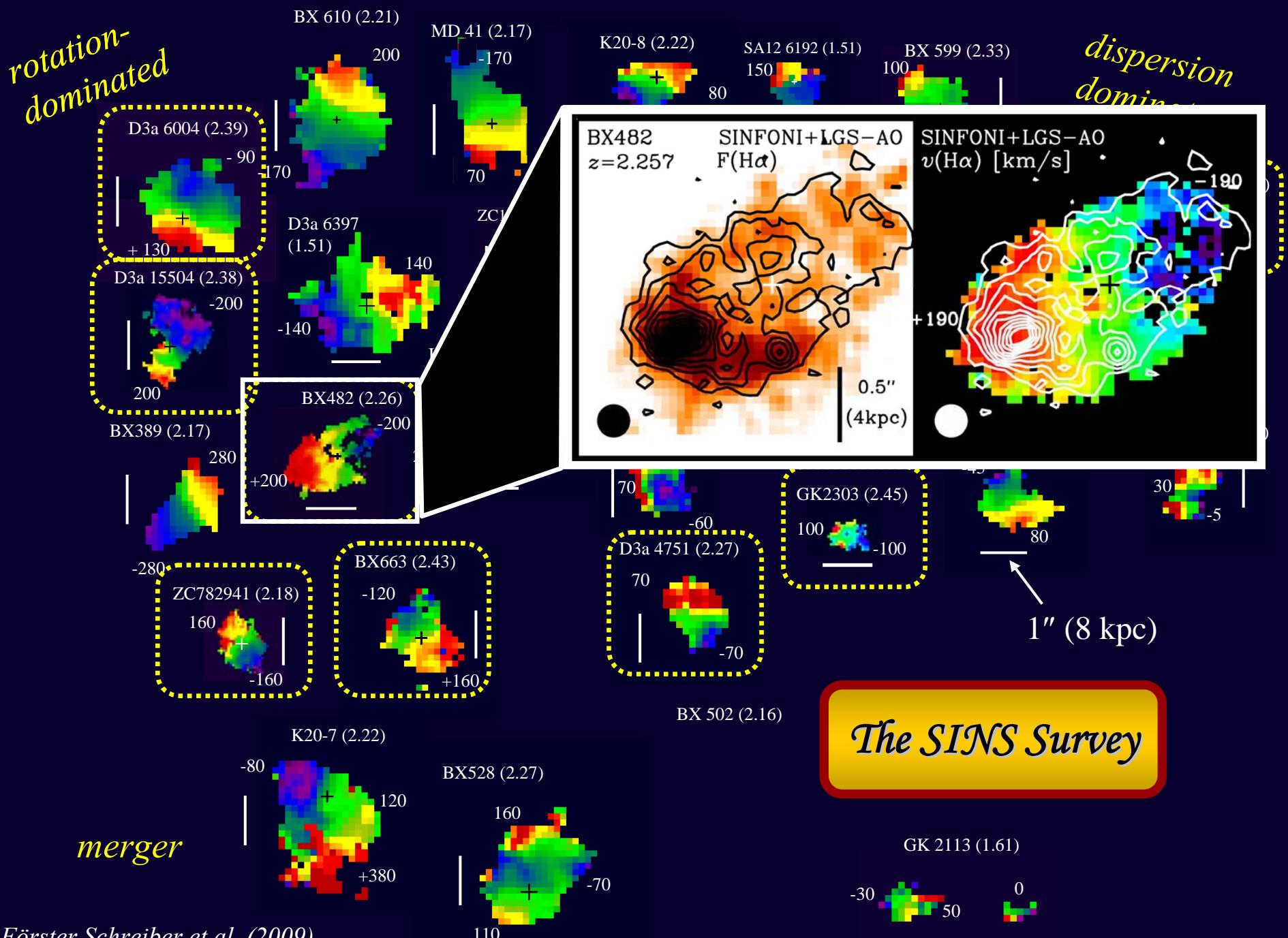
Förster Schreiber et al. 2006, 2009; Genzel et al. 2006, 2008; Bouché et al. 2007; Cresci et al. 2009; Shapiro et al. 2009
Also: e.g., Tecza et al. 2004; Nesvadba et al. 2006a,b; 2007; 2008; Swinbank et al. 2006, 2007; Wright et al. 2007, 2009;
Law et al. 2007, 2009; Stark et al. 2008; Bournaud et al. 2008; van Starkenburg et al. 2008; Epinat et al. 2009; Mannucci et al. 2009

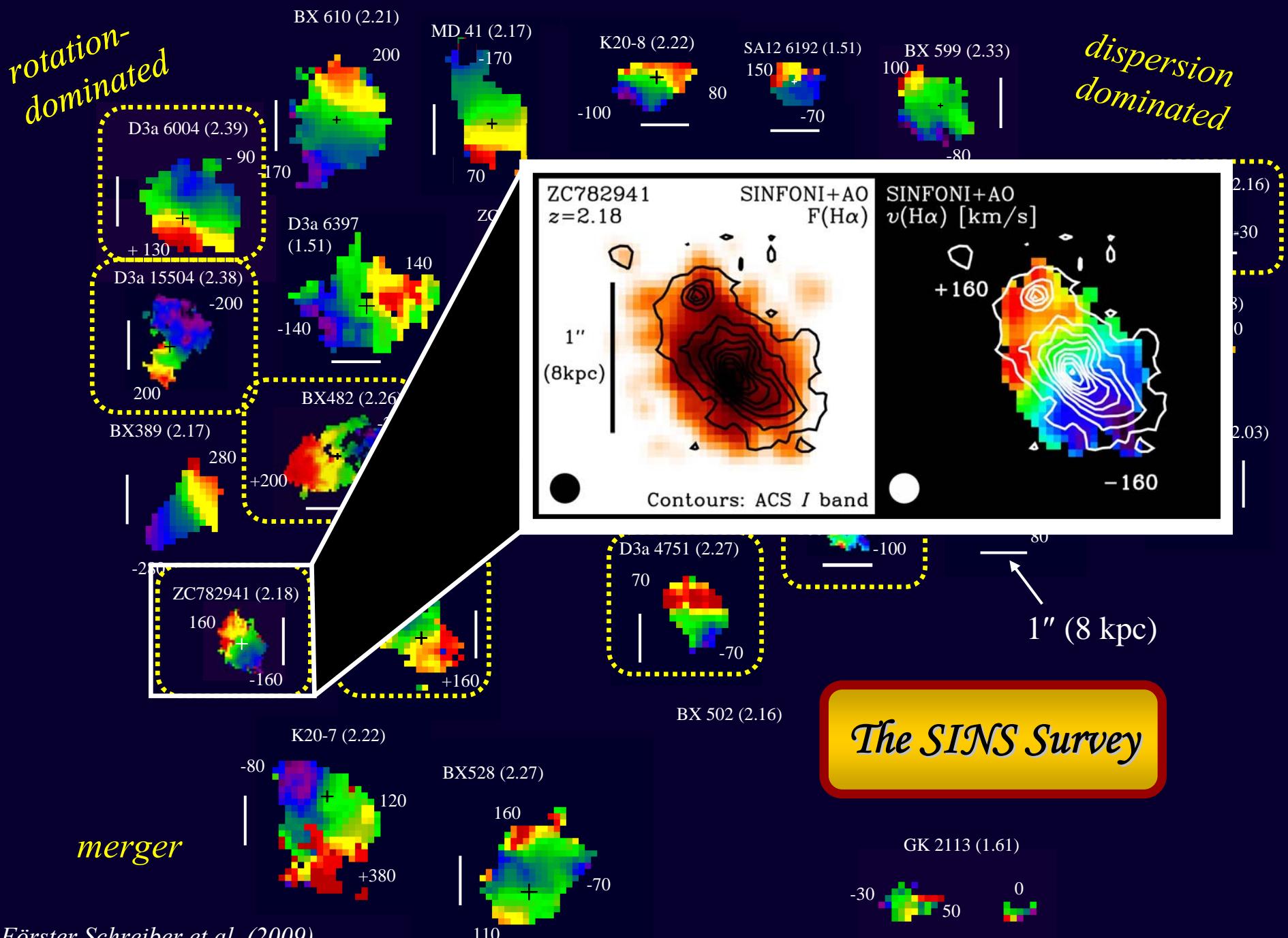
SINS H α Sample



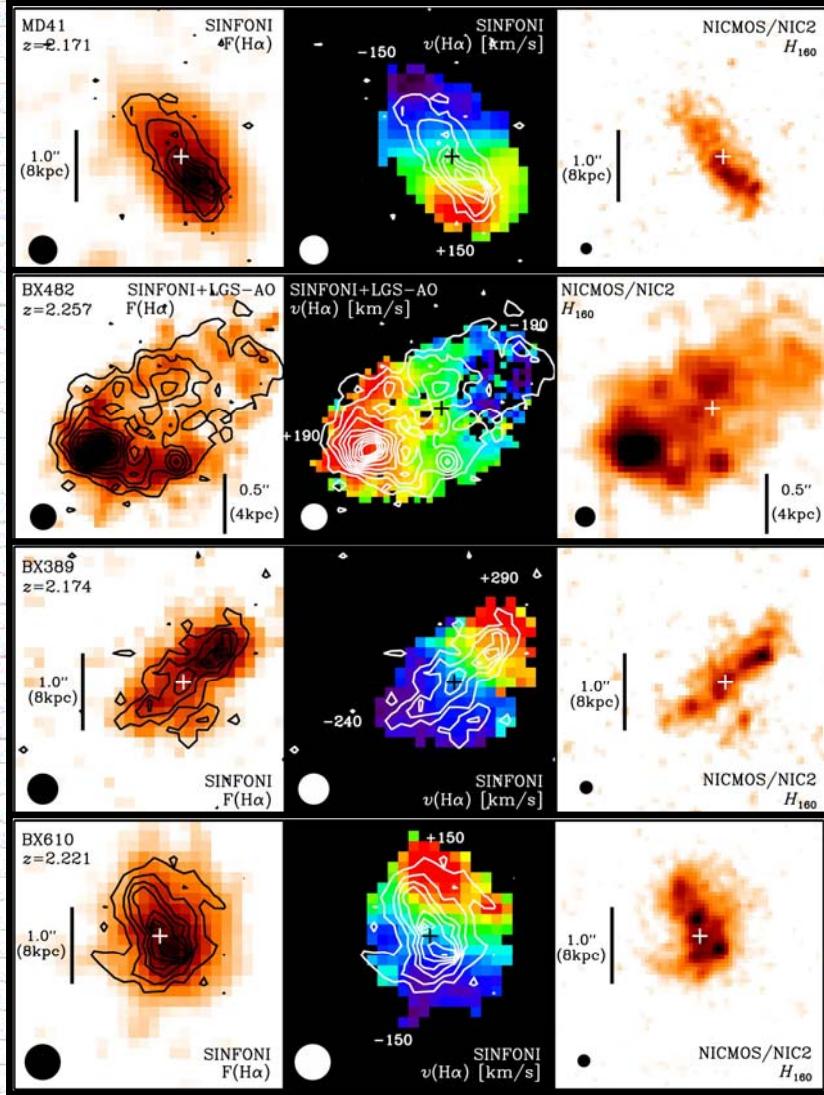






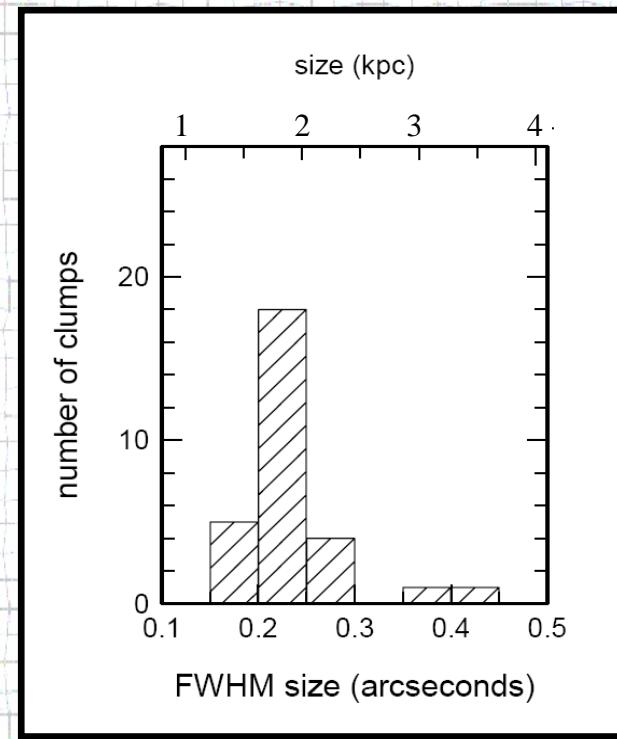


Dynamical Evolution of Gas-rich Disks



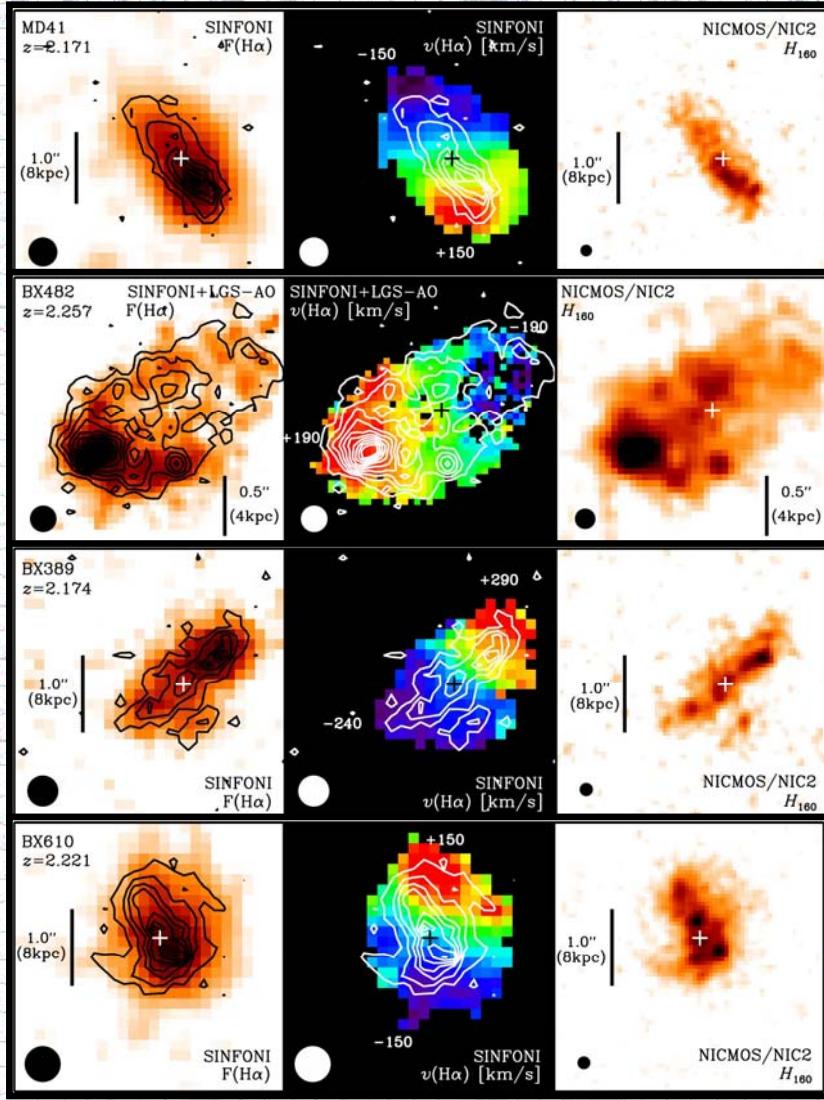
For $Q \approx 1$:

- $L_J \propto (\sigma_0/v_c)R_d \sim 2.5 \text{ kpc}$
- $M_{\text{cl}} \propto (L_{\text{cl}}v_c)^2/R_d \sim 10^{9.4} M_\odot$

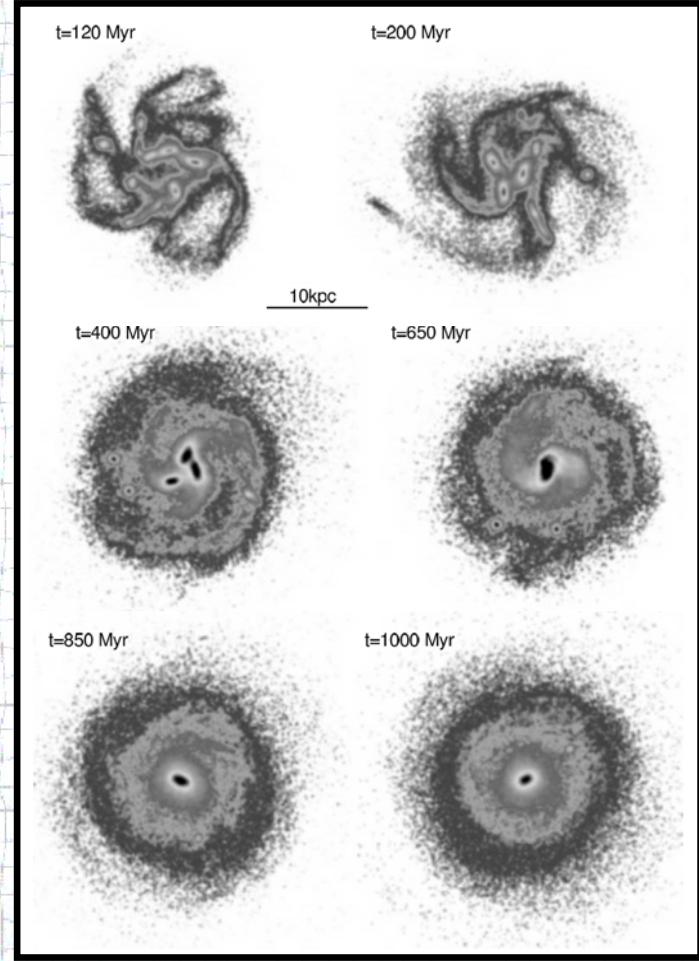


(Also, Cowie et al. 1995; van den Bergh et al. 1996;
Giavalisco et al. 1996; Conselice et al. 2004;
Lotz et al. 2004; Papovich et al. 2005;
Toft et al. 2007; Law et al. 2007;
Elmegreen, Elmegreen, et al. 2004-2009; and others)

Dynamical Evolution of Gas-rich Disks



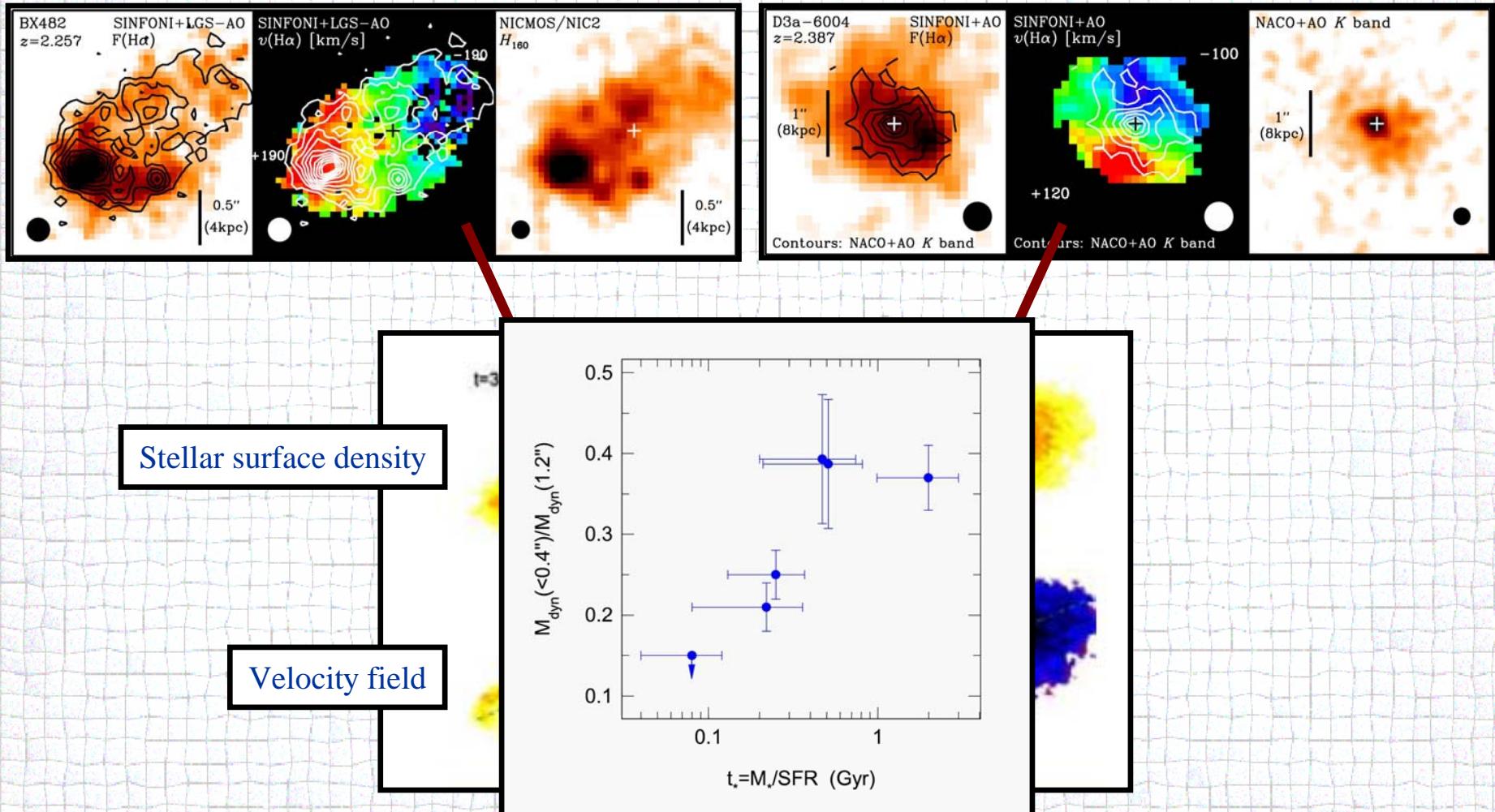
Förster Schreiber, Shapley, et al. (2009); Genzel et al. (2008)



Bournaud et al. (2007; 2008)

Also, e.g., Noguchi (1999); Immeli et al. (2004a, b); Semelin & Combes (2002); Naab et al. (in prep.)

Dynamical Evolution of Gas-rich Disks

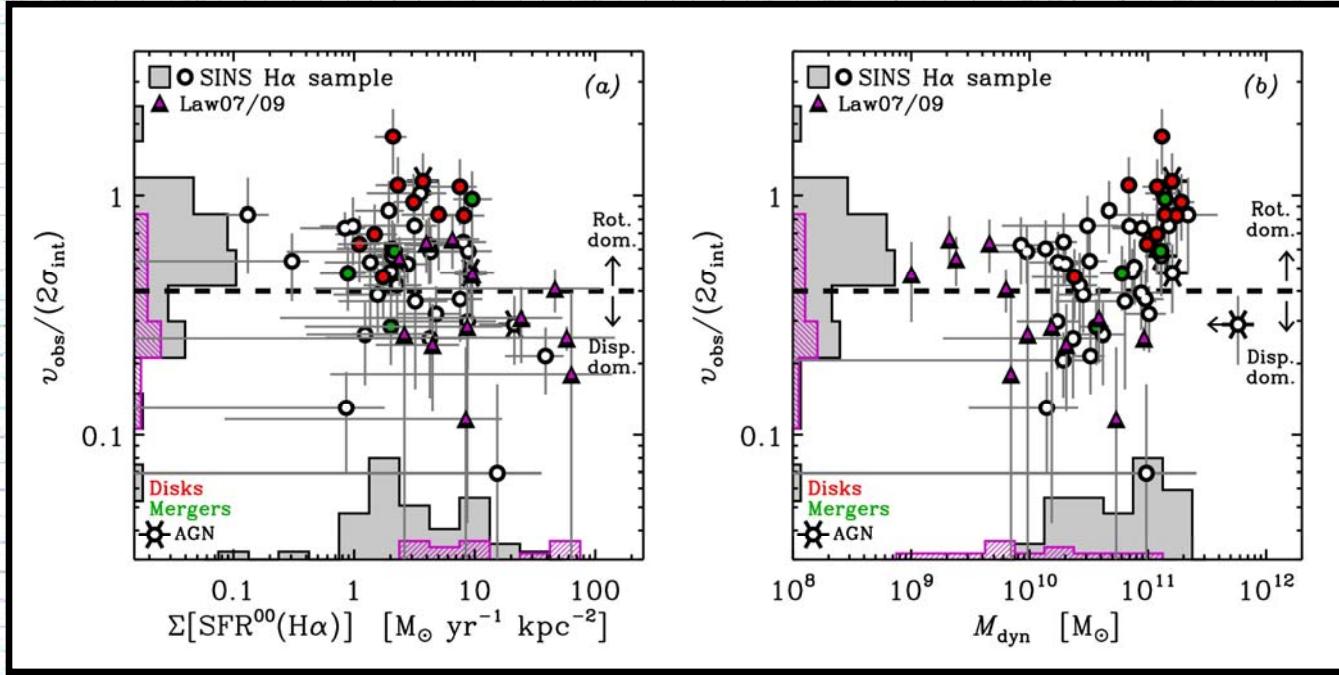


Förster Schreiber, Shapley, et al. (2009); Genzel et al. (2008)

Bournaud et al. (2007; 2008; 2009)
Also: Immeli et al. (2004a,b); Naab et al. (in prep.)

Large Velocity Dispersion of Star-forming Disks at $z \sim 2$

- Star formation feedback? Gas accretion? Internal dynamics?
- For $Q \approx 1$, $v_c/\sigma_0 \propto 1/f_{\text{gas}}$

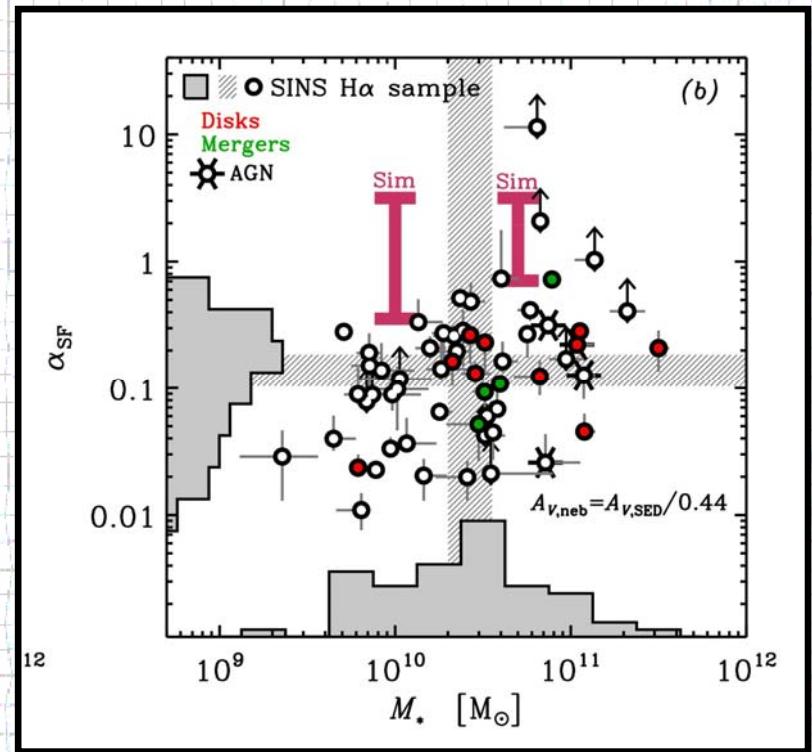
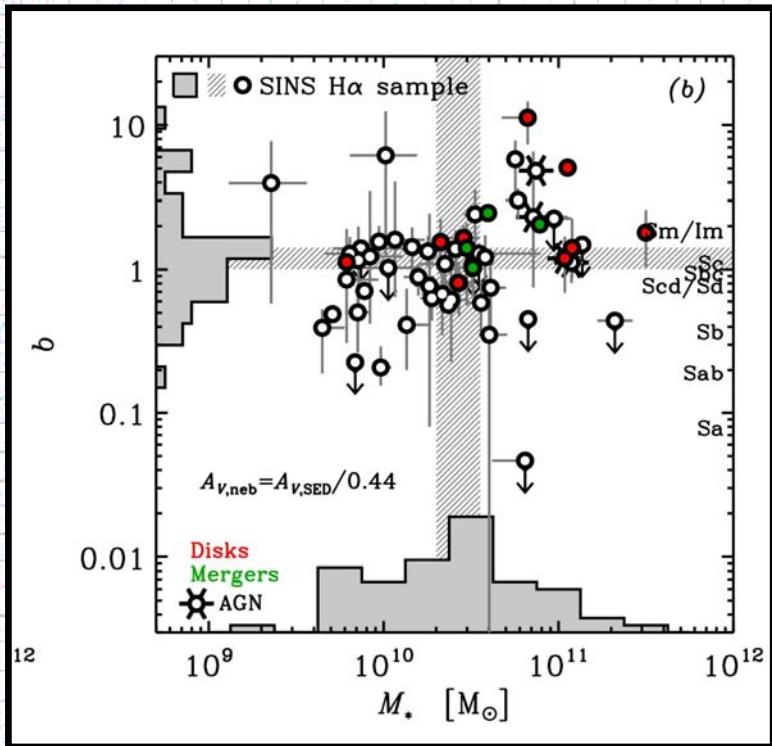


*Rotation- vs
Dispersion-
dominated:*
 $v_c/\sigma_0 \sim 1$
 \leftrightarrow
 $0.5v_{\text{obs}}/\sigma_{\text{int}} \sim 0.4$

Förster Schreiber et al. (2006; 2009); Genzel et al. (2006, 2008); Cresci et al. (2009)
also: Law et al. (2007, 2009); Wright et al. (2007); van Starkenburg (2008); Stark et al. (2008); Épinat et al. (2009)

Star Formation Activity and Timescales

$$\tau_{\star} \sim \tau_{\text{gas}} \sim 500 \text{ Myr} \sim \text{several } \tau_{\text{dyn}} \ll t_{\text{Hubble}}$$

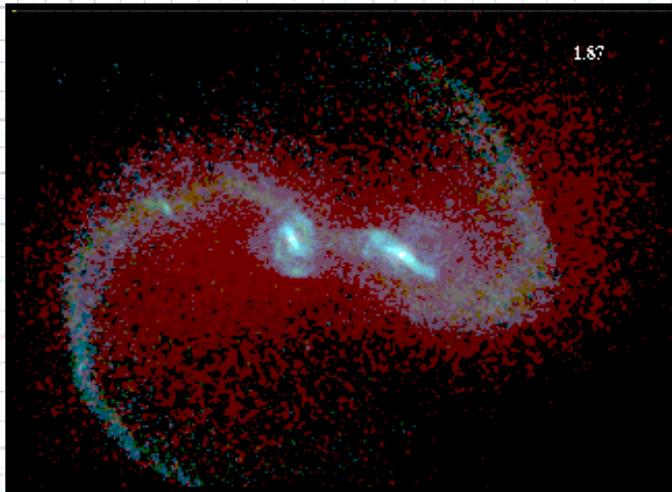


Förster Schreiber et al. (2006; 2009); Genzel et al. (2008); Bouché et al. (2009)
see also, e.g., Kennicutt et al. (1994); Daddi et al. (2007); Davé (2008); Chen et al. (2009)

Rapid Star Formation/Mass Accretion: Major Mergers or Smoother Infall?

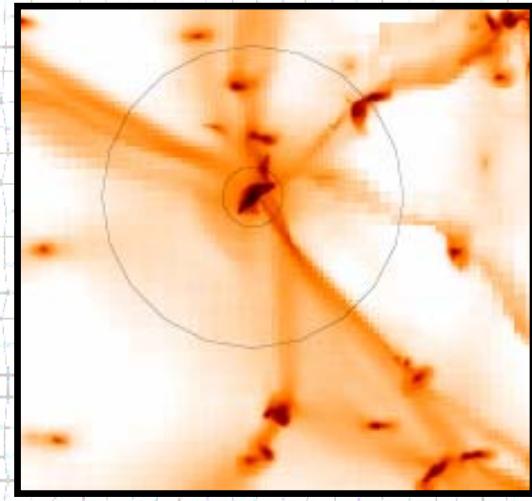
$$\tau_{\star} \sim \tau_{\text{gas}} \sim 500 \text{ Myr} \sim \text{several } \tau_{\text{dyn}} \ll t_{\text{Hubble}}$$

Major mergers



(e.g., Toomre & Toomre 1972; Barnes & Hernquist 1996;
Springel & Hernquist 2005; di Matteo et al. 2005;
Naab & Burkert 2003,2006; Hopkins et al. 2006;
Tacconi et al. 2006,2008; Swinbank et al. 2006
Robertson et al. 2008)

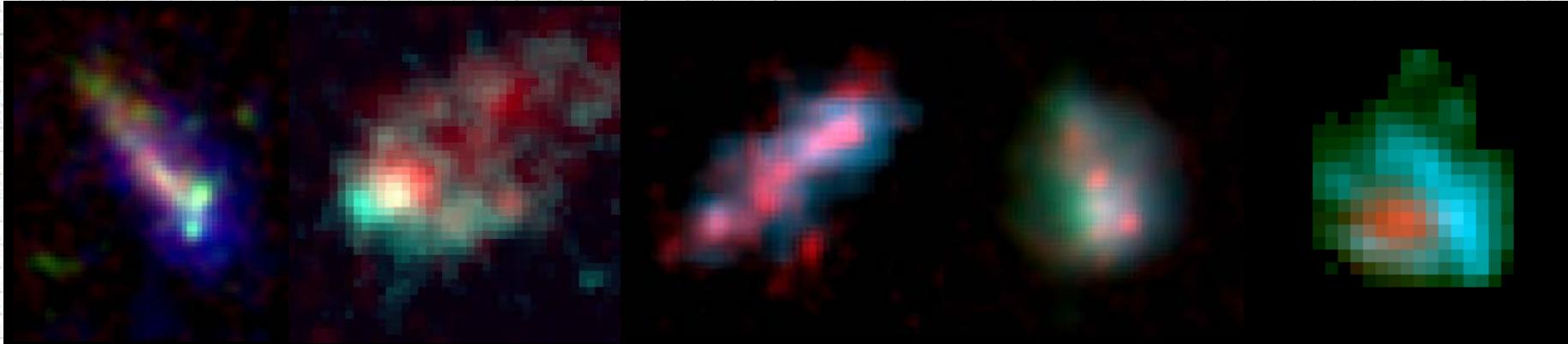
Cold flows/minor mergers



(e.g., Dekel & Birnboim 2003,2006; Kereš et al. 2005;
d'Onghia et al. 2006; Kitzbichler & White 2007;
Guo & White 2008; Davé 2008; Noeske et al. 2007;
Elbaz et al. 2007; Daddi et al. 2007;
Dekel et al. 2008, 2009; Genel et al. 2008, 2009)

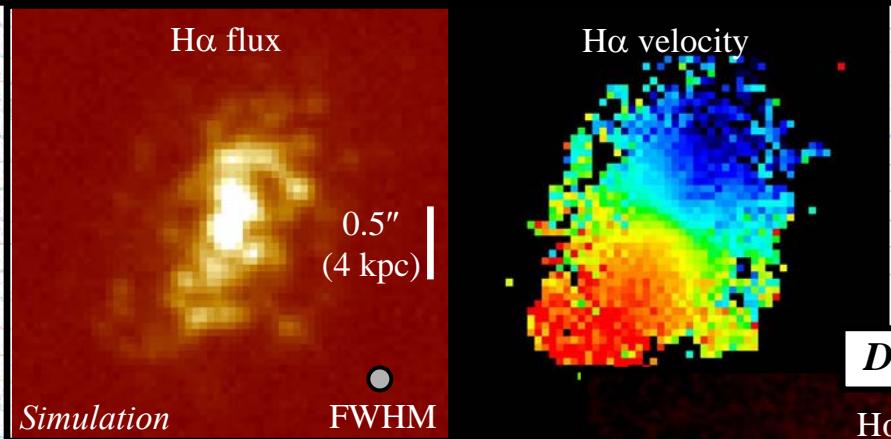
SINS Key Results

- ▶ Kinematics of SINS massive star-forming galaxies at $z \sim 2$
 $\sim 1/3$ rotation-dominated, $\sim 1/3$ compact dispersion-dominated, $\sim 1/3$ mergers
Fraction of rotation-dominated systems increase at higher masses
- ▶ Properties of massive $z \sim 2$ star-forming disks
Significantly more turbulent and gas-rich than local disks
Higher SFRs, large luminous/massive clumps
- ▶ Mass assembly, early evolution, and star formation activity
Evidence for smooth+rapid mass accretion via cold flows/minor mergers
Evidence for internal/secular processes in gas-rich disks and rapid bulge formation

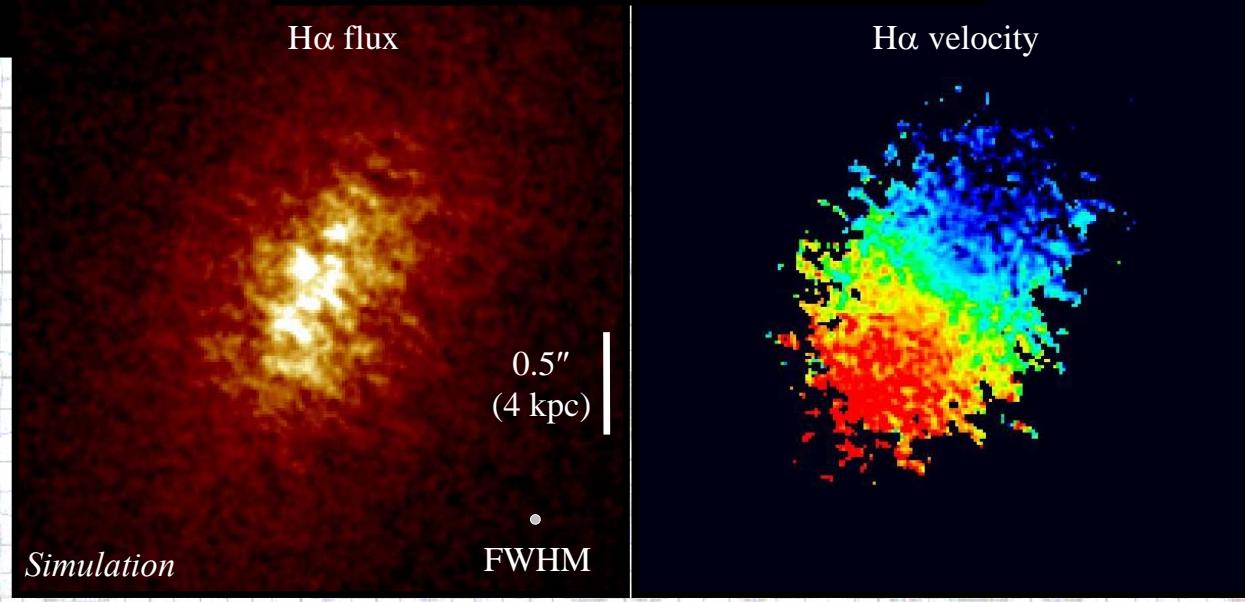


ELTs & ALMA: Resolving sub-kpc scales

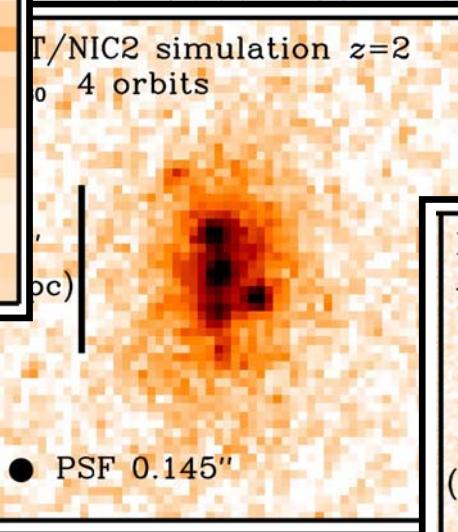
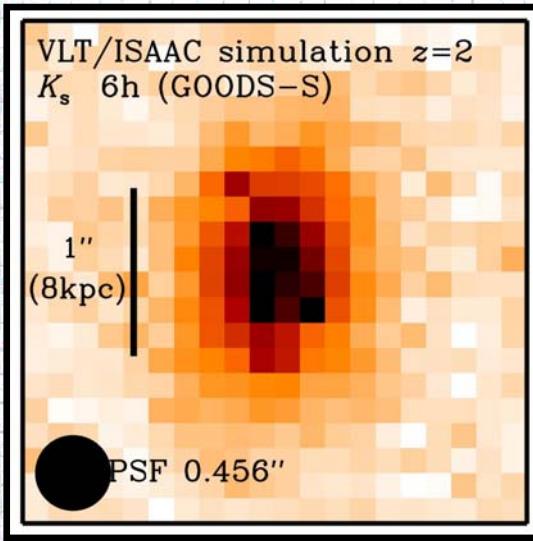
Disk galaxy at $z \sim 2$: IFS + AO on 8-10m telescopes



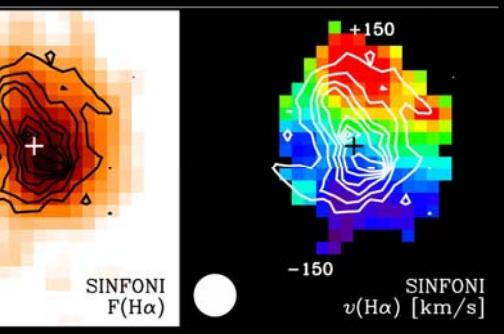
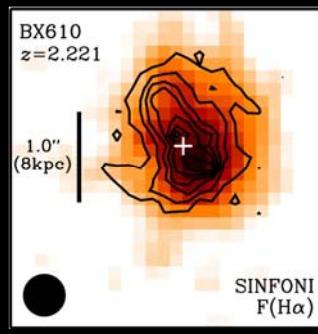
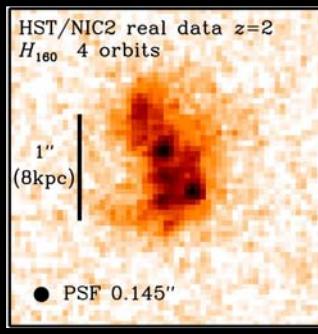
Disk galaxy at $z \sim 2$: IFS + AO on ELTs



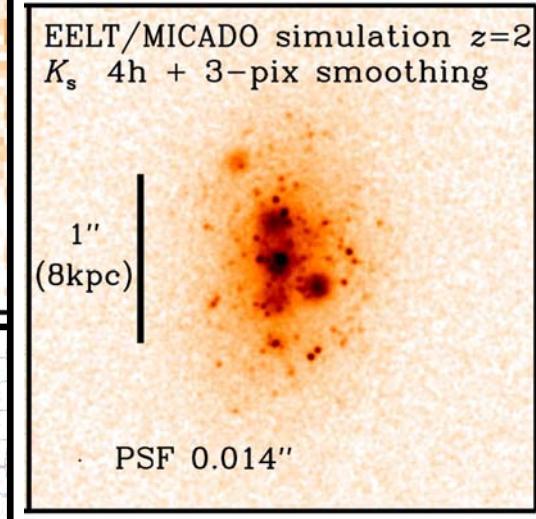
Simulations for E-ELT/MICADO



Mock disk galaxy at $z = 2.3$
 $R_{1/2} = 5 \text{ kpc}, K_{AB} = 21.3$



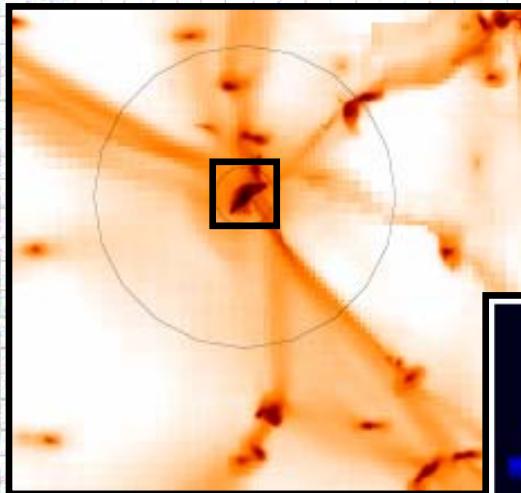
Real disk galaxy at $z = 2.3$
 $R_{1/2} = 5 \text{ kpc}, K_{AB} = 21.3$



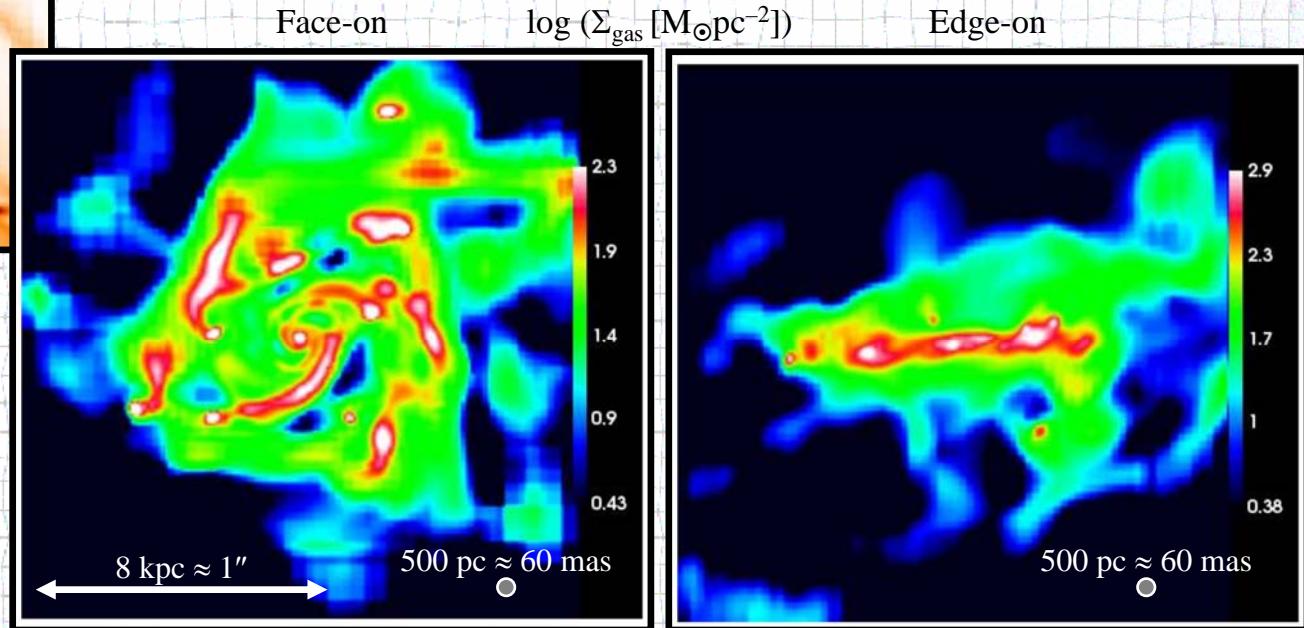
Resolution of $\sim 100 \text{ pc}$ at $z \sim 2$
Compact clusters detected to $K_{AB} \sim 28.5$

Sub-kpc structure at high-z: Simulations

Cold flows/Streams



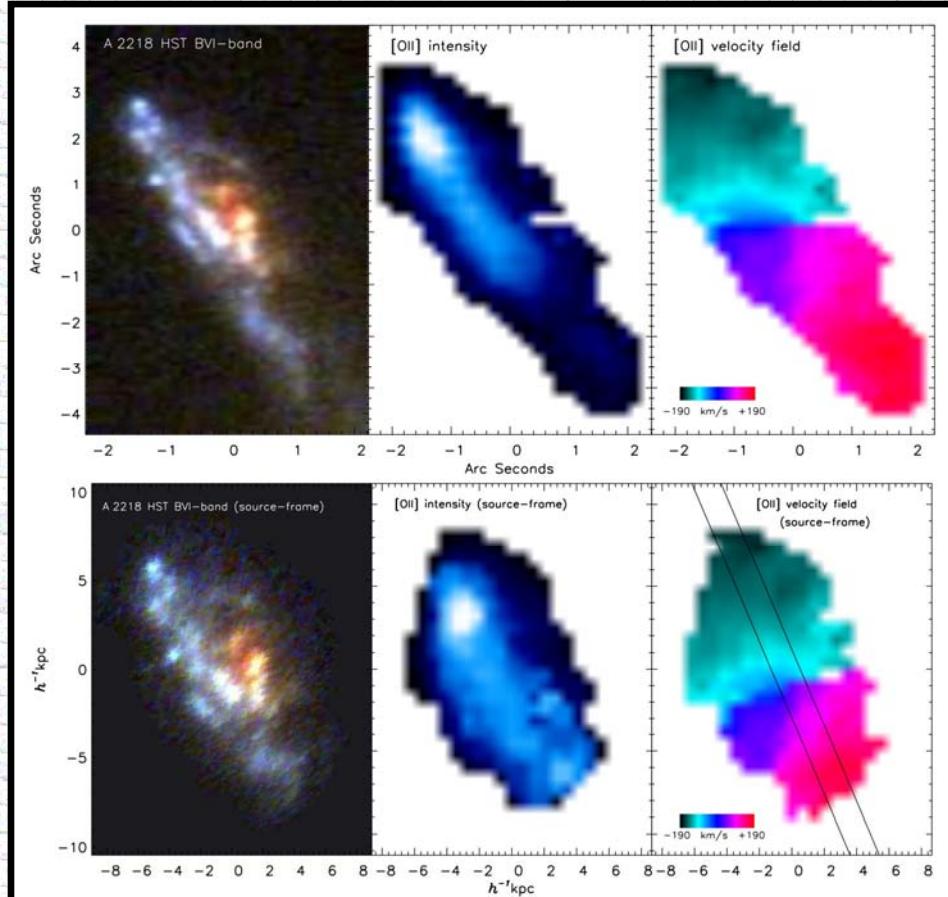
Example hydro-AMR cosmological simulation
Resolution: $70 \text{ pc} \approx 10 \text{ mas}$ at $z \sim 2$



Dekel et al. 2008; Dekel, Sari, & Ceverino 2009; Ceverino & Dekel, in prep.
also, e.g., Naab, Genel, et al., Bournaud et al. (in prep.), among others

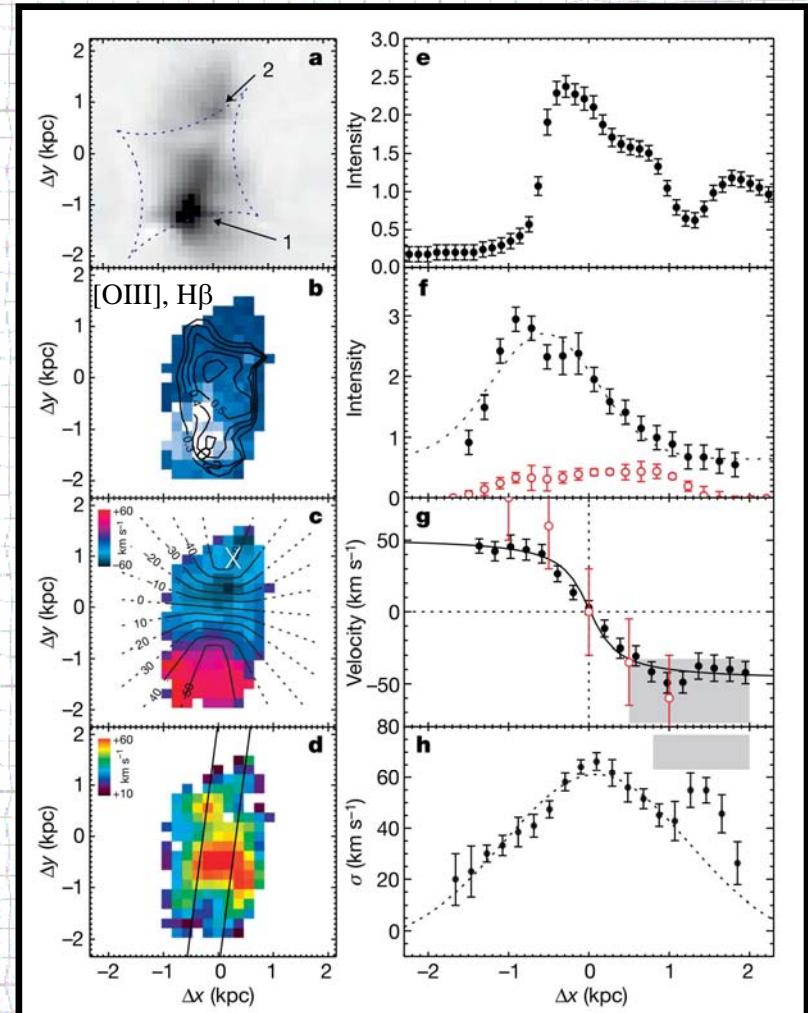
Sub-kpc structure at high z : Lensed Objects

Lensed galaxies at $z \sim 1$, $\mu \sim 10$;
HST resolution ~ 150 pc; GMOS resolution ~ 500 pc



Swinbank et al. (2003, 2006 + in prep.)

Lensed galaxy at $z \sim 3$, $\mu \sim 30$;
HST & OSIRIS+AO resolution $\sim 150 - 300$ pc



Stark et al. (2008)

Prospects for ELTs/ALMA

► Structural, dynamical, and stellar/physical properties on $\sim 100 - 500$ pc scales

Mass, density, metallicity, kinematics profiles

Disk instabilities, properties and substructure of clumps

Age, SFR, and dust distribution

Formation of bulges, globular clusters

Feedback from star formation and AGN

SMBH / IMBH formation

Push down the mass/luminosity function and to higher z

► Synergies ELTs/ALMA

Gas mass fraction is key parameter for evolution of galaxies

Cold gas distribution and dynamics

